

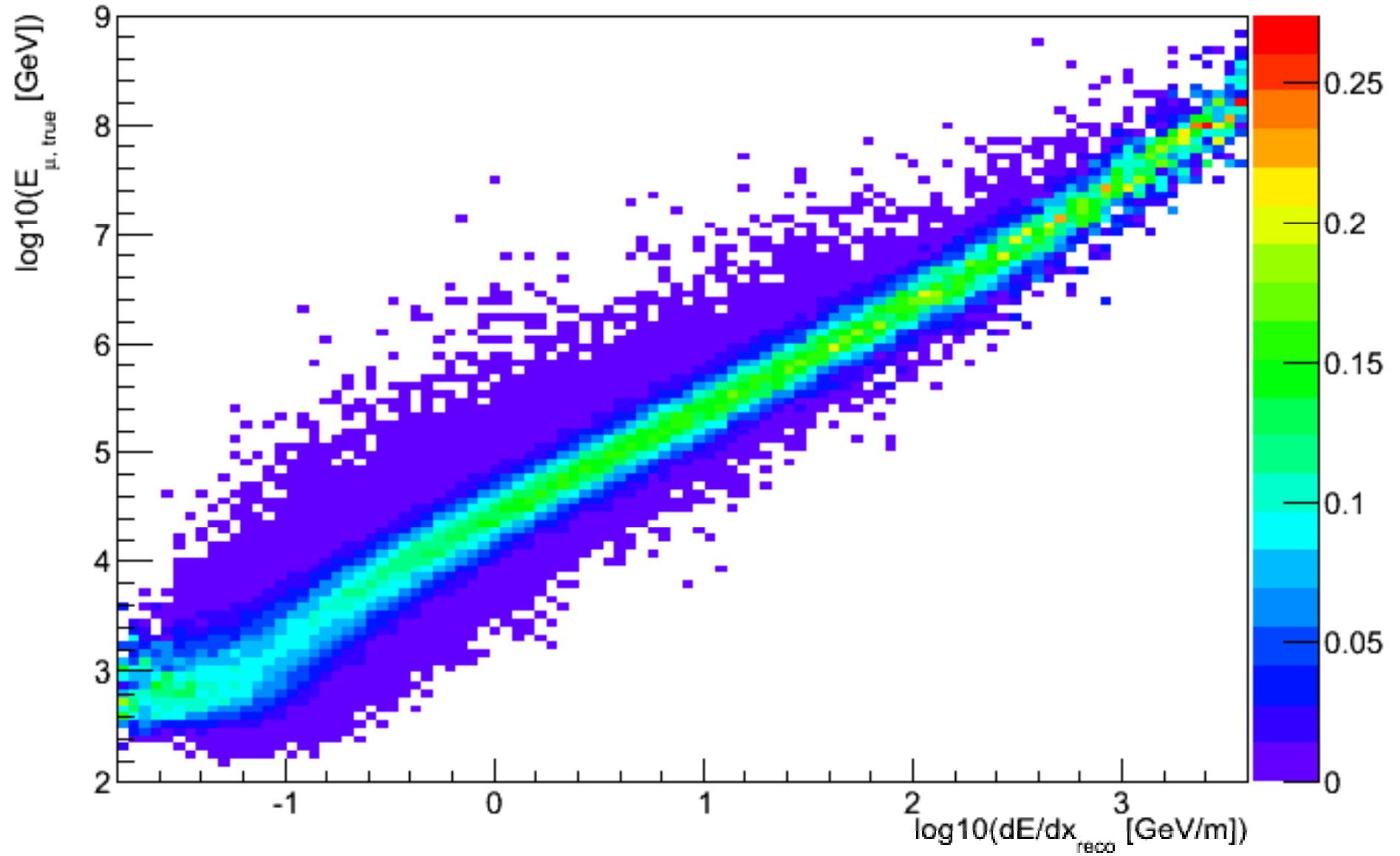
Apollon energy?

Anne Schukraft and Gary Hill

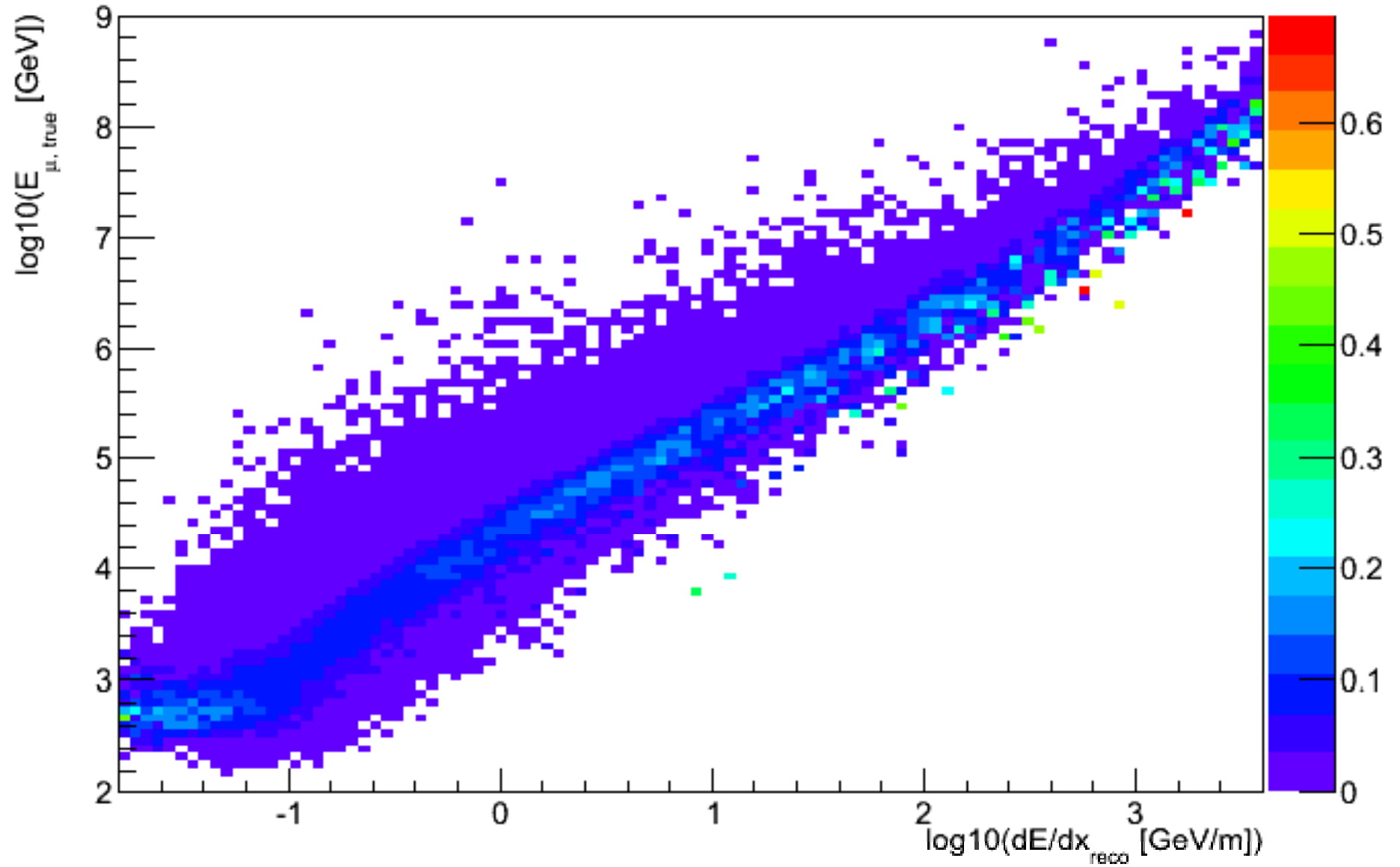
Note: all numbers are preliminary!

- How do we determine an energy?
- Know $\log_{10}(dE/dx) = 1.37$ for Apollon
- Look at a scatter plot of dE/dx vs true muon or neutrino energy
- Draw a vertical line at $\log_{10}(dE/dx) = 1.37$
- Integrate from bottom and top along curve to 16% to get a one-sigma central energy range
- However: this depends on how we weight the true energies... atmospheric, or E^{-2} ????

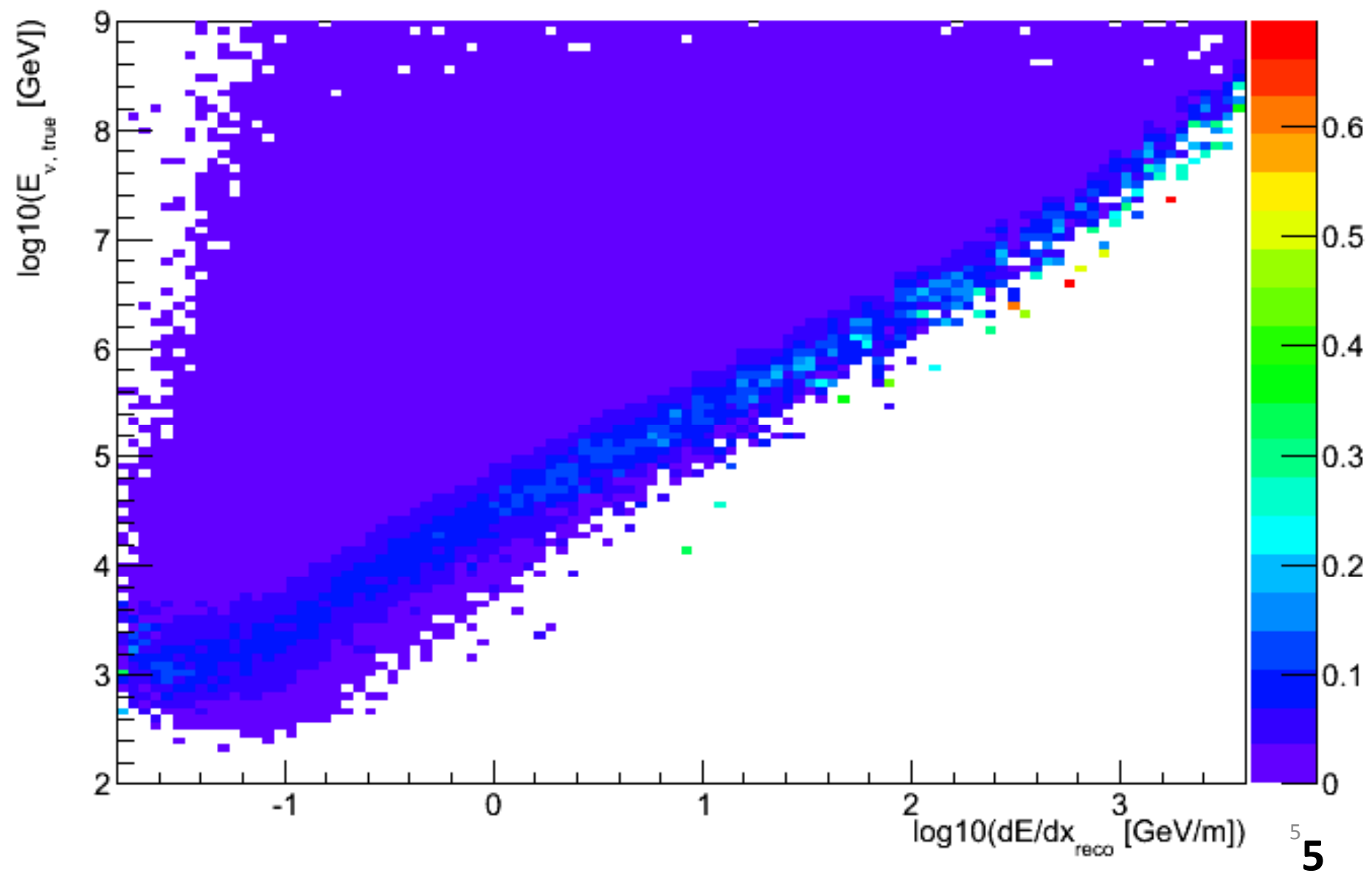
signal muon_allsky



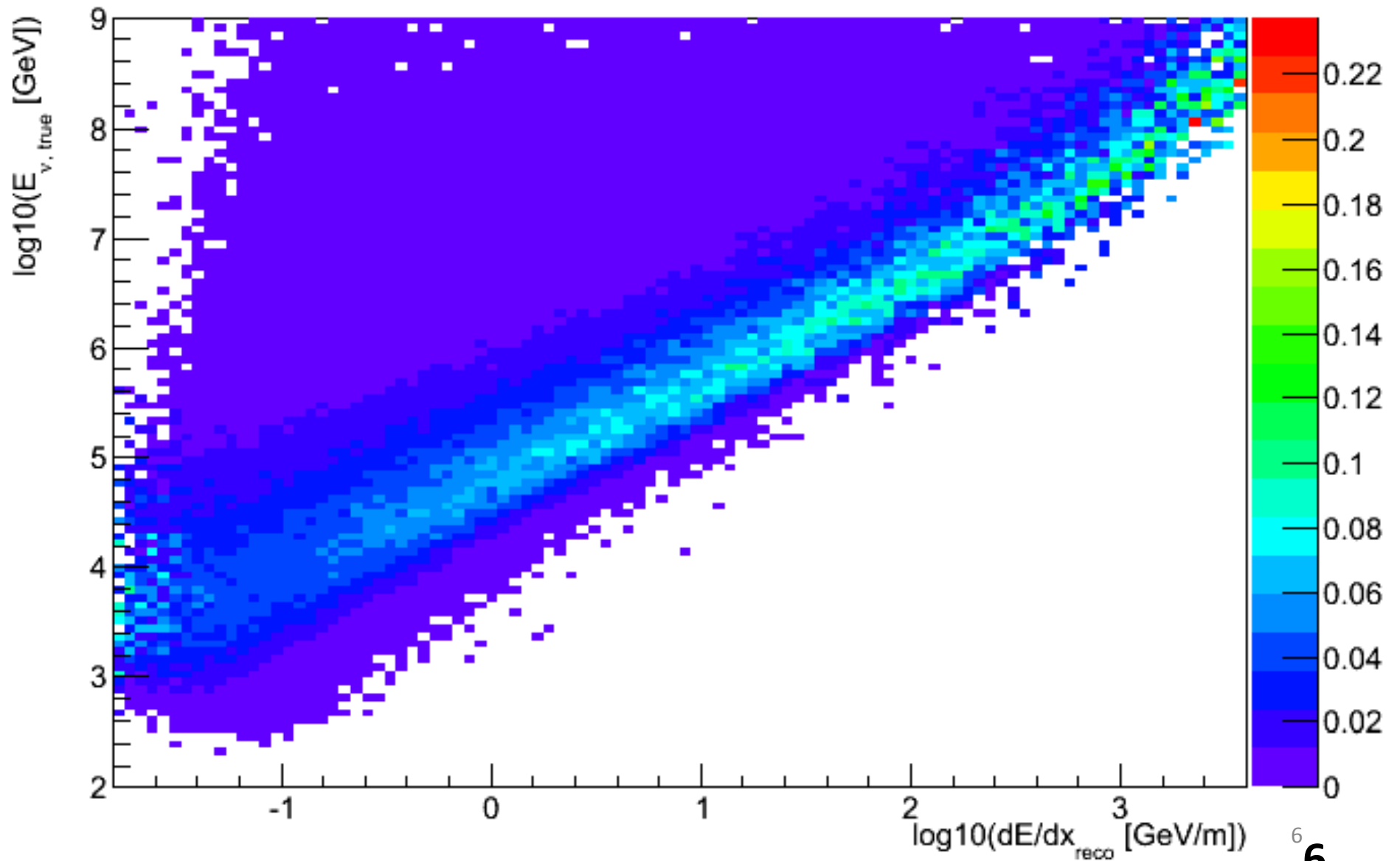
atmospheric muon_allsky



atmospheric neutrino_allsky

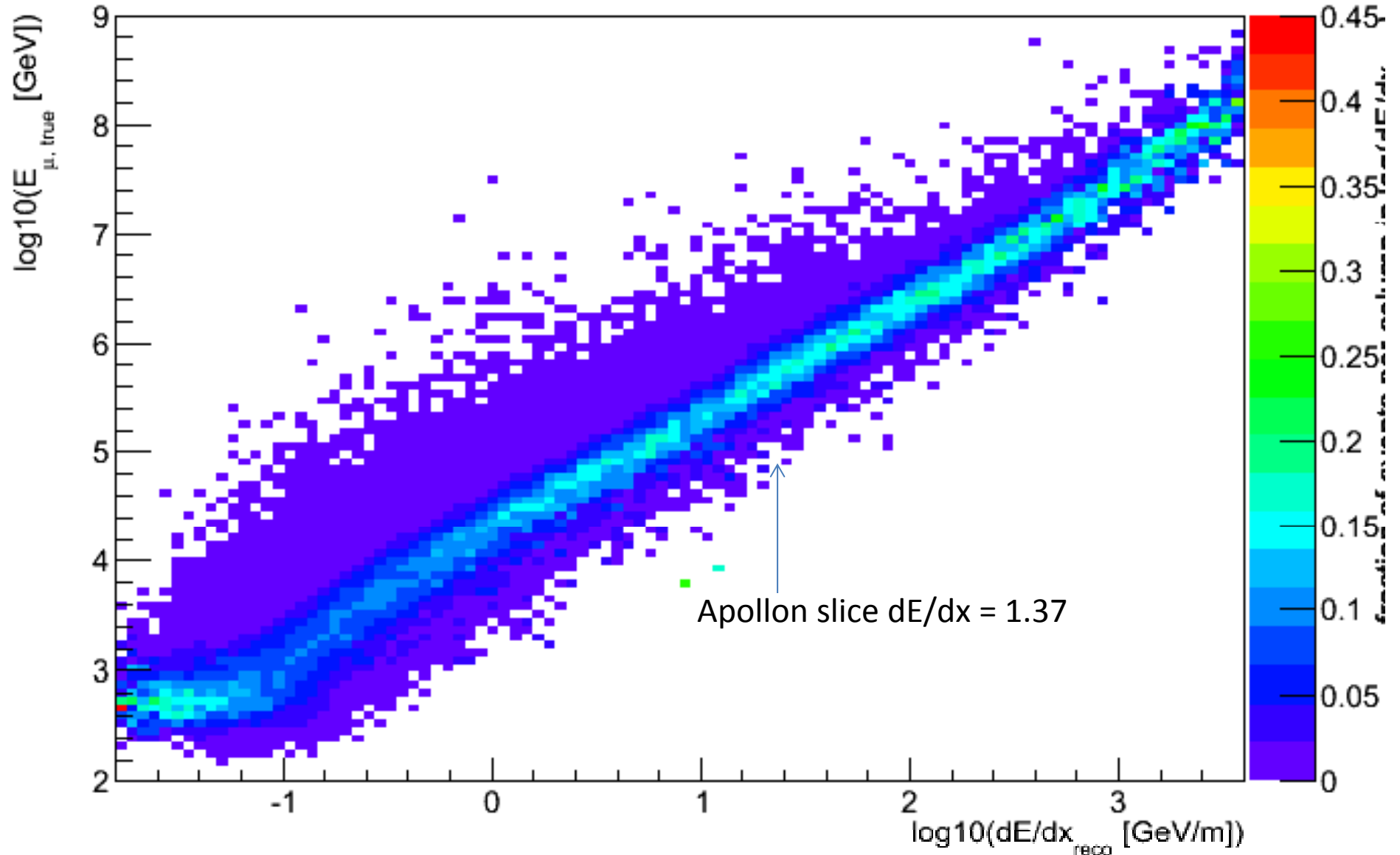


signal neutrino_allsky



- The choice of flux makes a difference:
- For Apollon:
 - 420 vs 580 TeV muon energy (median)
 - 600 TeV vs 2 PeV neutrino energy (median)
 - How do we choose?
- We don't have to choose:
- We know more: we fitted for both fluxes, so use the final best fit mixture as the weighting
- Weight to the best fit mixture from IC59 analysis

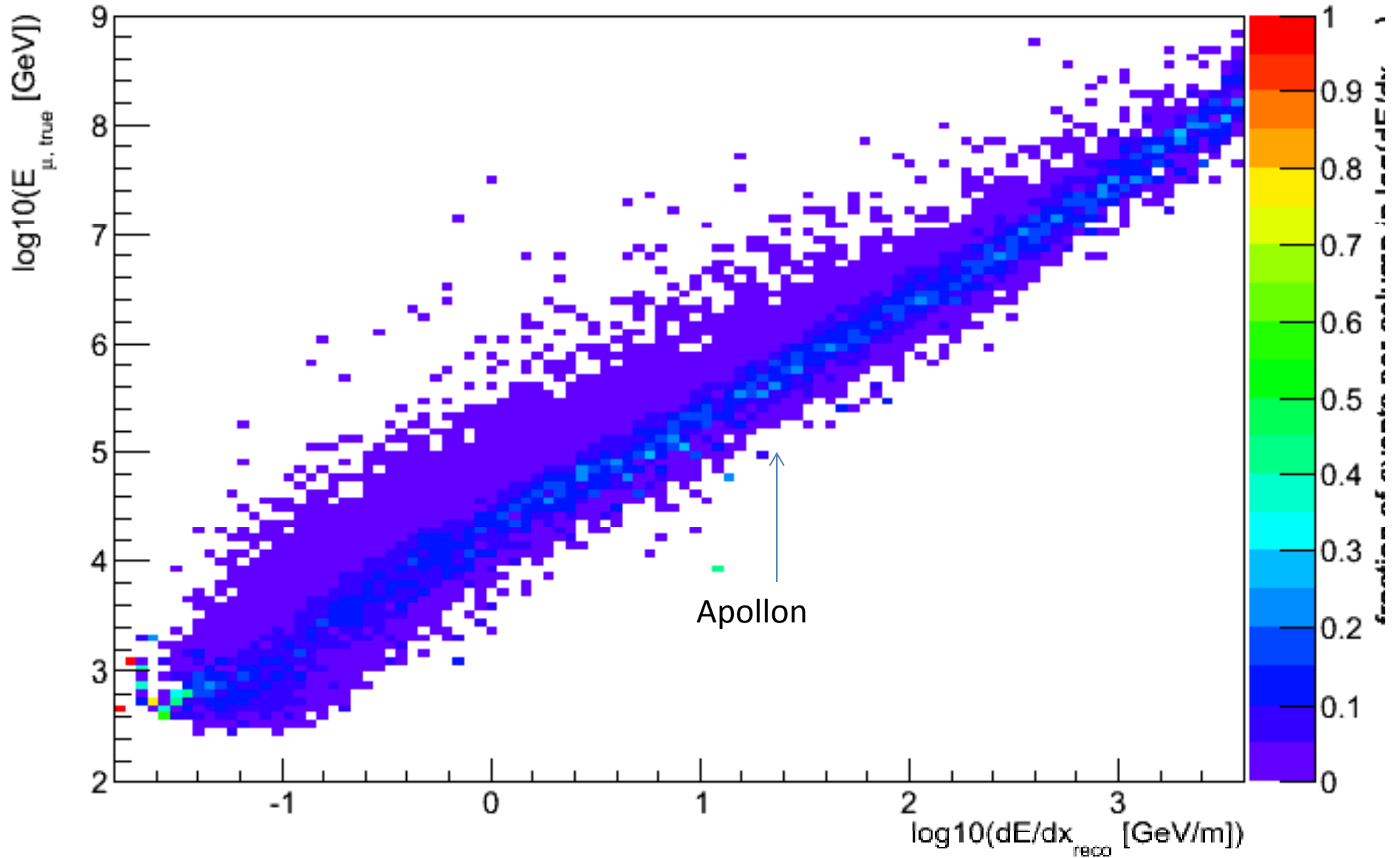
best_fit muon_allsky



Muon energy

Scatter plot for muons, weighted to best fit mixture of atmospheric and signal from the IC59 analysis

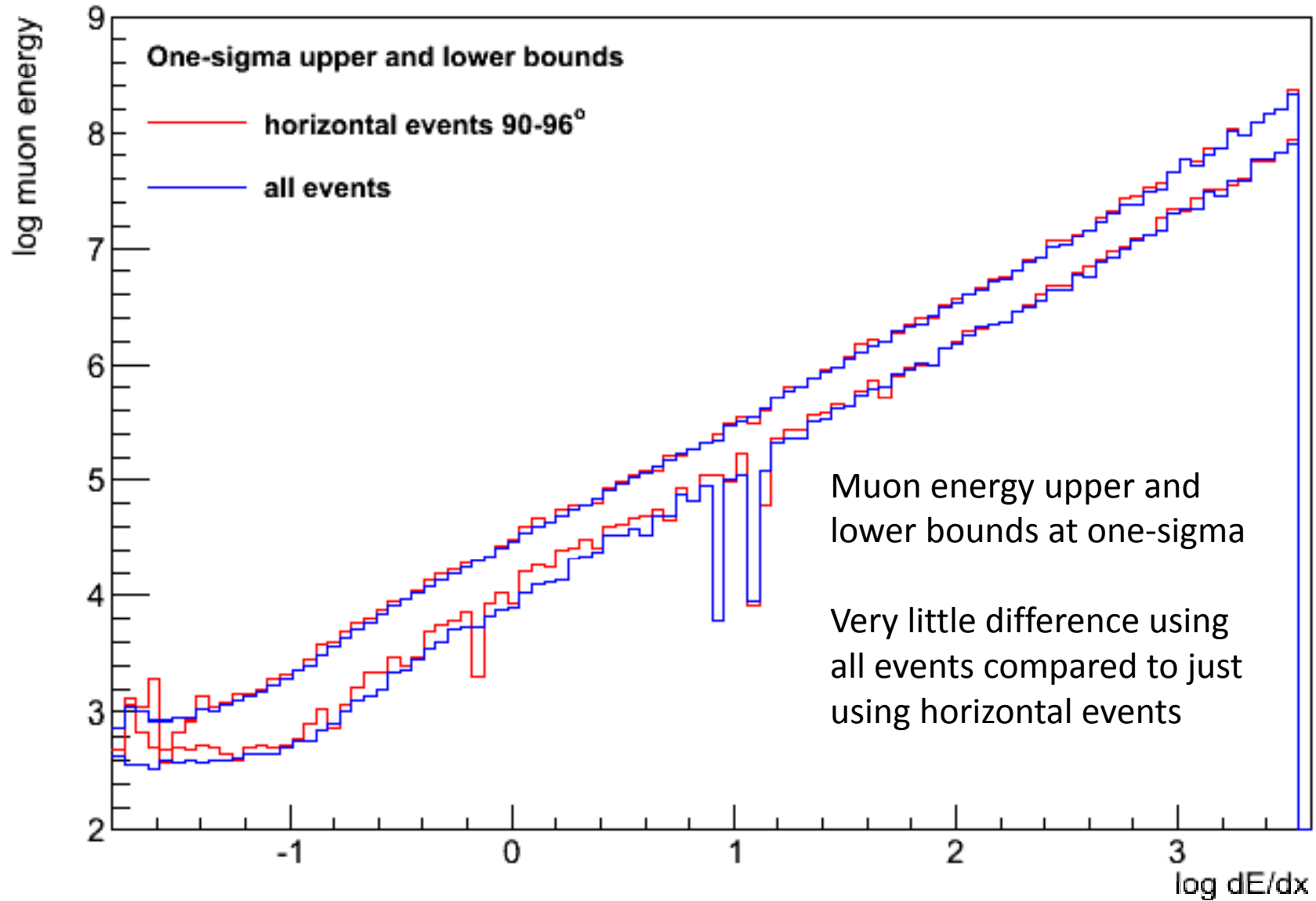
best_fit muon_horizon



Better: we know Apollon is near-horizontal, so take MC 90-96 degrees

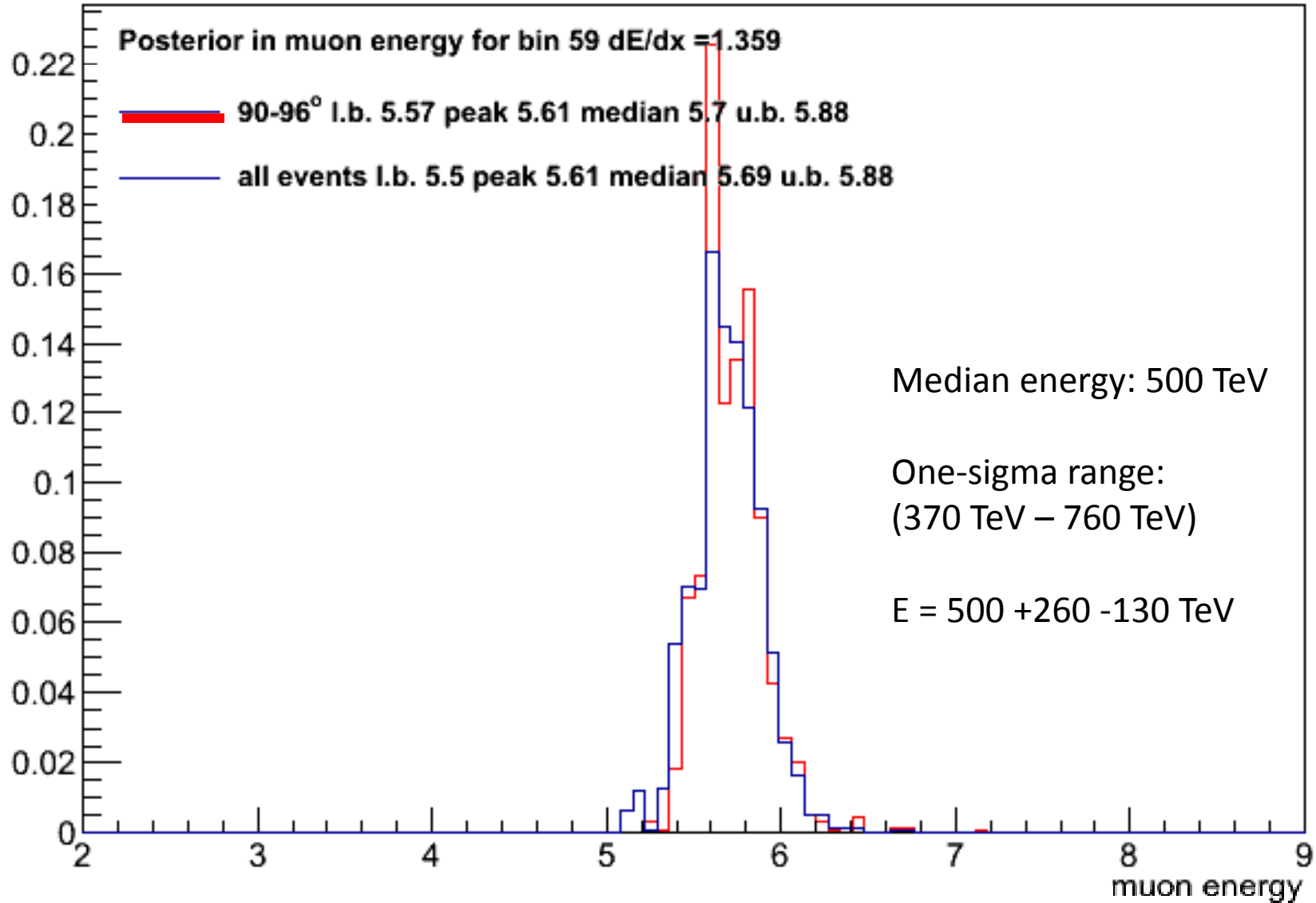
Muon energy

best_fit_muon_bounds



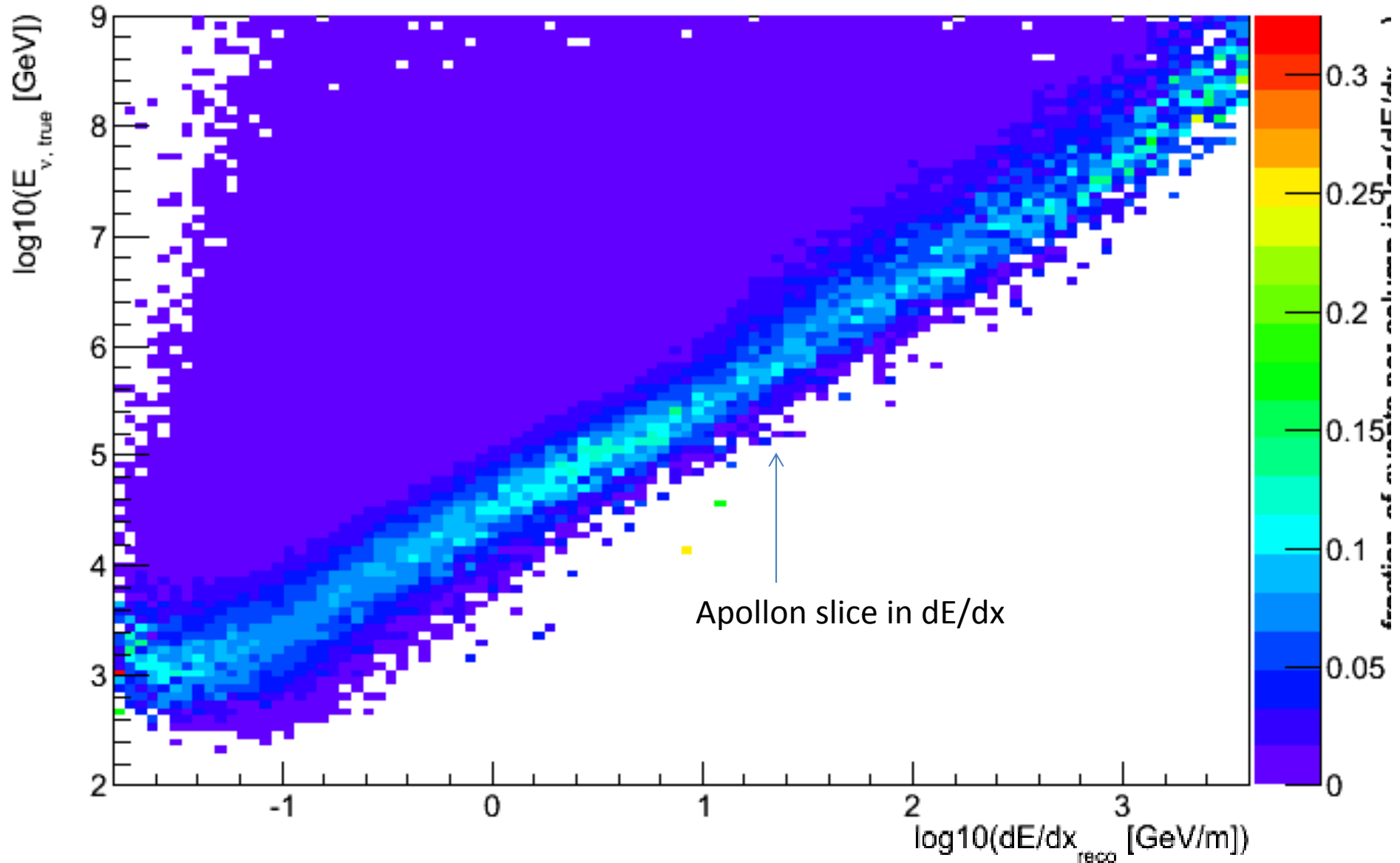
Muon energy

best_fit_muon_posterior for Apollon



Muon energy

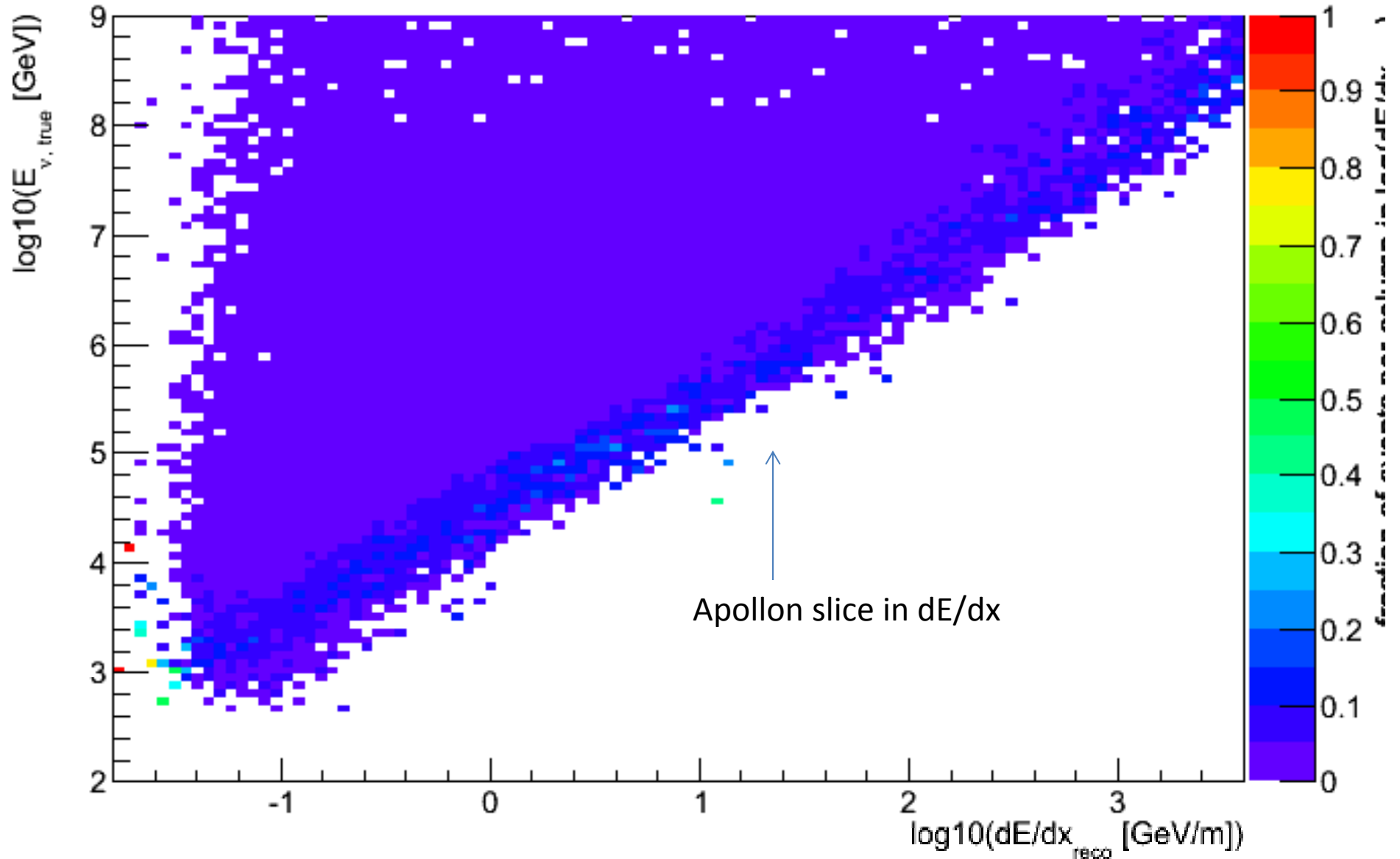
best_fit neutrino_allsky



Neutrino energy, all arrival directions

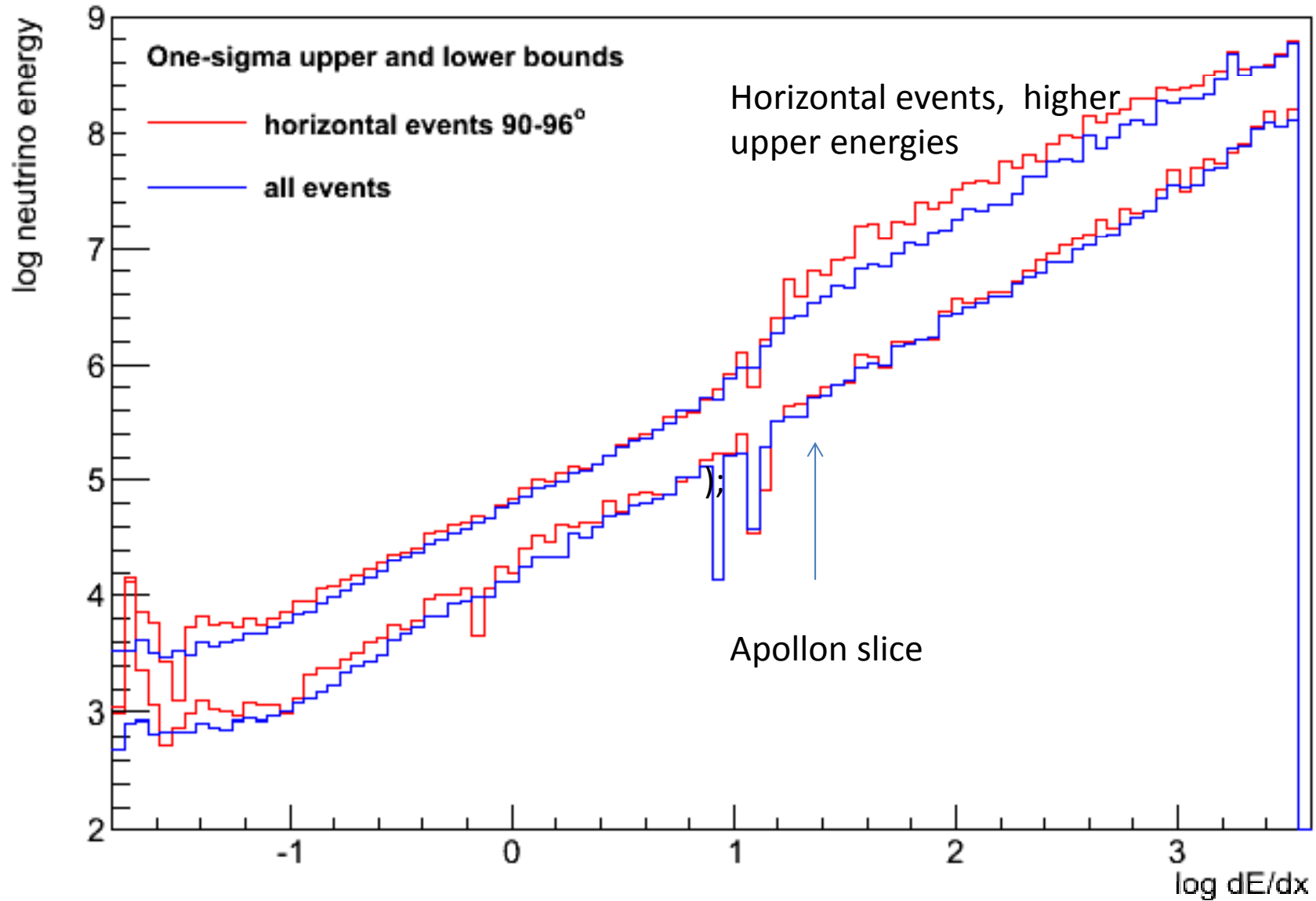
Exclude muon-less events

best_fit neutrino_horizon



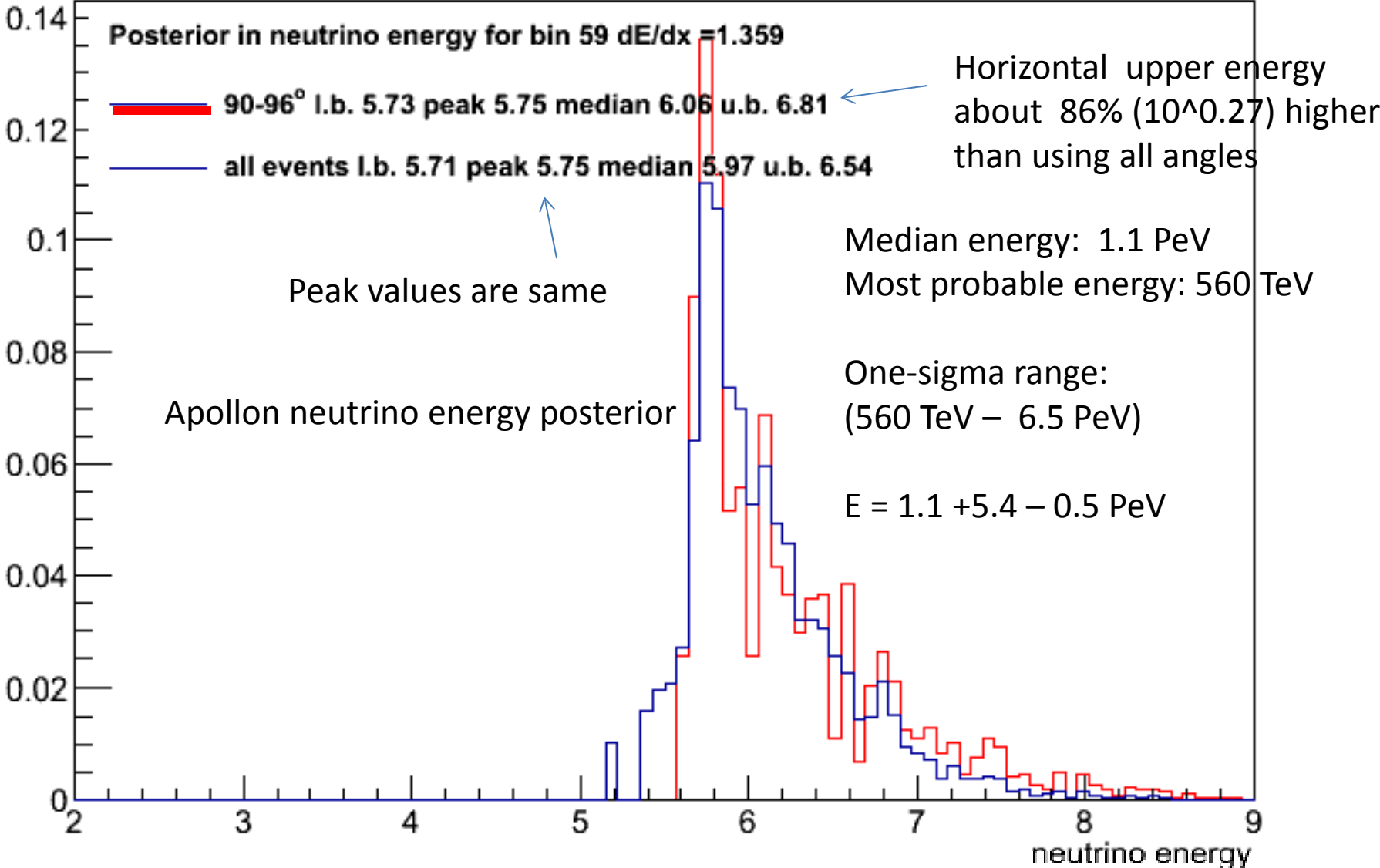
Neutrino energy, horizontal directions

best_fit_neutrino_bounds



See slightly higher neutrino energies for horizontal vs all directions – earth absorption

best_fit_neutrino_posterior



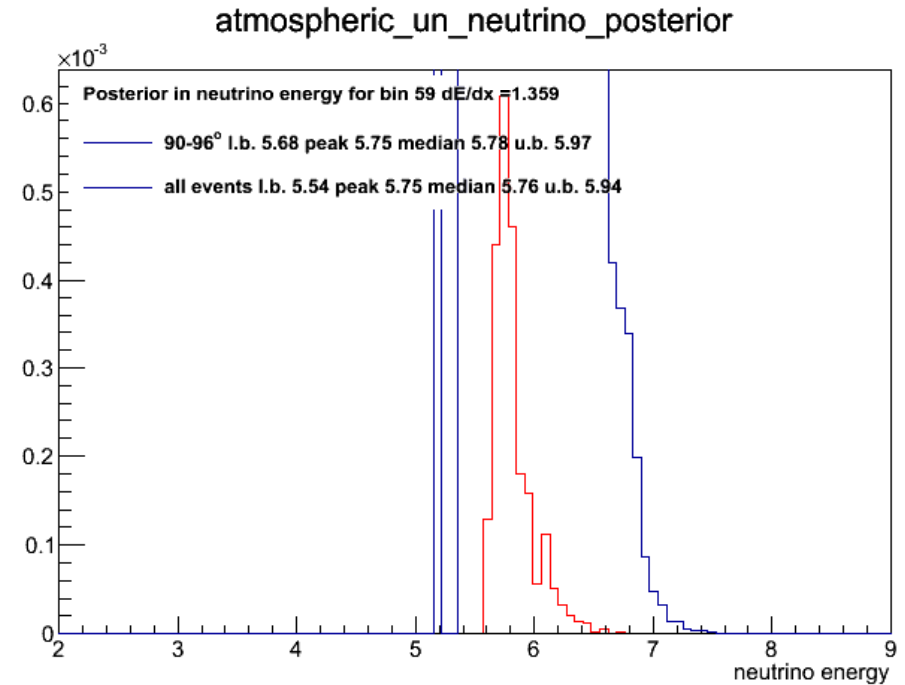
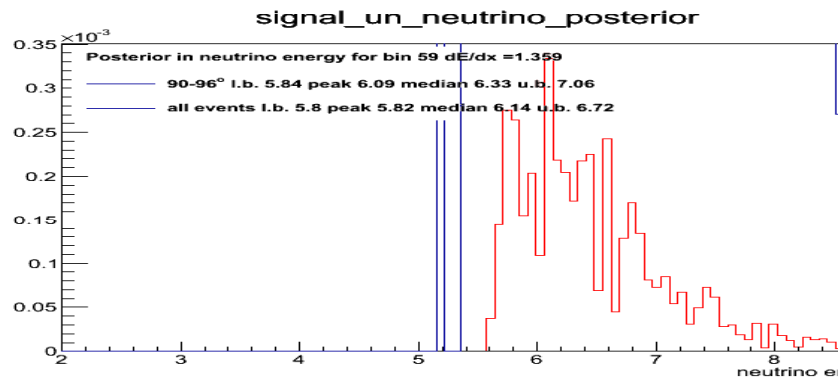
Neutrino energy

However...

- Apollon is either an atmospheric neutrino, or it is a signal neutrino
- The fit tells us the odds of each, for the observed dE/dx , and knowing it is horizontal

Look at the posteriors for each case

Areas are in the ratio 2:1 favouring a signal origin

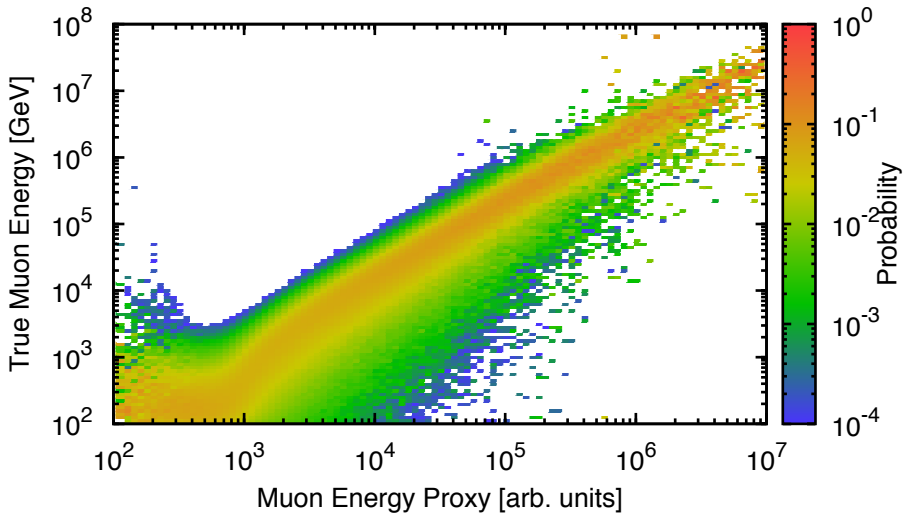


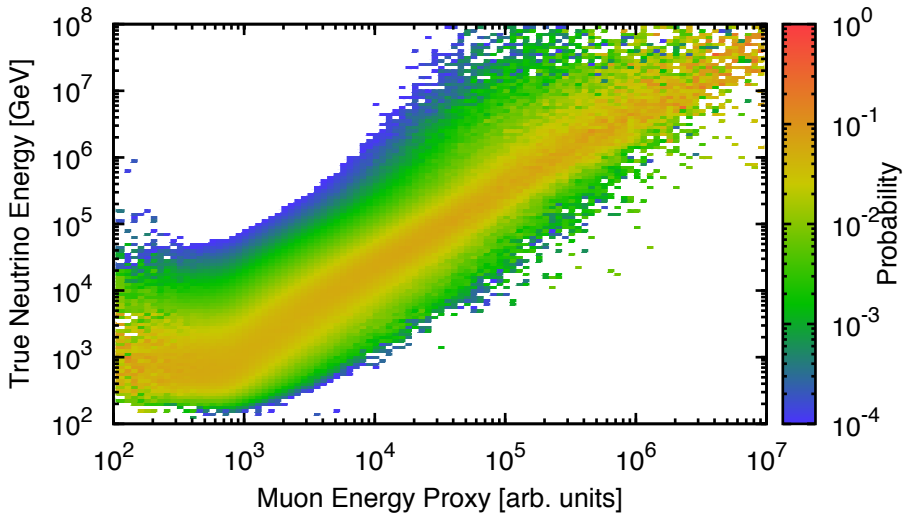
If it's signal (2/3), it's likely more than 1 PeV.

If it's atmospheric (1/3), it's likely less than 1 PeV.

Conclusions

- Apollon is likely higher energy than we have previously thought... sometimes quoting 100-300 TeV neutrino energy...
- 500 TeV muon energy
- 1 PeV neutrino energy
- All numbers are preliminary



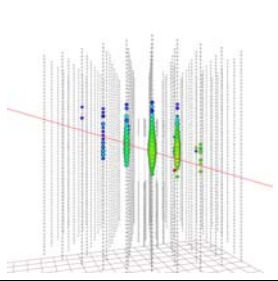
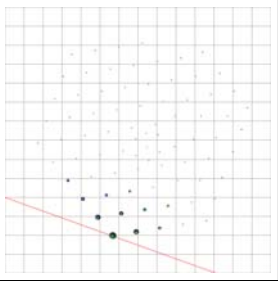
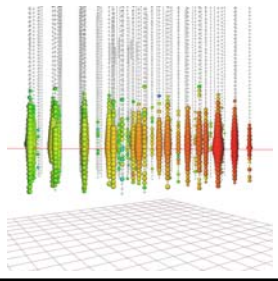
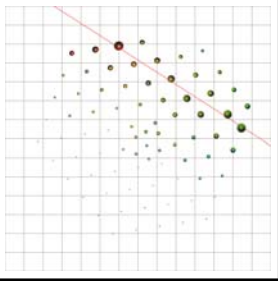
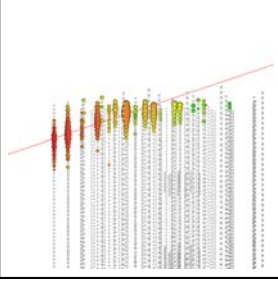
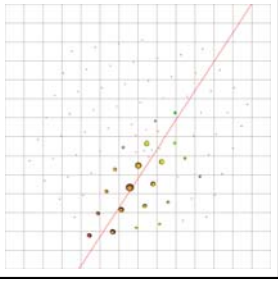


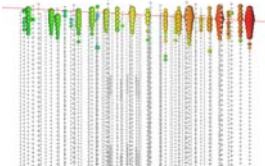
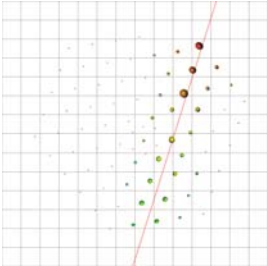
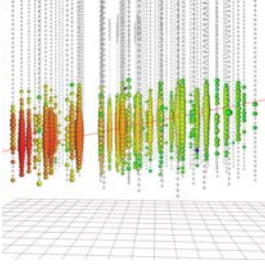
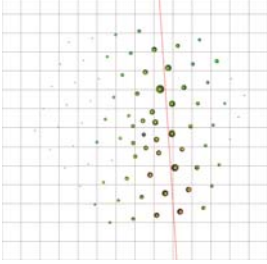
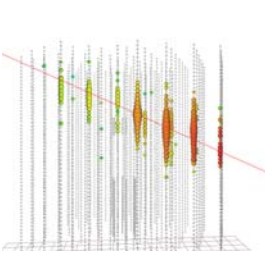
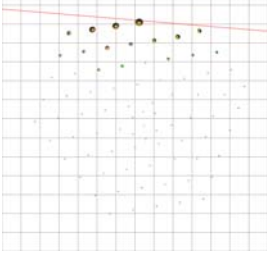
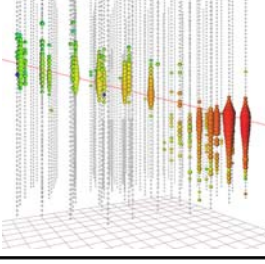
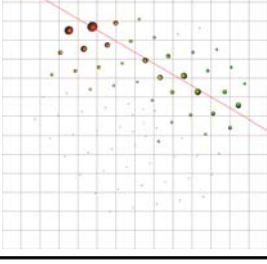
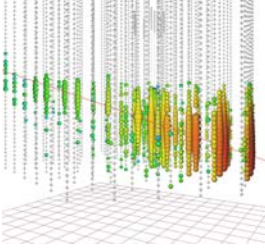
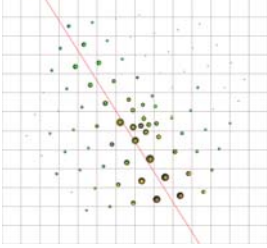
APPENDIX

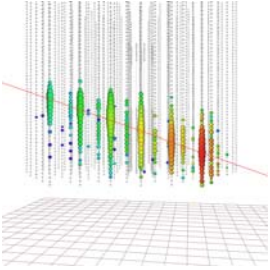
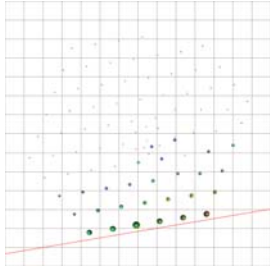
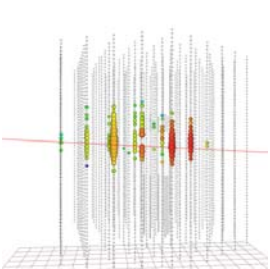
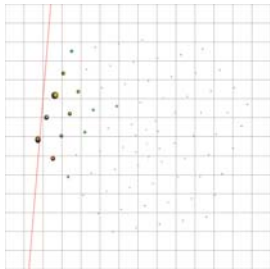
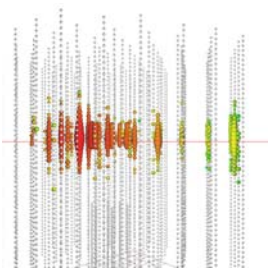
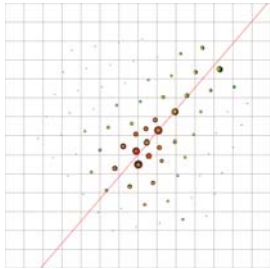
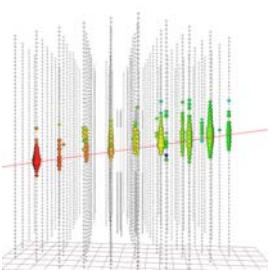
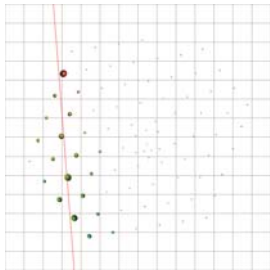
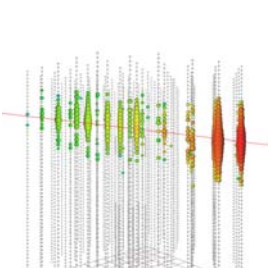
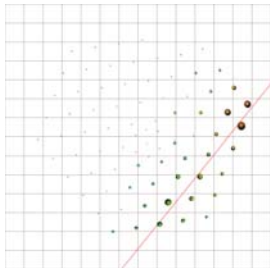
Details of Observed Events

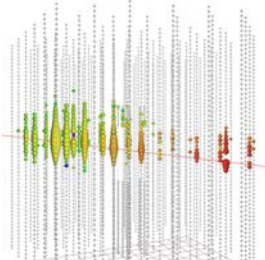
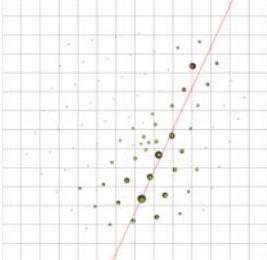
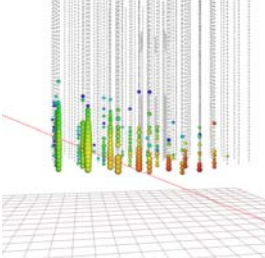
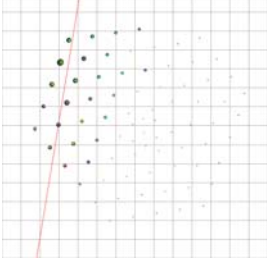
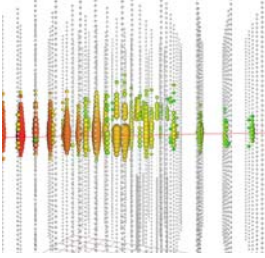
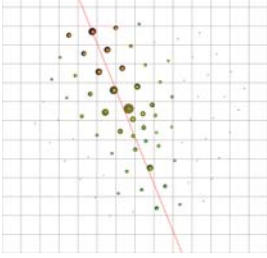
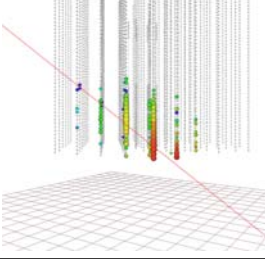
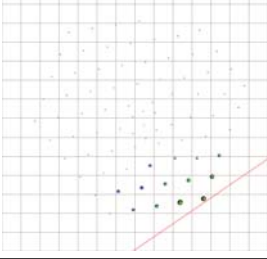
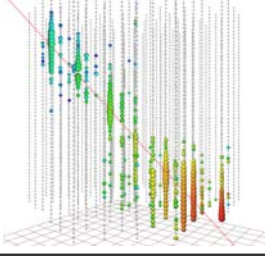
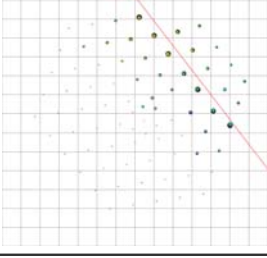
Table A.1 lists the 20 events which were found in the final data sample, which are those with energy proxy values greater than 50 TeV. The grid shown in the top view of the events has a spacing of 100 meters.

Table A.1: Images of observed events with highest estimated energies

Name	Side view	Top View	Energy Proxy Zenith Angle (Comments)
Dr. Heinrich Faust			290.1 TeV 106.3°
Dr. Hari Seldon			199.8 TeV 90.2°
Dr. Susan Calvin			160.2 TeV 108.7°

Captain Nemo			147.4 TeV 93.1°
Dr. Richard Seaton			141.9 TeV 101.1°
Dr. Henry Jekyll			133.2 TeV 114.1°
Dr. Emmet Brown			132.1 TeV 103.8°
Dr. Henry Walton Jones, Jr.			131.8 TeV 108.9°

Dr. Victor Frankenstein			109.3 TeV 106.3°
Dr. Strangelove			66.7 TeV 92.8°
Dr. Strangepork			65.1 TeV 90.0° (This is a clear starting event, also found by [103] and [104])
Impey Barbicane			64.0 TeV 97.9°
Professor Joseph Cavor			60.9 TeV 96.7°

Professor Abraham Van Helsing			60.2 TeV 96.8°
Dr. Moreau			55.2 TeV 112.0°
Buckaroo Banzai			53.4 TeV 91.0°
Dr. David Bowman			52.0 TeV 128.6°
Angus MacGyver			51.1 TeV 137.6°

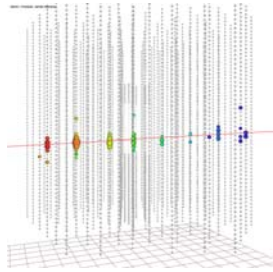
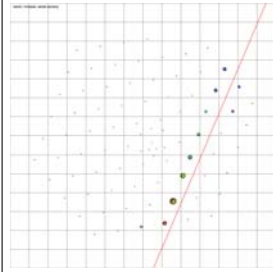
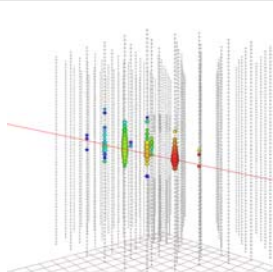
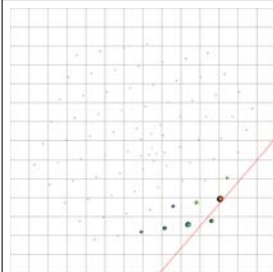
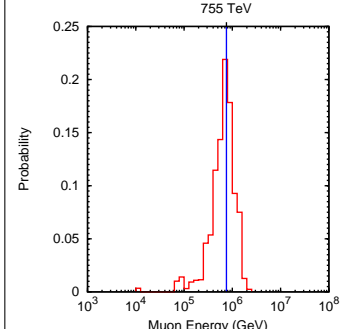
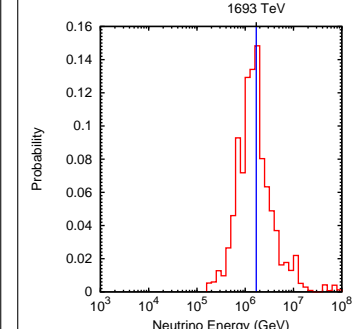
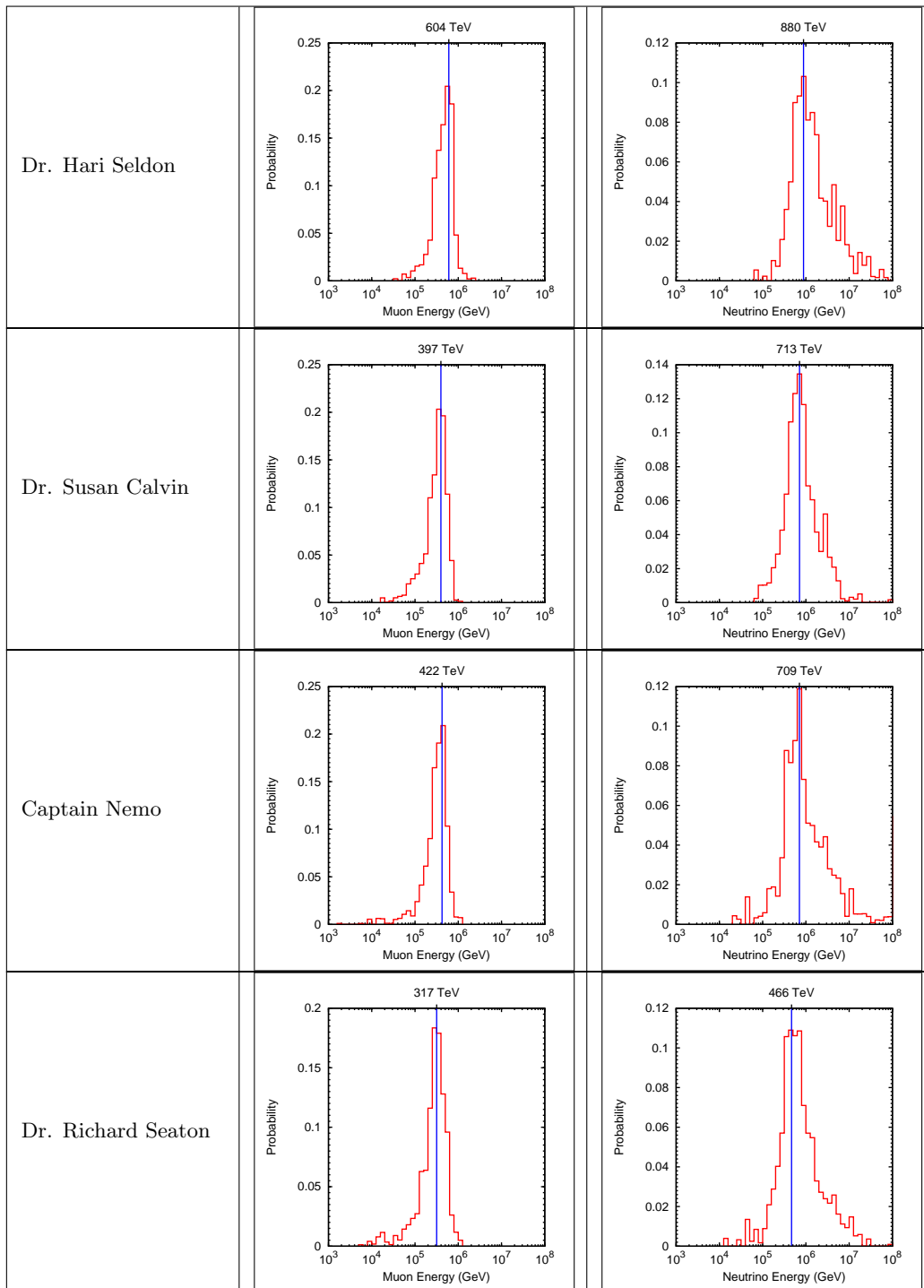
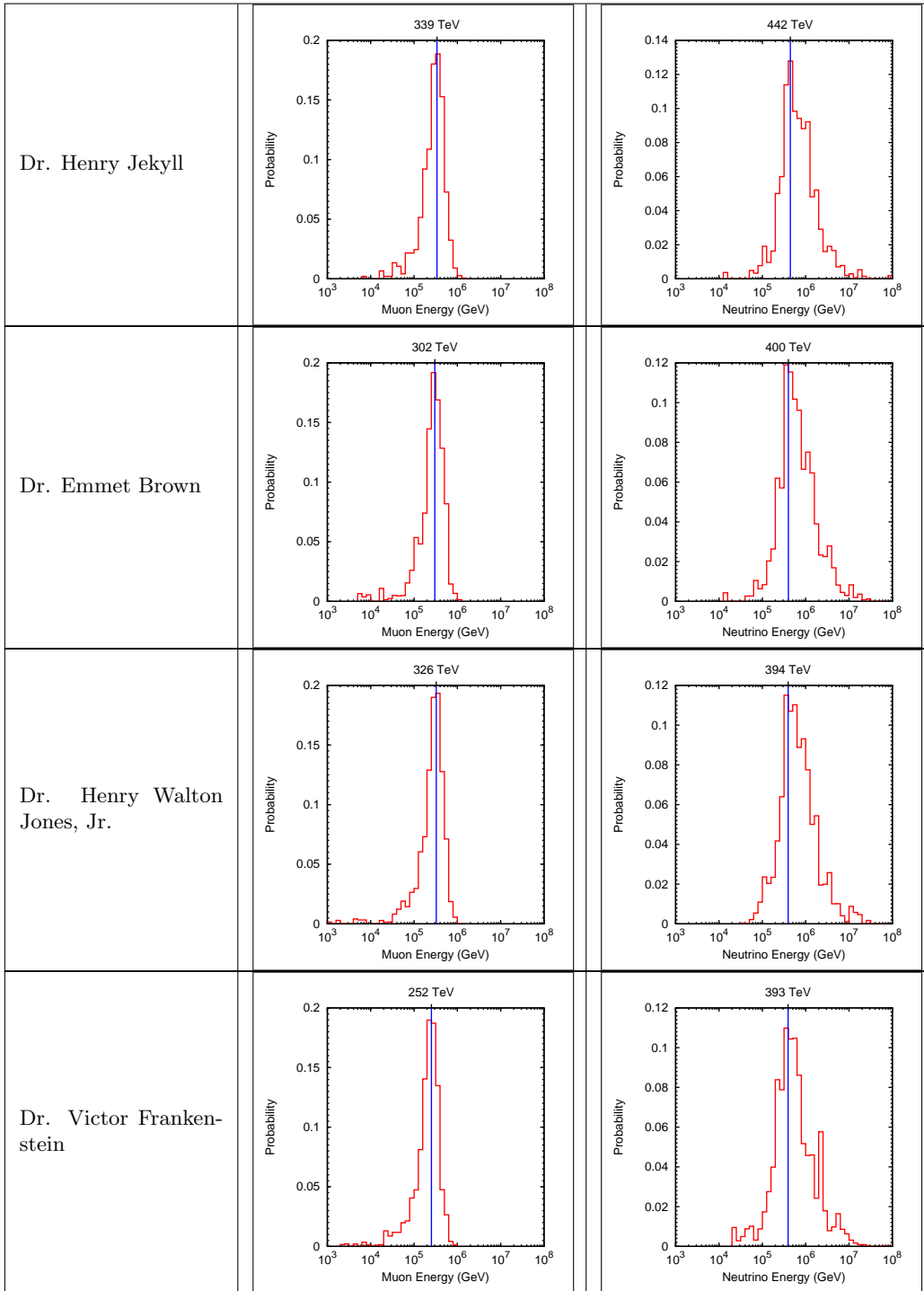
<p>Professor James Moriarty</p>			<p>50.7 TeV 93.2° (This event looks quite dim due to being entirely in the dust layer.)</p>
<p>Dr. Giacomo Rappacini</p>			<p>50.3 TeV 101.9°</p>

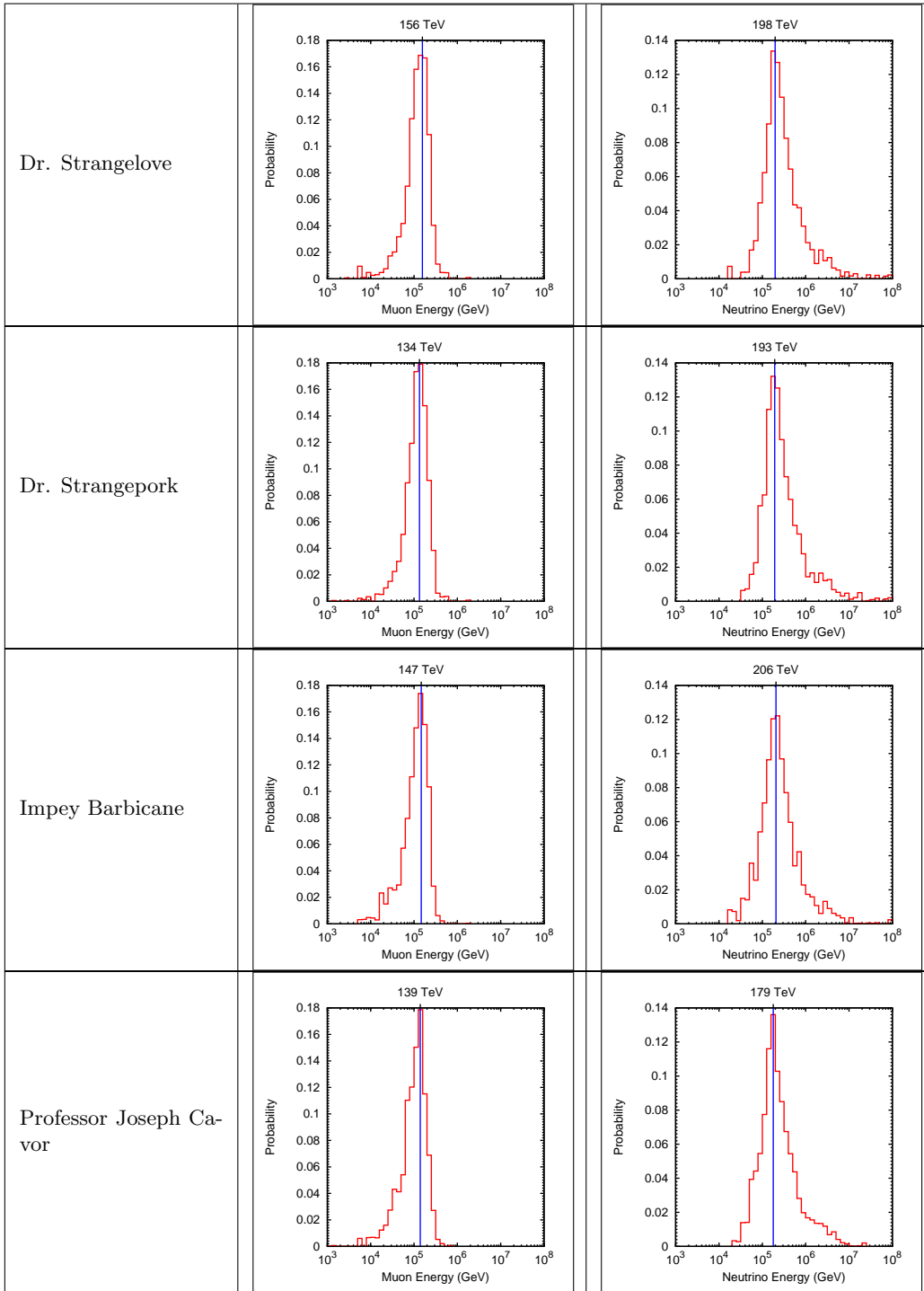
Table A.2 lists the same 20 events with their true muon and neutrino energy PDFs computed assuming the spectrum fitted in Section 6.3. Each PDF is constructed from the true properties of all simulated events which have energy proxies within 5% of the observed event’s value, and reconstructed zenith angles within 5° for events with energy proxies less than 100 TeV and 10° for those with larger energy proxies. Unfortunately, due to limited simulation statistics, many of these distributions still contain large fluctuations, particularly those with the highest energies. As a result, the estimated energies are rather imprecise and should be treated as ‘ballpark’ numbers only.

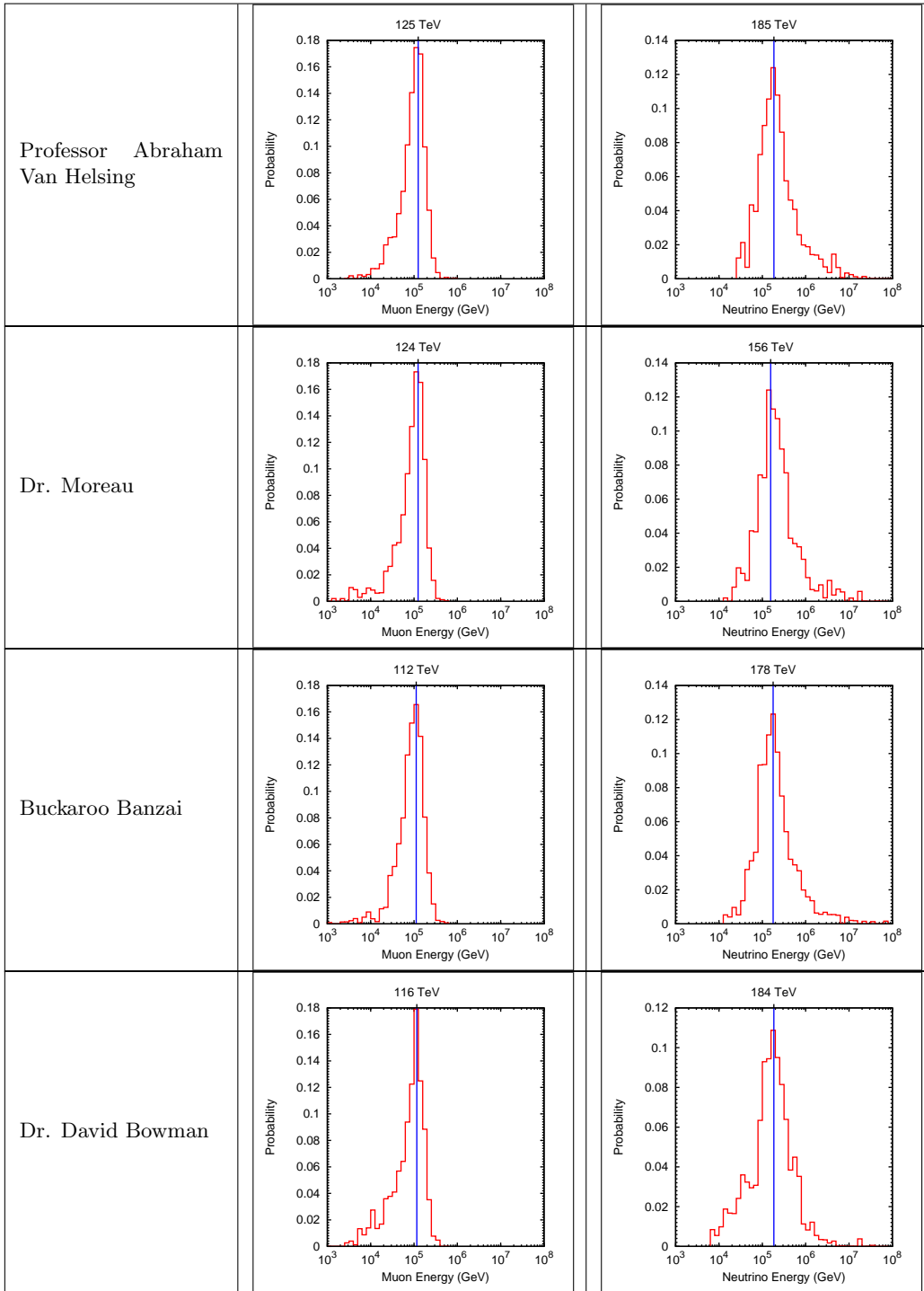
Table A.2: Inferred event energies

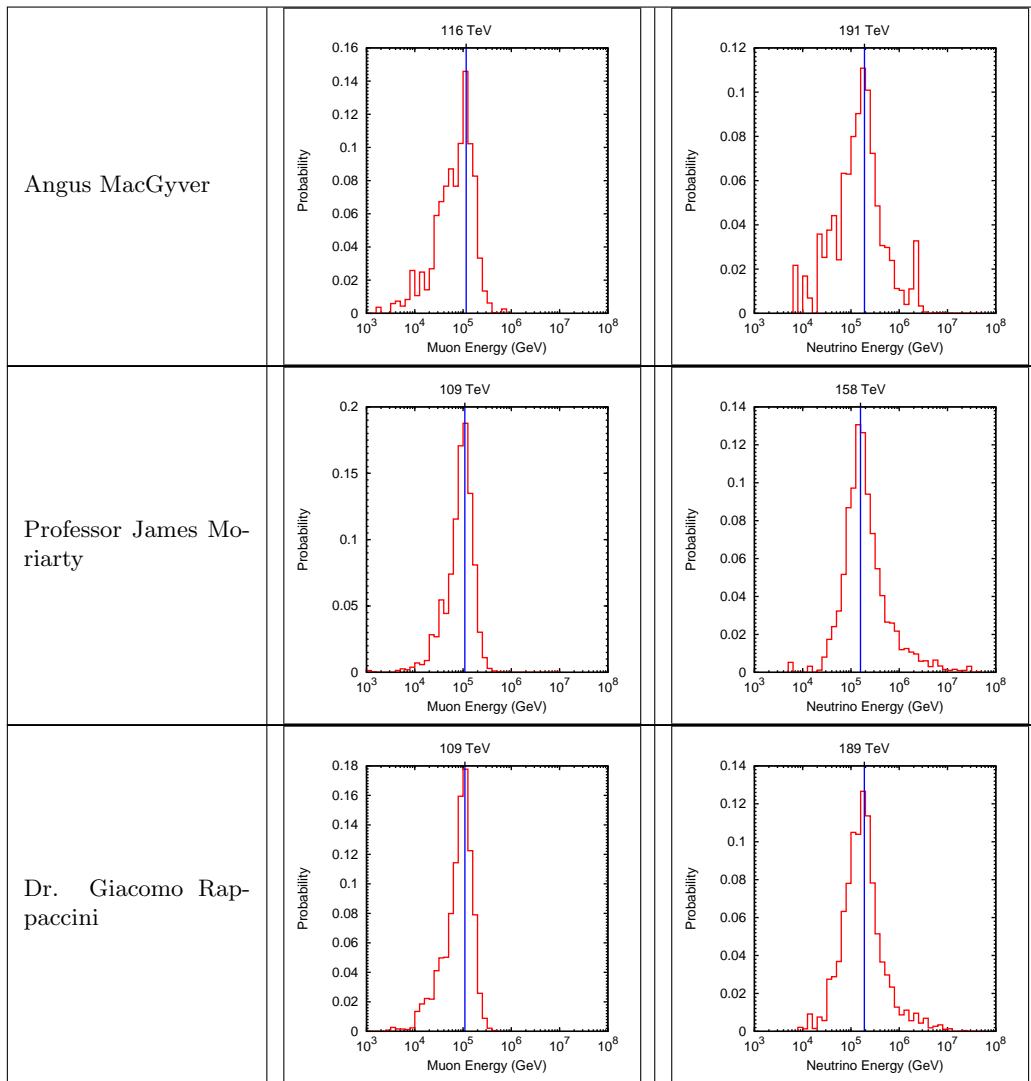
Name	Probable muon energy	Probable neutrino energy
<p>Dr. Heinrich Faust</p>	 <p>755 TeV</p>	 <p>1693 TeV</p>











APPENDIX

LIST OF REFERENCES

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