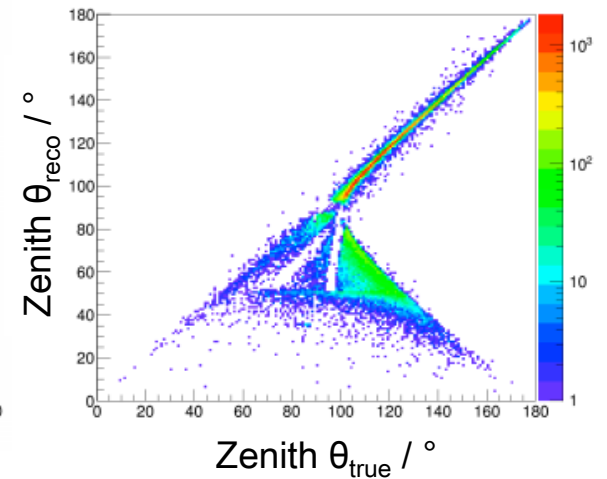
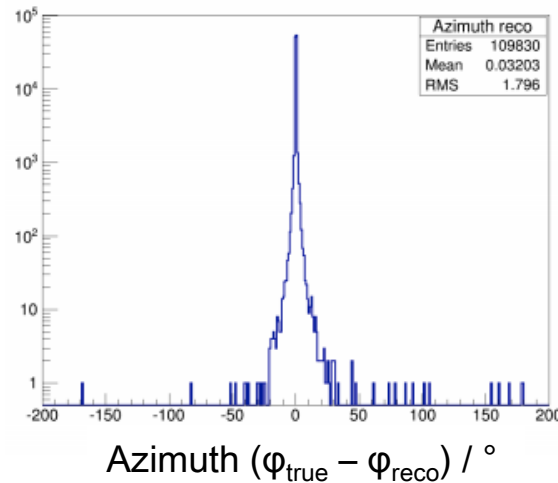
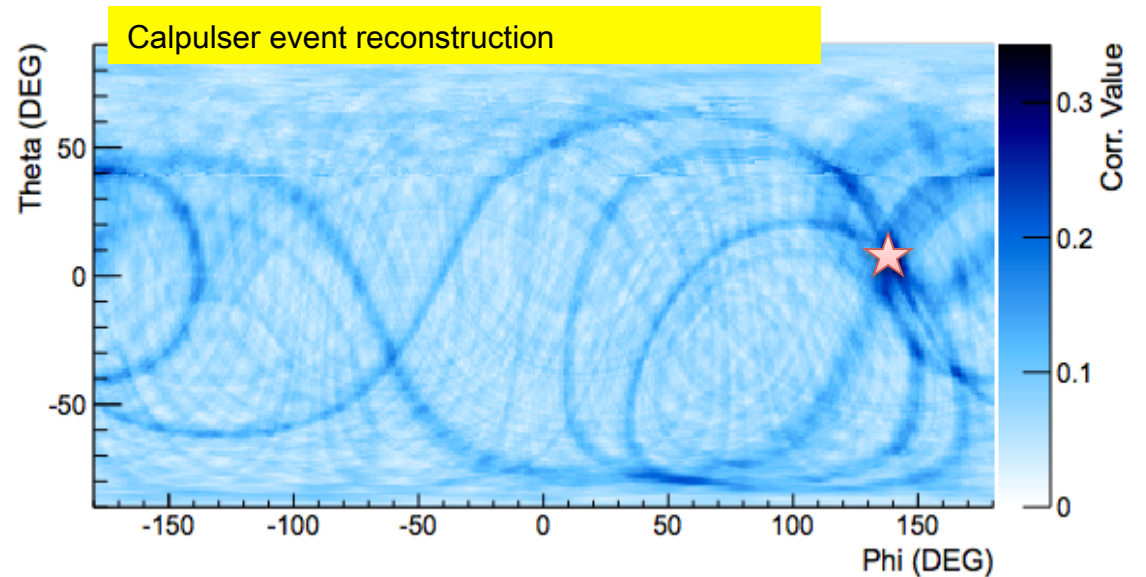


# Interferometric Reconstruction with Radiospline

Ming-Yuan Lu, UW-Madison  
2015 ARA Collaboration Meeting

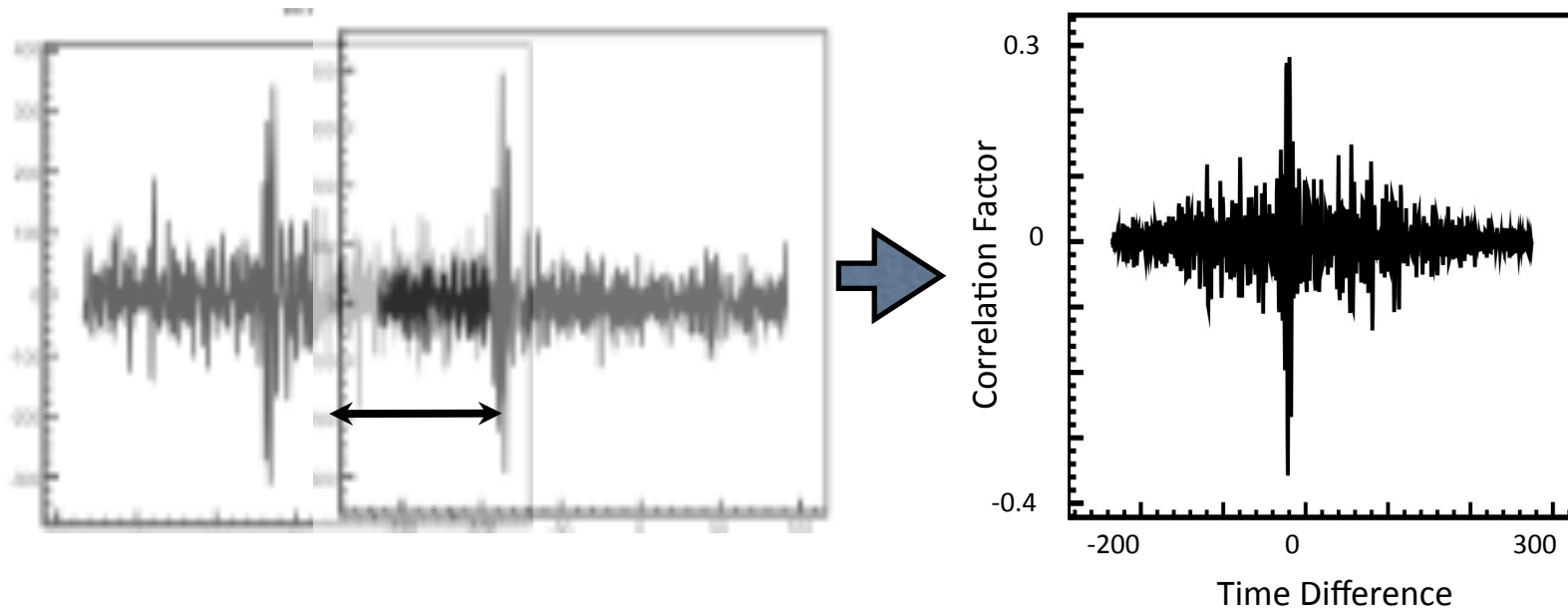
# Brief review of reconstruction methods so far

- Interferometric reconstruction with 2 hypothesized distances (30m & 3km)
- Matrix-based plane-wave reconstruction
- In general  $\sim 1^\circ$  angular resolution can be achieved by both methods



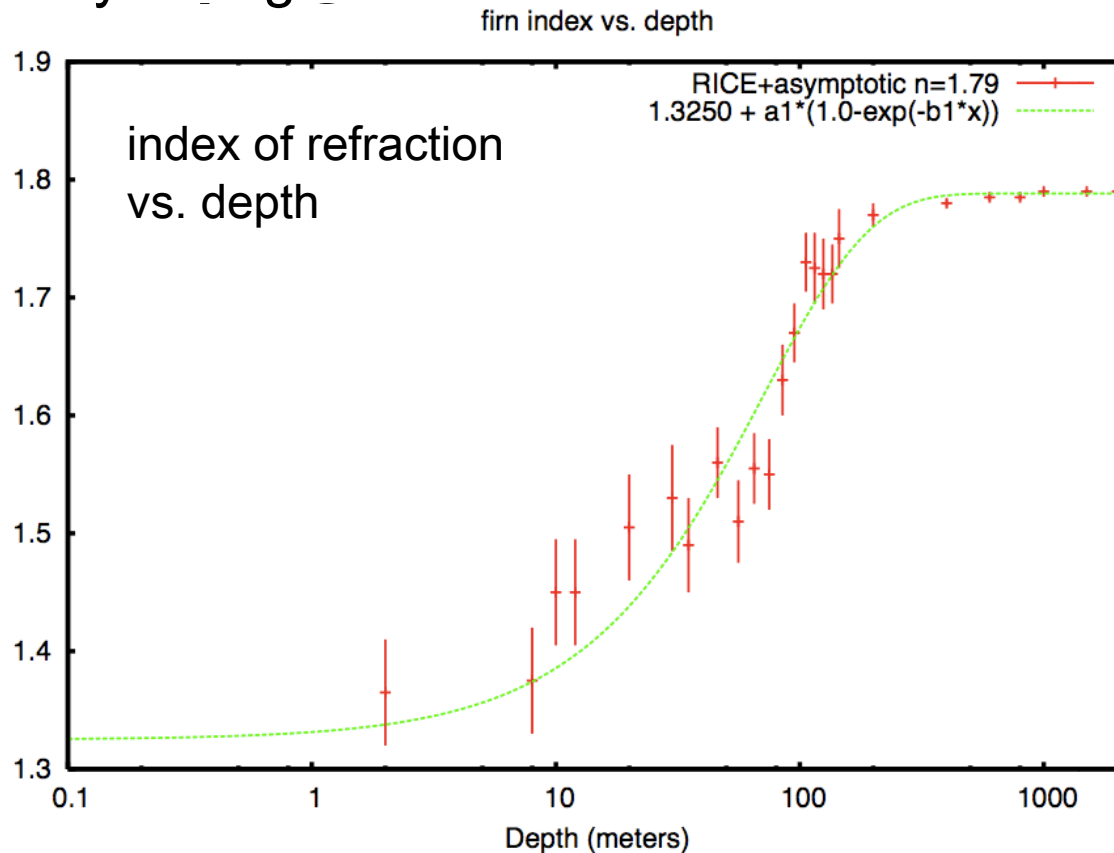
# Interferometric Reconstruction

- Attempt vertex reconstruction using relative timing information from hypothesized source positions
- Waveforms time-shifted according to computed delays



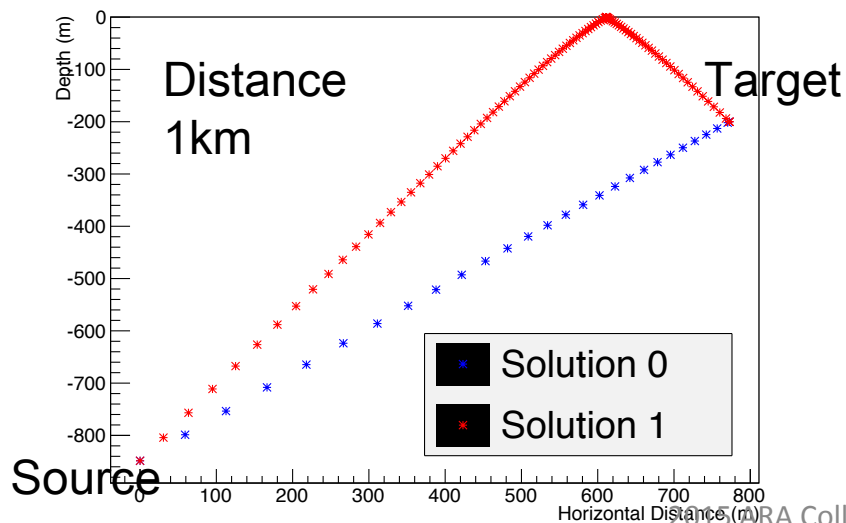
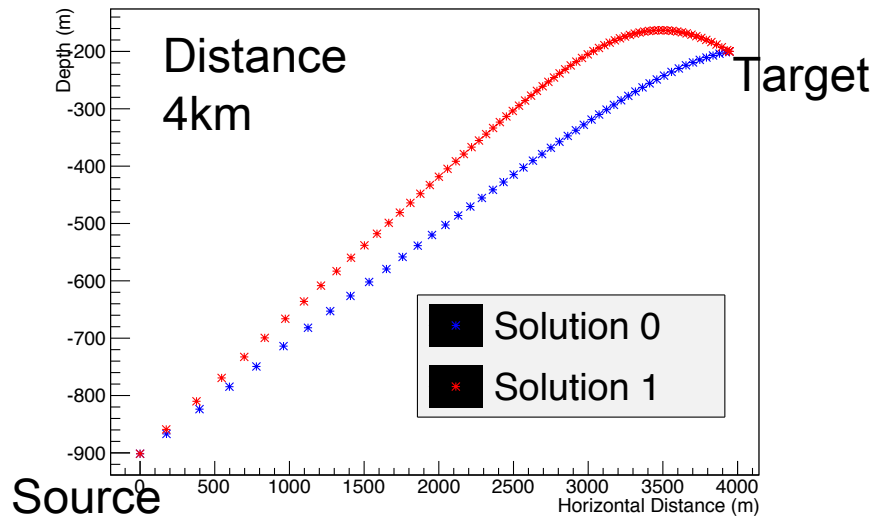
# Where to Get Delays?

- Ice index of refraction varies with depth. Change is most drastic near surface (firn). As a result, EM waves travel in curved paths – raytracing
- Ideal direction/distance reconstruction need to take into account raytracing effect



# Raytracing and Radiospline

- Semi-analytic approach to compute ray paths by C. Weaver. Delays computed by this approach tabulated and fitted with B-Spline

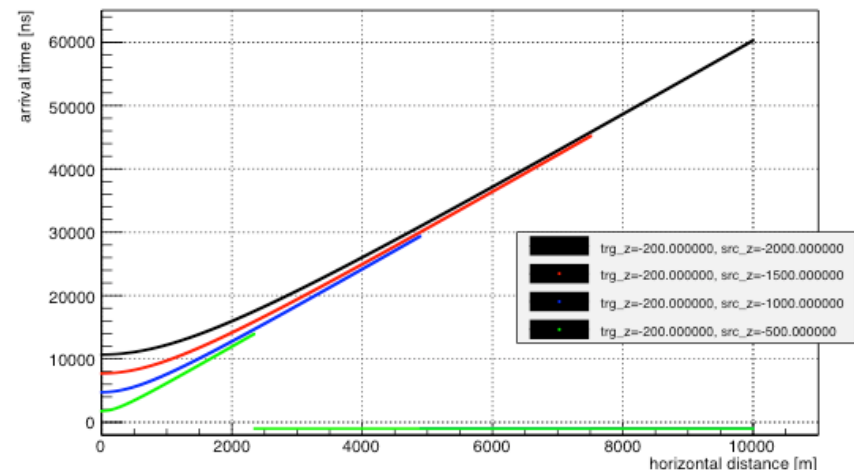


## Radiospline (See J. Kelley talk)

Determining shadow boundary

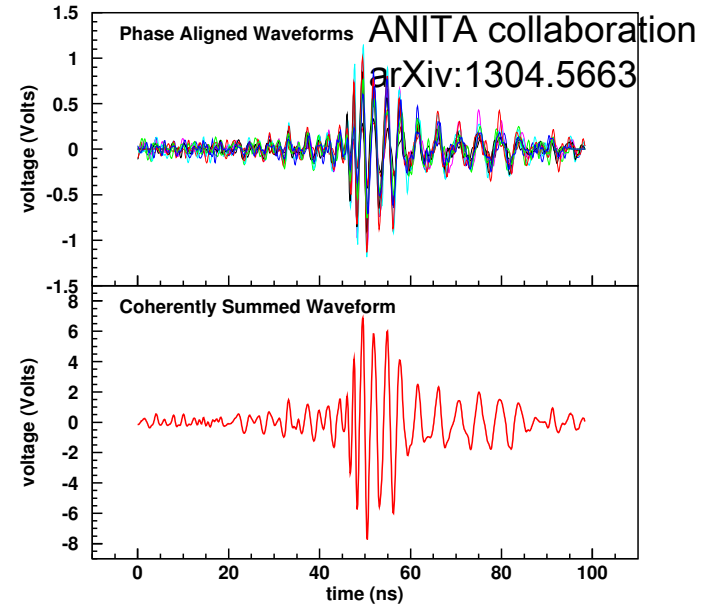


Determining delay



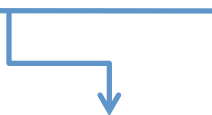
# Coherently Summed Waveforms

$$P_{\Sigma}(\hat{r}) = \frac{1}{Z_L T} \int_0^T \sum_{i=1}^{N_A} \sum_{j=1}^{N_A} dt \cdot v_i(t + \tau_i(\hat{r})) v_j(t + \tau_j(\hat{r}))$$

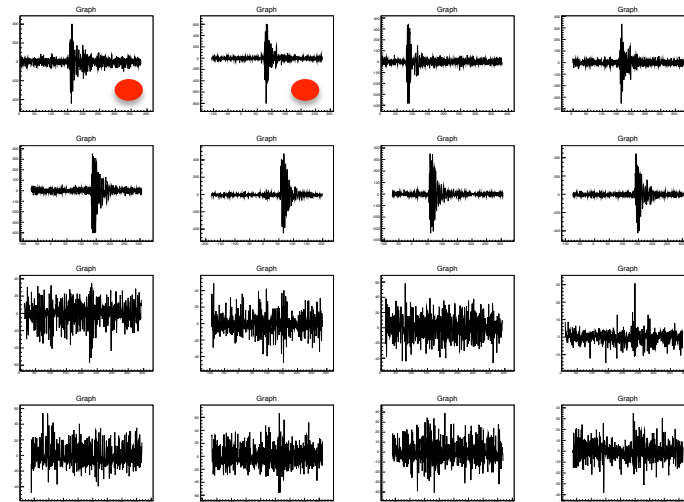
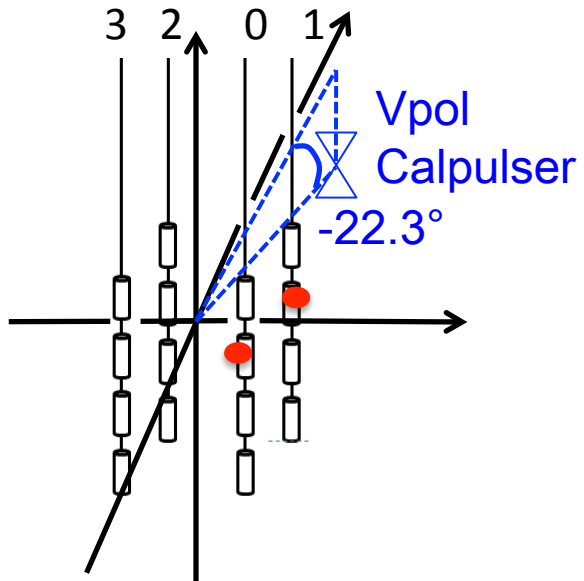


## Cross Correlation

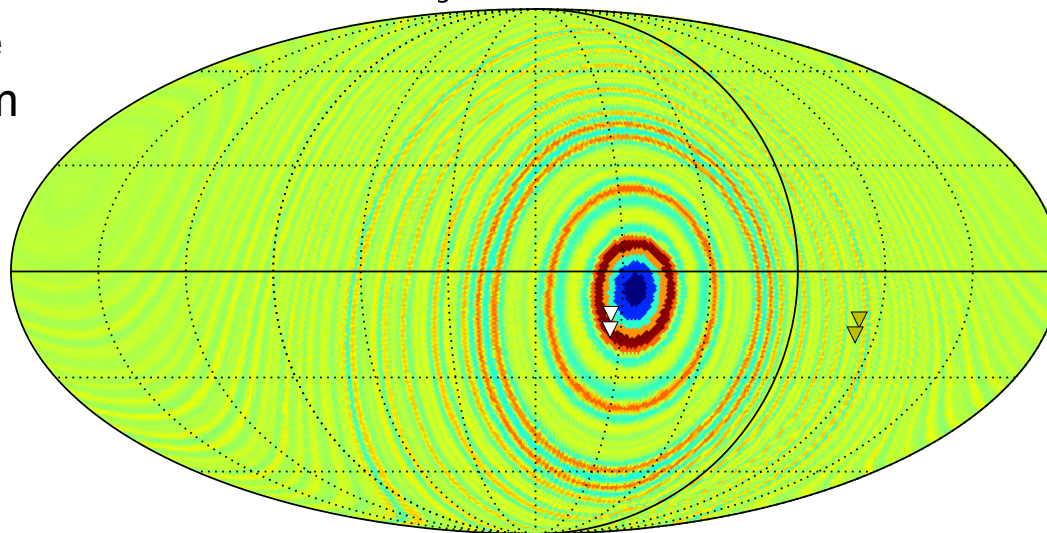
$$P_{\Sigma}(\hat{r}) = \sum_{i=1}^{N_A} P_i + \frac{1}{Z_L T} \sum_{i=1}^{N_A} \sum_{j \neq i}^{N_A} v_i \otimes v_j(\hat{r}), \quad P_i = \frac{1}{Z_L T} \int_0^T dt \cdot v_i^2(t)$$

  
 Directional information  
 “Coherence”

# Cross Correlation (A3 Calpulsar)

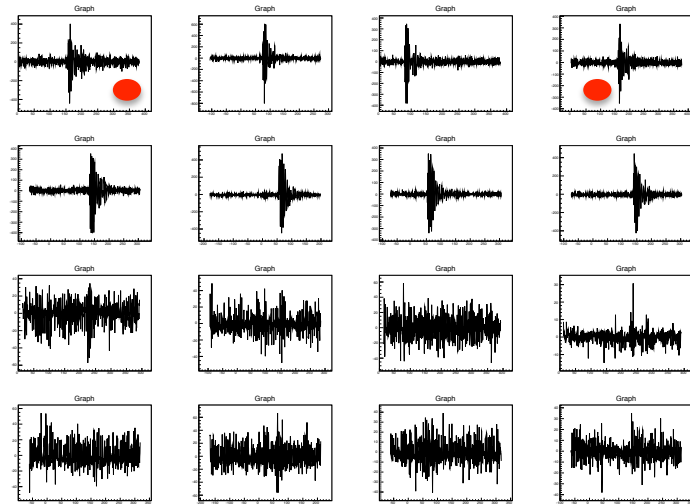
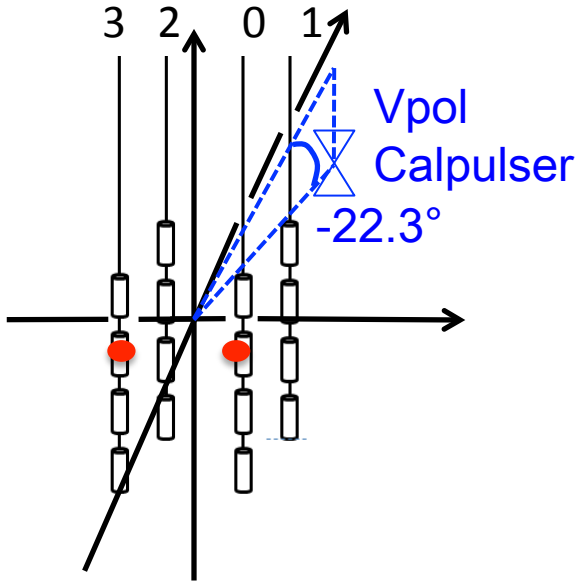


Assume true distance 42m

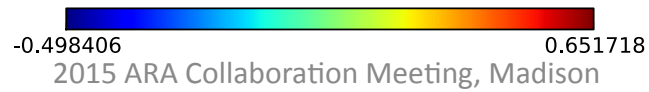
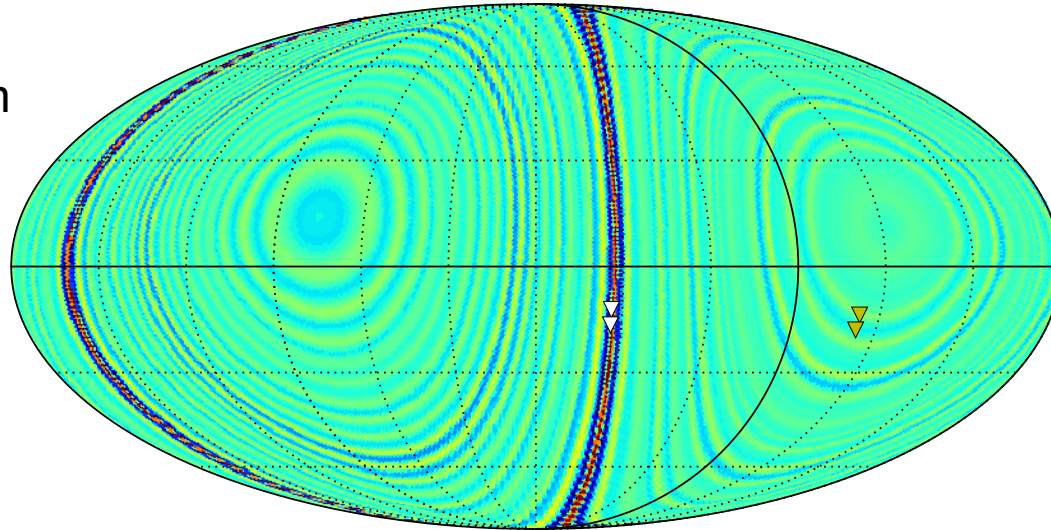


-0.366728 0.271287  
2015 ARA Collaboration Meeting, Madison

# Cross Correlation (A3 Calpulsar)

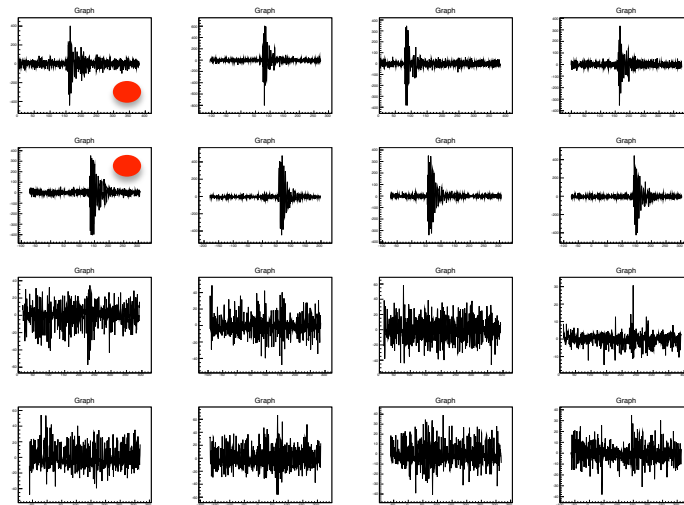
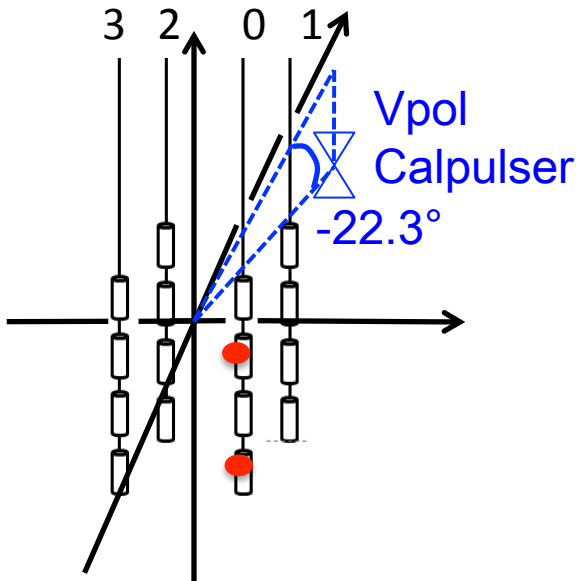


Assume true distance 42m

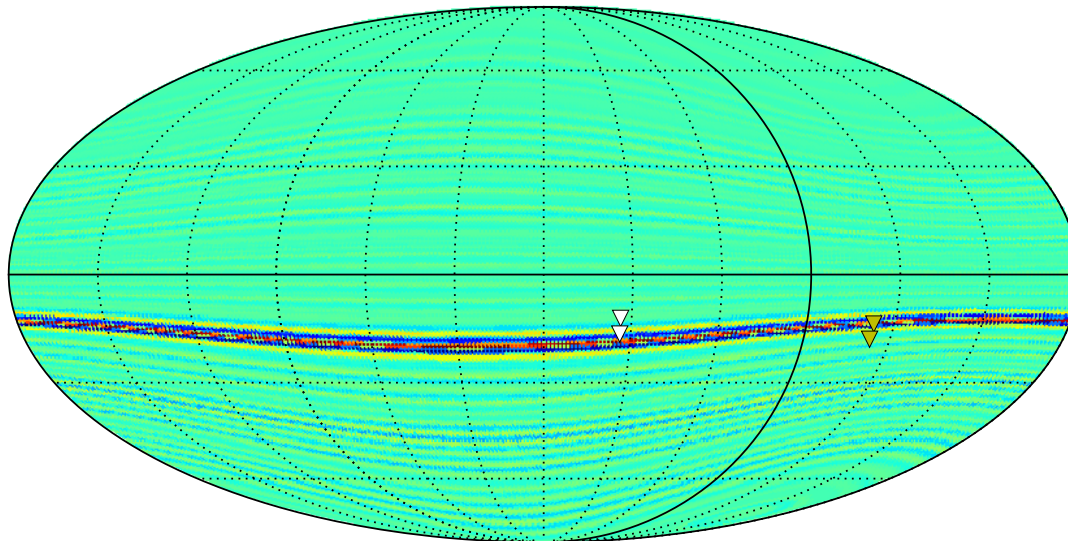




# Cross Correlation (A3 Calpulser)

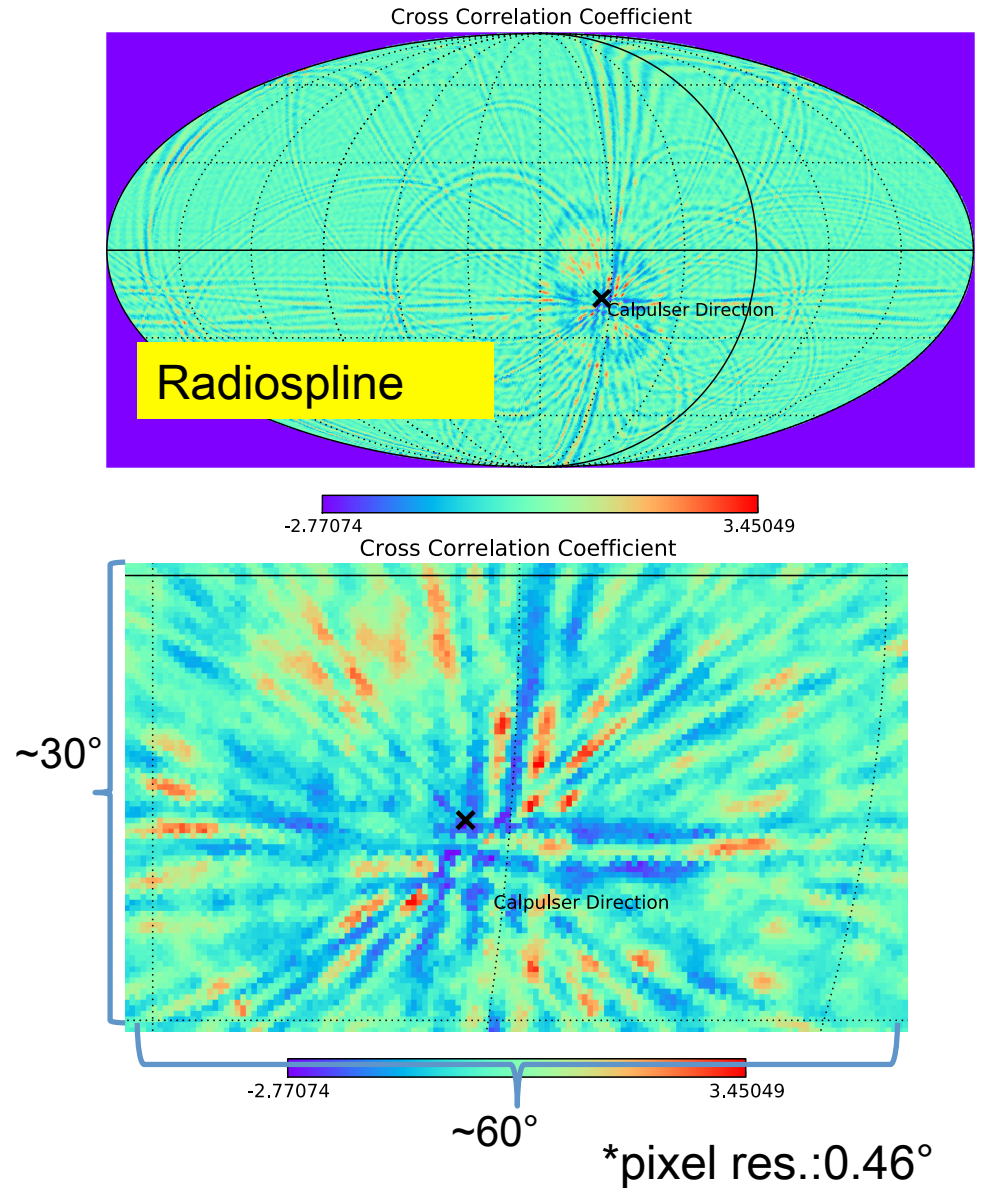


Assume true distance 42m



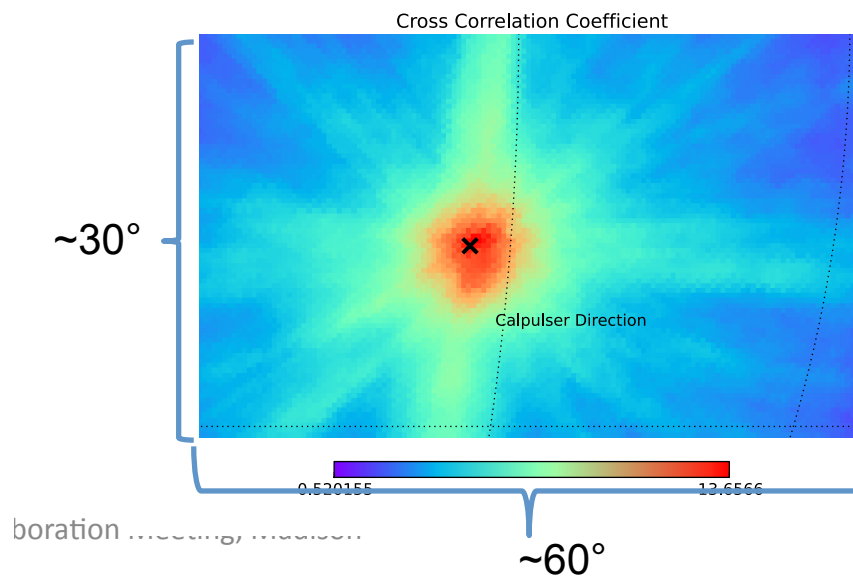
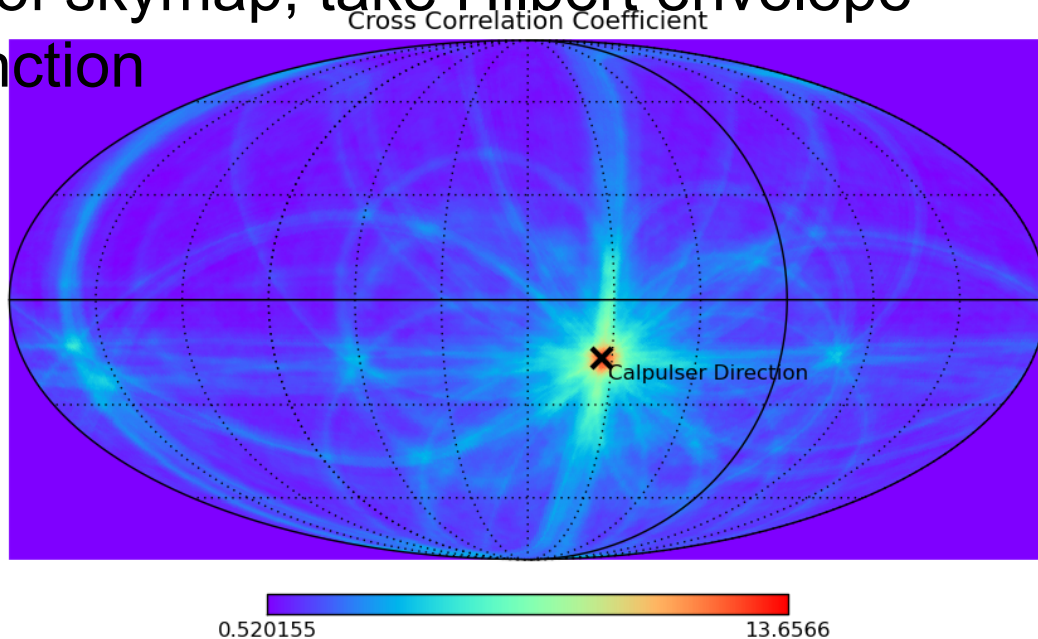
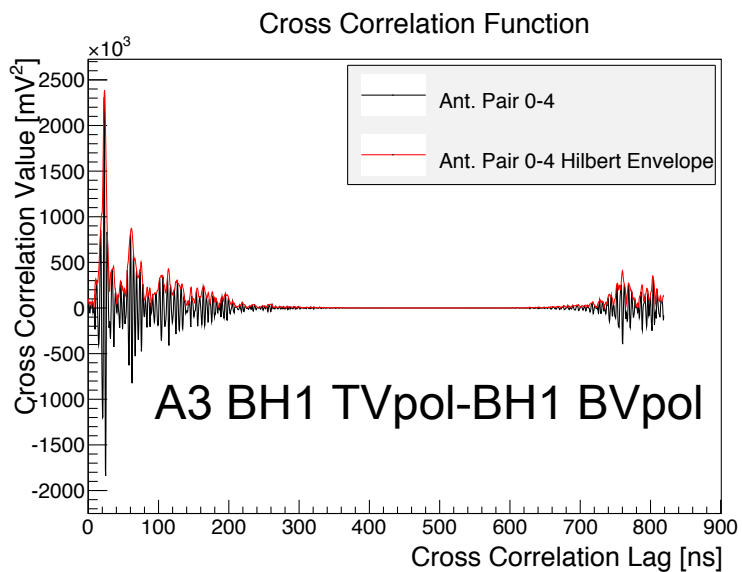
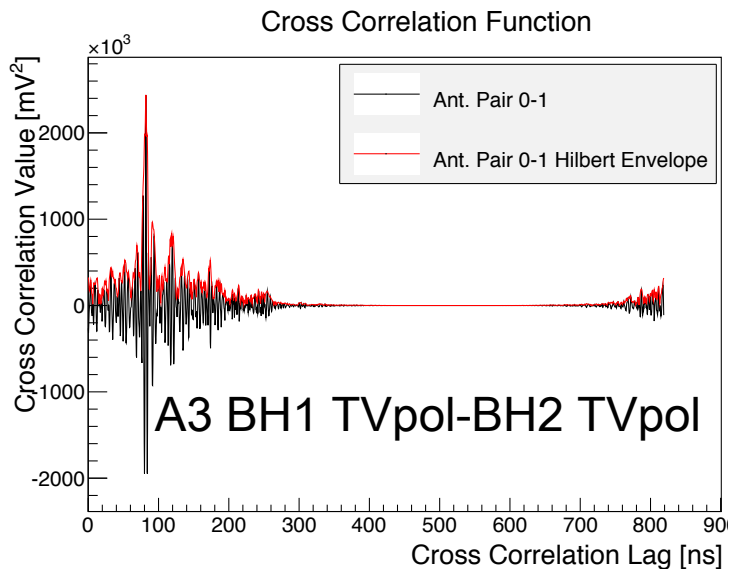
# Full Cross Correlation Skymap

- All 8 Vpol channels used
- ARA3 Vpol calpulers reconstruction assuming 42m distance (=true distance)
- 2013 run673 cal event #2



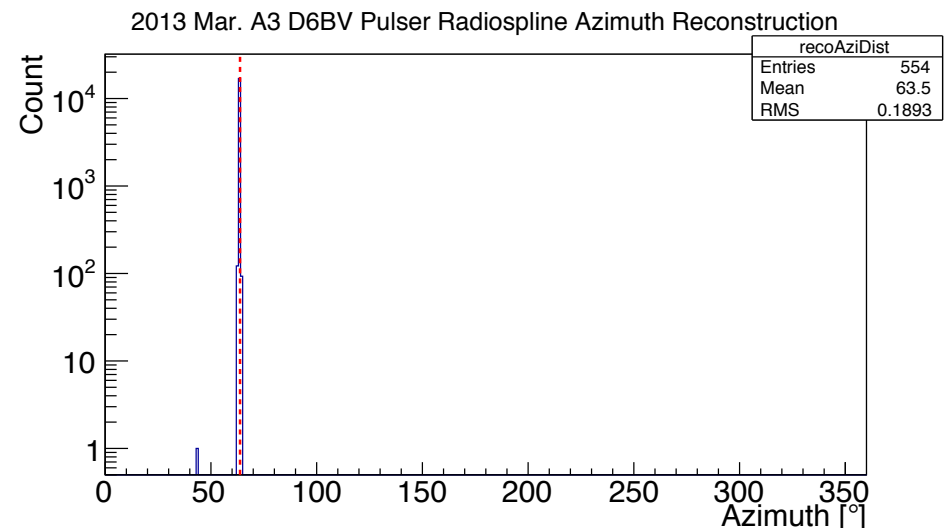
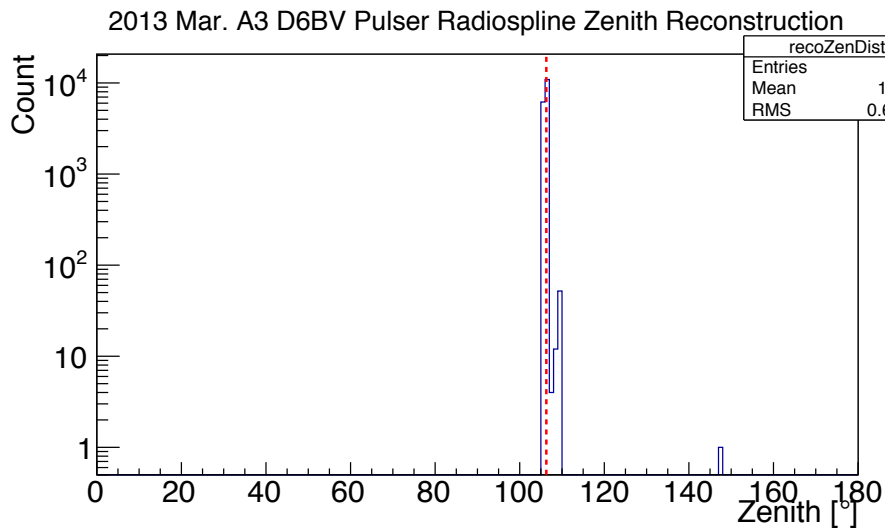
# Hilbert Envelope

- To remove graininess of skymap, take Hilbert envelope of cross-correlation function



# Calpulser Reconstruction

- A3 2013 March filtered events –D6BV pulser
- Calpulser distance fixed as known value (42m)



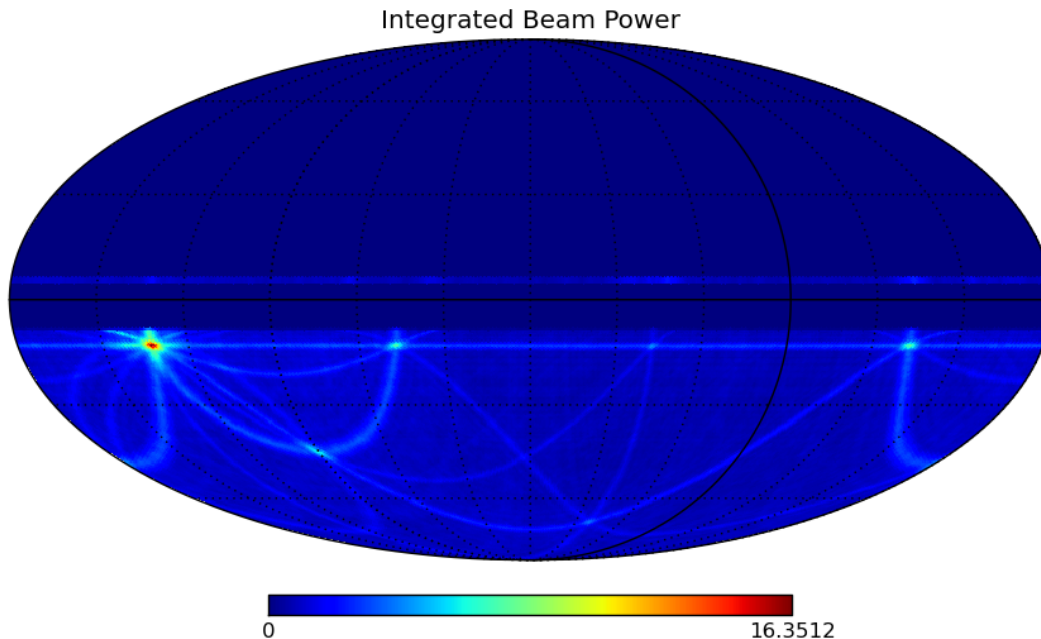
**True Pulser Distance: 42m**

	True	Reco Mean	Diff.
Zenith	106.3°	106.2°	-0.1°
Azimuth	63.8°	63.5°	-0.3°

# Angular Reconstruction

- Dataset:  $10^{18}$ eV neutrinos vertices randomly distributed up to 5km horizontal distance from station
- Trigger: default 3/8 multiplicity trigger
- Only allow single (1<sup>st</sup>) raytrace solution in simulation – no reflection!
- As starting point, place stringent event selection criteria
  - $N_{\text{chnl}}(V_{\text{peak}} > 5\sigma_{\text{noise}}) \geq 4$
  - Exclude saturated channels
- Assume simulation-true vertex distances in reconstruction

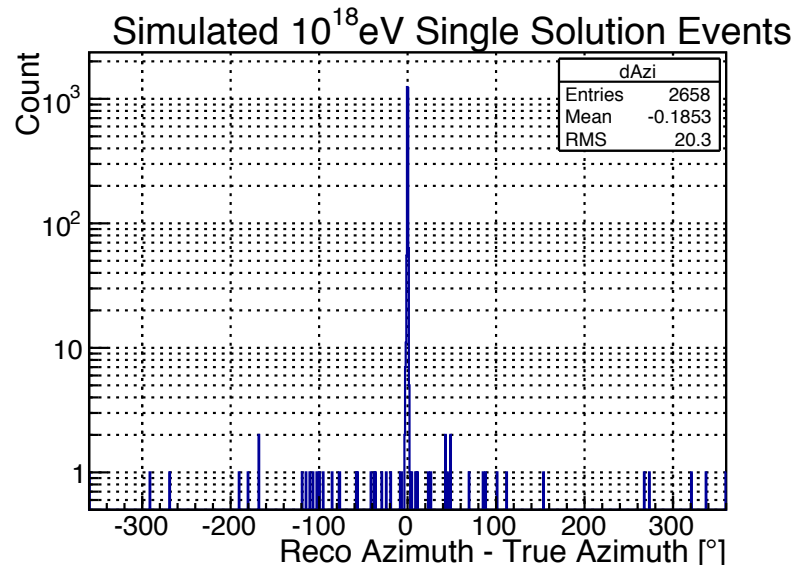
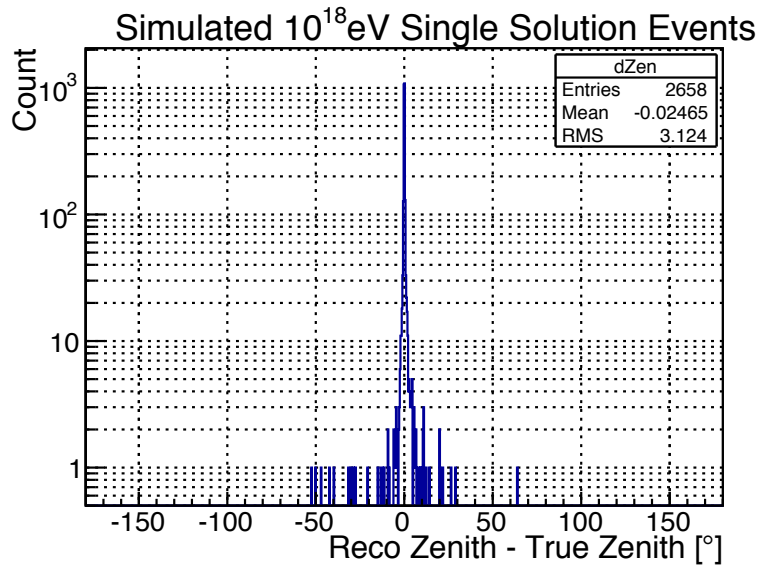
## Example event



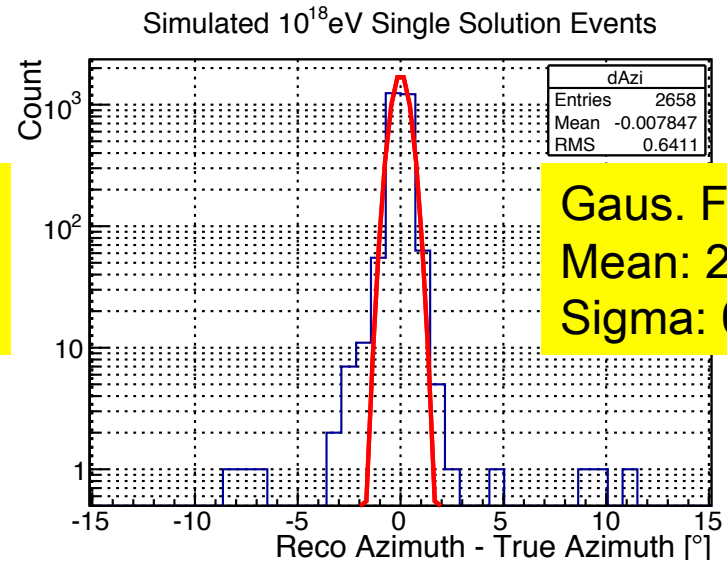
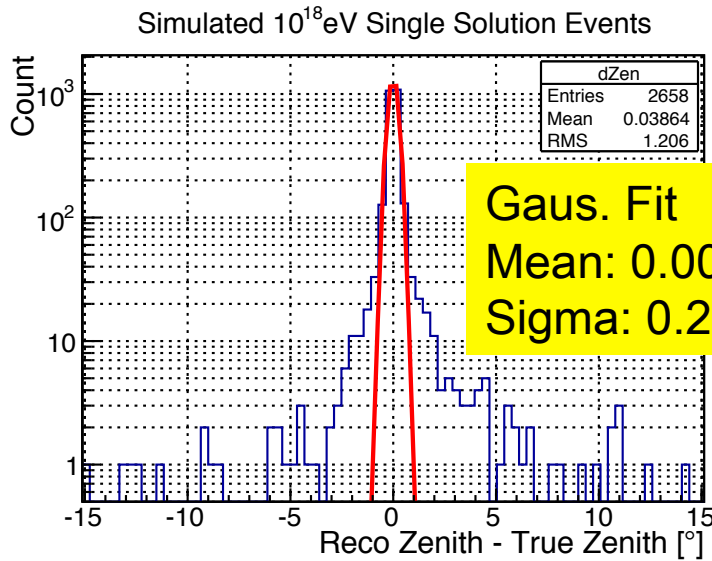
True Vertex Distance: 2581m

	True	Reco	Diff.
Zenith	102.74°	102.94°	0.2°
Azimuth	222.69°	222.89°	0.2°

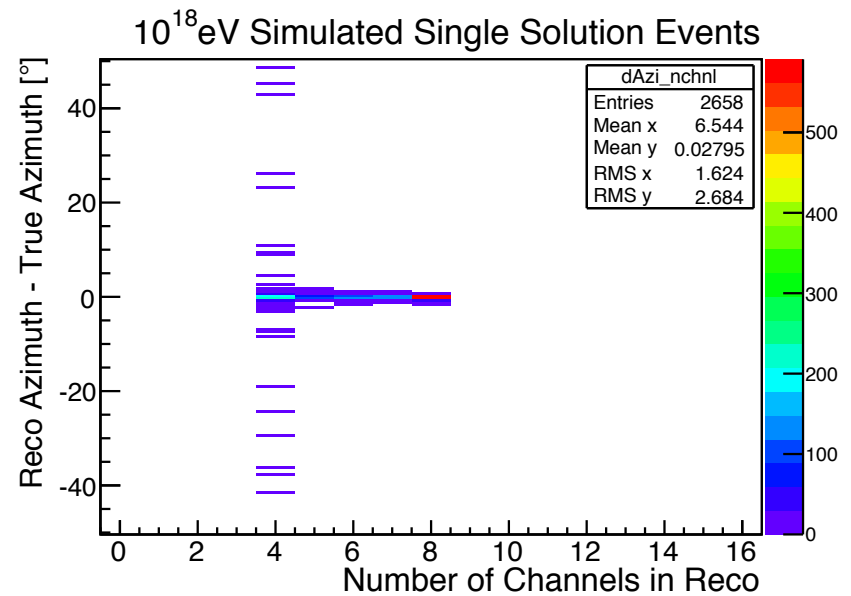
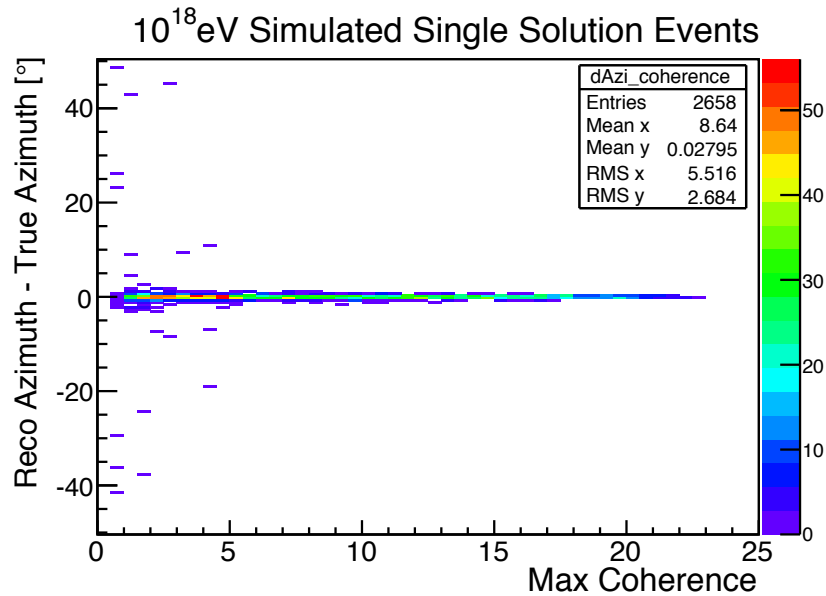
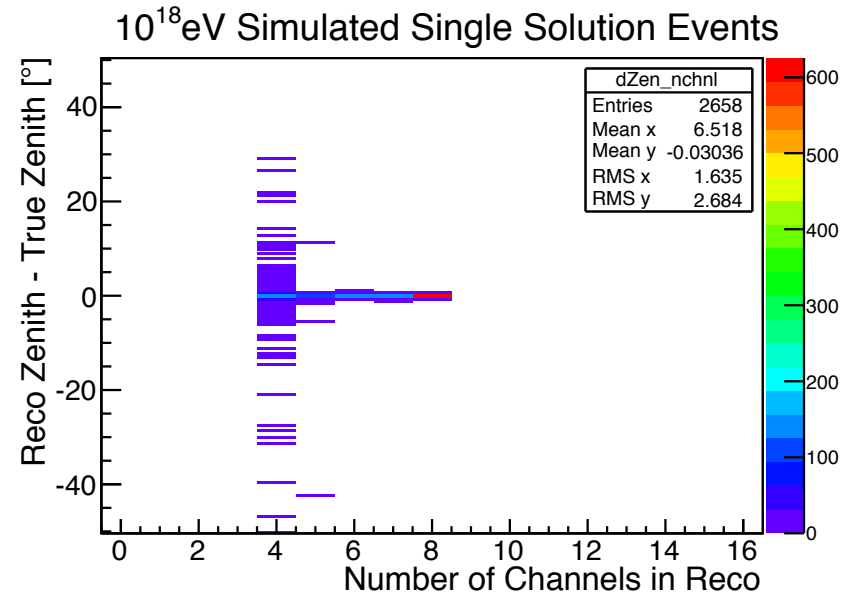
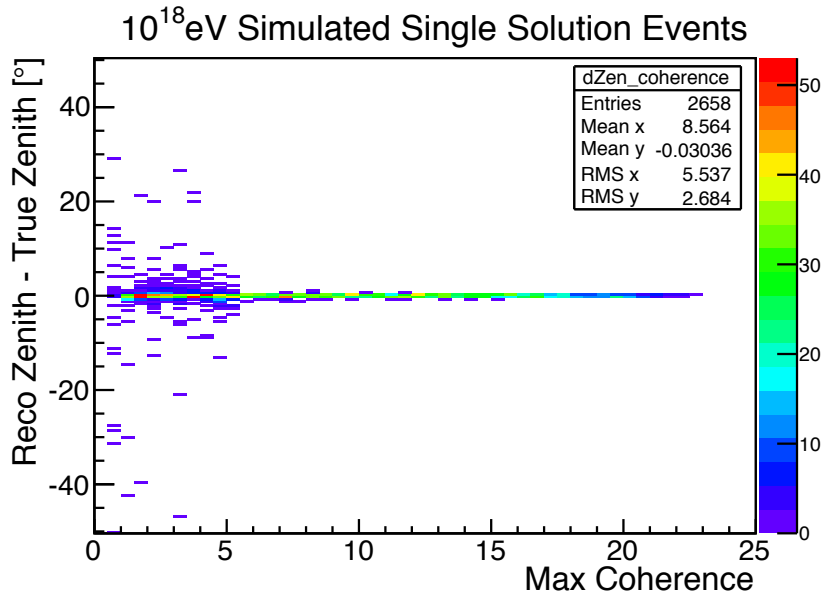
# Angular reconstruction



Zoom  
-15~15°



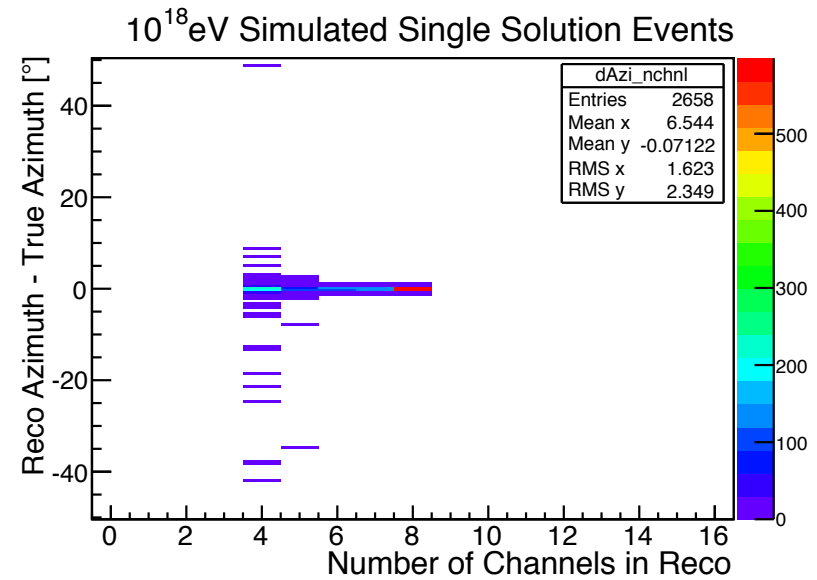
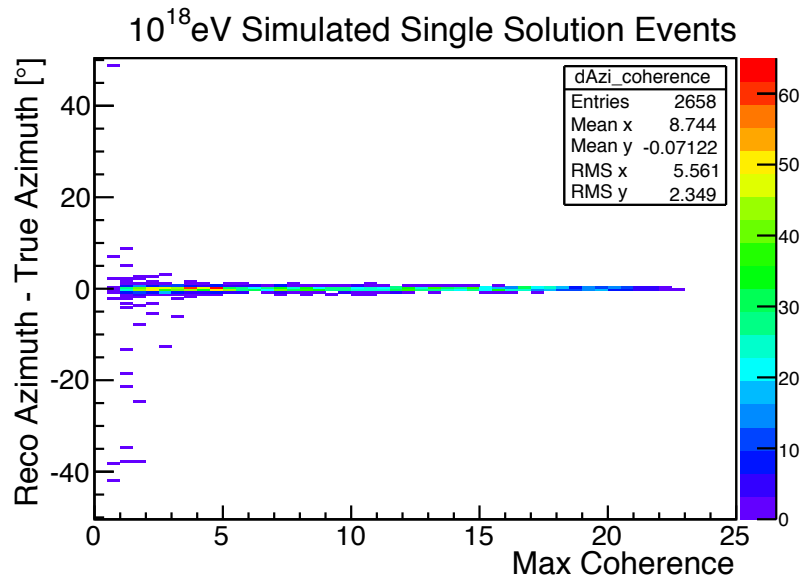
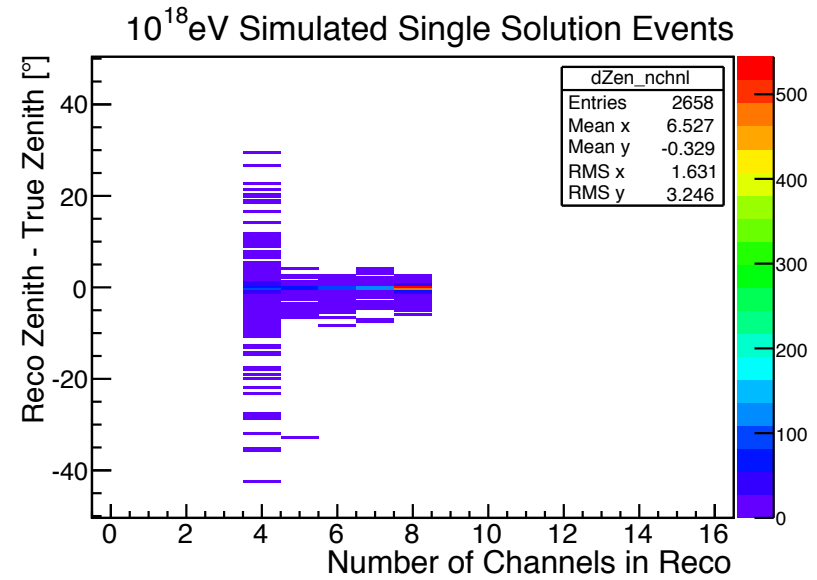
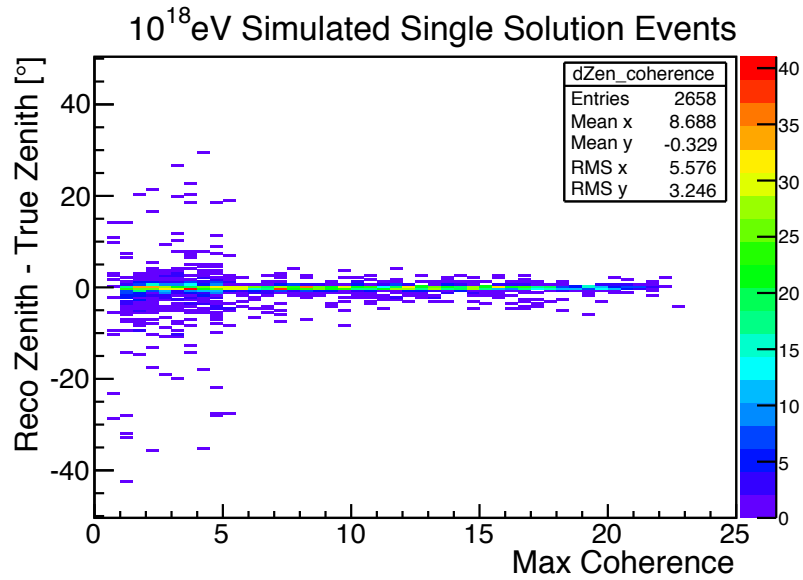
# Angular reconstruction



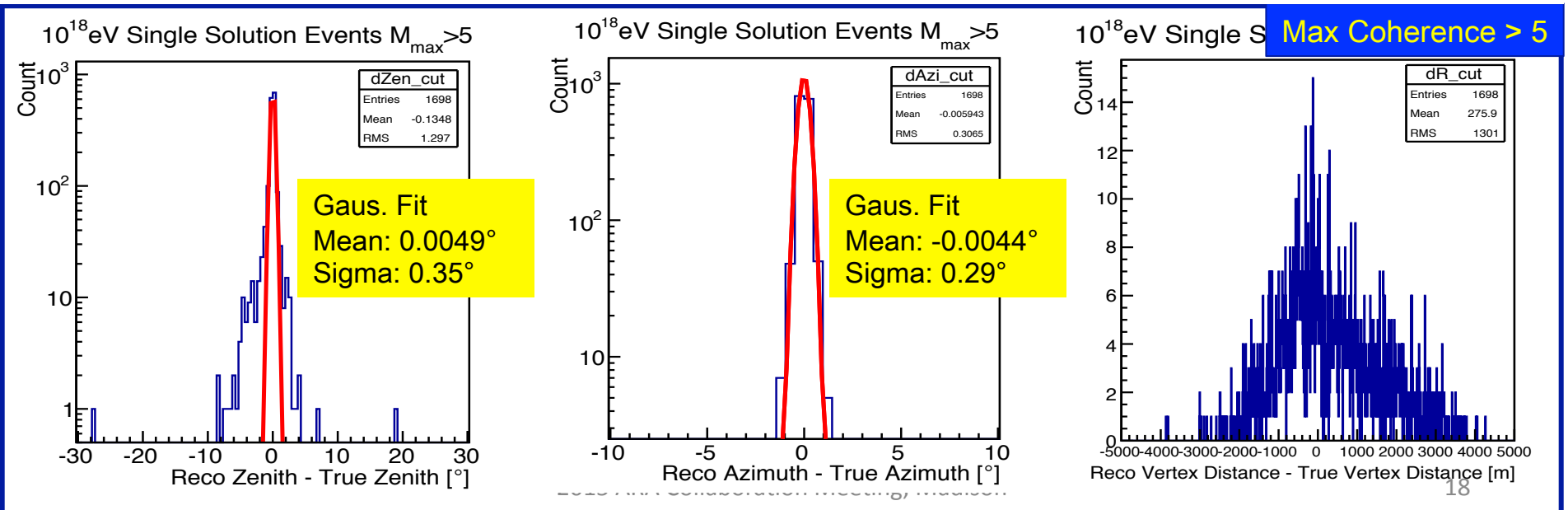
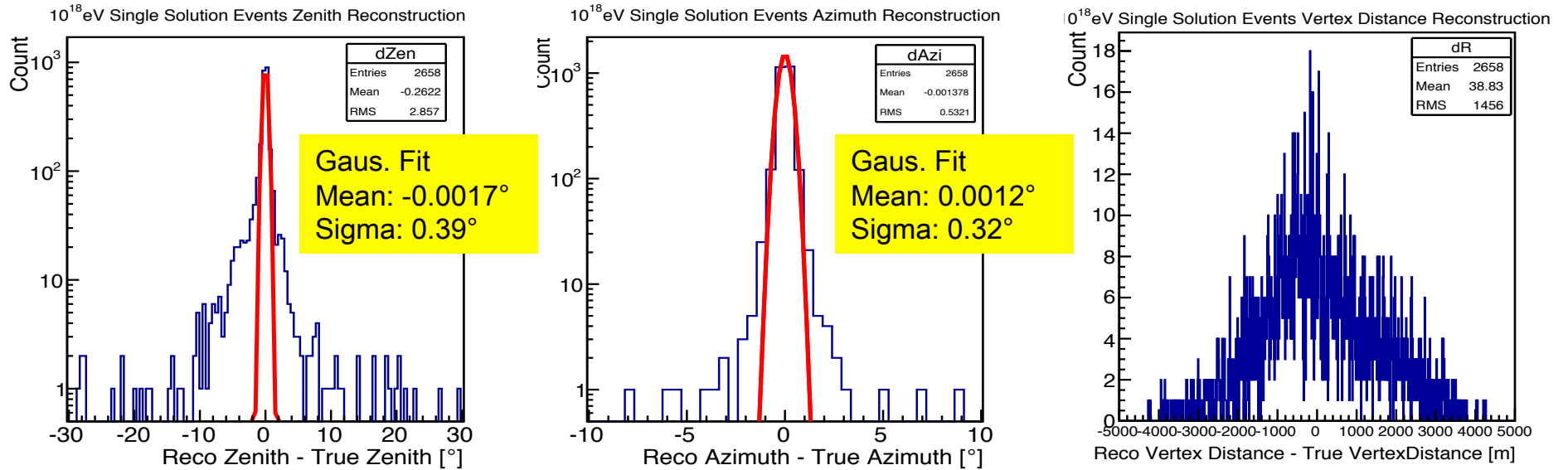




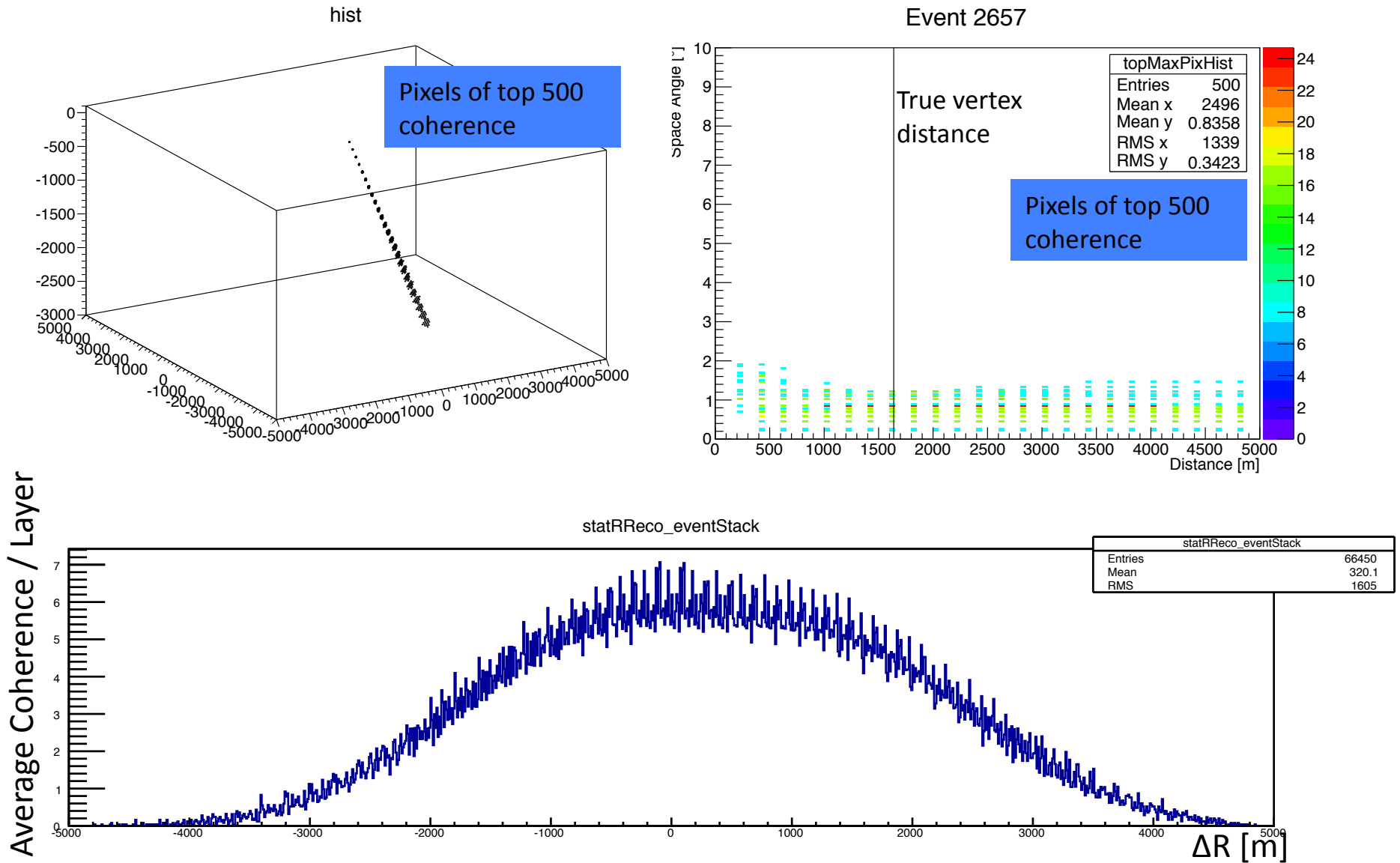
# Angular and Vertex Distance Reconstruction



# Angular and Vertex Distance Reconstruction



# Vertex Distance Reconstruction



Stacked  $\Delta R$  distribution for averaged coherence value on each layer

# Summary & Future Works

- An interferometric reconstruction framework with raytrace timing table was built. Vpol reconstruction of cal pulsars successful.
  - Framework tested with simulation. Distance as free parameter:  
Coherence cut ✘:  $\sigma_{\theta}:2.9^{\circ}$   $\sigma_{\phi}:0.53^{\circ}$   
Coherence cut ✔:  $\sigma_{\theta}:1.3^{\circ}$   $\sigma_{\phi}:0.31^{\circ}$
  - Vertex distance resolution is poor, possibly limited by our time resolution
- 
- Simulation with larger baseline could verify the above hypothesis, and point to future designs
  - Develop thermal / CW rejection methods

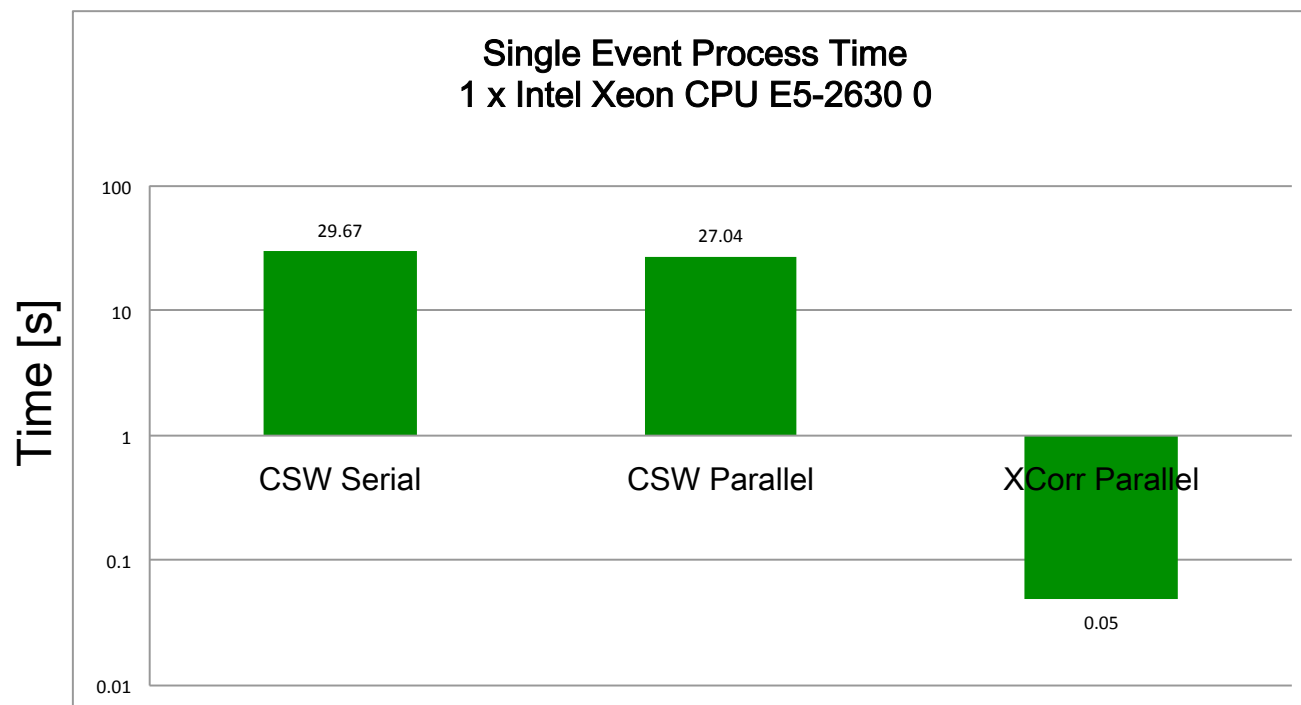
# Backup

# Event Vertex Reconstruction

- Direction reconstruction –  
Identify backgrounds/anthropogenic sources
- Distance reconstruction –  
Neutrino energy
- Polarization measurement –  
Neutrino direction

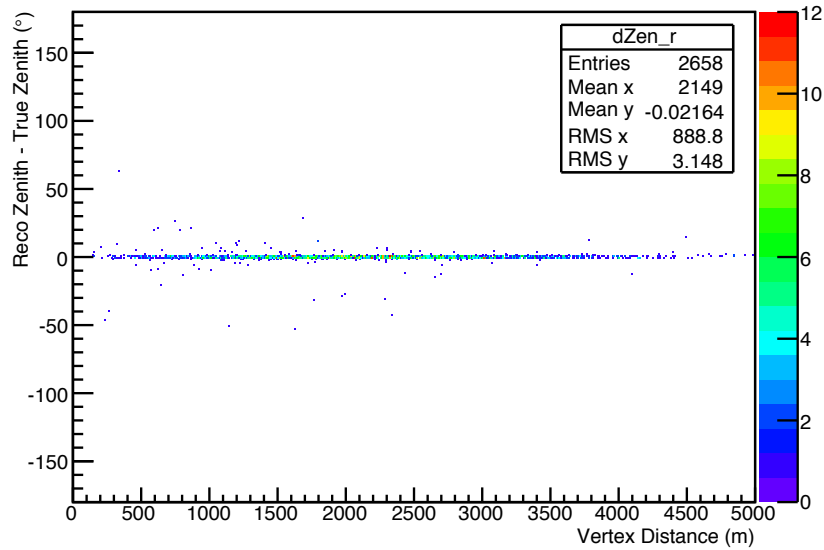
# Parallel Implementation

- OpenCL parallelism facilitates fast implementation of reconstruction
  - waveform-shifting, FFT, cross-correlating, waveform-summing
- Multiple GPU/CPU can be combined to maximize speed boost

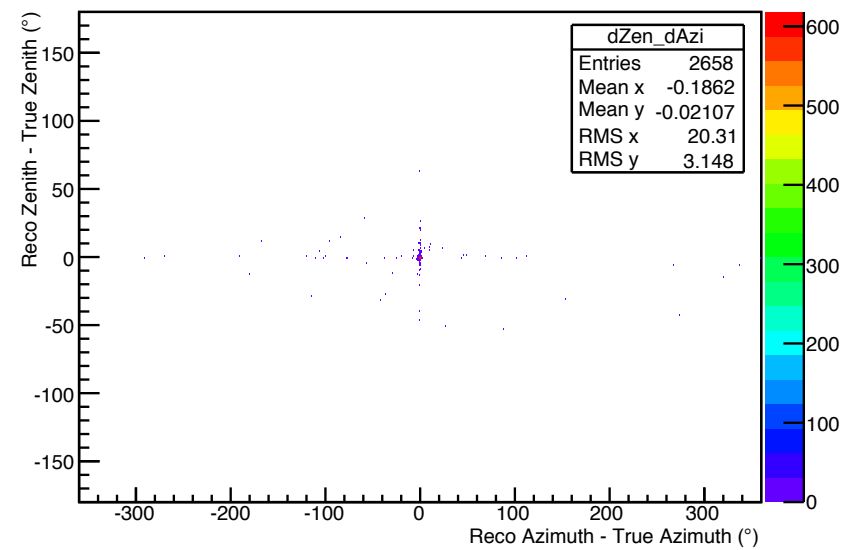


# 2D Angular Reconstruction

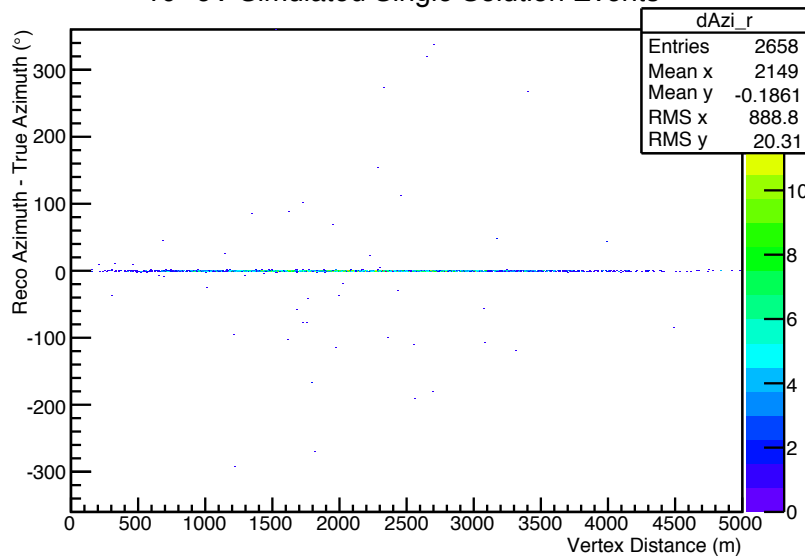
$10^{18}$  eV Simulated Single Solution Events



$10^{18}$  eV Simulated Single Solution Events



$10^{18}$  eV Simulated Single Solution Events





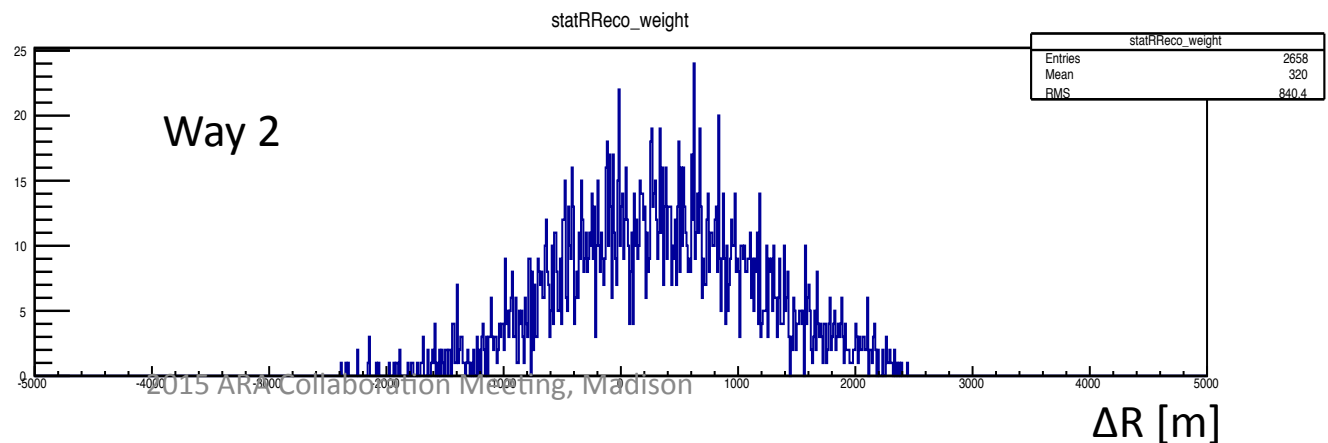
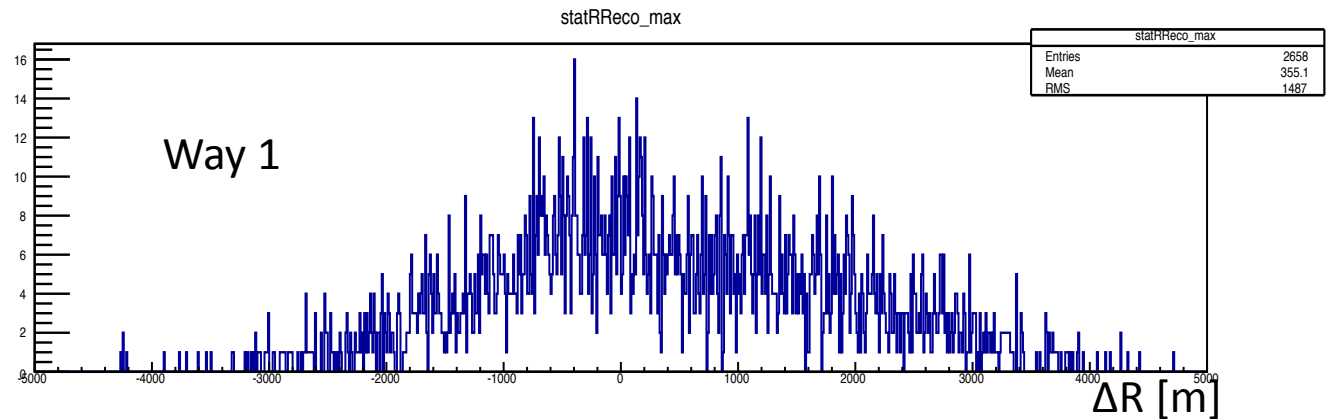
# Vertex Distance Reconstruction

*Question: with the ~flat distribution of coherence values around true vertex distance, can we use the distribution to get closer to the true distance?*

Simple way 1: treat the layer with the maximum averaged coherence as giving reconstructed distance

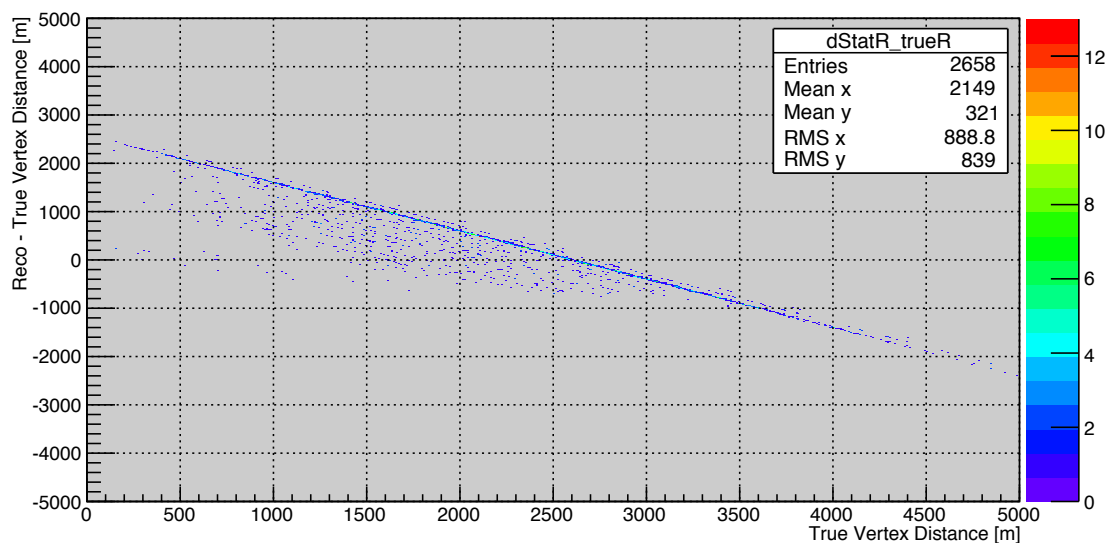
Simple way 2: use averaged coherence value on all layers to compute a weighted reconstructed distance

$$\bar{r} = \frac{\sum_{i=1}^{N_{layer}} r_i M_i}{\sum_{i=1}^{N_{layer}} M_i}$$



# Weighted Distance

10<sup>18</sup>eV Single Solution Events [Weighted Reco Radius]



Reconstruction tends to give result at the center of the distance phase space [20-5000] ->2600

10<sup>18</sup>eV Single Solution Events  $\Delta R$  [Weighted Reco Radius]

