



Stony Brook University

HEX Cascades

240m

300m

360m

NSString >= 3

Hans Niederhausen
Stony Brook University

01/28/2015

Cascade Reconstruction in HEX 360m vs 240m

Software + Codes:

- IceRec.release 04-08-00
- use Monopod (calibration scheme, calibration ..) from Jakob's analysis

`SVN/sandbox/jvansanten/projects/2013/monopod_recoverf/segments/BaseProc.py`

`SVN/sandbox/jvansanten/projects/2013/monopod_recoverf/segments/Calibration.py`

`SVN/sandbox/jvansanten/CascadeL3_IC79/trunk/resources/level4/redo_monopod_again.py`

Reconstruct 3 Cascade Samples:

Visible Energy: **1-50 TeV**, **50-200 TeV** (near threshold)
and **1-10 PeV** (target signal)

Containment requirements:

(well-contained cascades, not in dustlayer)

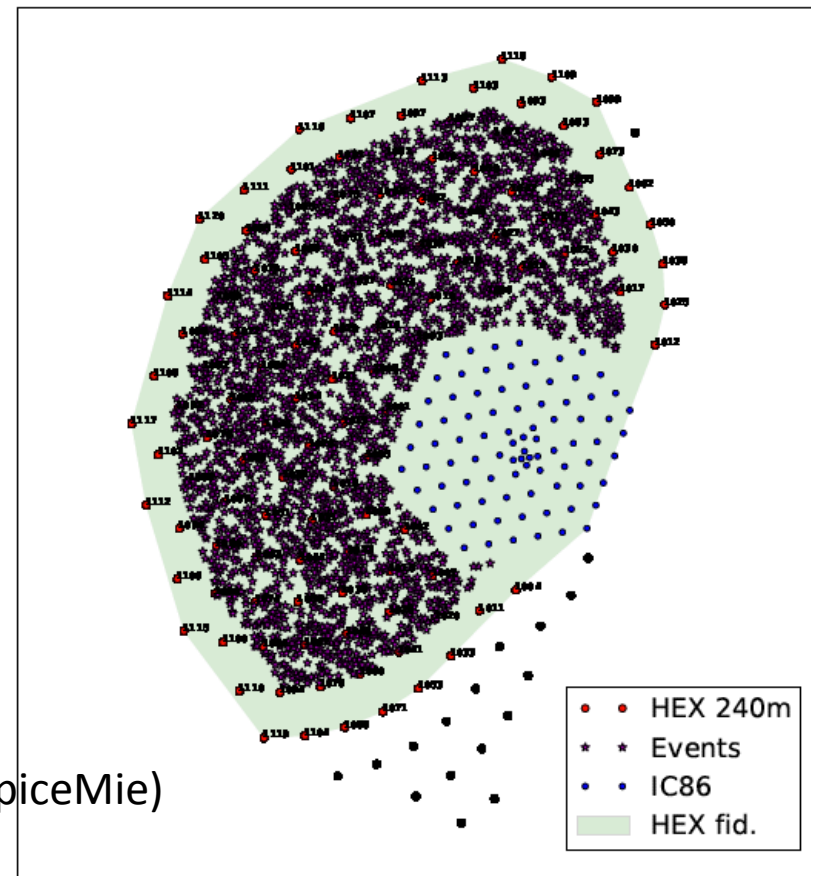
$\text{HEX_XYscale} < 0.9$

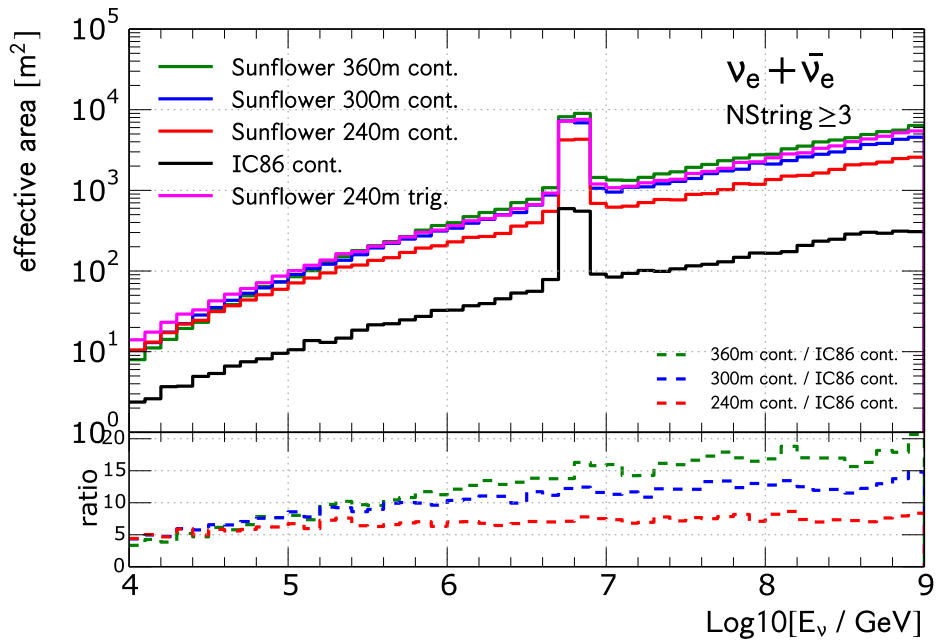
$\text{IC_XYScale} < 1.2$

$-450\text{m} < Z < -150\text{m}$ or $0\text{m} < Z < 450\text{m}$

Nstring ≥ 3.0

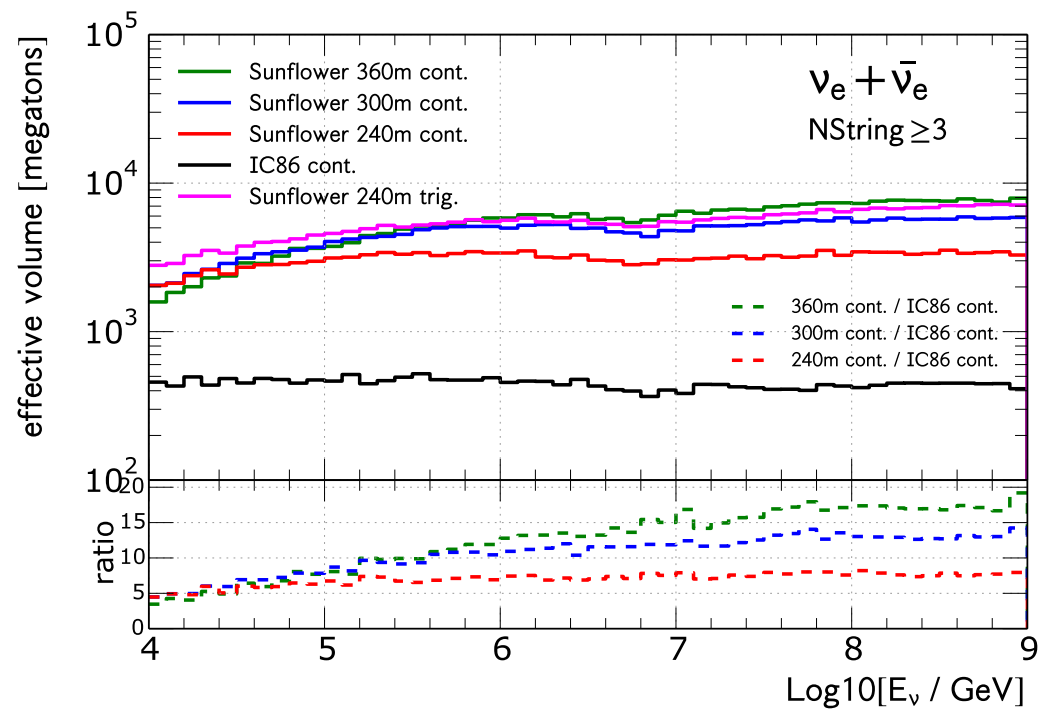
no icemodel systematics here. Use same icemodel (SpiceMie)
during reconstruction that was used for simulation.





Area

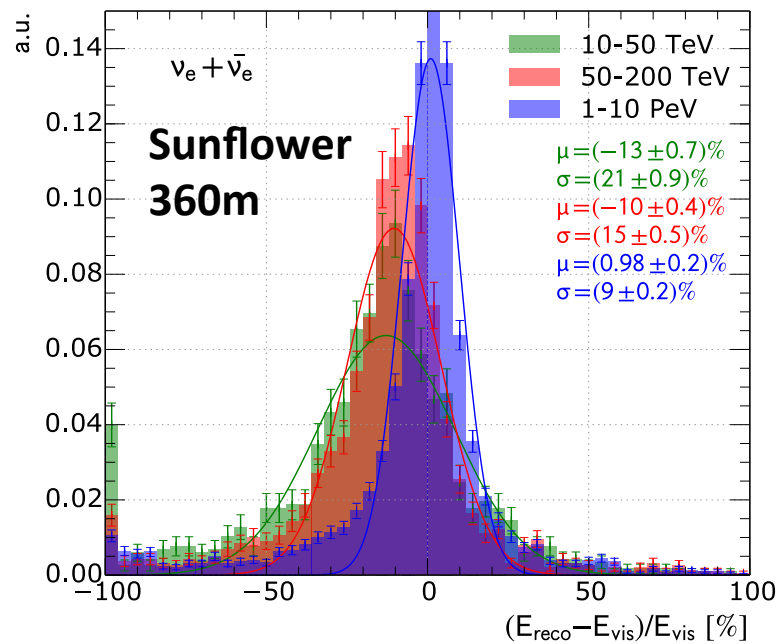
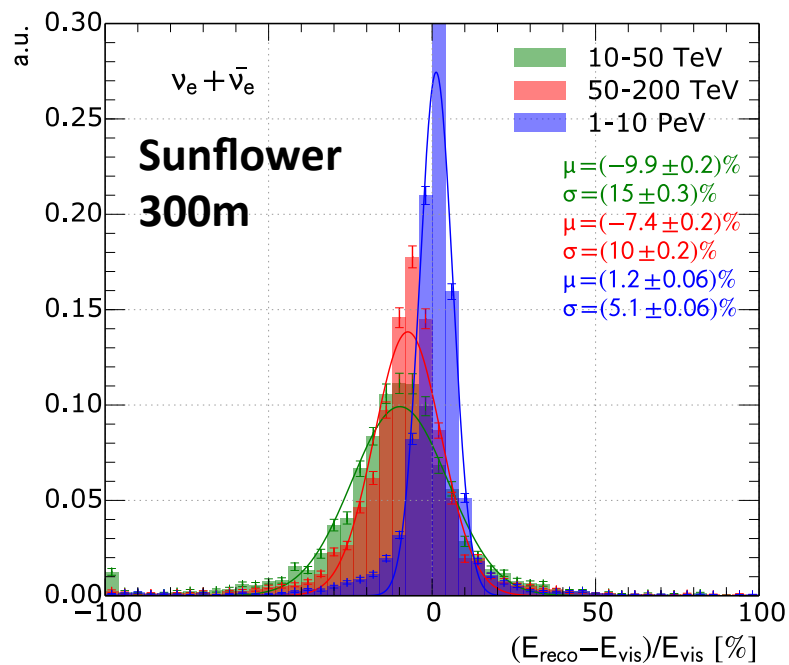
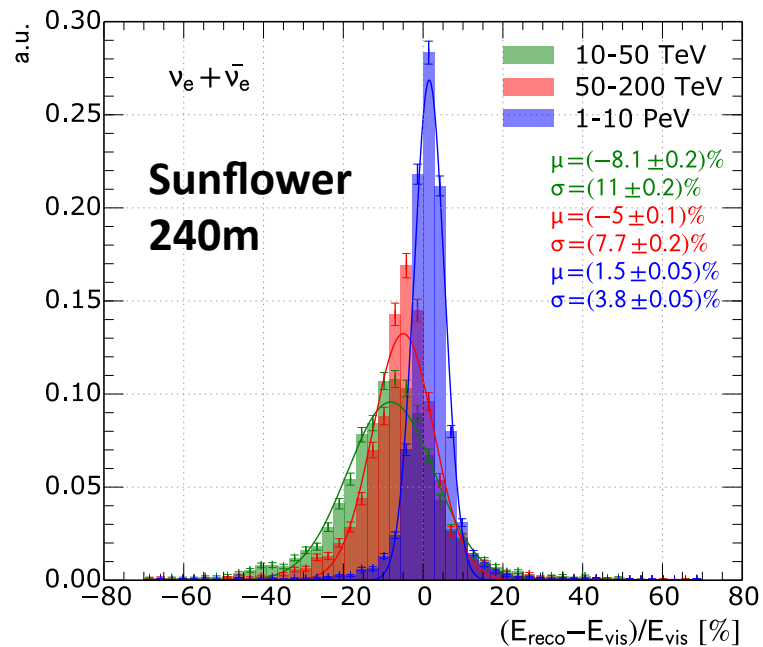
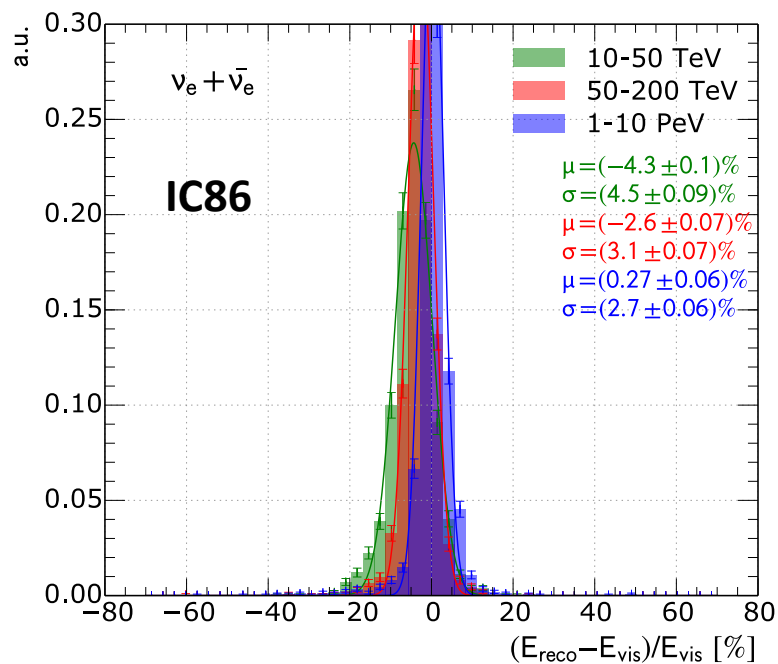
Volume

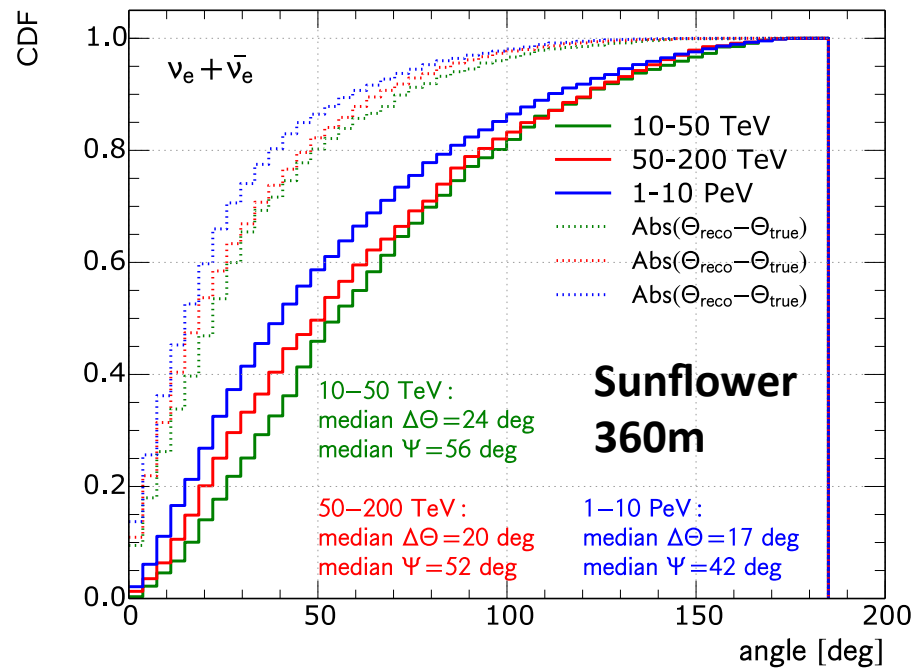
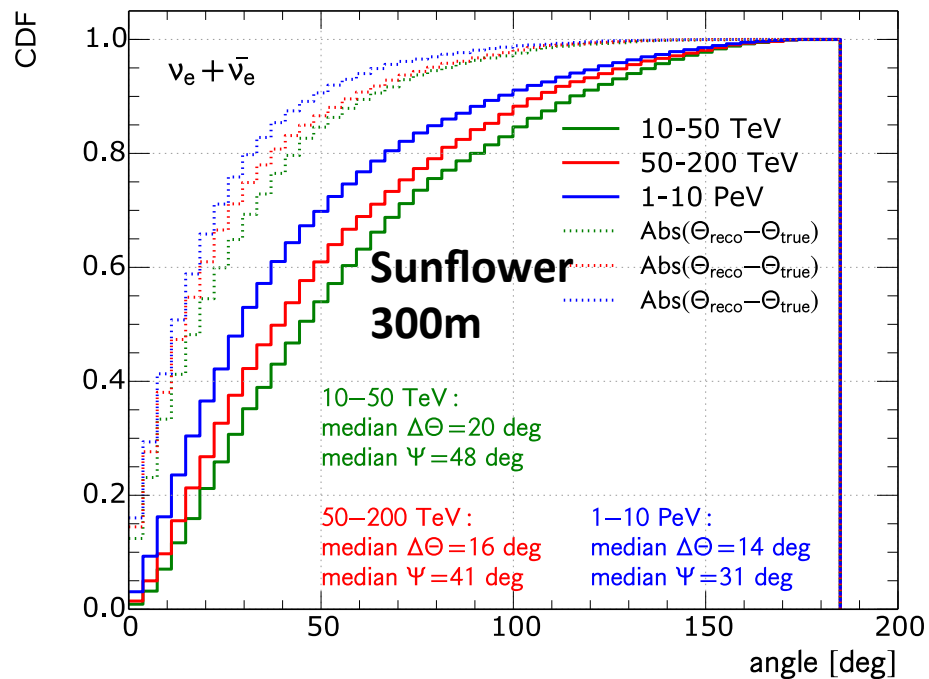
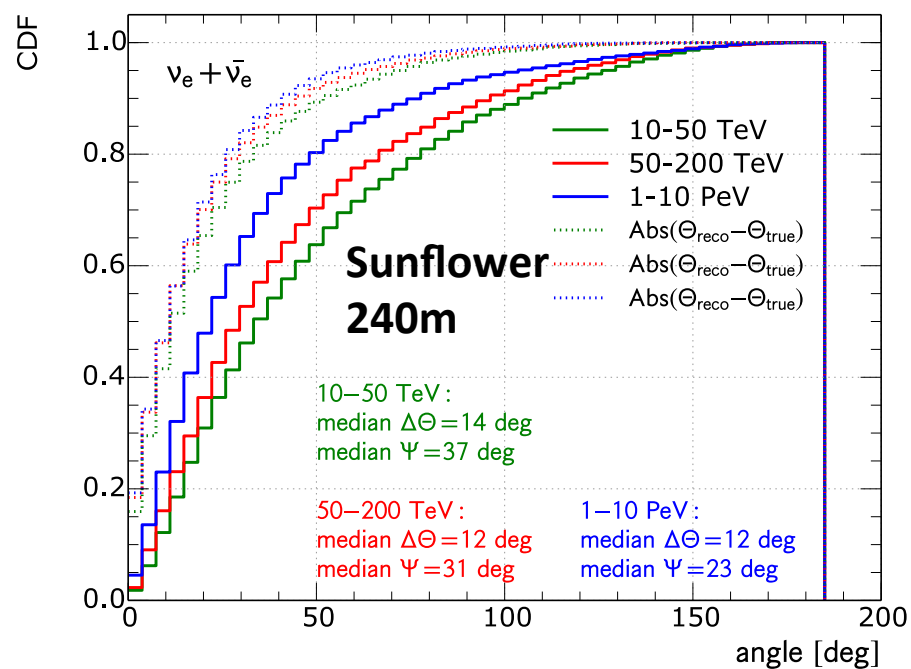
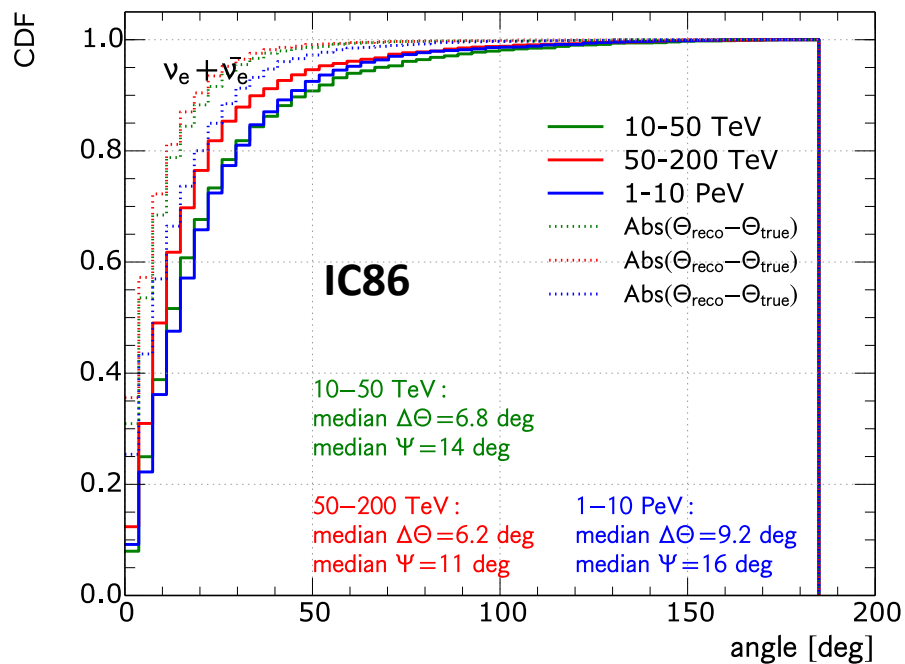


10 – 50 TeV (μ σ)	Sunflower 240m	Sunflower 300m	Sunflower 360m	IceCube 86
Energy [%]	(-8.1 11)	(-9.9 15)	(-13 21)*	(-4.3 4.5)
x [m]	(0.0 9.3)	(0.2 14)	(0.6 18)	(0.1 2.9)
y [m]	(-0.1 9.6)	(0.1 14)	(0.2 18)	(0.0 2.9)
z [m]	(-0.2 5.2)	(-0.2 8.6)	(-0.2 12)	(0.1 1.9)
Zenith [deg]	14	20	24	7
Opening Angle [deg]	37	48	56	14

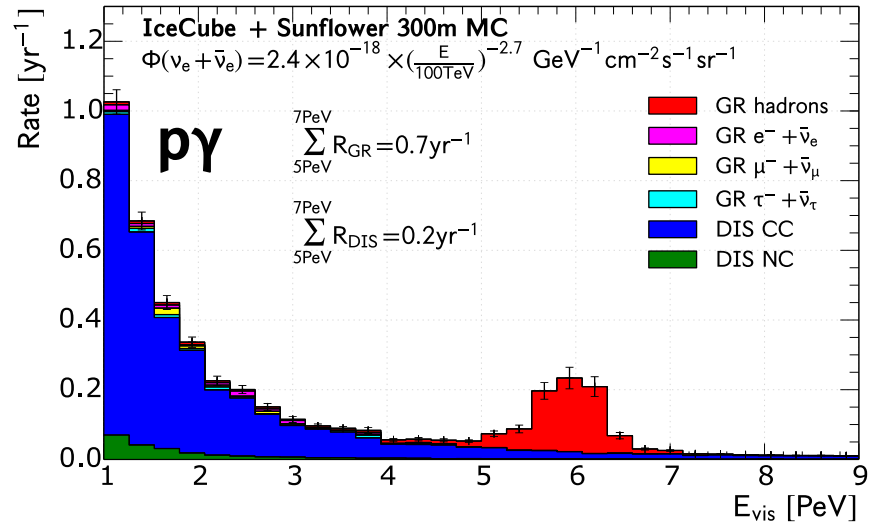
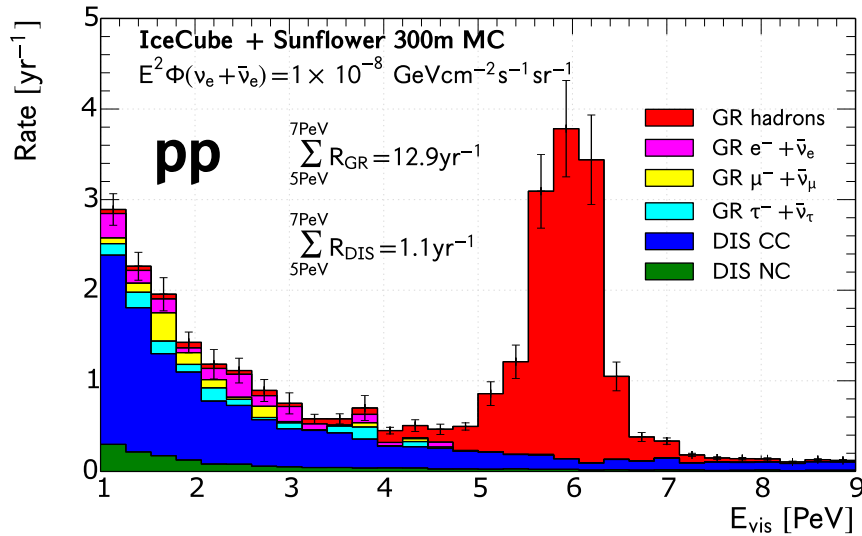
50 - 200 TeV (μ σ)	Sunflower 240m	Sunflower 300m	Sunflower 360m	IceCube 86
Energy [%]	(-5.0 7.7)	(-7.4 10)	(-10 15)	(-2.6 3.1)
x [m]	(-0.1 7.4)	(0.0 11)	(0.8 14)	(-0.1 2.0)
y [m]	(0.3 7.4)	(0.0 10)	(0.2 14)	(-0.1 2.0)
z [m]	(0.2 3.9)	(0.2 6.2)	(0.2 10)	(0.0 3.3)
Zenith [deg]	12	16	20	6.2
Opening Angle [deg]	31	41	52	11

1 – 10 PeV (μ σ)	Sunflower 240m	Sunflower 300m	Sunflower 360m	IceCube 86
Energy [%]	(1.5 3.8)	(1.2 5.1)	(1.0 9.0)	(0.3 2.7)
x [m]	(-0.1 4.3)	(0.1 7.0)	(0.3 9.8)	(-0.1 2.8)
y [m]	(-0.1 4.5)	(0.0 6.6)	(-0.2 10)	(0.1 2.6)
z [m]	(0.1 4.3)	(0.0 5.2)	(-0.1 6)	(0.0 3.3)
Zenith [deg]	12	14	17	9.2
Opening Angle [deg]	23	31	42	16





GR in 300m detector



Flux (pp source)	Sunflower 240m Rate [yr ⁻¹] (IC86)	Sunflower 300m Rate [yr ⁻¹]	Sunflower 360m Rate [yr ⁻¹]
std. E-2 g=2.0	7.2 (0.88)	12.9	16
HESE 3yr g=2.3	3.1 (0.38)	5.6	6.8
MLB g=2.7	0.9 (0.12)	1.7	2.1
Flux (pγ source)	Sunflower 240m	Sunflower 300m	Sunflower 360m
std. E-2	3.2 (0.39)	5.7	6.9
HESE 3yr	1.4 (0.17)	2.4	3.0
MLB	0.4 (0.05)	0.7	0.9

Φ_{ν_e} [GeV ⁻¹ cm ⁻² s ⁻¹ sr ⁻¹]	interaction type	pp source		p γ source	
		IC-86	HEX-240m	IC-86	HEX-240m
$1.0 \times 10^{-18}(E/100 \text{ TeV})^{-2.0}$	GR	0.88	7.2	0.39	3.2
	DIS	0.09	0.8	0.10	0.8
$1.5 \times 10^{-18}(E/100 \text{ TeV})^{-2.3}$	GR	0.38	3.1	0.17	1.4
	DIS	0.04	0.3	0.04	0.3
$2.4 \times 10^{-18}(E/100 \text{ TeV})^{-2.7}$	GR	0.12	0.9	0.05	0.4
	DIS	0.01	0.1	0.01	0.1

TABLE II. Expected number of contained neutrino-induced cascades [year⁻¹], for $5 \text{ PeV} < E_{vis} < 7 \text{ PeV}$.

Summary

- added Sunflower 300m and IC86@SMT8 to reconstruction performance comparison
- reconstruction of 300m not much worse compared to 240m (e.g. energy res. still much smaller than typical bin size in energy histograms)
- 300m has about $>\sim 50\%$ larger effective area/volume than 240m

- Depending on needs on directional resolution (atm. veto) and assuming background rejection above 100 TeV to be easy, **one may conclude that 300m is preferable**

- Location of files: /data/ana/HEX/NuE/floral_*.m/ (* = {240, 300, 360})