

The Astrophysical Multimessenger  
Observatory Network



# AMON – TRANSITION TO REALTIME OPERATIONS

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(for the AMON team)

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- Founded and hosted at Penn State
  - Internal initial funding
- Official NSF funded project as of 2014

## AMON development and advisory team

### Penn State

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 A. Keivani<sup>1,3</sup>, P. Mészáros<sup>1,2,3</sup>, C. Messick<sup>1</sup>, M. Mostafá<sup>1,3</sup>, C. Hanna<sup>1,3</sup>,  
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<sup>4</sup>Computer Science and Engineering

<sup>5</sup>Institute for CyberScience

### Others

S. Barthelmy<sup>1</sup>, I. Bartos<sup>2</sup>, F. Feroz<sup>3</sup>, M. Smith<sup>4</sup>, I. Taboada<sup>5</sup>

<sup>1</sup>NASA GSFC

<sup>2</sup>Columbia University, Dept of Physics

<sup>3</sup>Cambridge University

<sup>4</sup>NASA JPL

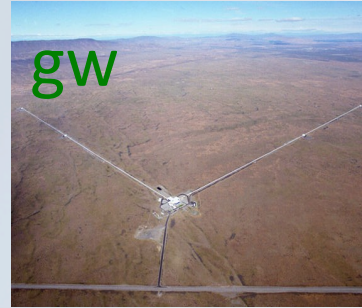
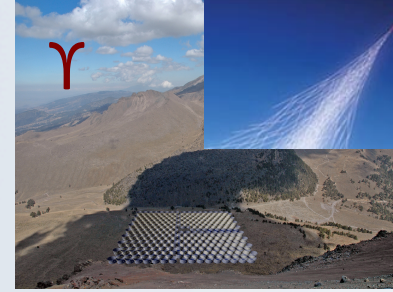
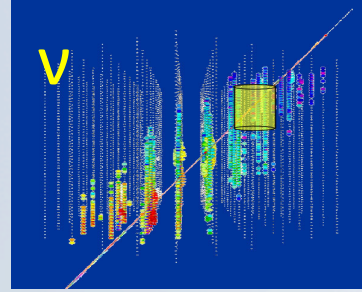
<sup>5</sup>Georgia Institute of Technology

# The AMON Idea

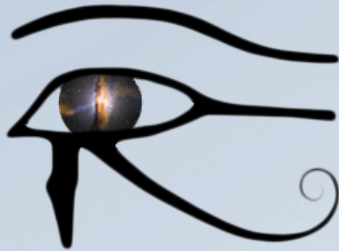
Use messenger particles of all four fundamental forces

## Triggering observatories

- Provide **sub-threshold** candidate events to AMON in real time



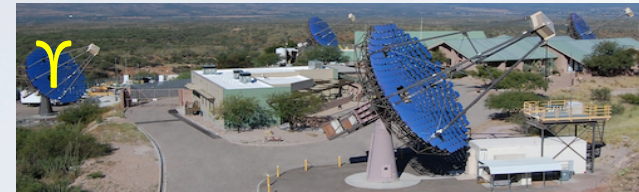
## AMON



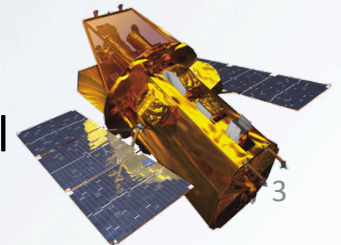
- Seeks **coincidences** in time and space
- Generates **alerts** - broadcast and archived
- Enables archival analyses

## Follow-up observatories

- respond to AMON alerts

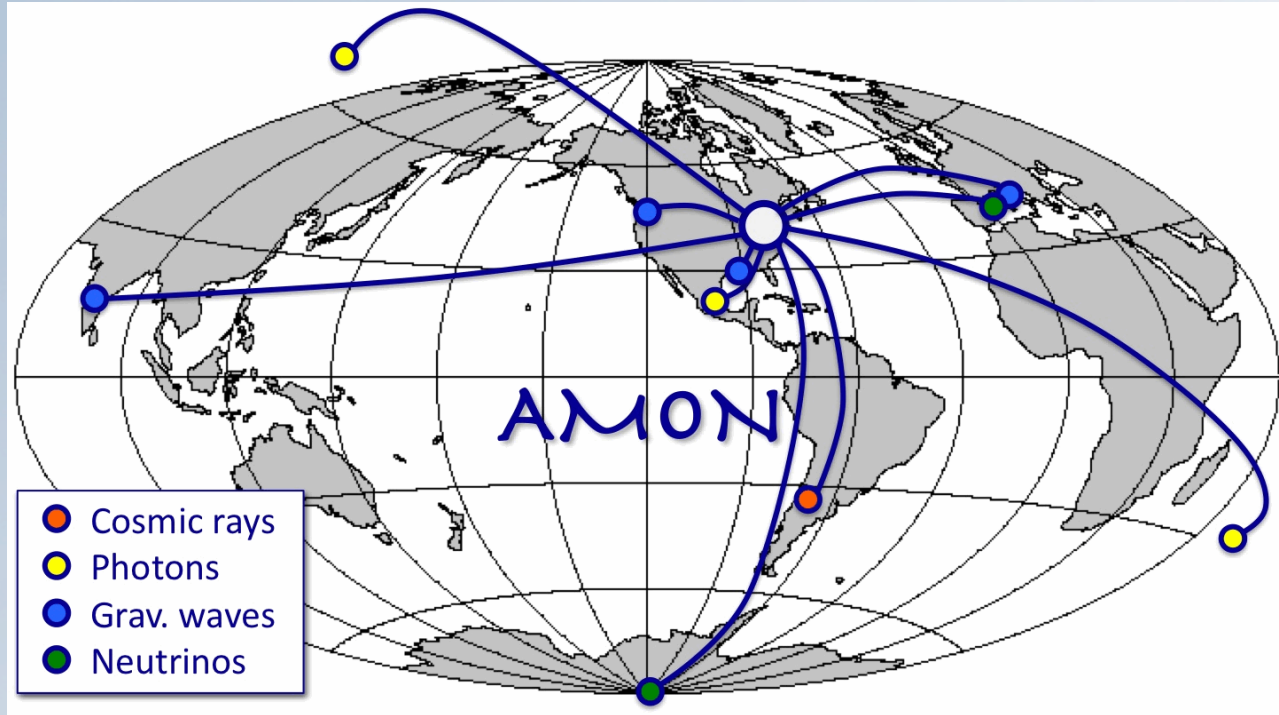


x, UV, optical





6 Memoranda of Understandings (MoUs) signed  
 + 1 MoU in review, + 2 more letter of collaboration + many more in the future



**Astrop.Phys. Vol. 45, 56–70, 2013**

**Triggering:**

**Follow-up:**

**Pending:**

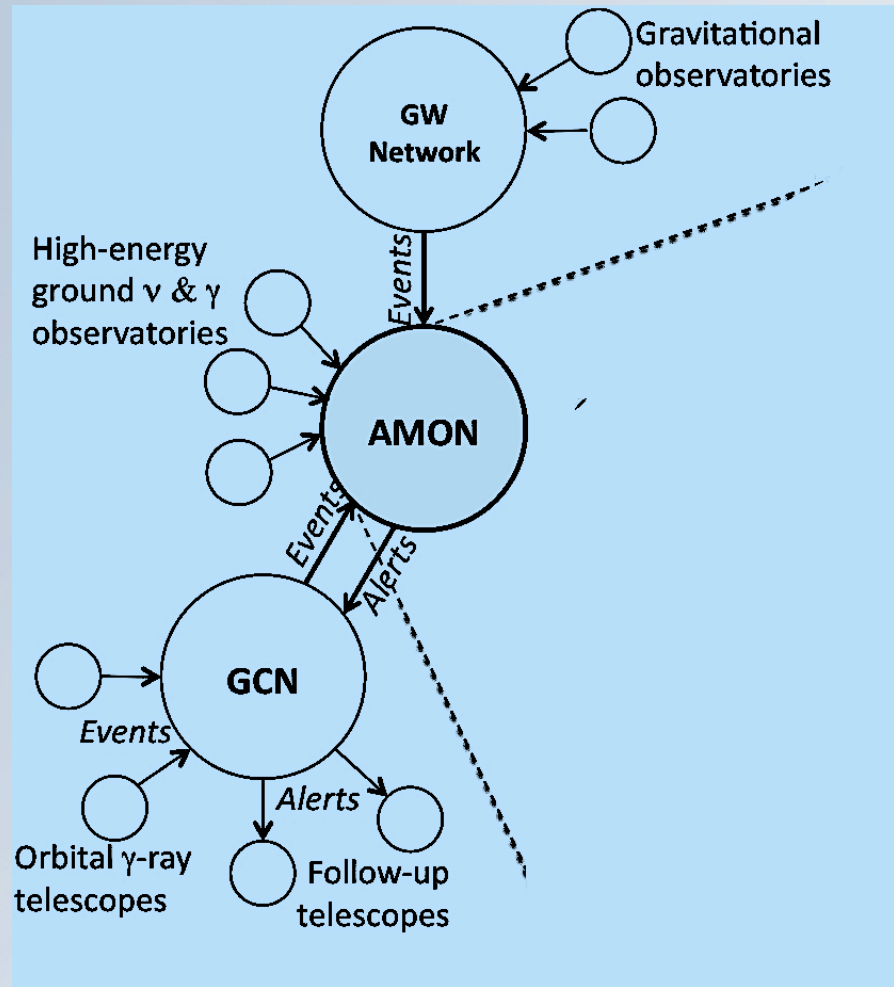
**IceCube**  
**ANTARES**  
**Auger**  
**HAWC**  
**VERITAS**  
**Swift BAT**

**Swift XRT & UVOT**  
**VERITAS**

**LIGO**  
**MAGIC**  
**FACT**  
**H.E.S.S.**  
**PTF**  
**TA...**

15-05-03

# AMON system - data flow



AMON will utilize existing:

- Gamma-ray Coordinates Network (GCN)**
- Gravitational Wave (GW) Network
- Open to other networks (e.g. SNEWS)

# NETWORK STATUS



First full version of AMON database designed and implemented, now being used and tested:

- Data from triggering observatories inserted
  - done: **IC-40, IC-59, Swift, Fermi** [public]
  - done: **ANTARES 2008** [private]
  - in progress: **IceCube, HAWC, VERITAS, ANTARES** [private – pending permissions],  
**Auger** [private - approved]  
**LIGO S5** [public]
- Real-time test with fake and real (IC scrambled) data performed



## Archival analysis using public data:

- IC40 and Fermi LAT (done, see talk by A. Keivani)
- IC40/59 and Swift BAT sub threshold (in progress)
- IC40/59 and Fermi LAT (in progress)
- Swift and LIGO S5 (in progress)



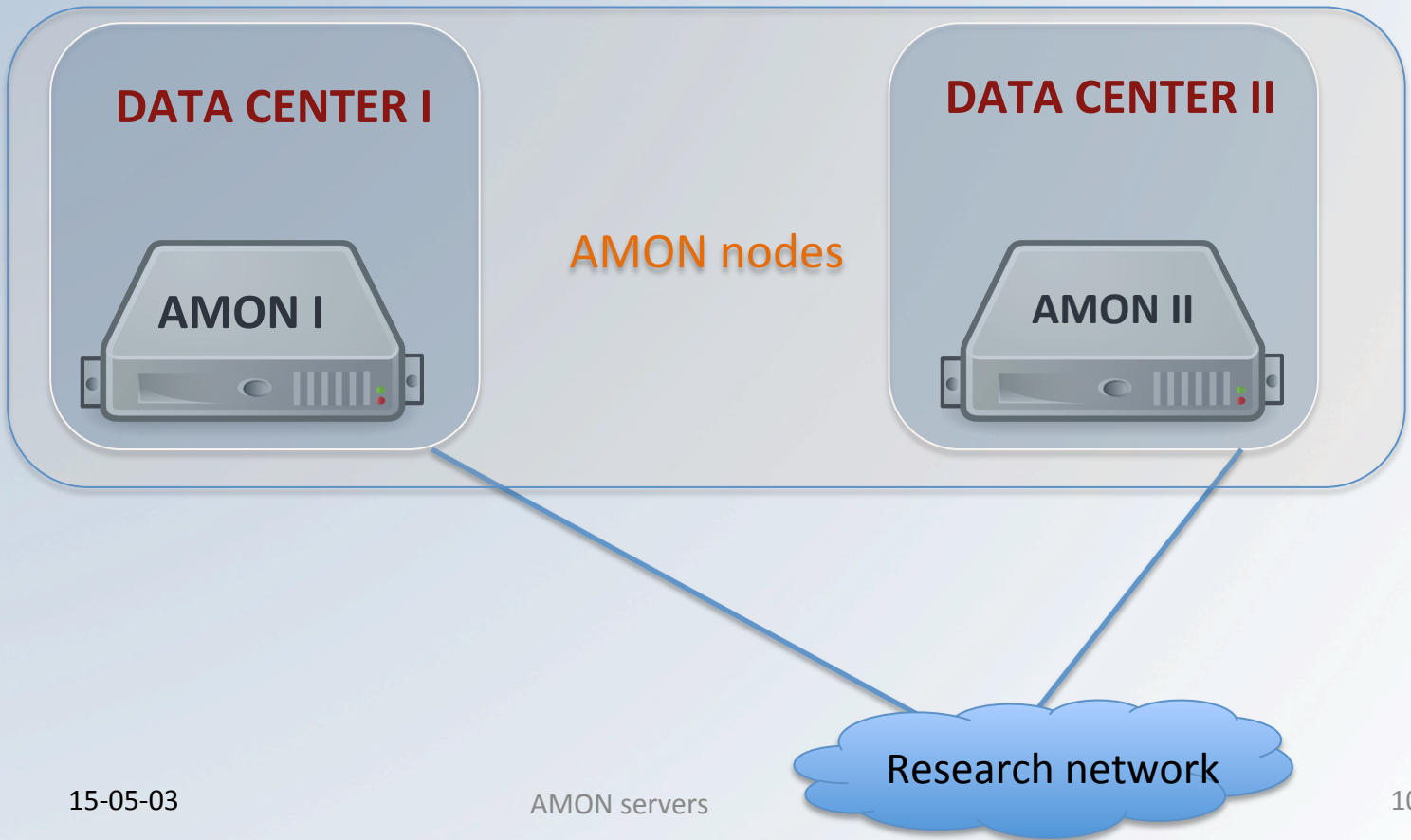


# AMON application server is up and running since August 2014!

- built using Python/Twisted, asynchronous, tested with several simulated clients
- Accepts HTTP POST requests (Twisted client available, but accepts other clients)
- Open for authorized connections (TLS certificates)
- Start issuing alerts (VOEvents) in August 2015



- Deployment of the two new high-uptime servers [in progress]
  - systems are physically and cyber secure
  - hardware and power redundant
  - memory mirroring
- Operational by June 30, 2015





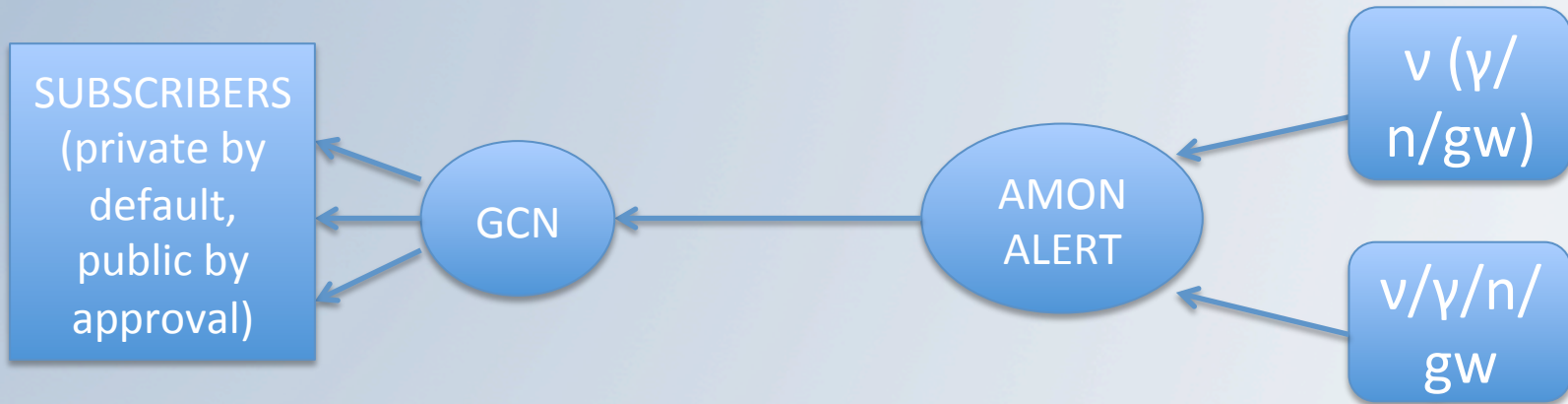
## Several efforts toward real-time analysis:

- IceCube  $\nu_{\mu}$  singlet stream – **done**
- IceCube high-energy starting events (HESE) – **underway**
- IC extremely high-energy (EHE) events – **underway**
- Auger events – **underway**
- Ongoing efforts with other member collaborations on getting their real-time subthreshold streams





## Subthreshold streams – used in AMON coincidence analysis



- GCN client directly connects to AMON by using vTCP protocol and gets AMON alerts
- Subscribers can choose to get original AMON VOEvent format or any other standard GCN formats (e.g. email)
- AMON receiver program is already built by GCN (S. Barthelmy)
- To get connection running within next few weeks



# Steps needed to set up a real-time stream with AMON

Observatory links to AMON

Observatory	Stream content & format	TLS certificate	Test stream (fake data)	Test steam (real data scrambled)	Real data stream
IC Singlet	✓	✓	✓	✓	In progress
IC HESE	✓	✓	In progress		
IC EHE	✓	✓			
Auger	✓	✓	In progress		
HAWC	In progress				
VERITAS	In progress				
Swift	✓	Not needed	Not needed	Not needed	In progress
Fermi	✓	Not needed	Not needed	Not needed	In progress





# Conclusions

- AMON has made a significant progress toward real-time and archival analysis
- AMON application server is online - open for authorized connections!
- New high-uptime dual hardware deployment underway
- Ongoing realtime stream with the scrambled neutrino singlet events from IceCube
- Expecting soon realtime new streams
- AMON will distribute alerts via GCN (private streams by default, public if approved— ready within few weeks)
- AMON will start issuing alerts in August 2015





## EXTRA SLIDES





## Event content common to each observatory :

stream number,  
id number,  
revision number  
trigger time  
position  
positional error  
number of events  
time window  
error on time  
false positive rate density  
p-value  
type of the event  
pointing  
observatory location  
type of the PSF

## Event content specific to each observatory :

parameter name: (*energy, SNR, etc.*).  
value of the parameter  
units (*TeV etc.*)





## AMON Alert content:

stream number

id number

revision number

time

position of the best fit

positional error

number of events

time window

error on time

false positive rate density

experiments observing

experiments triggered

type of the alert

skymap



# AMON will receive events and send alerts in VOEvent format

- Standardized data packet format simplifies protocols for data handling (e.g. adding new observatory will not require new methods for injection of data into database and analysis stream)
- VOEvent is used by larger astronomical community i.e. became a standard for **real-time** event distribution (e.g. GCN notices, Swift, Fermi, LIGO, AMON etc.)
- Well structured in XML format with simple schema
- Easily interpreted by software, can be read by robotic telescopes (important for real-time analysis and near real-time follow-up)

