

Particle Physics in Icecube

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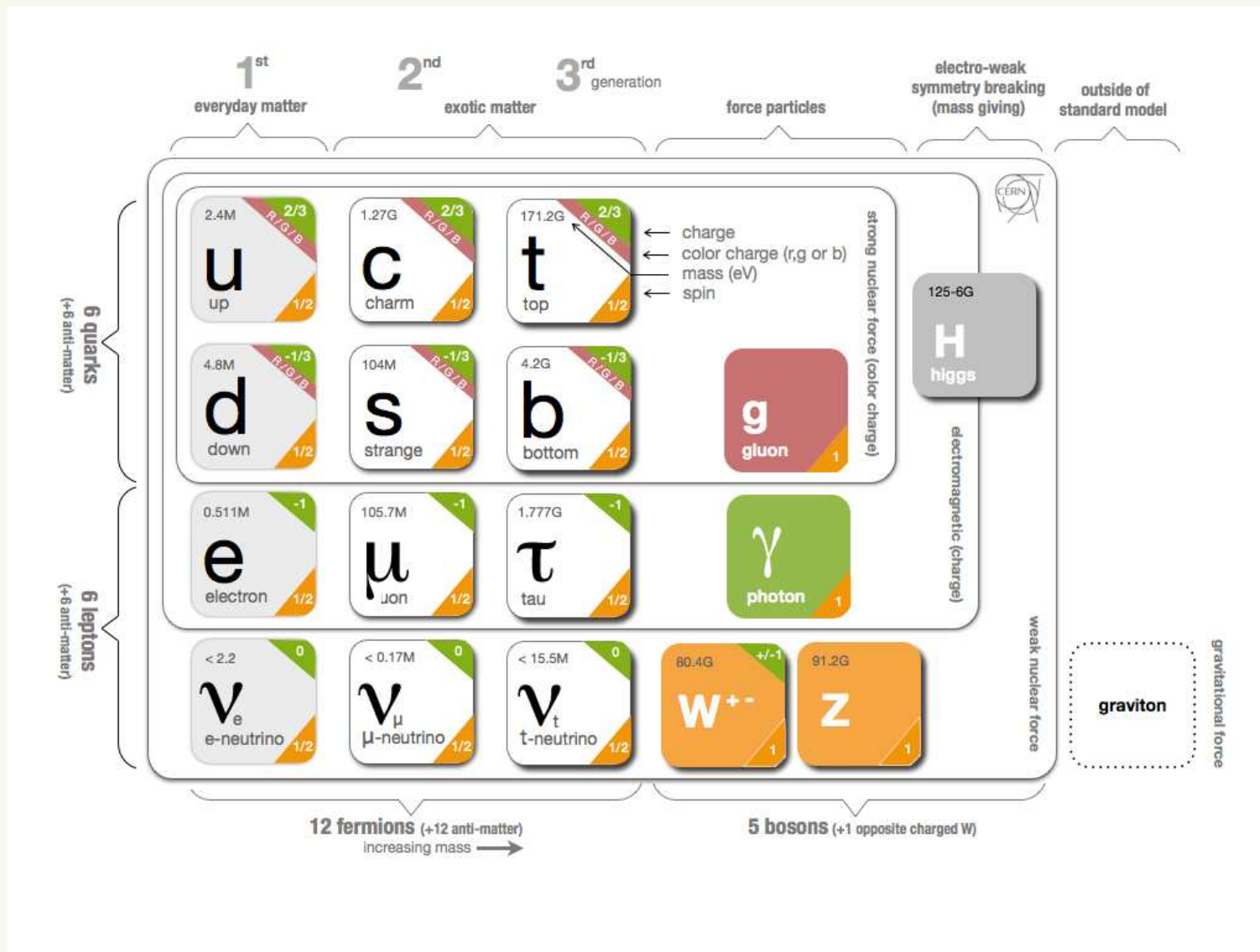
Department of Physics, University of Wisconsin and Wisconsin IceCube Particle Astrophysics Center



Bootcamp 2014

- Standard Model of Particle Physics
- Reasons to go beyond
- What Icecube can do?
 - Indirect signal for the decay or annihilation of new **particles**.
 - Direct detection of new **particles**.
 - Neutrino oscillations and neutrino oscillations as a window to new physics.

Standard Model of Particle Physics



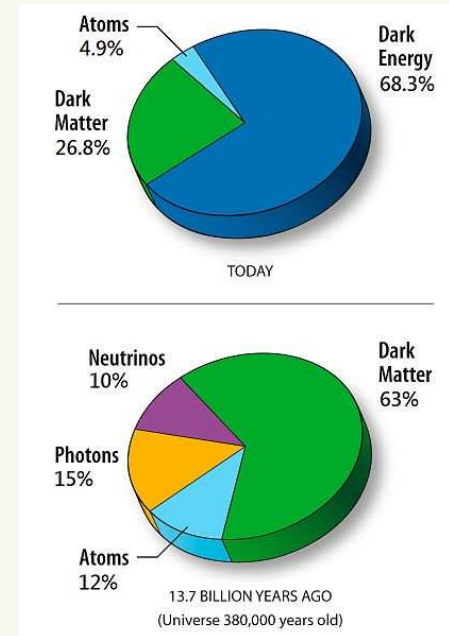
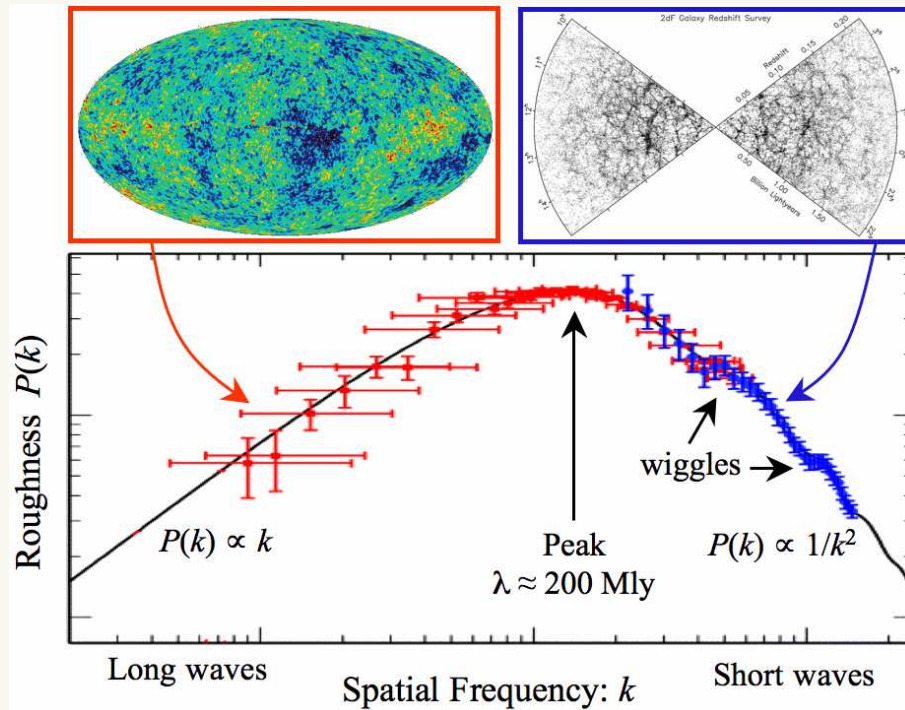
Standard Model of Particle Physics

$$\begin{aligned}
& -\frac{1}{2}\partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_s f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\nu^c - \frac{1}{4}g_s^2 f^{abc} f^{ade} g_\mu^b g_\nu^c g_\mu^d g_\nu^e + \\
& \frac{1}{2}ig_s^2 (\bar{q}_i^\sigma \gamma^\mu q_j^\sigma) g_\mu^a + \bar{G}^a \partial^2 G^a + g_s f^{abc} \partial_\mu \bar{G}^a G^b g_\mu^c - \partial_\nu W_\mu^+ \partial_\nu W_\mu^- - \\
& M^2 W_\mu^+ W_\mu^- - \frac{1}{2}\partial_\nu Z_\mu^0 \partial_\nu Z_\mu^0 - \frac{1}{2c_w^2} M^2 Z_\mu^0 Z_\mu^0 - \frac{1}{2}\partial_\mu A_\nu \partial_\mu A_\nu - \frac{1}{2}\partial_\mu H \partial_\mu H - \\
& \frac{1}{2}m_h^2 H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - M^2 \phi^+ \phi^- - \frac{1}{2}\partial_\mu \phi^0 \partial_\mu \phi^0 - \frac{1}{2c_w^2} M \phi^0 \phi^0 - \beta_h \left[\frac{2M^2}{g^2} + \right. \\
& \left. \frac{2M}{g} H + \frac{1}{2}(H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-) \right] + \frac{2M^4}{g^2} \alpha_h - igc_w [\partial_\nu Z_\mu^0 (W_\mu^+ W_\nu^- - \\
& W_\nu^+ W_\mu^-) - Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + Z_\mu^0 (W_\nu^+ \partial_\nu W_\mu^- - \\
& W_\nu^- \partial_\nu W_\mu^+)] - ig s_w [\partial_\nu A_\mu (W_\mu^+ W_\nu^- - W_\nu^+ W_\mu^-) - A_\nu (W_\mu^+ \partial_\nu W_\mu^- - \\
& W_\mu^- \partial_\nu W_\mu^+) + A_\mu (W_\nu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\nu W_\mu^+)] - \frac{1}{2}g^2 W_\mu^+ W_\mu^- W_\nu^+ W_\nu^- + \\
& \frac{1}{2}g^2 W_\mu^+ W_\nu^- W_\mu^+ W_\nu^- + g^2 c_w^2 (Z_\mu^0 W_\mu^+ Z_\nu^0 W_\nu^- - Z_\mu^0 Z_\nu^0 W_\mu^+ W_\nu^-) + \\
& g^2 s_w^2 (A_\mu W_\mu^+ A_\nu W_\nu^- - A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_w c_w [A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - \\
& W_\nu^+ W_\mu^-) - 2A_\mu Z_\mu^0 W_\nu^+ W_\nu^-] - g\alpha [H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-] - \\
& \frac{1}{8}g^2 \alpha_h [H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2] - \\
& g M W_\mu^+ W_\mu^- H - \frac{1}{2}g \frac{M}{c_w^2} Z_\mu^0 Z_\mu^0 H - \frac{1}{2}ig [W_\mu^+ (\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - \\
& W_\mu^- (\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)] + \frac{1}{2}g [W_\mu^+ (H \partial_\mu \phi^- - \phi^- \partial_\mu H) - W_\mu^- (H \partial_\mu \phi^+ - \\
& \phi^+ \partial_\mu H)] + \frac{1}{2}g \frac{1}{c_w} (Z_\mu^0 (H \partial_\mu \phi^0 - \phi^0 \partial_\mu H) - ig \frac{s_w}{c_w} M Z_\mu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \\
& ig s_w M A_\mu (W_\mu^+ \phi^- - W_\mu^- \phi^+) - ig \frac{1-2c_w^2}{2c_w} Z_\mu^0 (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + \\
& ig s_w A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \frac{1}{4}g^2 W_\mu^+ W_\mu^- [H^2 + (\phi^0)^2 + 2\phi^+ \phi^-] - \\
& \frac{1}{4}g^2 \frac{1}{c_w^2} Z_\mu^0 Z_\mu^0 [H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2 \phi^+ \phi^-] - \frac{1}{2}g^2 \frac{s_w^2}{c_w} Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + \\
& W_\mu^- \phi^+) - \frac{1}{2}ig^2 \frac{s_w^2}{c_w} Z_\mu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2}g^2 s_w A_\mu \phi^0 (W_\mu^+ \phi^- + \\
& W_\mu^- \phi^+) + \frac{1}{2}ig^2 s_w A_\mu H (W_\mu^+ \phi^- - W_\mu^- \phi^+) - g^2 \frac{s_w}{c_w} (2c_w^2 - 1) Z_\mu^0 A_\mu \phi^+ \phi^- - \\
& g^1 s_w^2 A_\mu A_\mu \phi^+ \phi^- - \bar{e}^\lambda (\gamma \partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda \gamma \partial \nu^\lambda - \bar{u}_j^\lambda (\gamma \partial + m_u^\lambda) u_j^\lambda - \\
& \bar{d}_j^\lambda (\gamma \partial + m_d^\lambda) d_j^\lambda + ig s_w A_\mu [-(\bar{e}^\lambda \gamma^\mu e^\lambda) + \frac{2}{3}(\bar{u}_j^\lambda \gamma^\mu u_j^\lambda) - \frac{1}{3}(\bar{d}_j^\lambda \gamma^\mu d_j^\lambda)] + \\
& \frac{ig}{4c_w} Z_\mu^0 [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{e}^\lambda \gamma^\mu (4s_w^2 - 1 - \gamma^5) e^\lambda) + (\bar{u}_j^\lambda \gamma^\mu (\frac{4}{3}s_w^2 - \\
& 1 - \gamma^5) u_j^\lambda) + (\bar{d}_j^\lambda \gamma^\mu (1 - \frac{8}{3}s_w^2 - \gamma^5) d_j^\lambda)] + \frac{ig}{2\sqrt{2}} W_\mu^+ [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + \\
& (\bar{u}_j^\lambda \gamma^\mu (1 + \gamma^5) C_{\lambda\kappa} d_j^\kappa)] + \frac{ig}{2\sqrt{2}} W_\mu^- [(\bar{e}^\lambda \gamma^\mu (1 + \gamma^5) e^\lambda) + (\bar{d}_j^\kappa C_{\lambda\kappa}^\dagger \gamma^\mu (1 + \\
& \gamma^5) u_j^\lambda)] + \frac{ig}{2\sqrt{2}} \frac{m_h^2}{M} [-\phi^+ (\bar{\nu}^\lambda (1 - \gamma^5) e^\lambda) + \phi^- (\bar{e}^\lambda (1 + \gamma^5) \nu^\lambda)] - \\
& \frac{g}{2} \frac{m_h^2}{M} [H (\bar{e}^\lambda e^\lambda) + i\phi^0 (\bar{e}^\lambda \gamma^5 e^\lambda)] + \frac{ig}{2M\sqrt{2}} \phi^+ [-m_d^\kappa (\bar{u}_j^\lambda C_{\lambda\kappa} (1 - \gamma^5) d_j^\kappa) + \\
& m_u^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 + \gamma^5) d_j^\kappa) + \frac{ig}{2M\sqrt{2}} \phi^- [m_d^\lambda (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 + \gamma^5) u_j^\kappa) - m_u^\kappa (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 - \\
& \gamma^5) u_j^\kappa) - \frac{g}{2} \frac{m_h^2}{M} H (\bar{u}_j^\lambda u_j^\lambda) - \frac{g}{2} \frac{m_h^2}{M} H (\bar{d}_j^\lambda d_j^\lambda) + \frac{ig}{2} \frac{m_h^2}{M} \phi^0 (\bar{u}_j^\lambda \gamma^5 u_j^\lambda) - \\
& \frac{ig}{2} \frac{m_h^2}{M} \phi^0 (\bar{d}_j^\lambda \gamma^5 d_j^\lambda) + \bar{X}^+ (\partial^2 - M^2) X^+ + \bar{X}^- (\partial^2 - M^2) X^- + \bar{X}^0 (\partial^2 - \\
& \frac{M^2}{c_w^2}) X^0 + \bar{Y} \partial^2 Y + igc_w W_\mu^+ (\partial_\mu \bar{X}^0 X^- - \partial_\mu \bar{X}^+ X^0) + ig s_w W_\mu^+ (\partial_\mu \bar{Y} X^- - \\
& \partial_\mu \bar{X}^+ Y) + igc_w W_\mu^- (\partial_\mu \bar{X}^- X^0 - \partial_\mu \bar{X}^0 X^+) + ig s_w W_\mu^- (\partial_\mu \bar{X}^- Y - \\
& \partial_\mu \bar{Y} X^+) + igc_w Z_\mu^0 (\partial_\mu \bar{X}^+ X^+ - \partial_\mu \bar{X}^- X^-) + ig s_w A_\mu (\partial_\mu \bar{X}^+ X^+ - \\
& \partial_\mu \bar{X}^- X^-) - \frac{1}{2}g M [\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{c_w^2} \bar{X}^0 X^0 H] + \\
& \frac{1-2c_w^2}{2c_w} ig M [\bar{X}^+ X^0 \phi^+ - \bar{X}^- X^0 \phi^-] + \frac{1}{2c_w} ig M [\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + \\
& ig M s_w [\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + \frac{1}{2}ig M [\bar{X}^+ X^+ \phi^0 - \bar{X}^- X^- \phi^0]
\end{aligned}$$

Reasons to go beyond

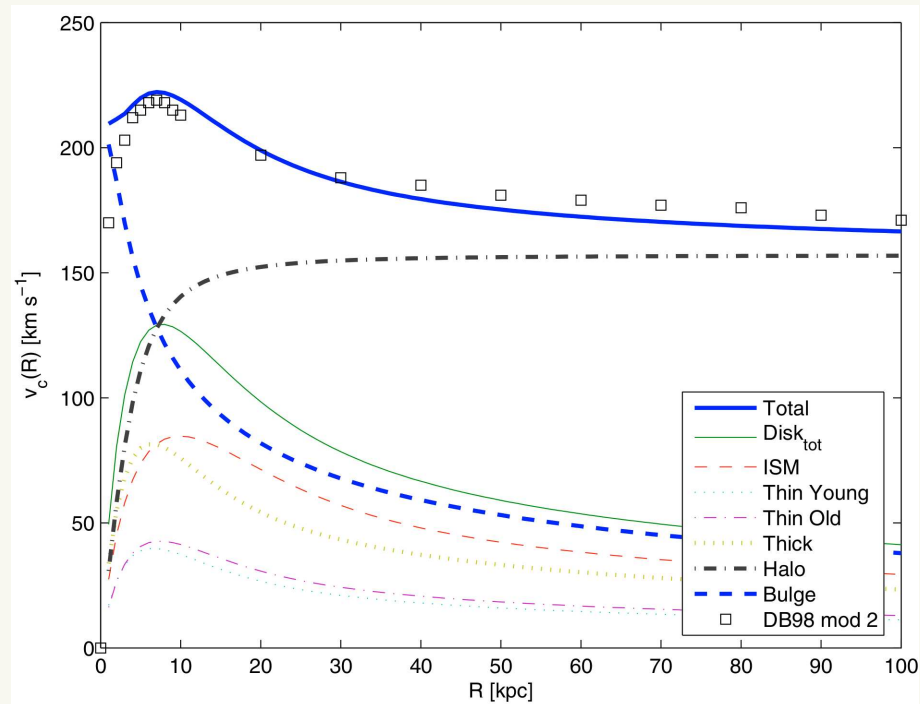
- Neutrinos oscillations(Massive ν)!!
- Gravity is not included in a consistent way
- The SM doesn't look very natural($m_{top} \approx 10^{11} m_\nu$)
The same ratio as a human and Mont Everest mass!
- Cosmology can not be explained with just Gravity + SM
 - The prediction for the Vacuum energy is 10^{120} orders of magnitude off!
 - Apparently only 4.9% of the energy in the universe is from the SM.

Reason to go beyond



- Large scale cosmological observations tells us that the baryonic matter is only **15%** of the total matter in the Universe.

Reason to go beyond

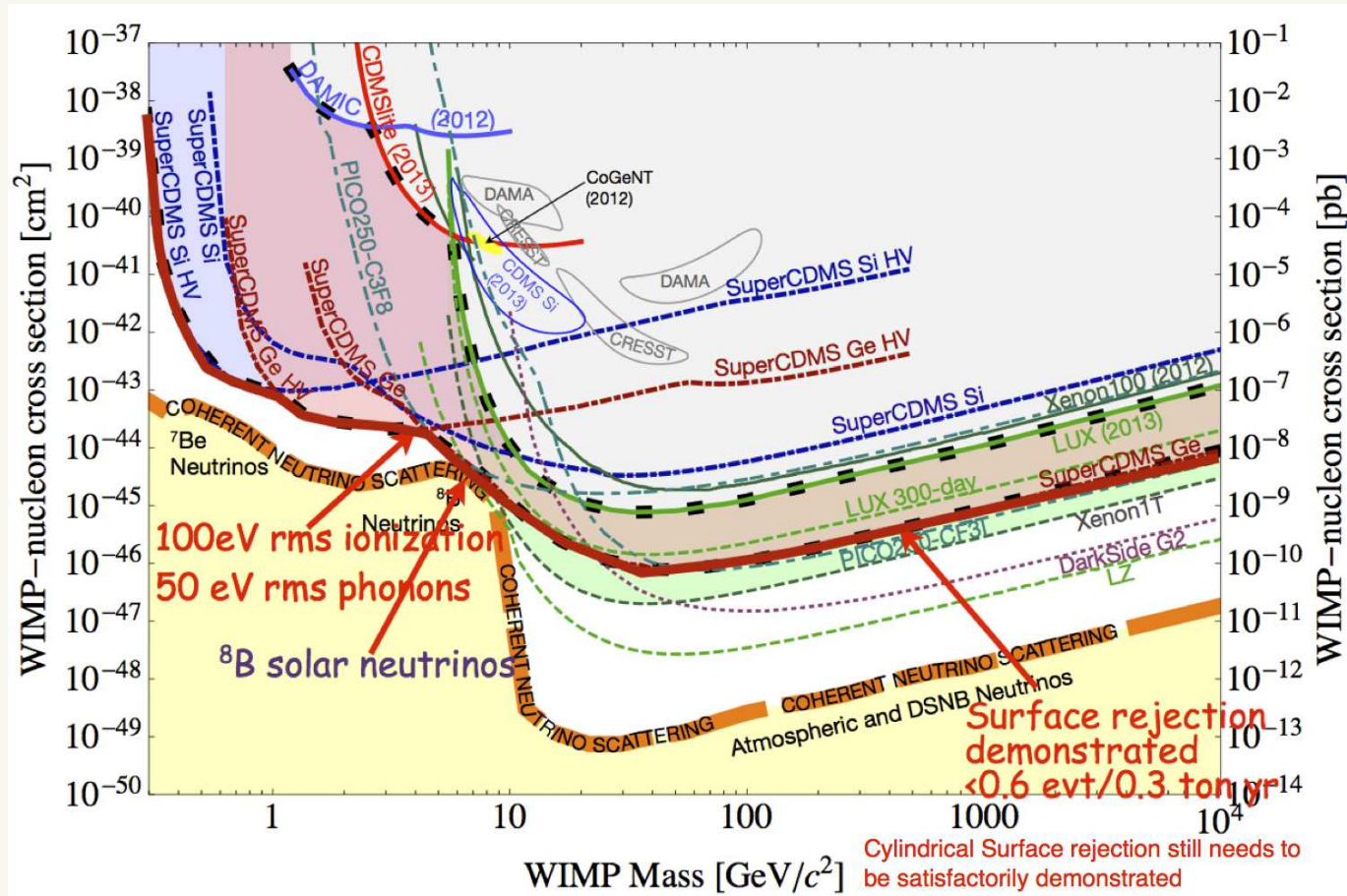


A.Vallenari et.al. A&A 451, 125–139 (2006)

- Local measurements of the galaxy kinematics suggest an important contribution from the halo

Reason to go beyond

WIMP?



The WIMP Miracle

- Heavy particle initially in thermal equilibrium
- The thermal relic density is

$$\Omega_x \propto \frac{1}{\langle \sigma v \rangle} \approx \frac{m_X^2}{g_X^4}$$

- But we get the right number with a physical scale already in the standard model!! Electroweak Scale

$$m_X \approx 100\text{GeV}, g_X \approx 0.6 \rightarrow \Omega_X \approx 0.1$$

Reason to go beyond

No WIMPS are also possible, even in the standard model (QCD?)



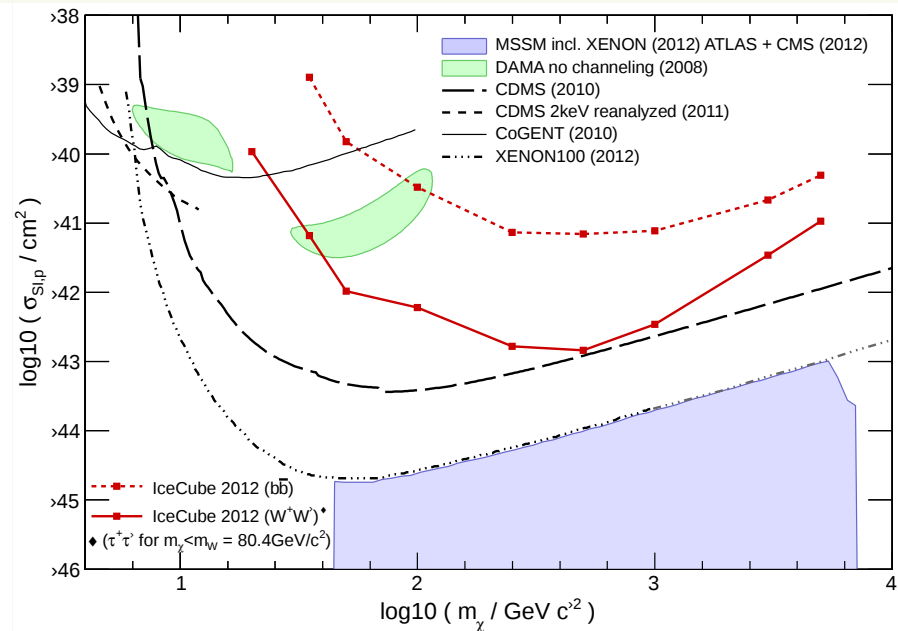
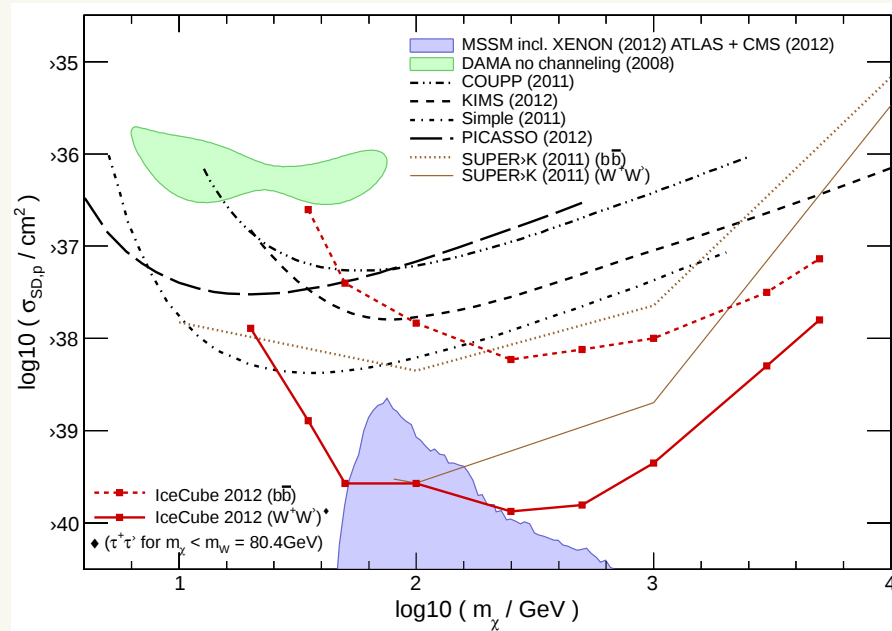
photo of a non trivial gauge configuration produced during the QCD phase transition in the early universe

Indirect signal for the decay of the new particles

- The DM particles may decay in SM particles(may be ν)
- Possible signals:
 - Galactic Center
 - Sun(High density because the capture) If there is not an extra dark mediator is a low energy search.

Indirect signal for the decay of the new **particles**

90% exclusion for the SD and SI cross sections



Indirect signal for the decay of the new **particles**

- May be we are already measuring new physics, where the HESE ν com from?:
 - Superheavy Particle Origin of IceCube PeV Neutrino Events (s-channel enhancement of neutrino-quark scattering)
(1305.6907) *Vernon Barger & Wai-Yee Keung*
 - Neutrinos at IceCube from Heavy Decaying Dark Matter (Explore models focused on the 2 events at 1PeV, mono-energetic line)
(1303.7320) *Brian Feldstein et.al.*
 - Are IceCube neutrinos unveiling PeV-scale decaying dark matter? (Generic signatures in the energy spectra)
(1308.1105) *Arman Esmaili & Pasquale Dario Serpico*
 - Geometric Compatibility of IceCube TeV-PeV Neutrino Excess and its Galactic Dark Matter Origin (Spatial distribution analysis and spectra)
(1311.5864) *Yang Bay, Ran Lu & Jordi Salvado*

Direct detection of new **particles**

- Relativistic Monopoles
- Mildly Relativistic Monopoles
- Non-Relativistic GUT Monopoles, Nuclearites (Strangelets), Q-balls, etc...
- SUSY searches (double Sleptons)
- TeV Gravity

Direct detection of new particles

- Monopoles are topological defects of the gauge field that appear typically in phase transitions(big bang).
- The classical equations for a gauge field are:

$$F = dA$$

$$d * dA = J$$

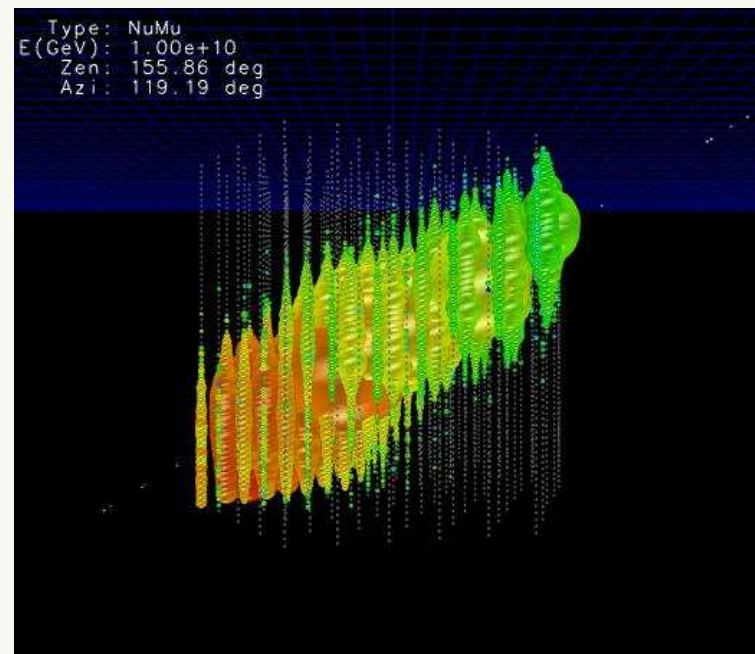
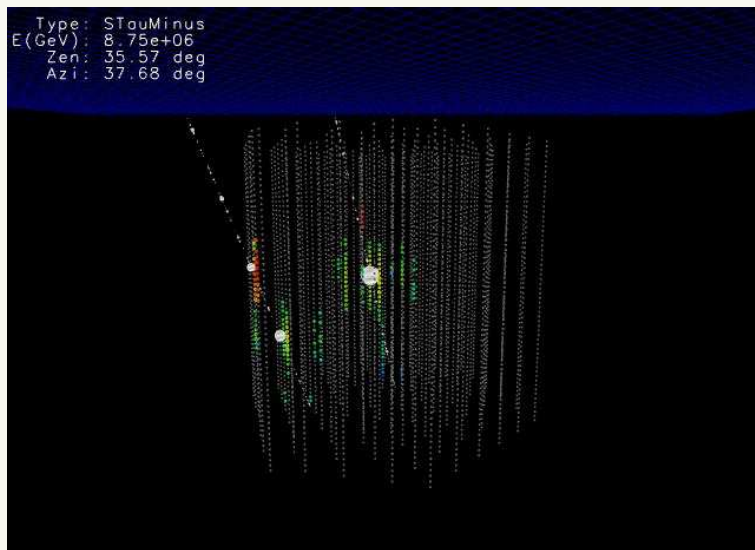
- $ddw = 0$ for all w
- They are expected to be produced in the early universe during the GUT phase transitions

Direct detection of new particles

- Ultra high energies CR may produce new particles(SUSY, TeV Gravity)
- SUSY a possible signal is two parallel track-like events
- With extra dimensions the effective 4D G_N becomes smaller, making possible to produce BH from CR

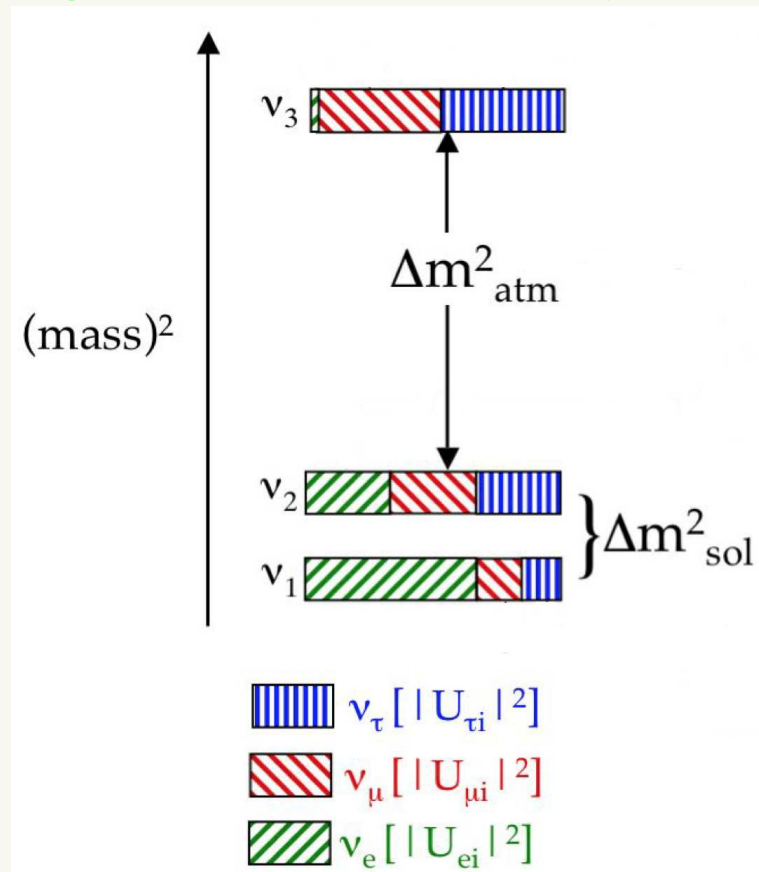
Direct detection of new **particles**

SUSY double stau and BH signal



Neutrino oscillations

Neutrino oscillations : **mass eigenstates** ($\nu_i; i = 1, 2, 3$) and **flavor eigenstates** ($\nu_\alpha; \alpha = e, \mu, \tau$) are not the same.



[B. Kayser, hep-ph/0506165 (2004)]

[C. Gonzalez-Garcia et al., JHEP 12 (2012)]

$$\Delta m_{\text{sol}}^2 = 7.5 \times 10^{-5} \text{eV}^2$$

$$|\Delta m_{\text{atm}}^2| = 2.4 \times 10^{-3} \text{eV}^2$$

$$\nu_i = \sum_{\beta} U_{\beta i} \nu_{\beta}$$

$$U = U(\theta_{12}, \theta_{23}, \theta_{13})$$

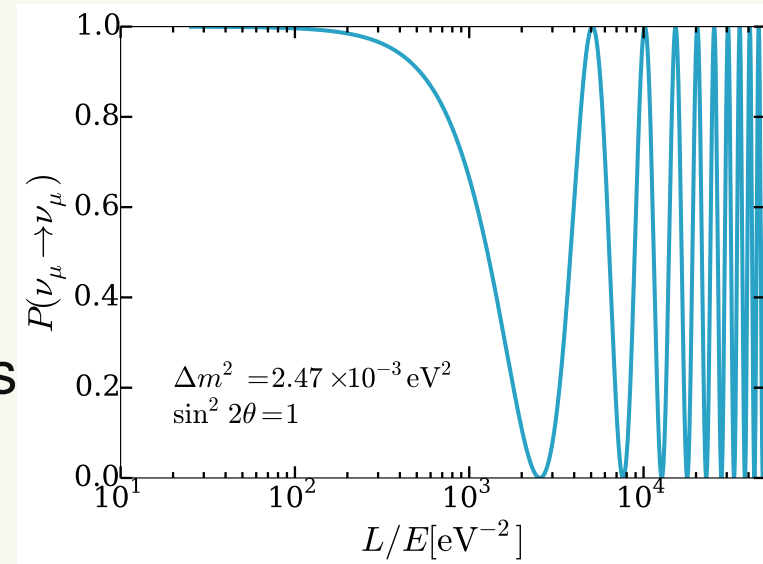
$$|U| \simeq \begin{pmatrix} 0.8 & 0.5 & 0.1 \\ 0.3 & 0.7 & 0.6 \\ 0.4 & 0.5 & 0.8 \end{pmatrix}$$

Neutrino oscillations

In two generations the oscillation probability at a given distance L and energy E in vacuum

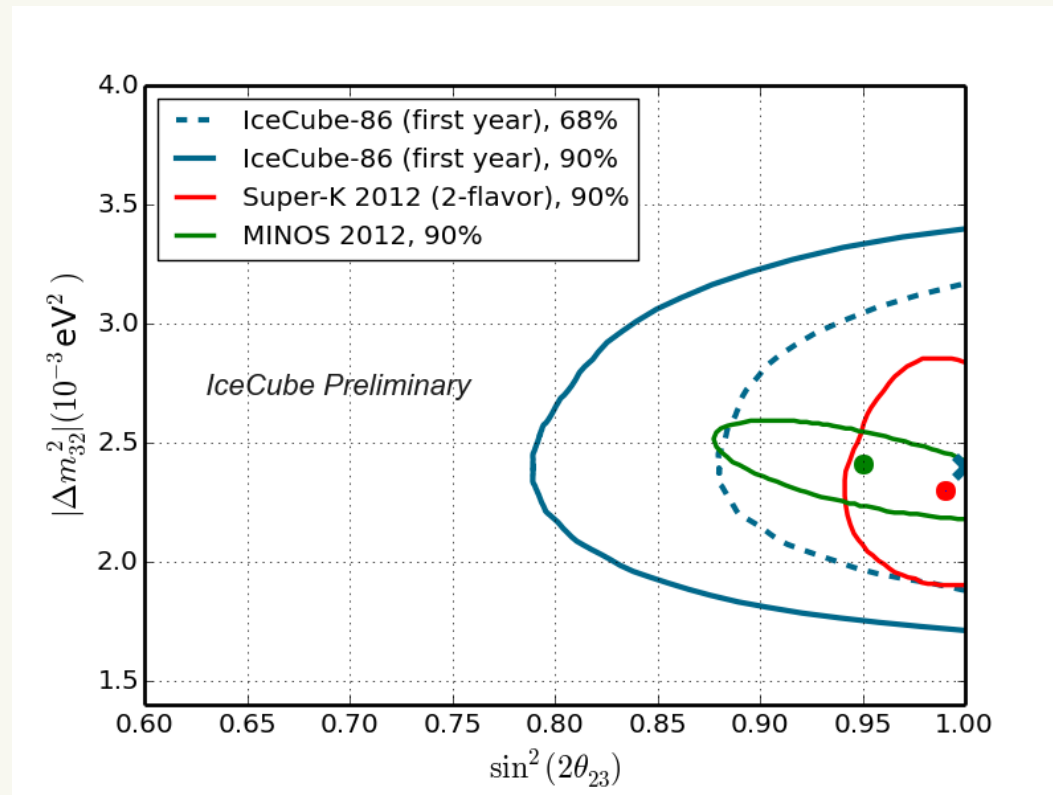
$$P_{\nu_\alpha \rightarrow \nu_\alpha} \left(\frac{L}{E} \right) = 1 - \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E} \right)$$

- $\sin^2 2\theta$: oscillation amplitude
- Δm^2 : oscillation frequency
 - $L/E \ll 1/\Delta m^2 \rightarrow$ no oscillations
 - $L/E \sim 1/\Delta m^2 \rightarrow$ oscillations
 - $L/E \gg 1/\Delta m^2 \rightarrow$ fast oscillations ("averaged")



Neutrino oscillations

Icecube is able to measure atmospheric neutrinos produced by the CR in the atmosphere



Shows that the oscillation picture is consistent also at high energies

Neutrino oscillations(**Sterile** ν)

- **LSND** found $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ **oscillation** with $\Delta m^2 \sim 1\text{eV}^2$ and $\sin^2 2\theta \sim 0.003$
- MiniBoone $\nu_\mu \rightarrow \nu_e$ and $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ appearance
 - No significant excess at high energies ($E > 475$ MeV)
 - Unexplained events at low energies, interpretation as oscillations similar to LSND: $\Delta m^2 \sim 1\text{eV}^2$
- Gallium Anomaly, SAGE and GALLEX event rates lower than expected, can be explained by ν_e disappearance with $\Delta m^2 \geq 1\text{eV}^2$
- New reactor flux calculation (Mueller et al., 1101.2663, P. Huber, 1106.0687) 3% higher, tension in short-baseline ($L \leq 100\text{m}$) experiments, can be explained by ν_e disappearance with oscillation with $\Delta m^2 \sim 1\text{eV}^2$.

Neutrino oscillations(Sterile ν)

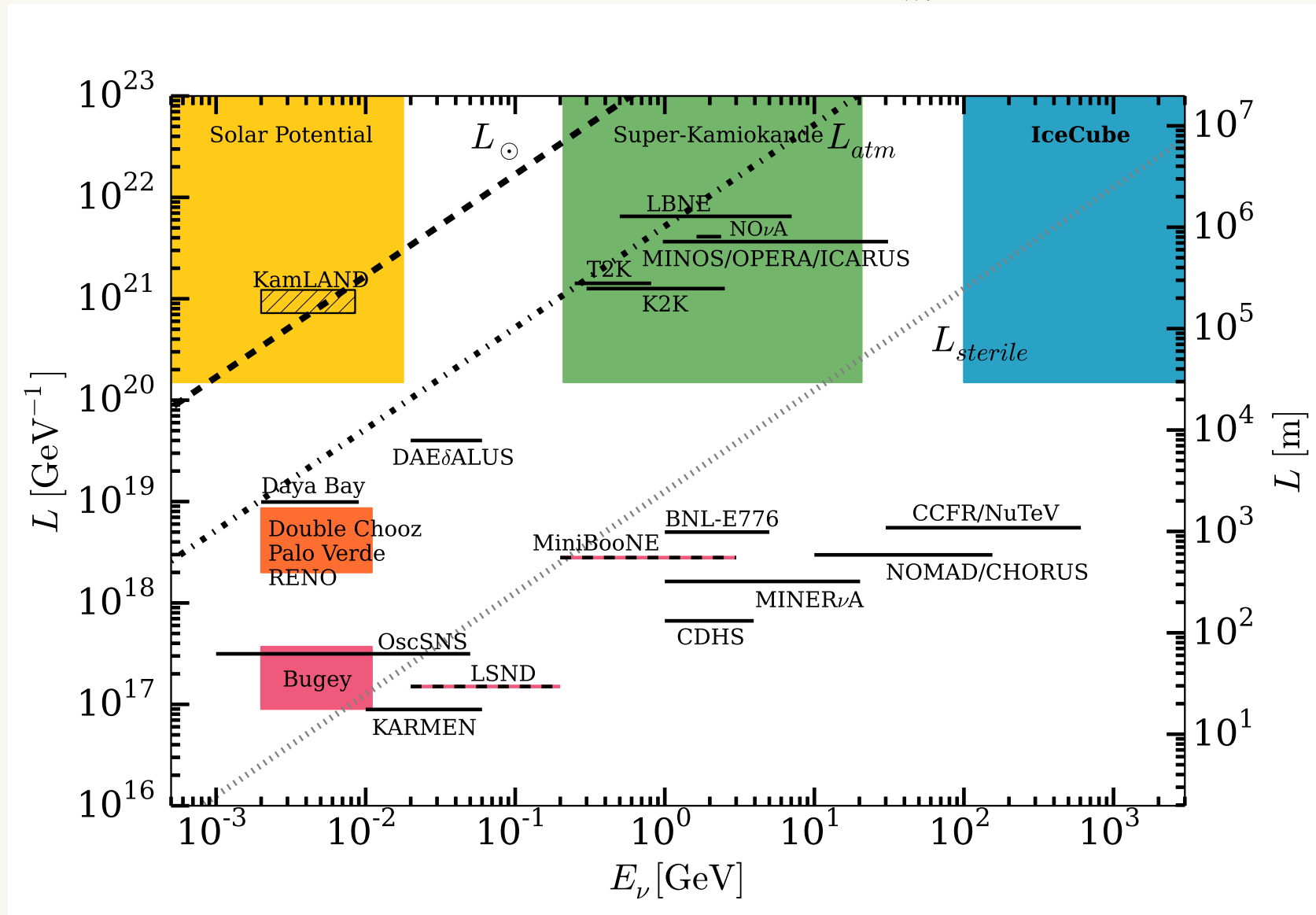
More Motivation? ... cosmology ...

- Λ CDM+ N_{eff} with WMAP9 + spt + act $N_{eff} = 3.89 \pm 0.67(68\%CL)$
- Λ CDM+ m_ν with WMAP9 + spt + act + SN1a $\sum m_\nu < 0.56eV(95\%CL)$
- $w\Lambda$ CDM+ m_ν with WMAP9 + spt + act + SN1a $\sum m_\nu < 1.2eV(95\%CL)$
- Λ CDM+ $m_\nu + N_{eff}$ with Planck + WP +spt + act
 $\sum m_\nu < 0.6eV \quad N_{eff} = 3.29^{+0.67}_{-0.64}(95\%CL)$
- Λ CDM+ $m_\nu + N_{eff}$ with Planck + WP +spt + act +BAO
 $\sum m_\nu < 0.28eV \quad N_{eff} = 3.32^{+0.54}_{-0.52}(95\%CL)$
- Λ CDM+ $m_\nu + N_{eff}$ with Planck + WP + Bicep2
 $\sum m_\nu < 0.81eV \quad \Delta N_{eff} = 1.08^{+0.49}_{-0.61}$ Maria Archidiacono et.al. arXiv:1404.1794

[Planck Collaboration, arXiv:1303.5076]

Neutrino oscillations(**Sterile** ν)

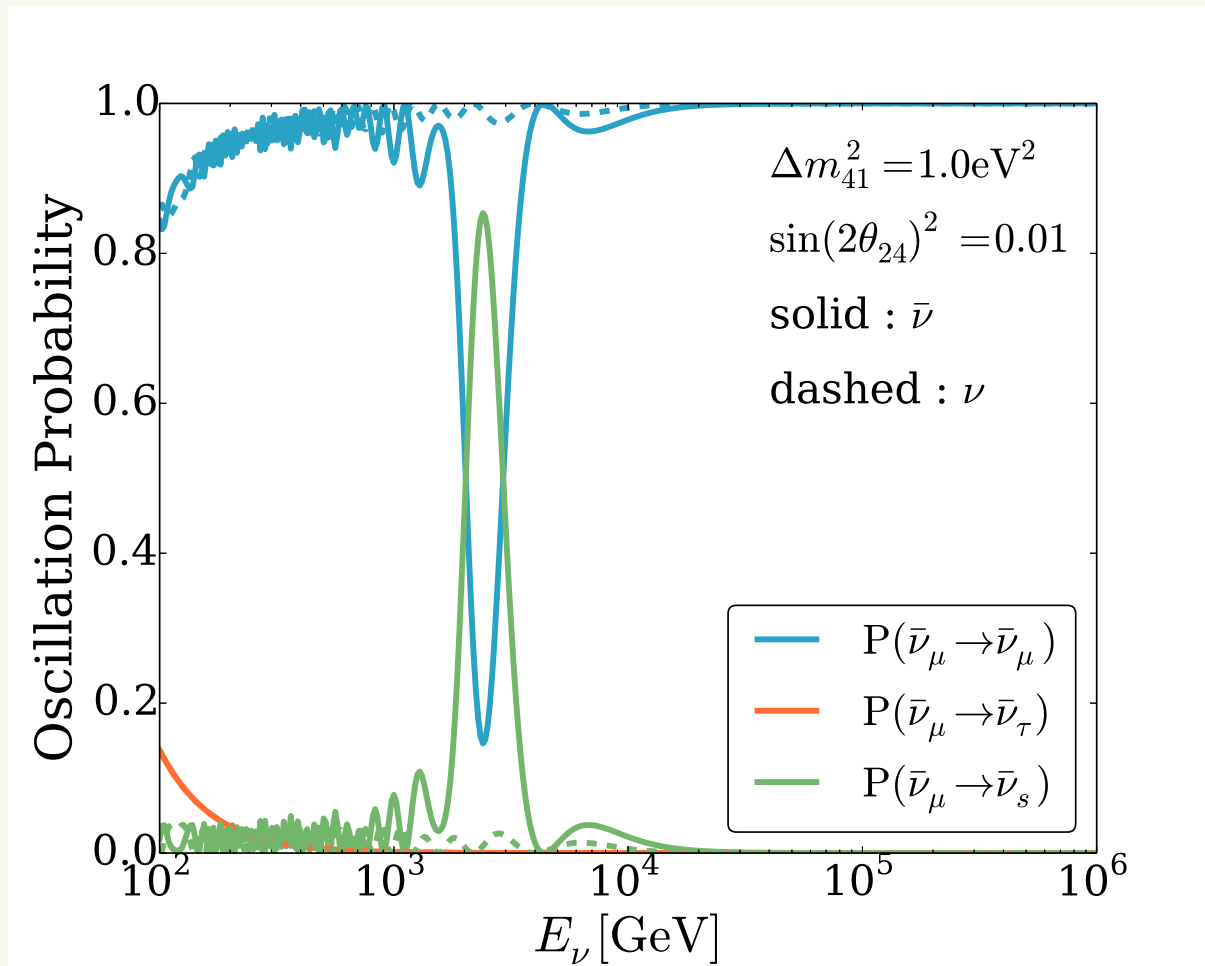
$$\text{Experiments : } L_{\text{osc}} = 2\pi \frac{E}{\Delta m^2}$$



Neutrino oscillations(**Sterile** ν)

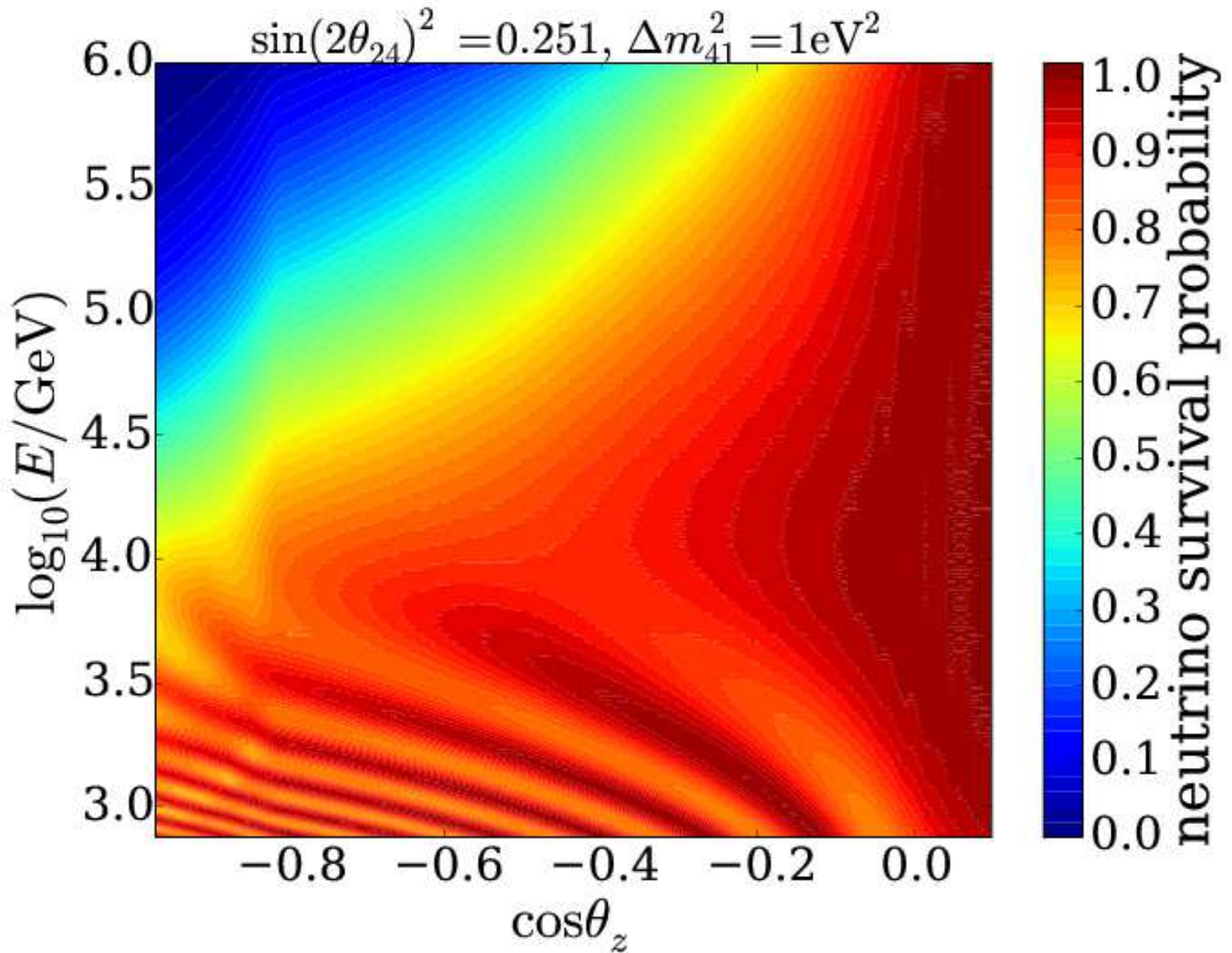
In the **Earth** for sterile neutrino $\Delta m^2 = O(1eV^2)$ the MSW effect happens when

$$E_\nu = \frac{\Delta m^2 \cos 2\theta}{2\sqrt{2}G_F N} \sim O(TeV)$$



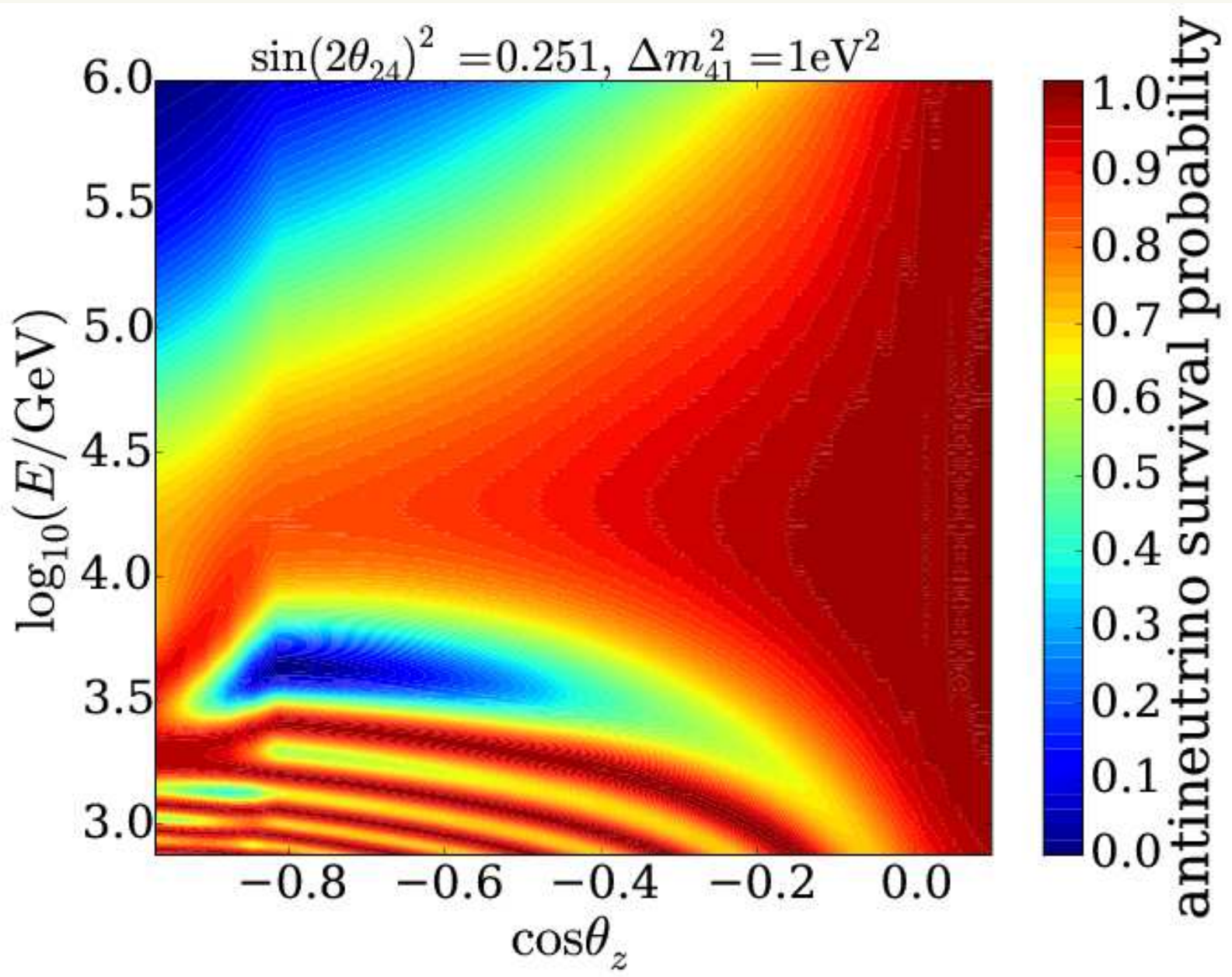
Neutrino oscillations(**Sterile** ν)

Oscillogram $\phi_\mu^{3+1}/\phi_\mu^{ini}$ for $\Delta m_{41}^2 = 1eV^2$, $\sin(2\theta_{24})^2 = 0.251$



Neutrino oscillations(Sterile ν)

Oscillogram $\phi_\mu^{3+1}/\phi_\mu^{ini}$ for $\Delta m_{41}^2 = 1\text{eV}^2$, $\sin(2\theta_{24})^2 = 0.251$



Neutrino oscillations(New Physics)

- Due to the coherent interactions Neutrino oscillations can be very sensitive to new physics at higher scales
- Lorentz Violation
- NSI

Neutrino oscillations(New Physics)

$$H_{\pm} \equiv \frac{\Delta m^2}{4E} \mathbf{U}_{\theta} \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix} \mathbf{U}_{\theta}^{\dagger} + \sum_n \sigma_n^{\pm} \frac{\Delta \delta_n E^n}{2} \mathbf{U}_{\xi_n, \pm \eta_n} \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix} \mathbf{U}_{\xi_n, \pm \eta_n}^{\dagger}$$

$\sigma_n^{\pm}$ sign for ν $\bar{\nu}$

$$\xi_1 = \xi_{vep} \quad \Delta \delta_1 = 2|\phi|(\gamma_1 - \gamma_2) \equiv 2|\phi|\Delta\gamma \leq 1.6 \times 10^{-24}, \text{ for VEP}$$

$$\xi_1 = \xi_{vli}, \quad \Delta \delta_1 = (c_1 - c_2) \equiv \delta c/c \leq 1.6 \times 10^{-24}, \text{ for VLI}$$

$$\xi_0 = \xi_Q, \quad \Delta \delta_0 = Q(k_1 - k_2) \leq 6.3 \times 10^{-23} \text{ GeV}, \text{ for coupling to torsion}$$

$$\xi_0 = \xi_{\cancel{C}PT}, \quad \Delta \delta_0 = b_1 - b_2 \leq 5.0 \times 10^{-23} \text{ GeV}, \text{ for } \cancel{C}PT, \text{ VLI}$$

[M.C. Gonzalez-Garcia, F.Halzen, M.Maltoni, ArXiv:0502223]

Neutrino oscillations(New Physics)

