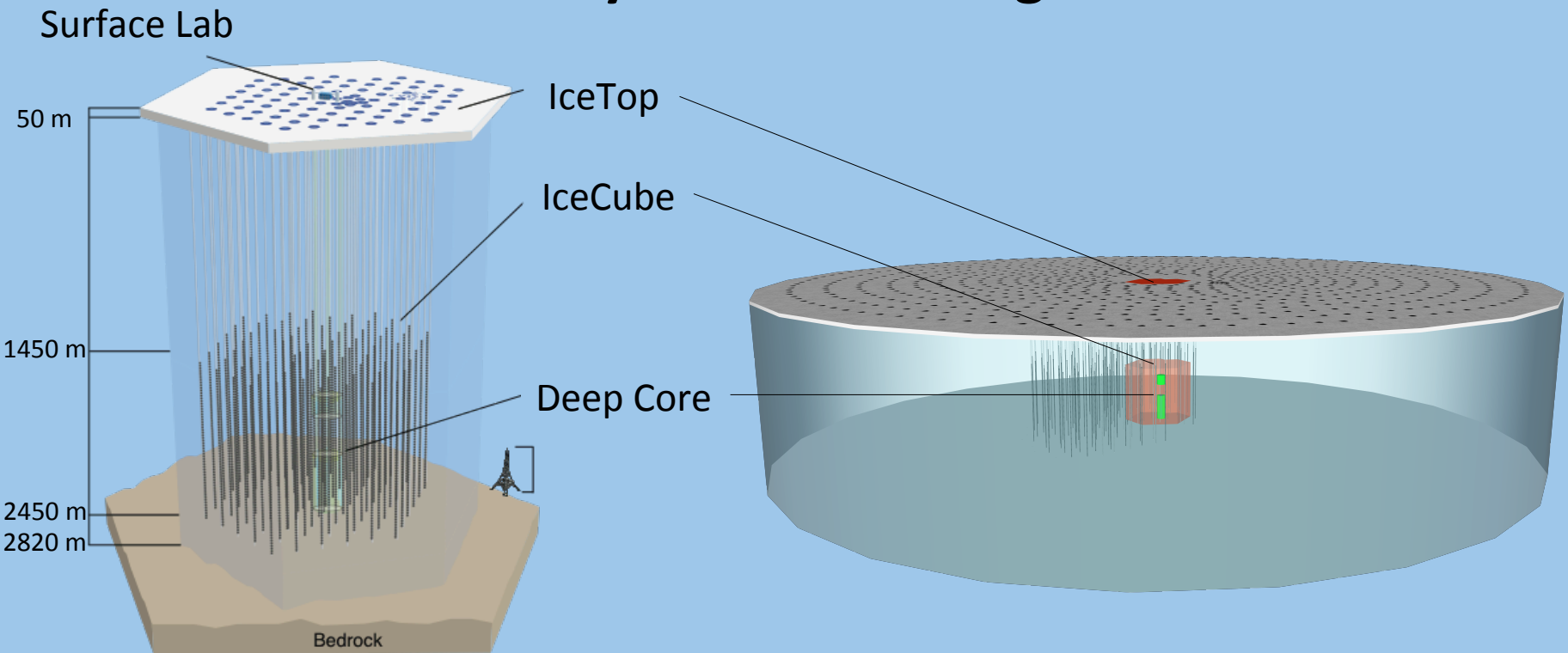


IceVeto

A high-energy extension for IceCube.

by Jan Auffenberg



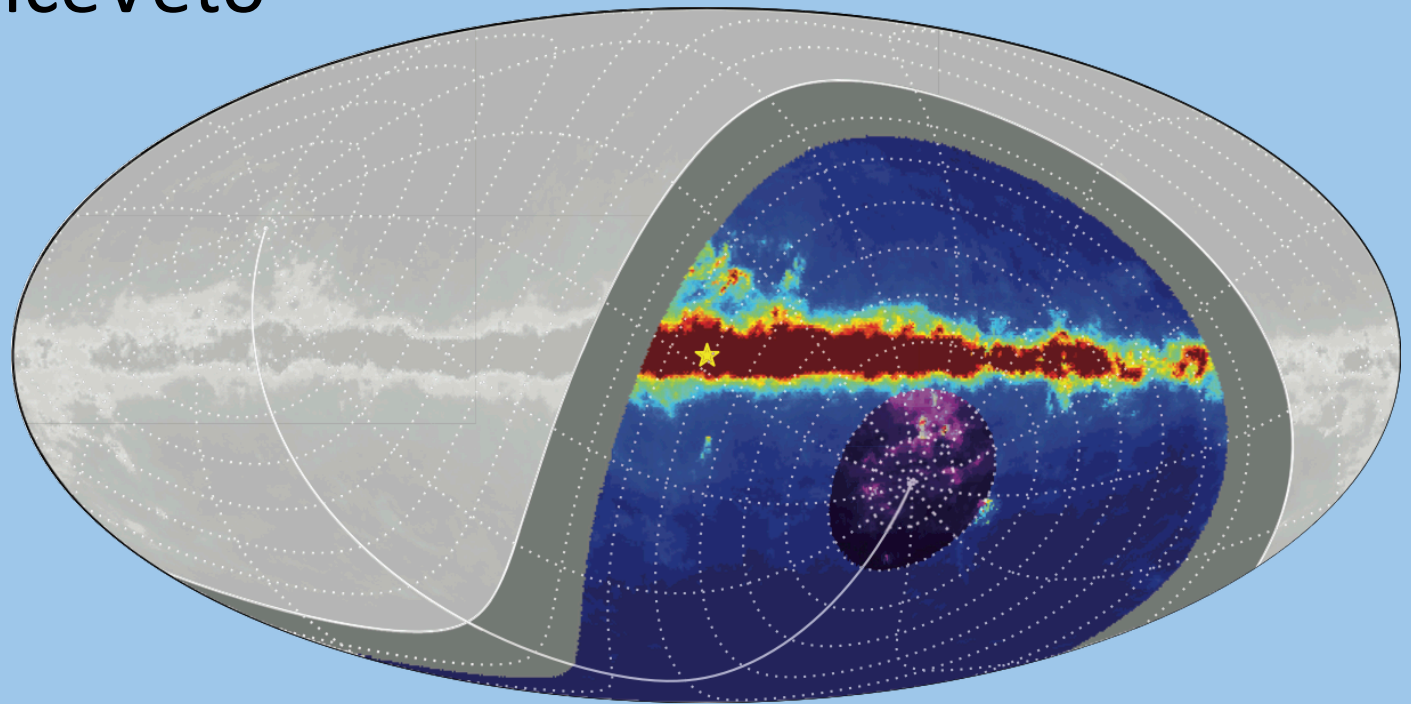
- Where do **high-energy cosmic-rays** come from?

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- Where do **astrophysical neutrinos** point?

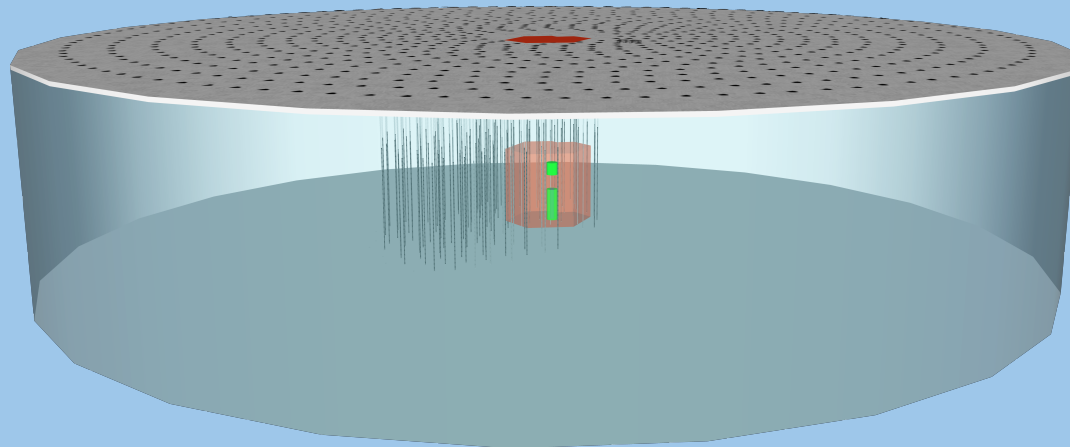
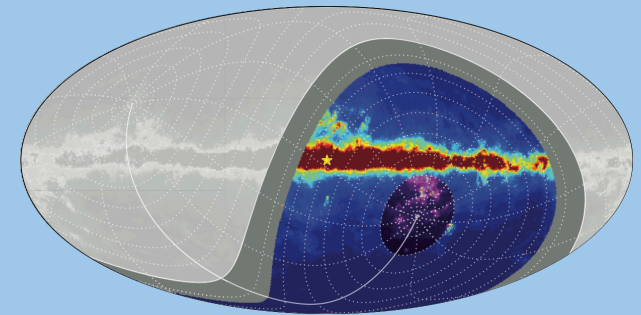
R&D Proposal: IceVeto



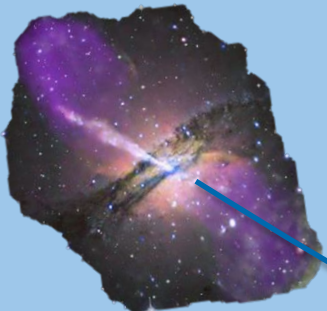
- Where do **high-energy cosmic-rays** come from?
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- R&D of IceVeto
 - Why



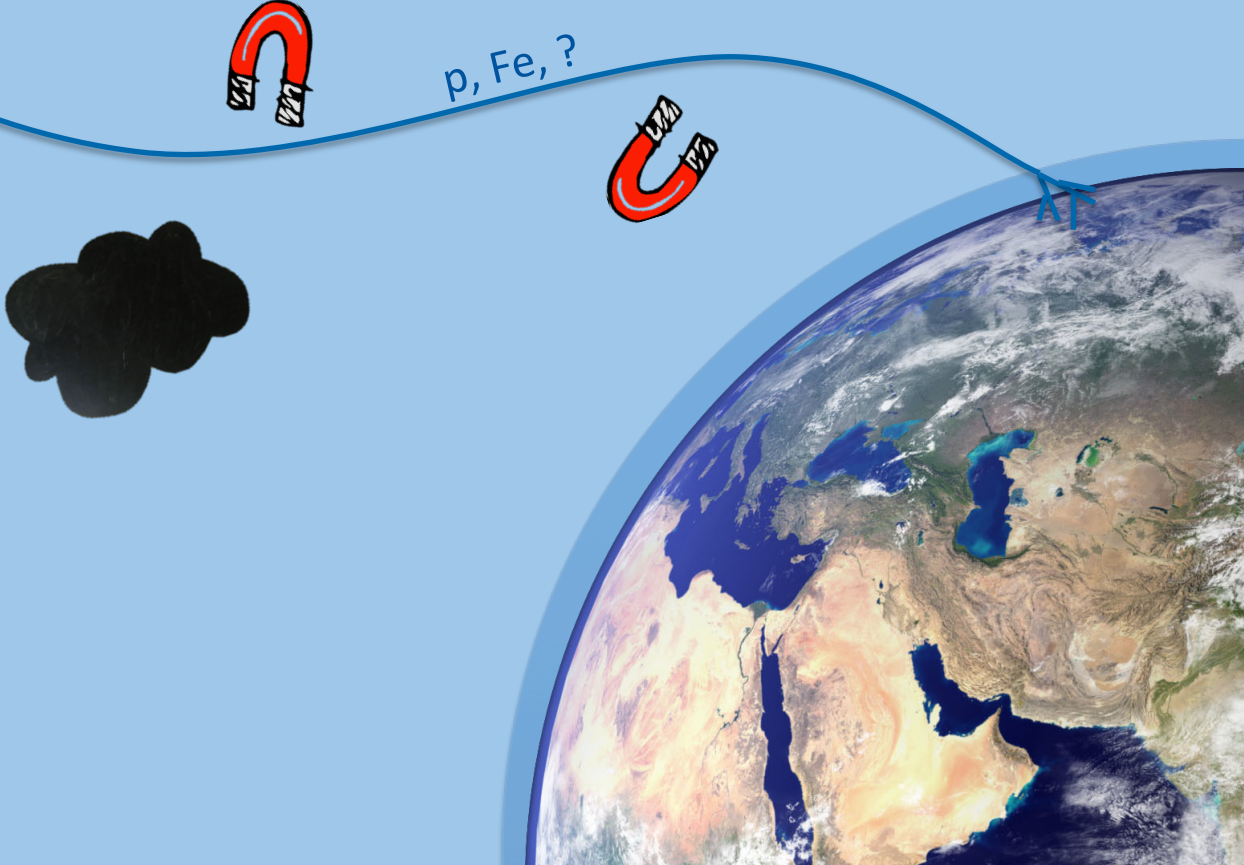
- Where do **high-energy cosmic-rays** come from?
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Why astrophysical neutrino search?



Cosmic-Rays: Unknown origin as they get bent in magnetic fields

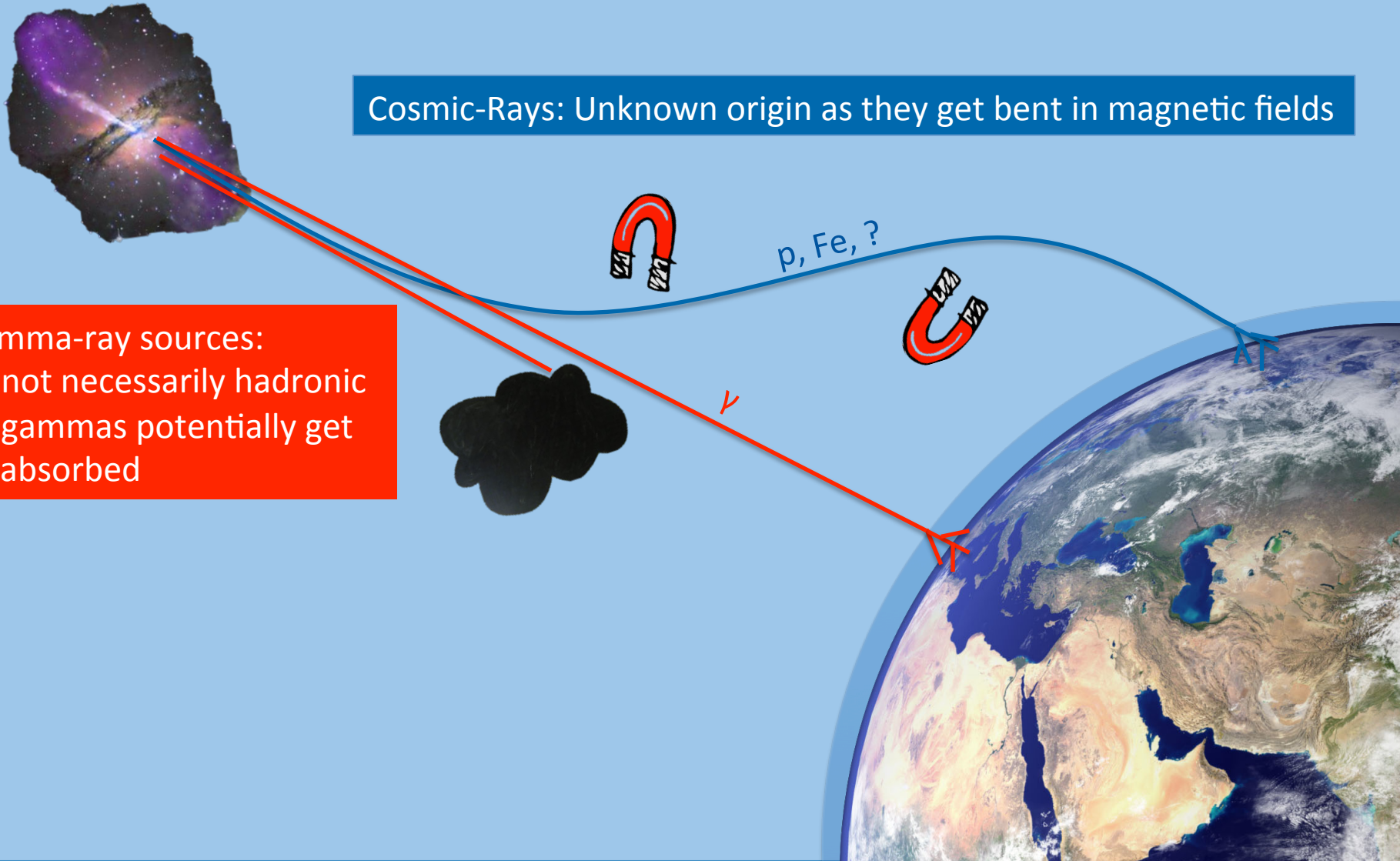


Why astrophysical neutrino search?

Cosmic-Rays: Unknown origin as they get bent in magnetic fields

Gamma-ray sources:

- not necessarily hadronic
- gammas potentially get absorbed



Why astrophysical neutrino search?

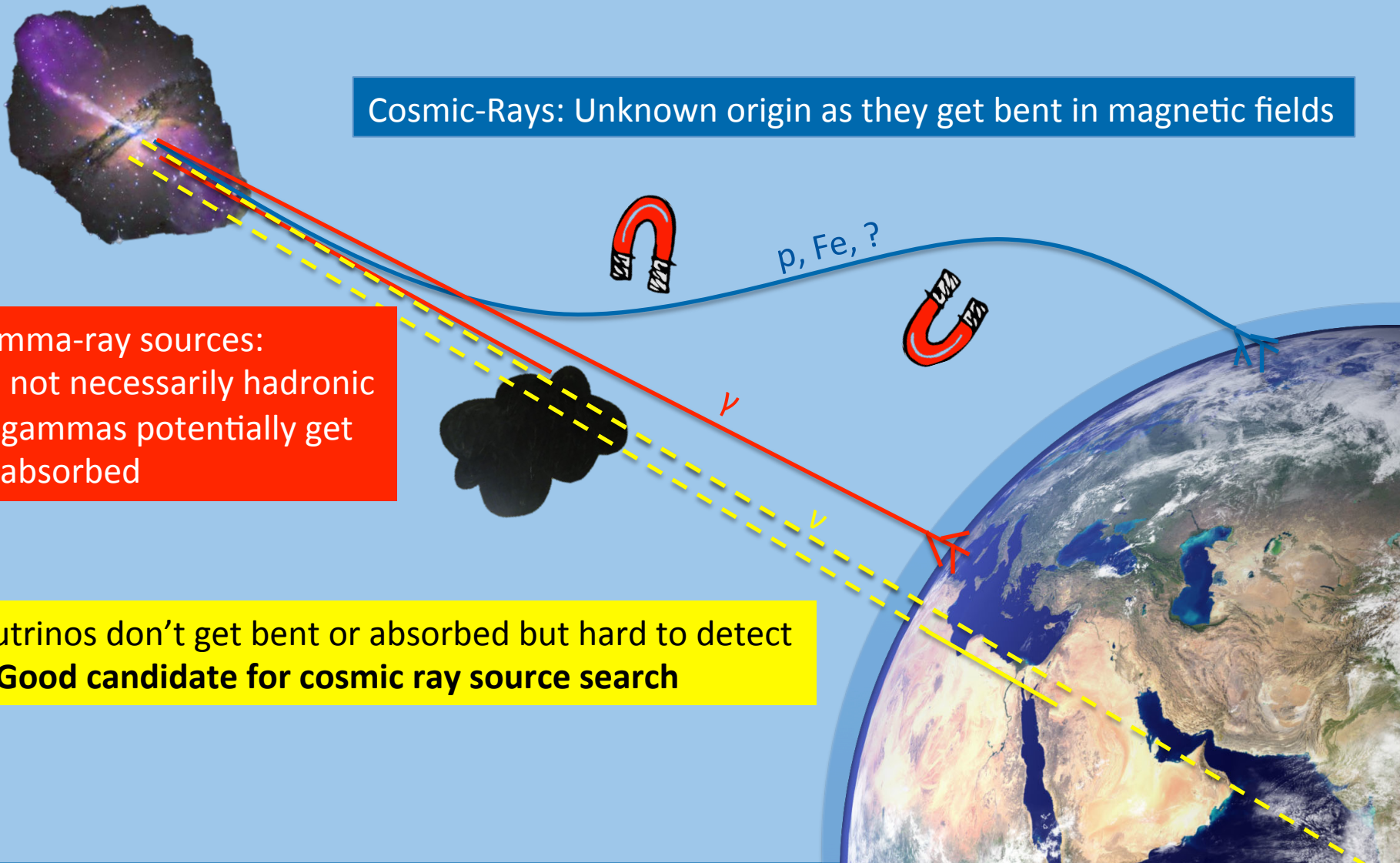
Cosmic-Rays: Unknown origin as they get bent in magnetic fields

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Neutrinos don't get bent or absorbed but hard to detect

- **Good candidate for cosmic ray source search**



Why astrophysical neutrino search?

Cosmic-Rays: Unknown origin as they get bent in magnetic fields

Gamma-ray sources:

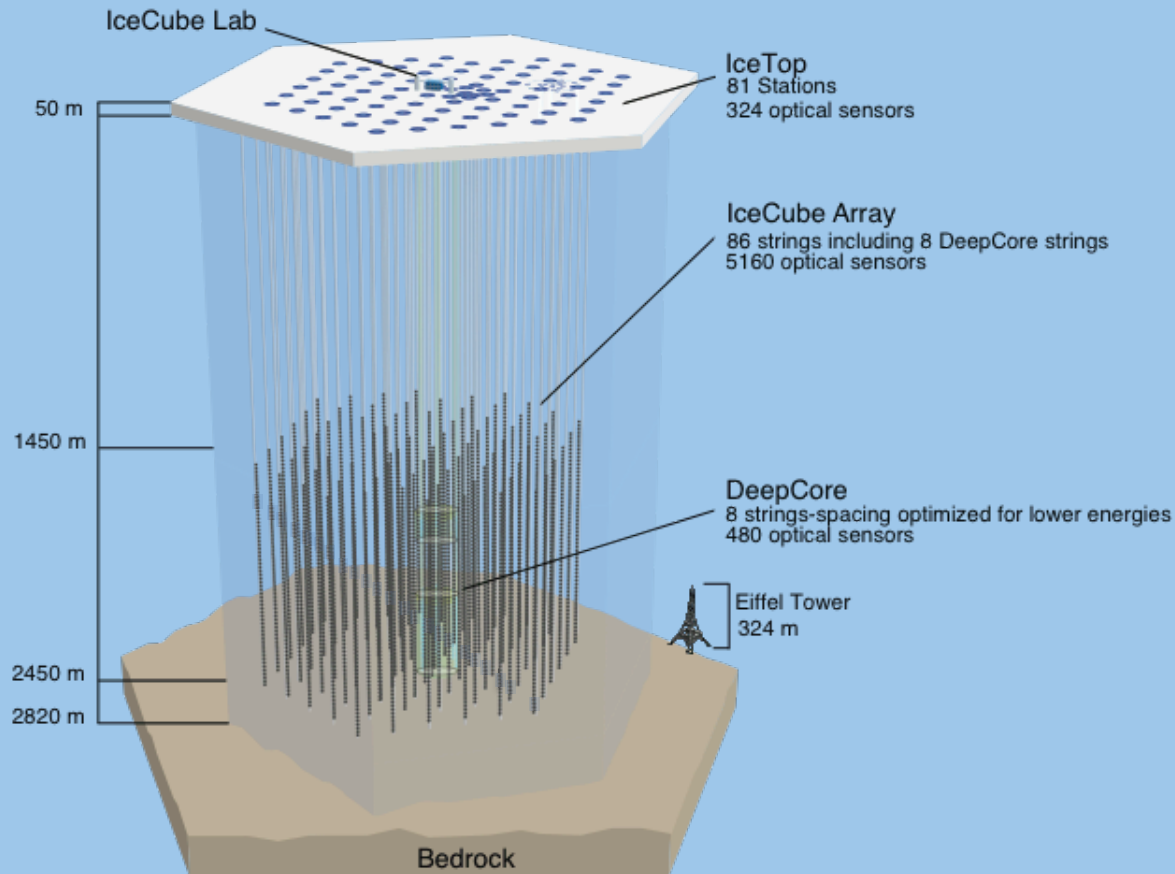
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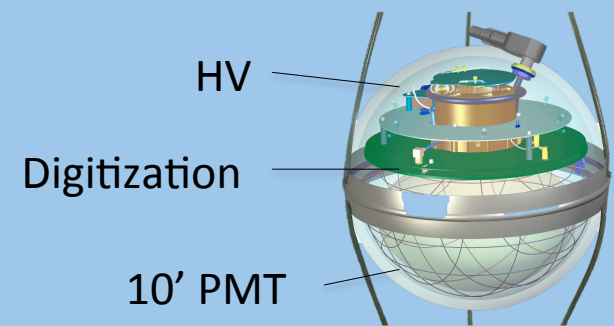
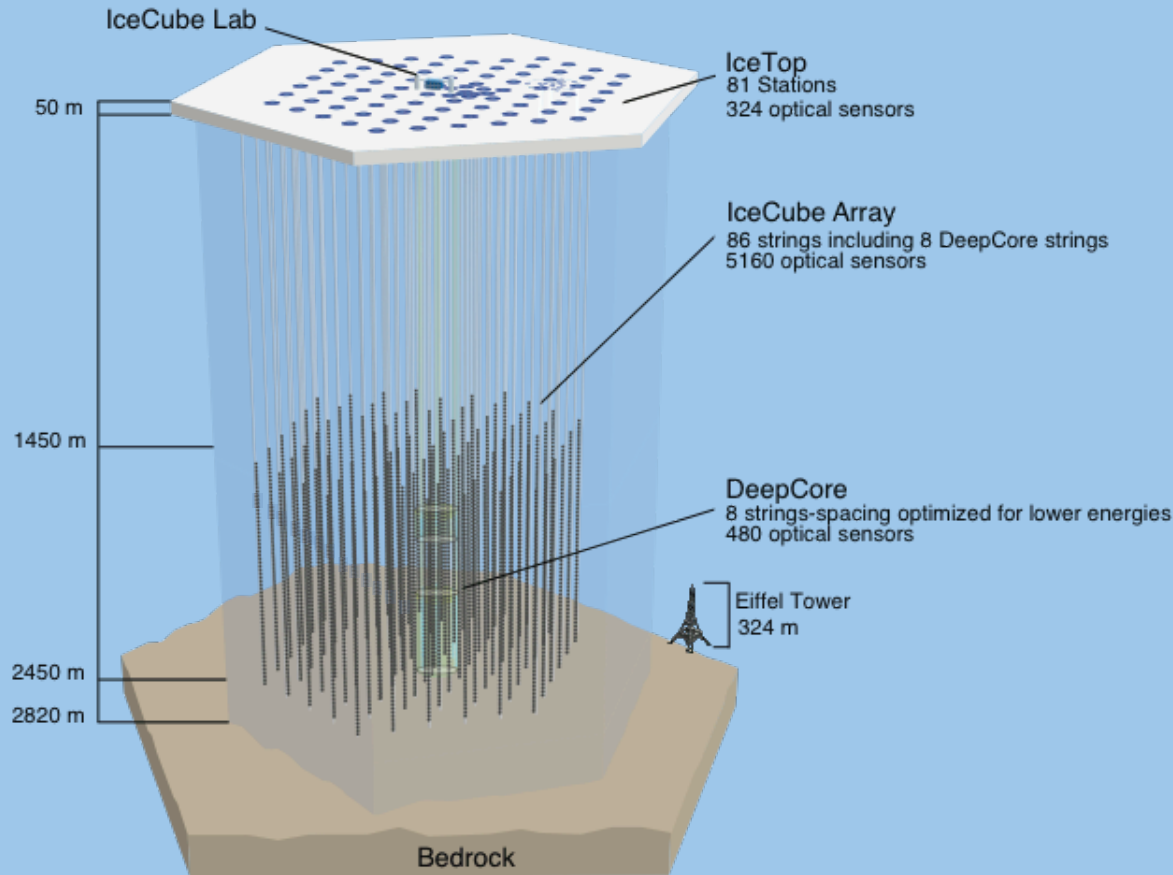
- **Good candidate for cosmic ray source search**

Neutrinos point back to their sources !

IceCube today

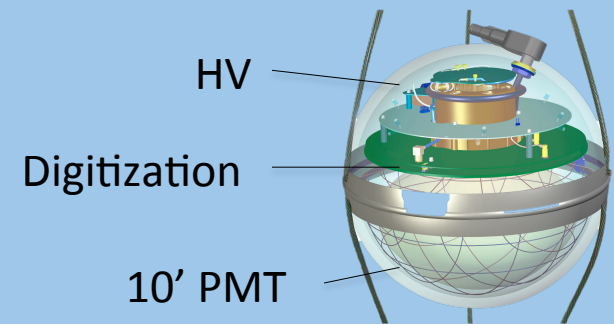
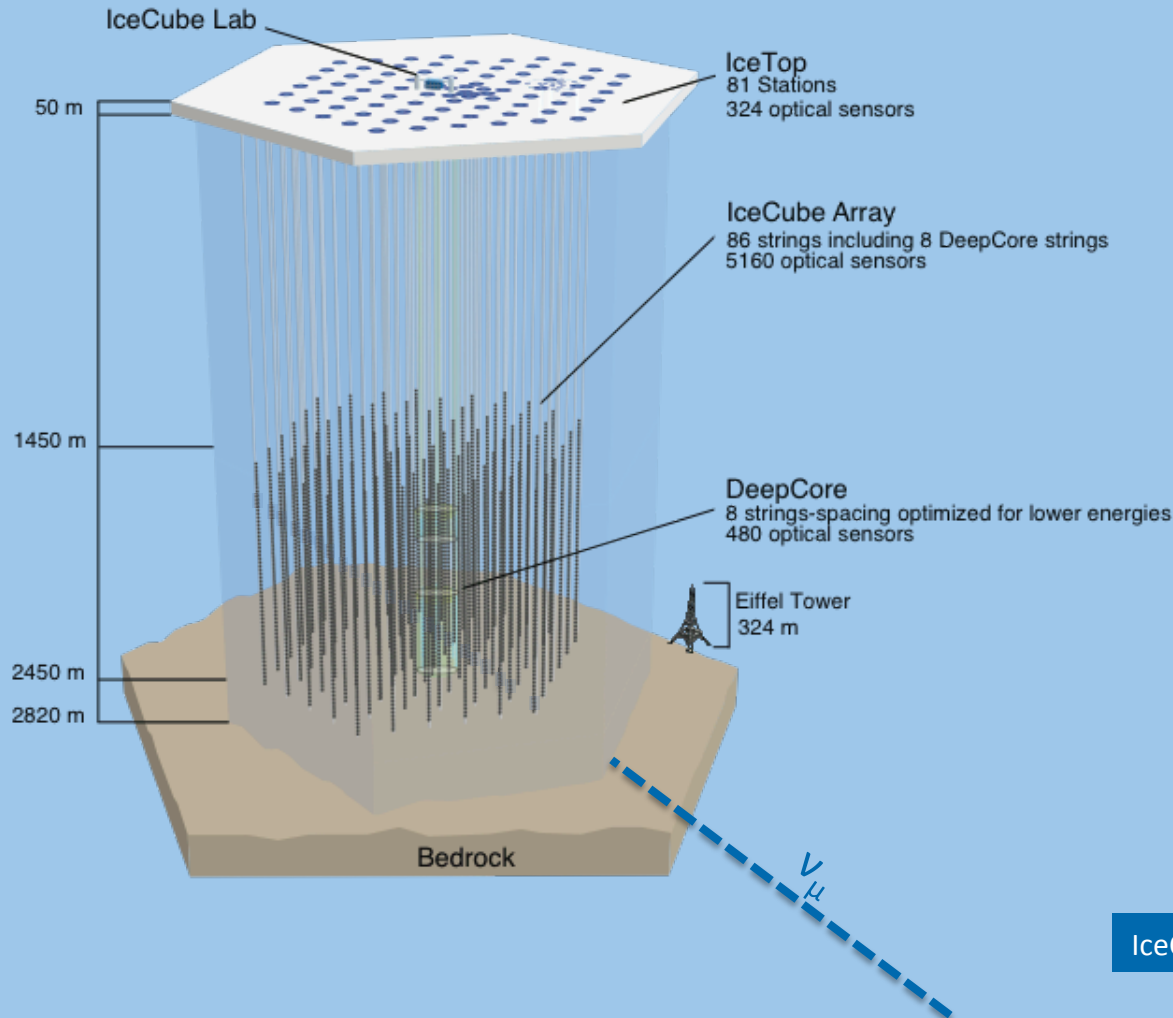


IceCube today



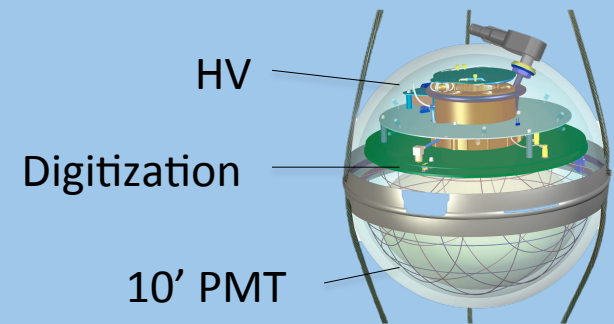
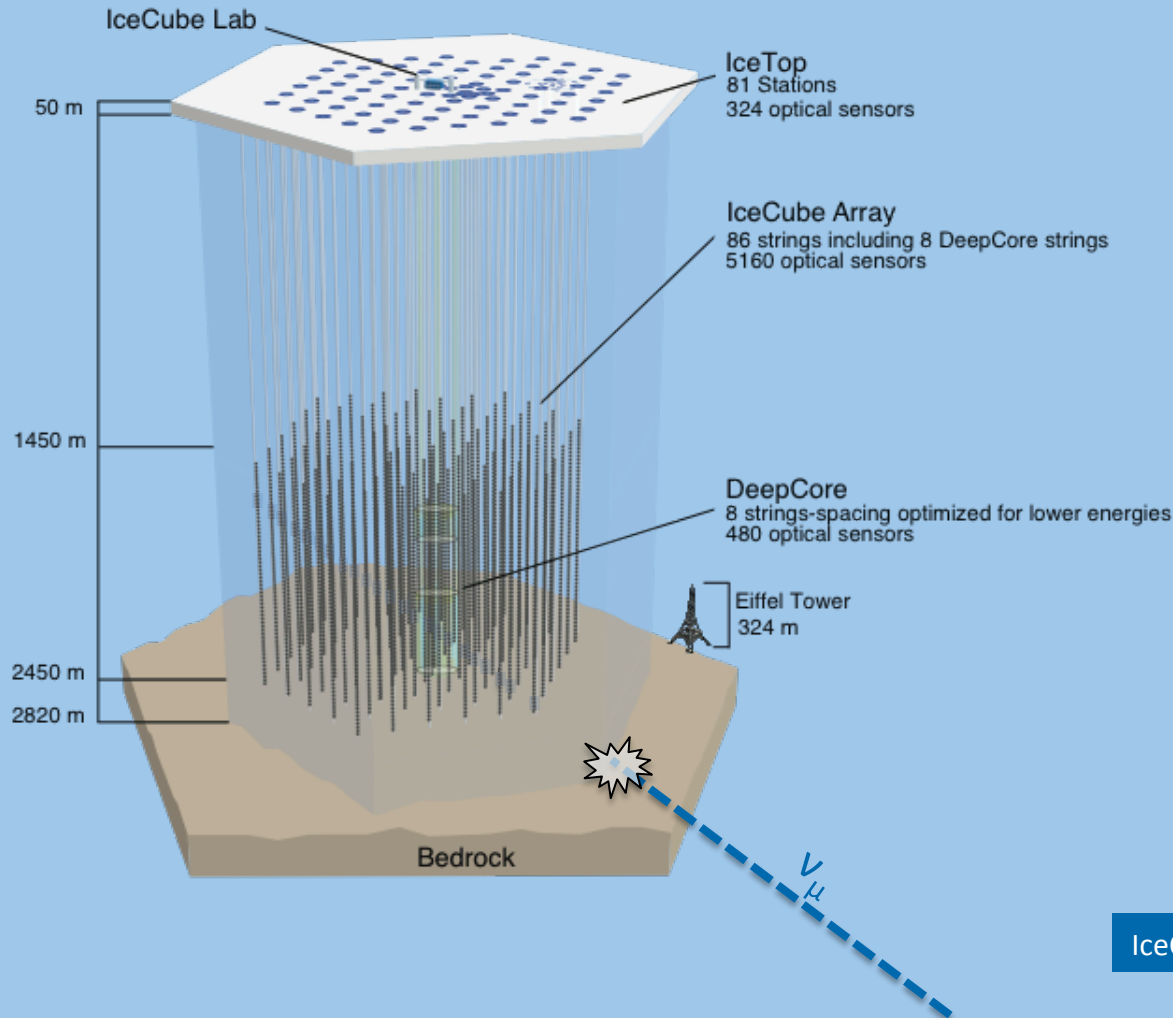
IceCube et al. Nucl. Inst. Meth. A 618 (2010) 139-152

IceCube today



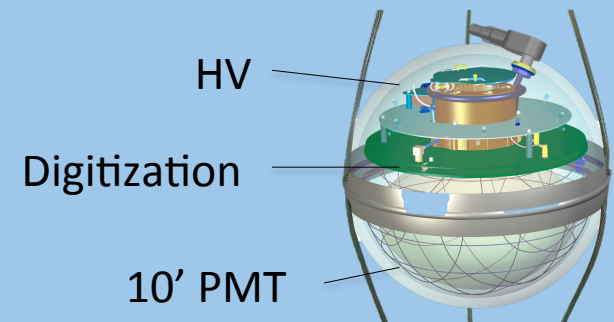
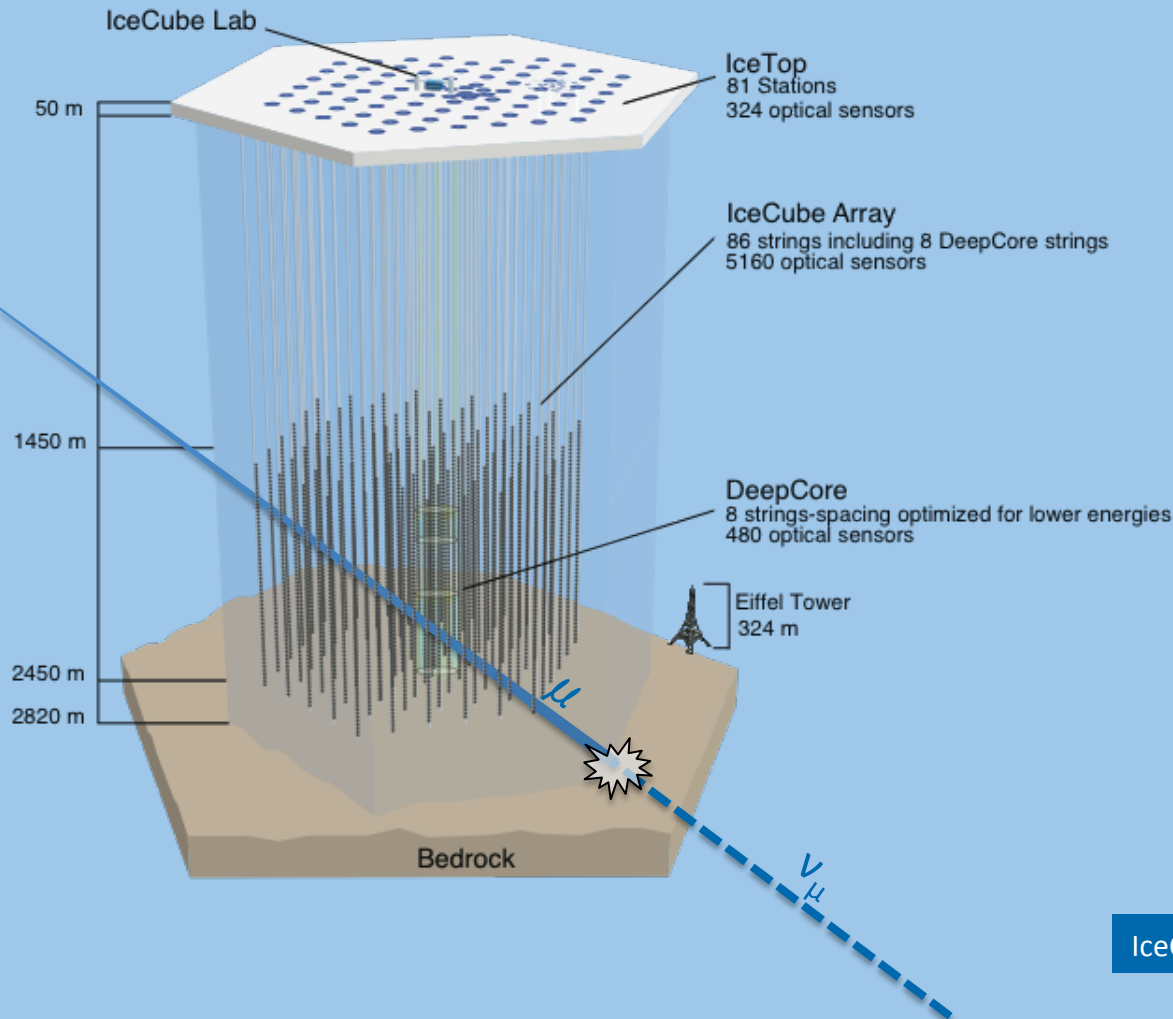
IceCube et al. Nucl. Inst. Meth. A 618 (2010) 139-152

IceCube today



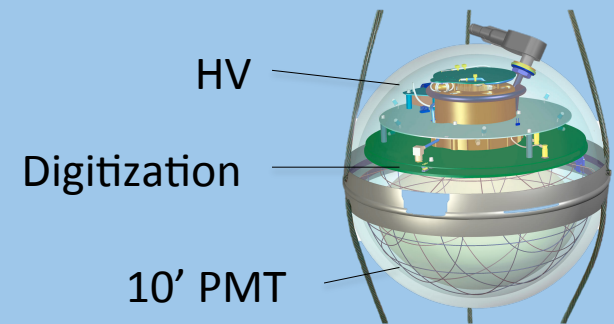
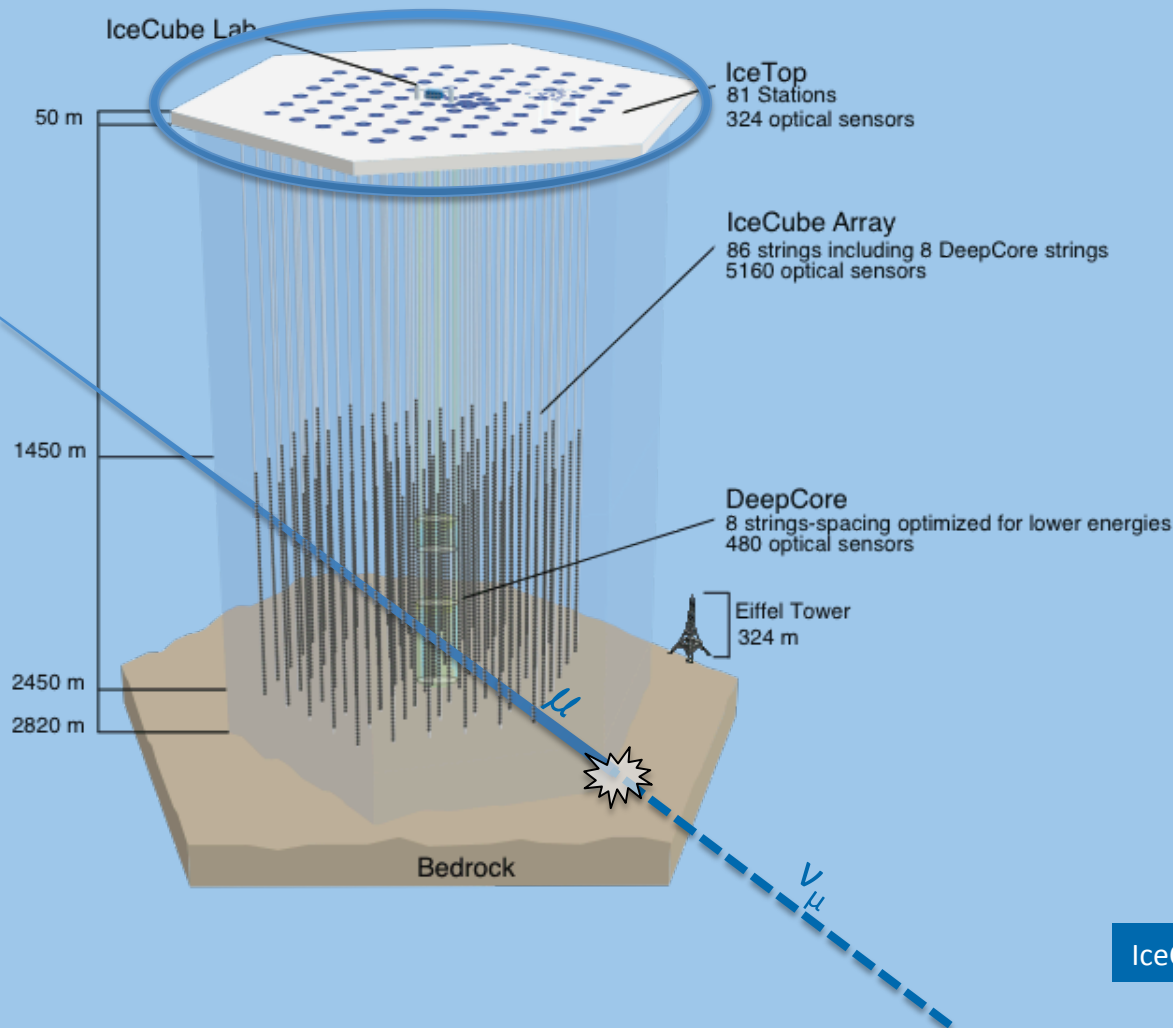
IceCube et al. Nucl. Inst. Meth. A 618 (2010) 139-152

IceCube today



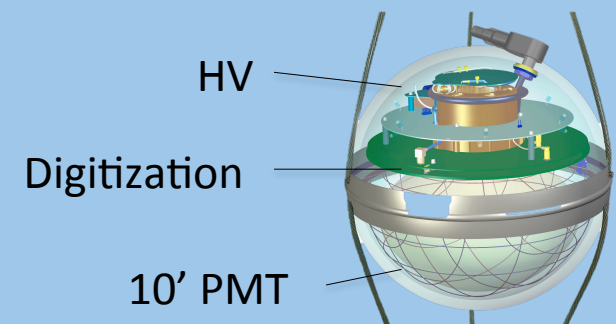
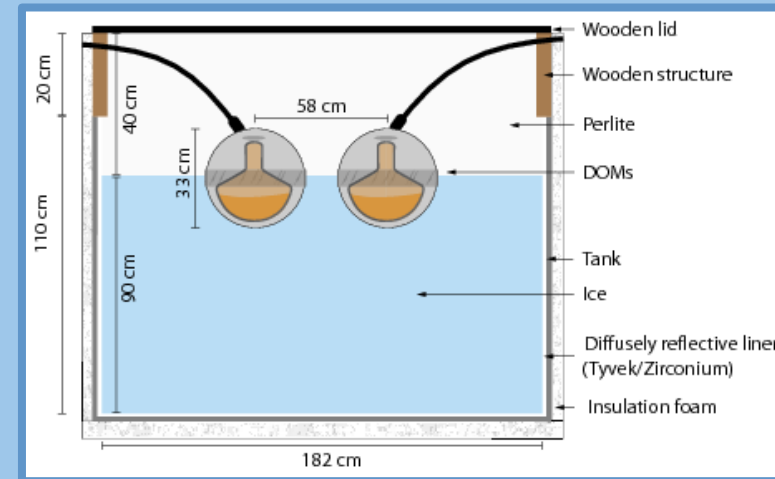
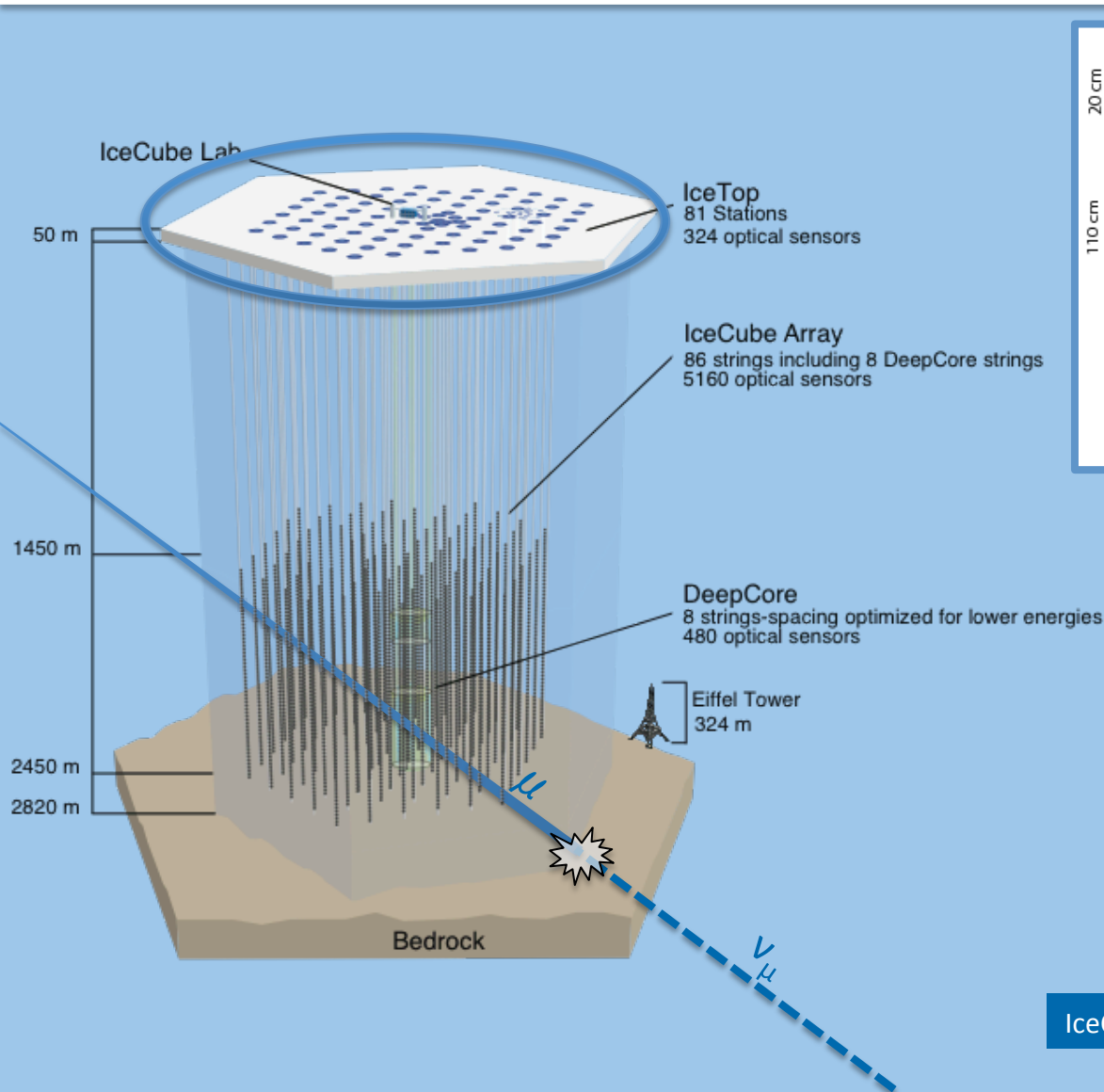
IceCube et al. Nucl. Inst. Meth. A 618 (2010) 139-152

IceCube today



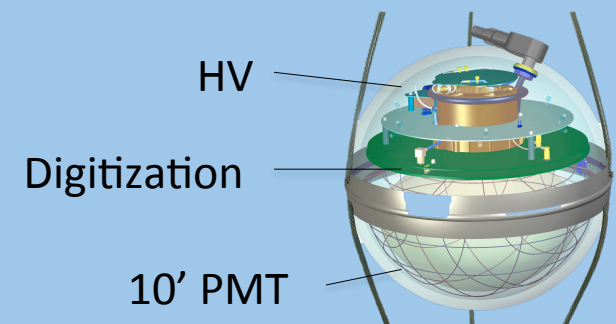
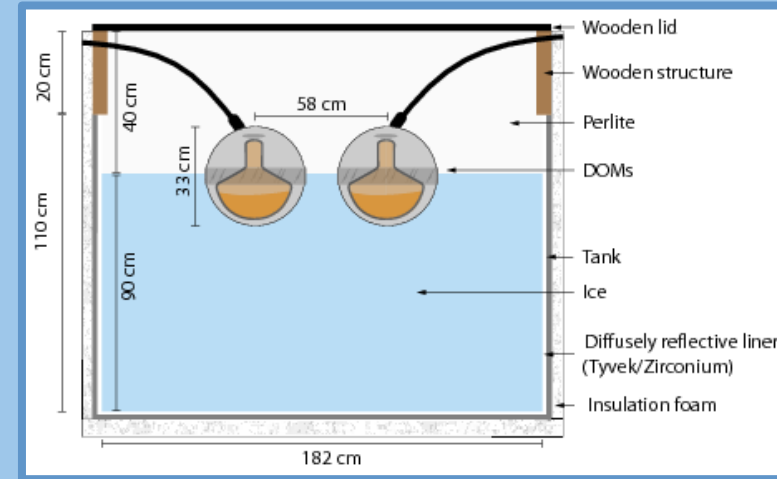
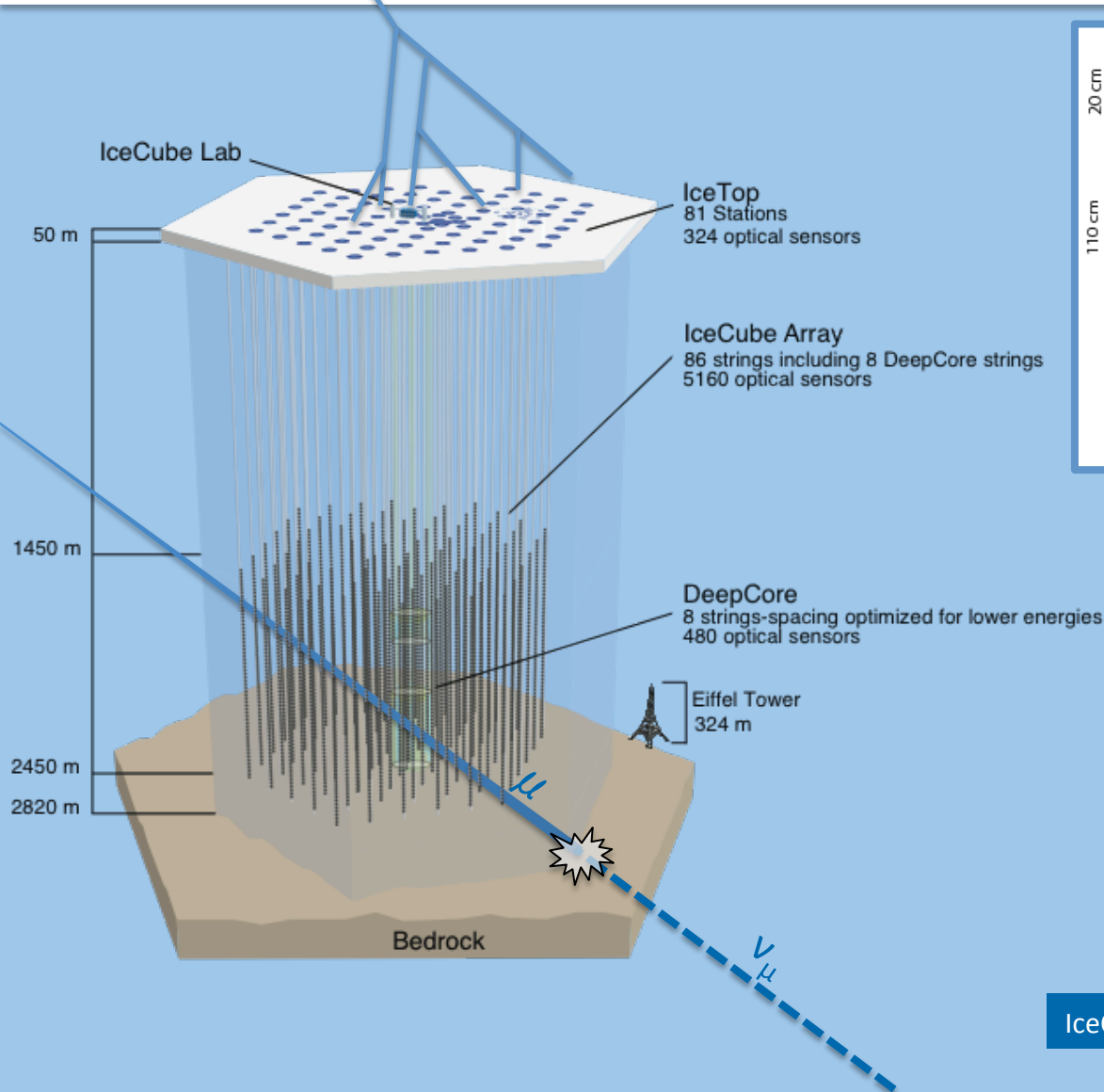
IceCube et al. Nucl. Inst. Meth. A 618 (2010) 139-152

IceCube today



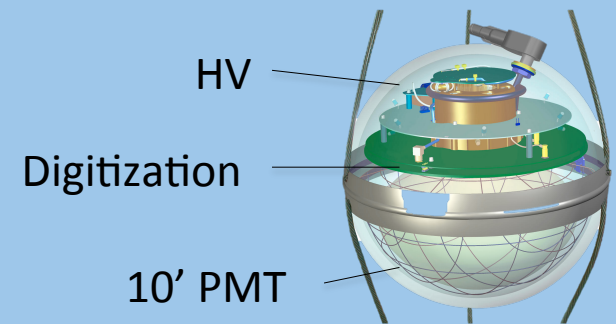
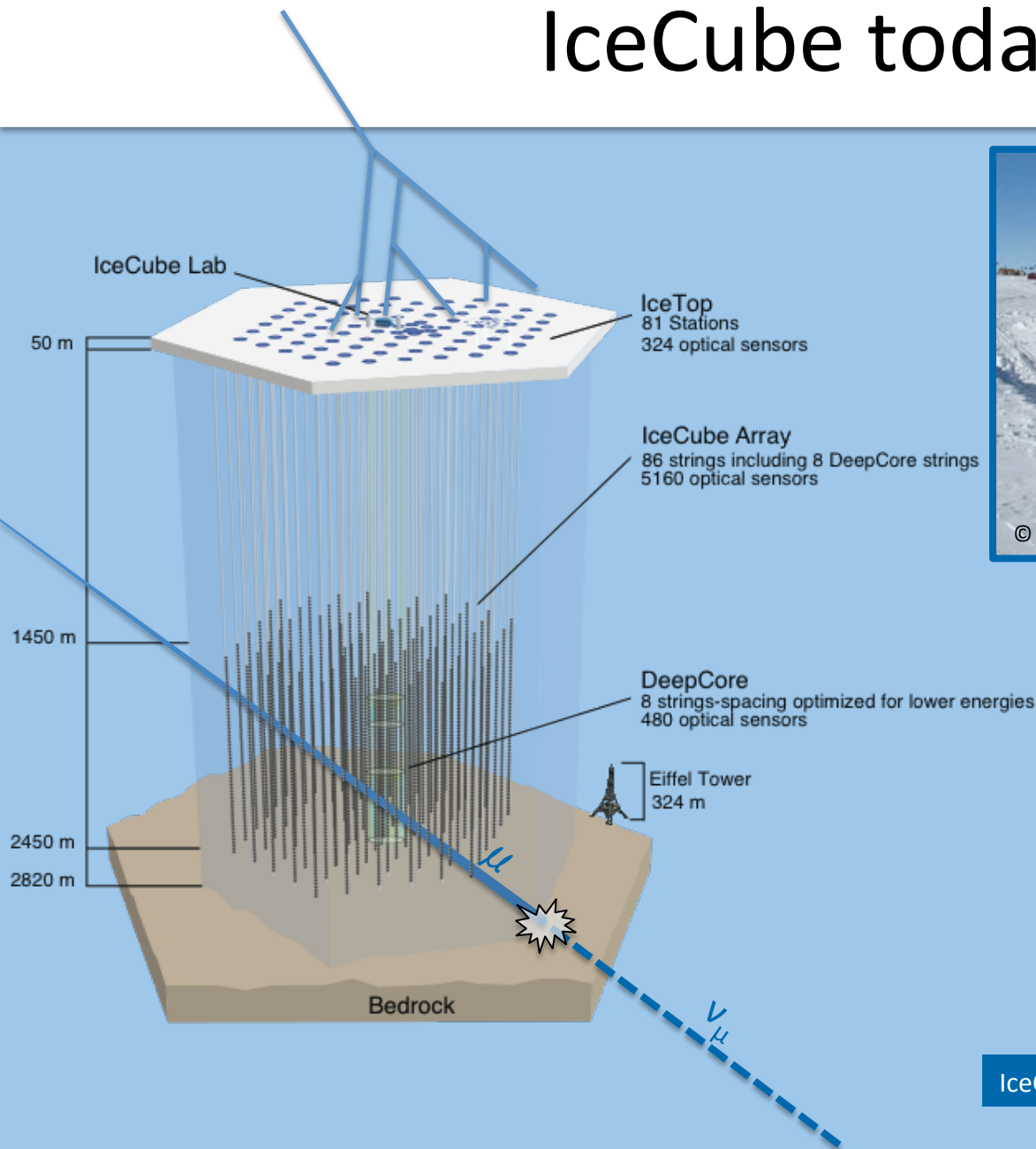
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IceCube today



IceCube et al. Nucl. Inst. Meth. A 618 (2010) 139-152

IceCube today



IceCube et al. Nucl. Inst. Meth. A 618 (2010) 139-152

Observed Neutrino Signatures ICECUBE

Observed Neutrino Signatures

Neutral Current /Electron
Neutrino
so called “**shower**”

Observed Neutrino Signatures

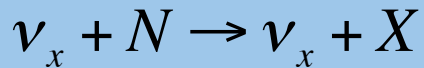
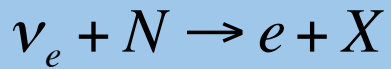
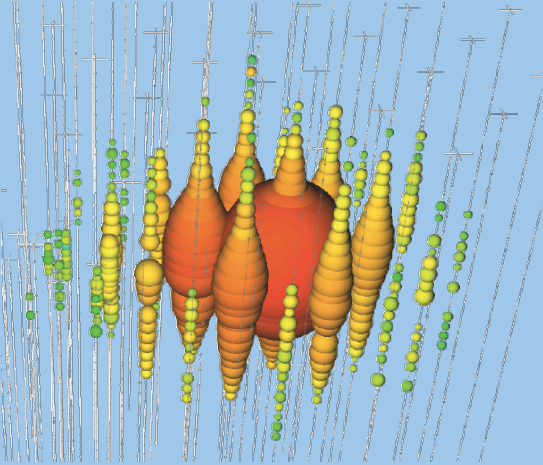
Neutral Current /Electron
Neutrino
so called “**shower**”

CC Muon Neutrino
so called “**track**”

Observed Neutrino Signatures



Neutral Current /Electron
Neutrino
so called “**shower**”

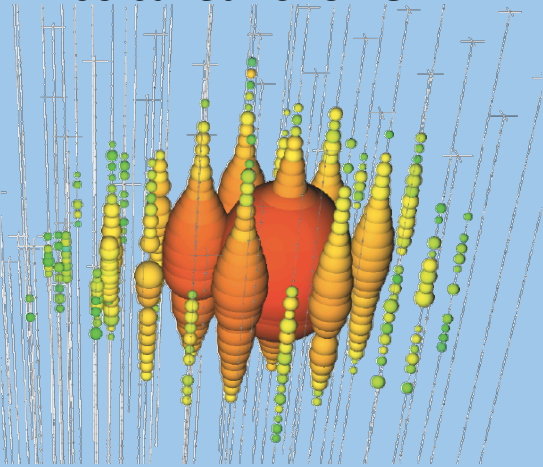


CC Muon Neutrino
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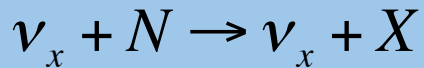
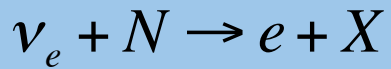
Observed Neutrino Signatures



Neutral Current /Electron
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CC Muon Neutrino
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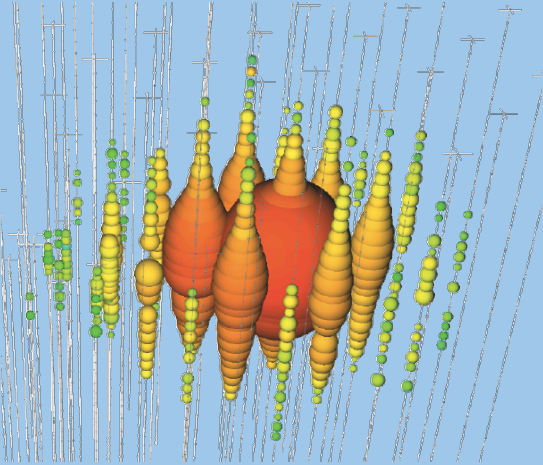


- Good Energy resolution

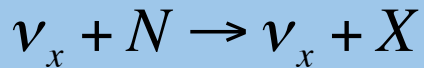
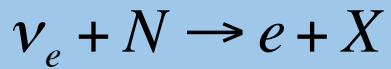
Observed Neutrino Signatures



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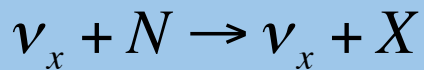
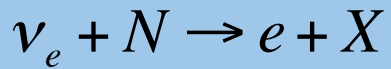
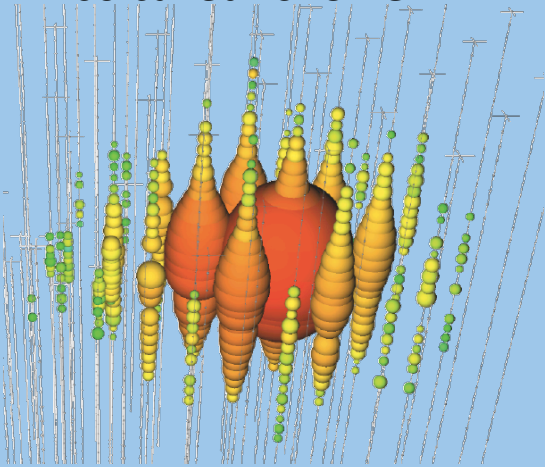


- Good Energy resolution
- Bad angular resolution

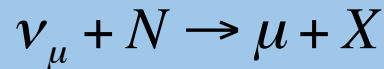
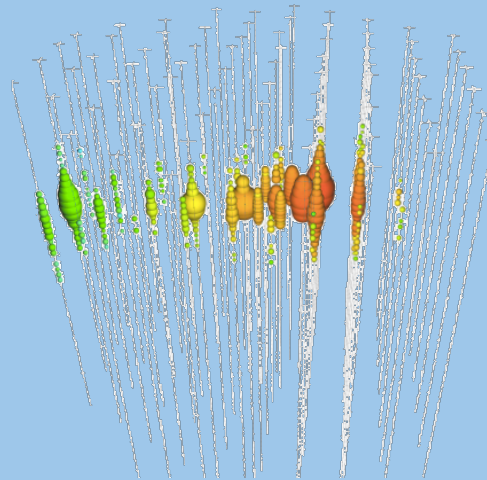
Observed Neutrino Signatures



Neutral Current /Electron
Neutrino
so called “**shower**”



CC Muon Neutrino
so called “**track**”

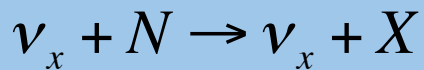
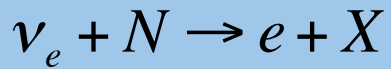
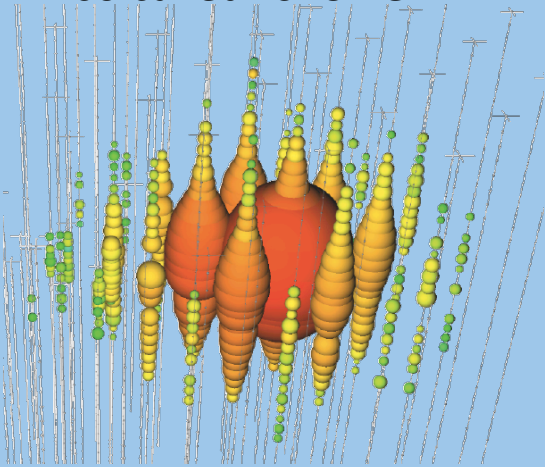


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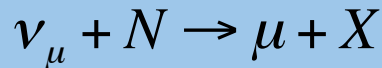
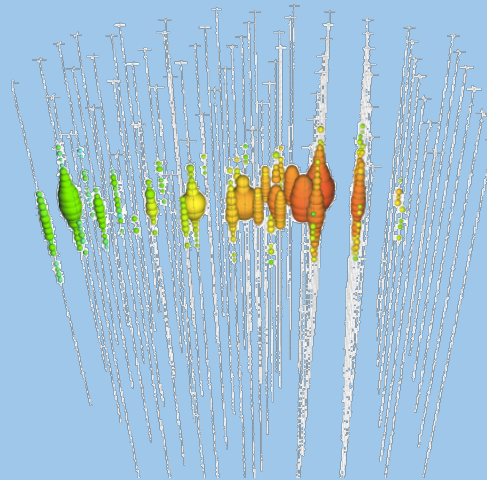
Observed Neutrino Signatures



Neutral Current /Electron
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CC Muon Neutrino
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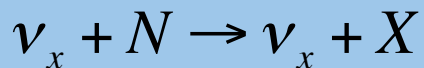
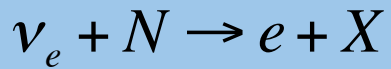
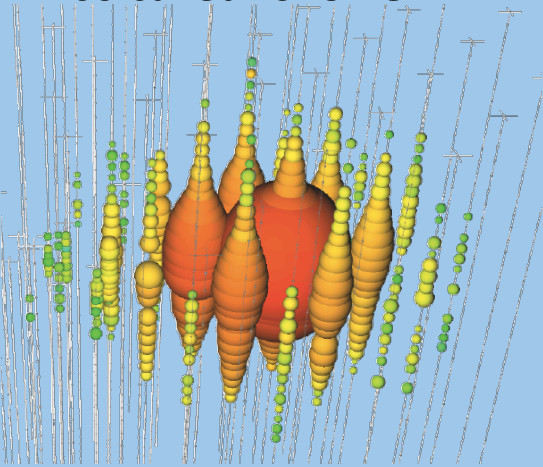


- Good Energy resolution
- Bad angular resolution
- Bad Energy resolution

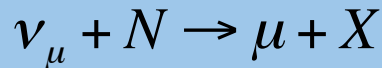
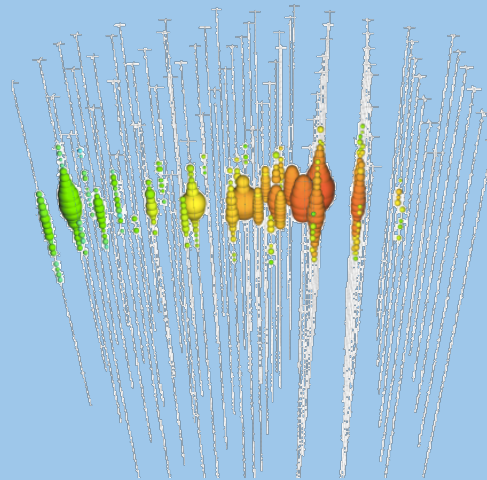
Observed Neutrino Signatures



Neutral Current /Electron
Neutrino
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CC Muon Neutrino
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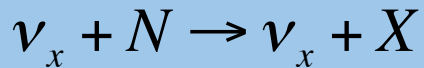
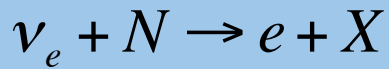
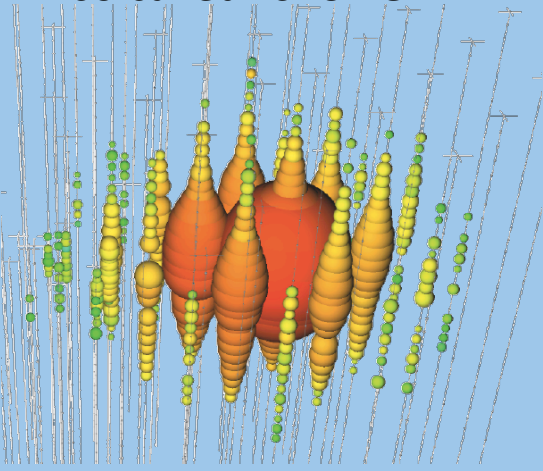


- Good Energy resolution
- Bad angular resolution
- Bad Energy resolution
- **Good angular resolution**

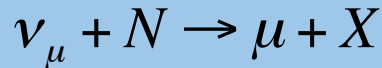
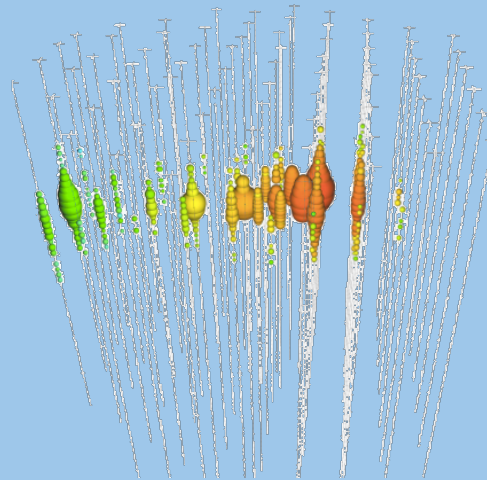
Observed Neutrino Signatures



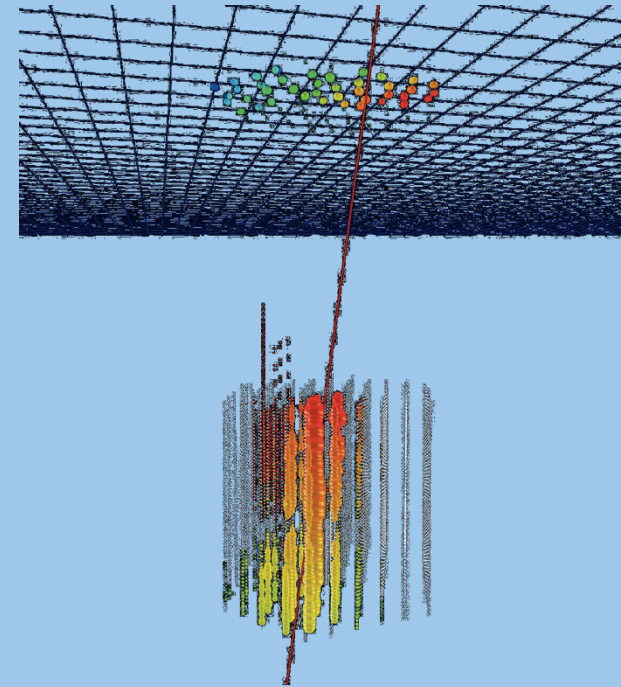
Neutral Current /Electron
Neutrino
so called “**shower**”



CC Muon Neutrino
so called “**track**”



Cosmic-Ray background

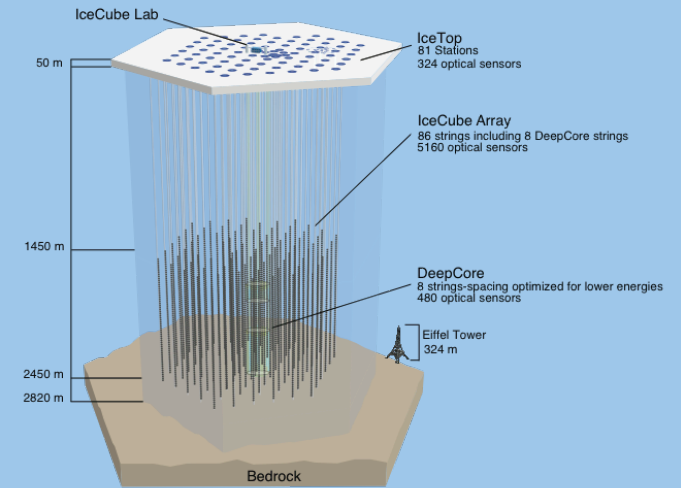


- Good Energy resolution
- Bad angular resolution

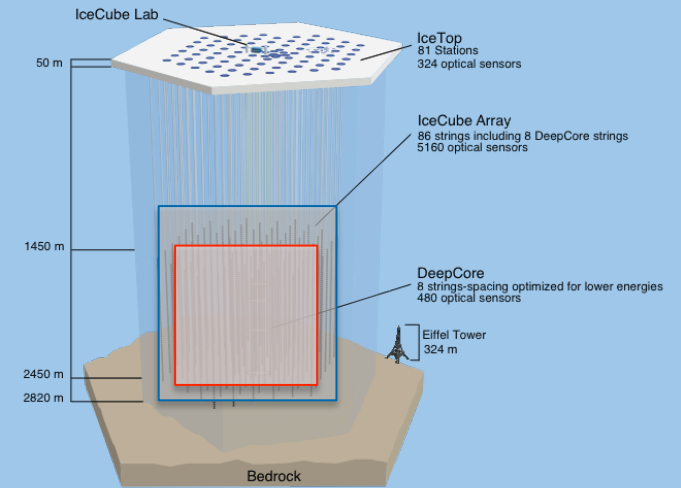
- Bad Energy resolution
- **Good angular resolution**

- Bad Energy resolution
- Good angular resolution

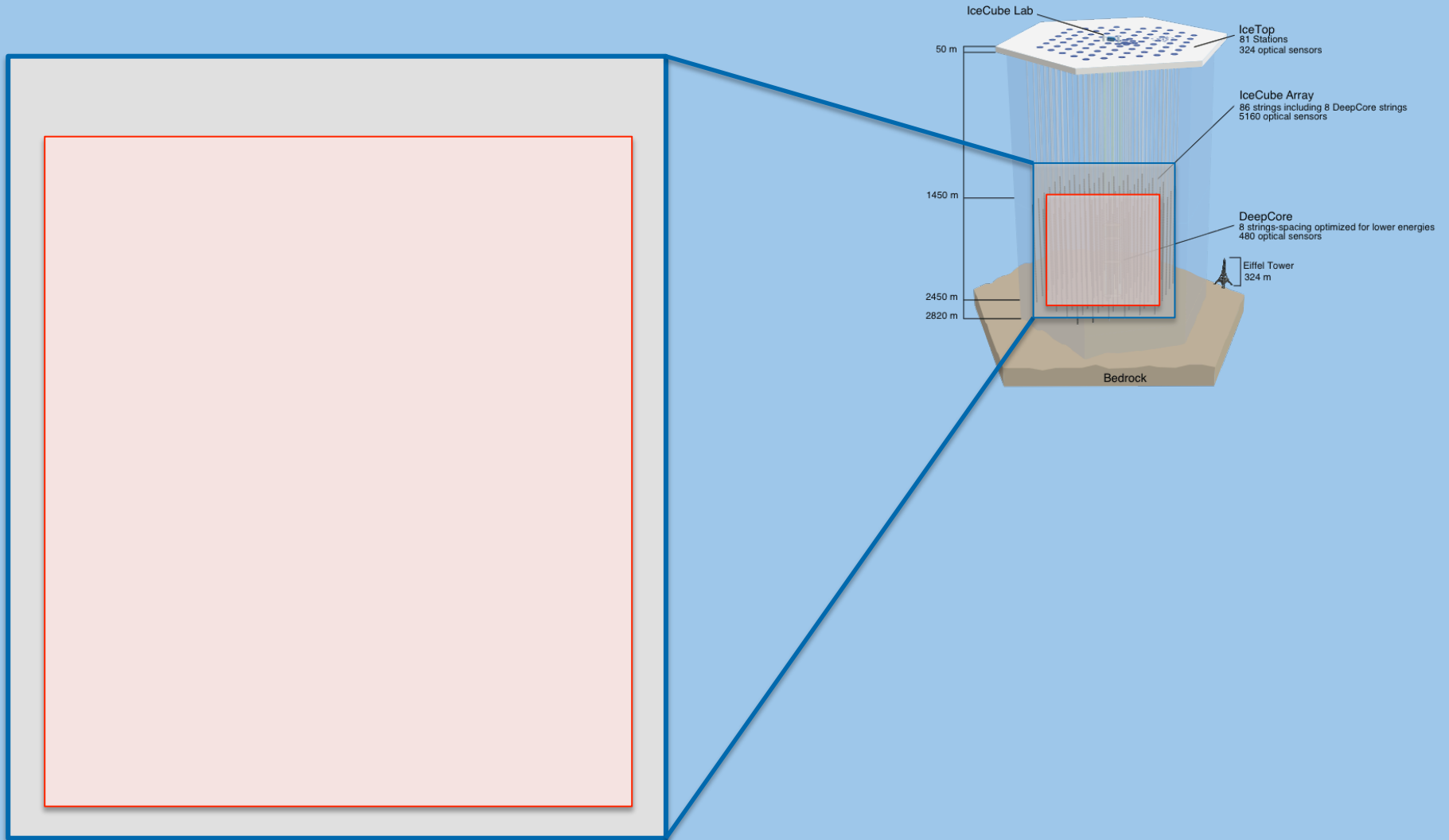
IceCube Results: Starting Events



IceCube Results: Starting Events



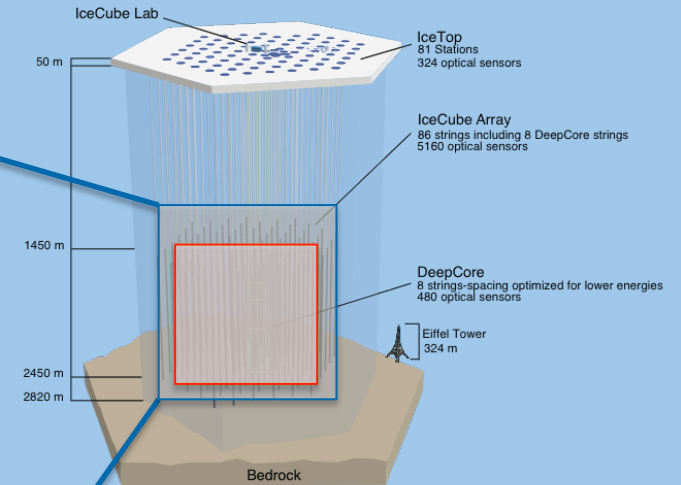
IceCube Results: Starting Events



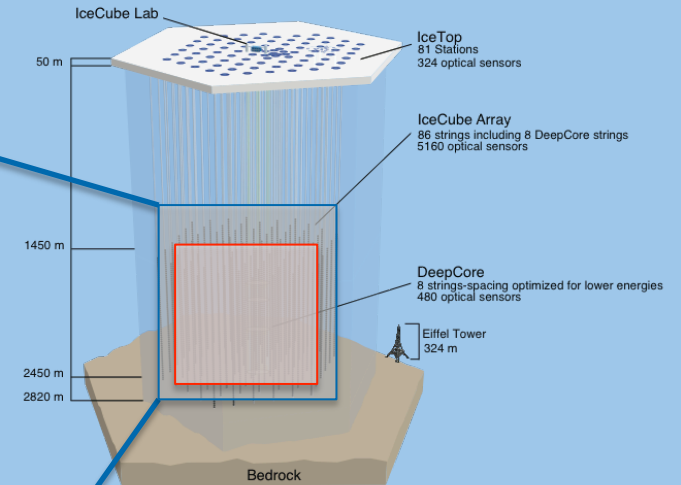
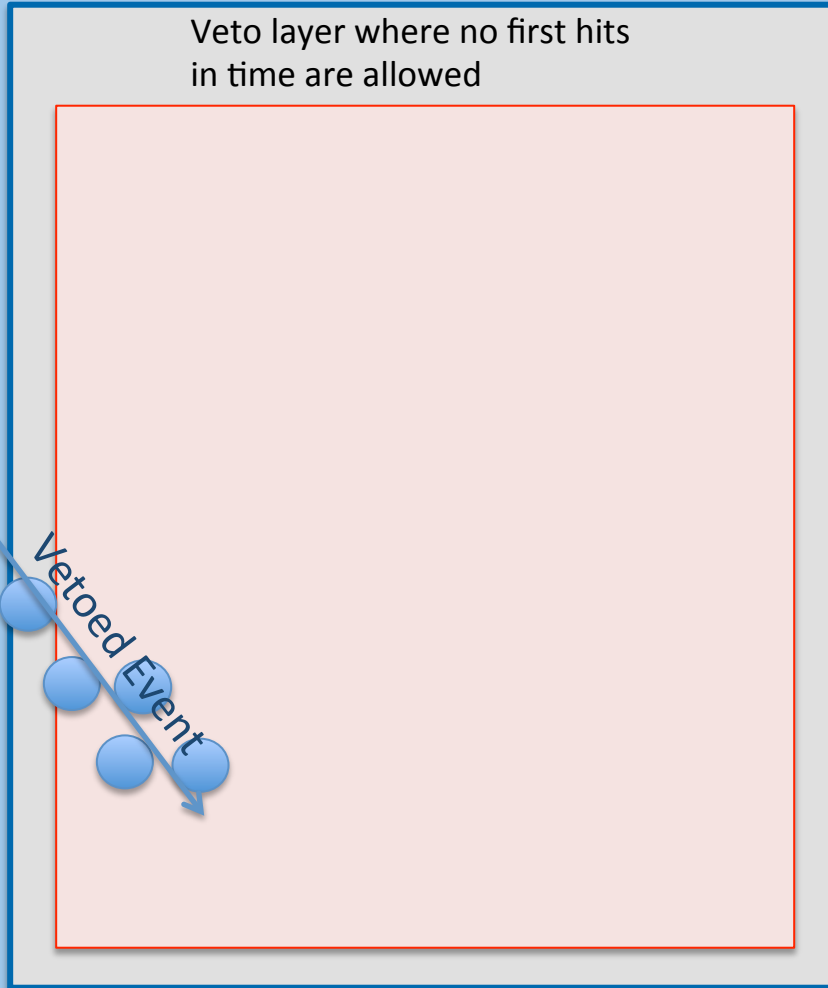
IceCube Results: Starting Events



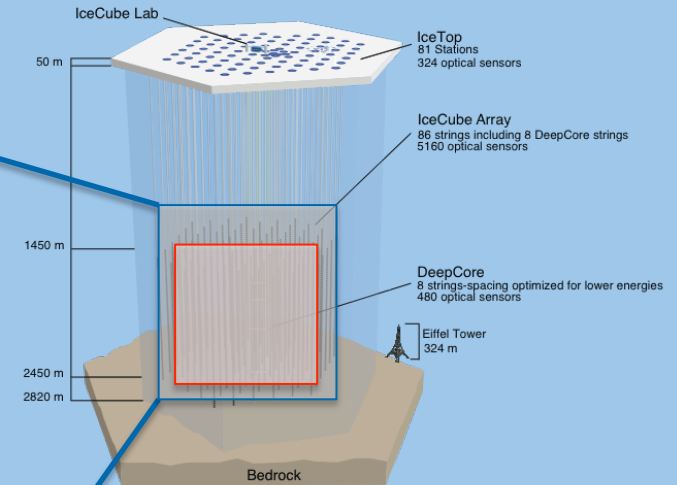
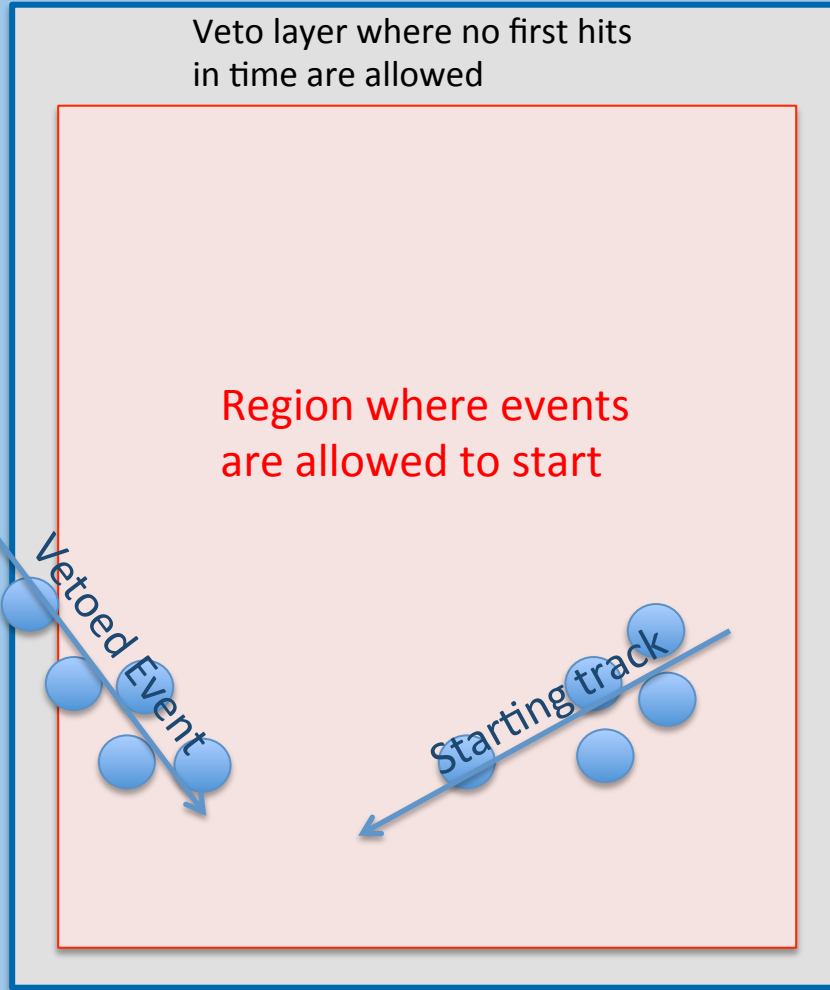
Veto layer where no first hits
in time are allowed



IceCube Results: Starting Events

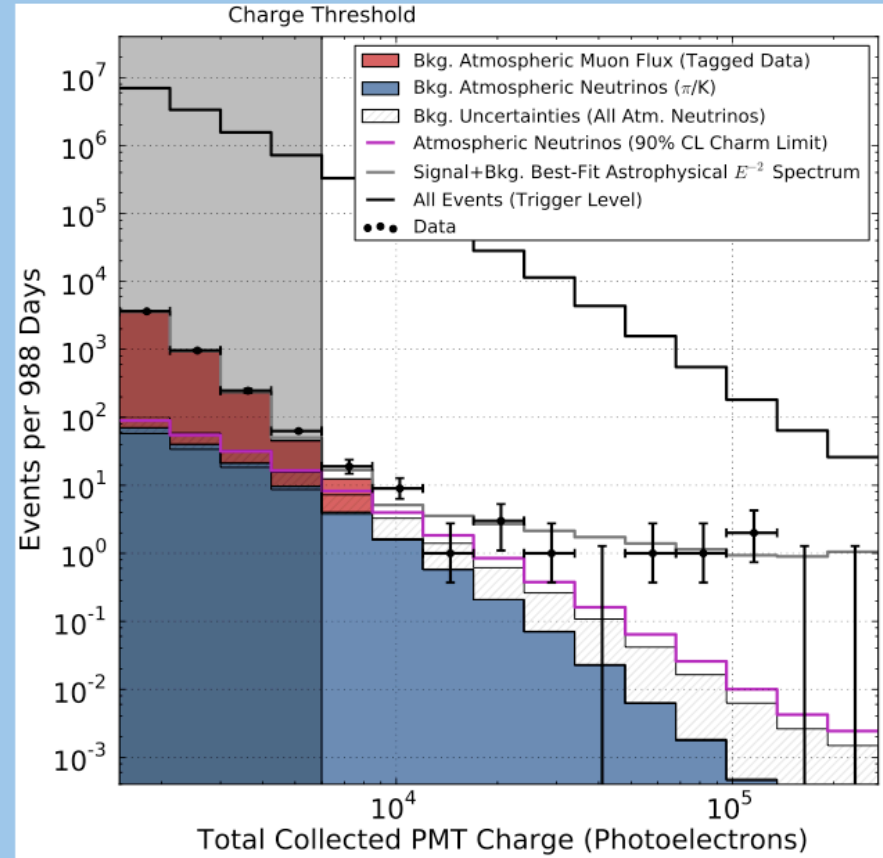
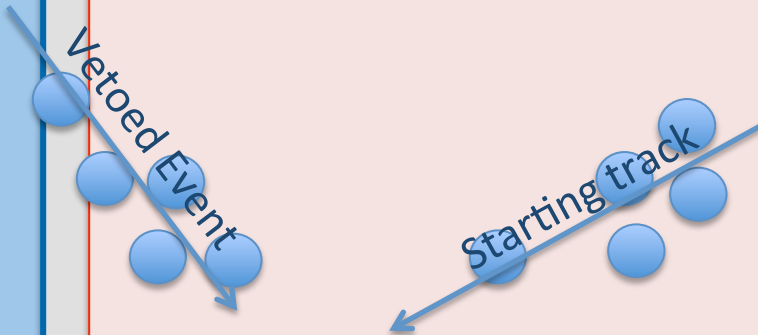


IceCube Results: Starting Events



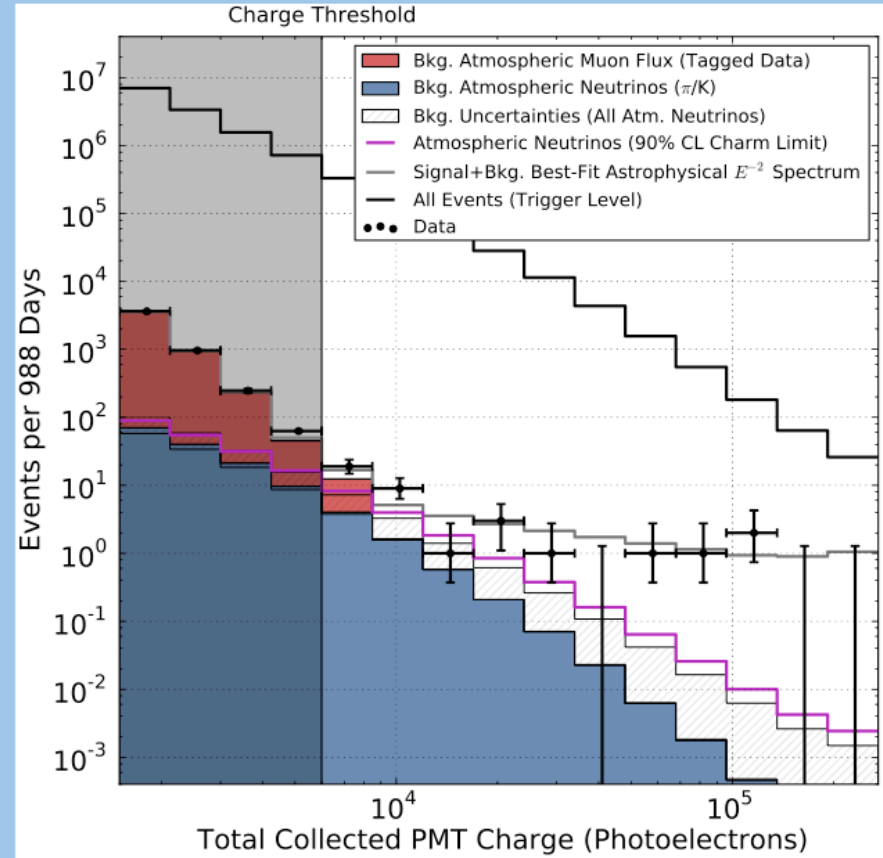
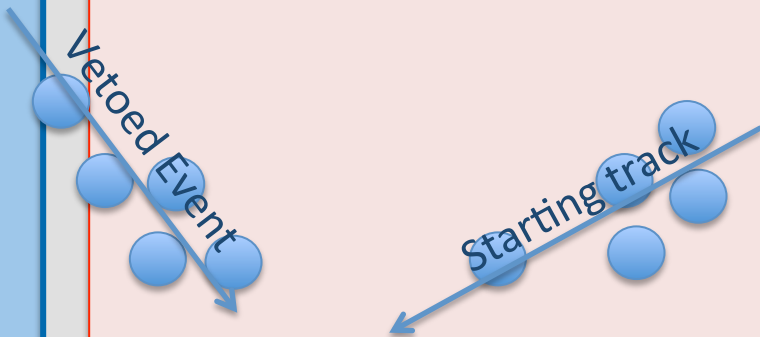
Veto layer where no first hits
in time are allowed

Region where events
are allowed to start



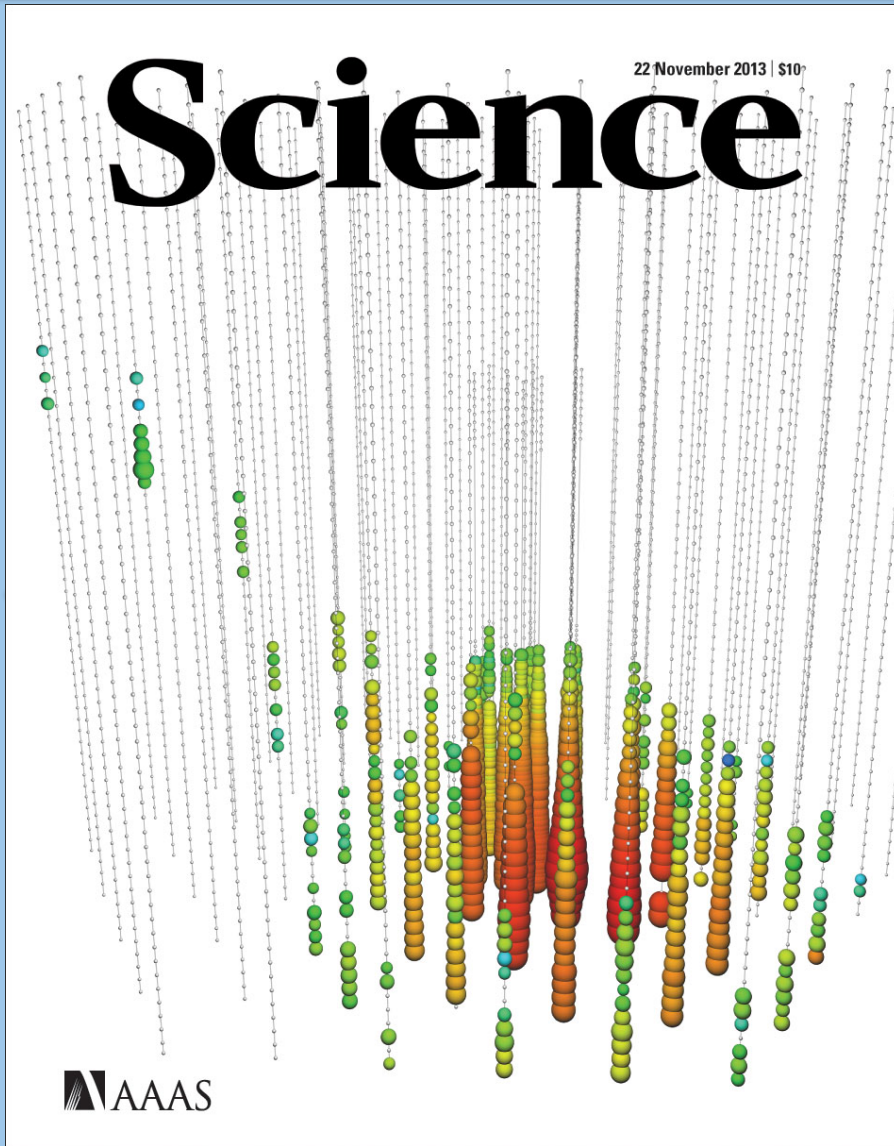
Veto layer where no first hits
in time are allowed

Region where events
are allowed to start



- null hypothesis rejected with 5.7σ after three years of data taking

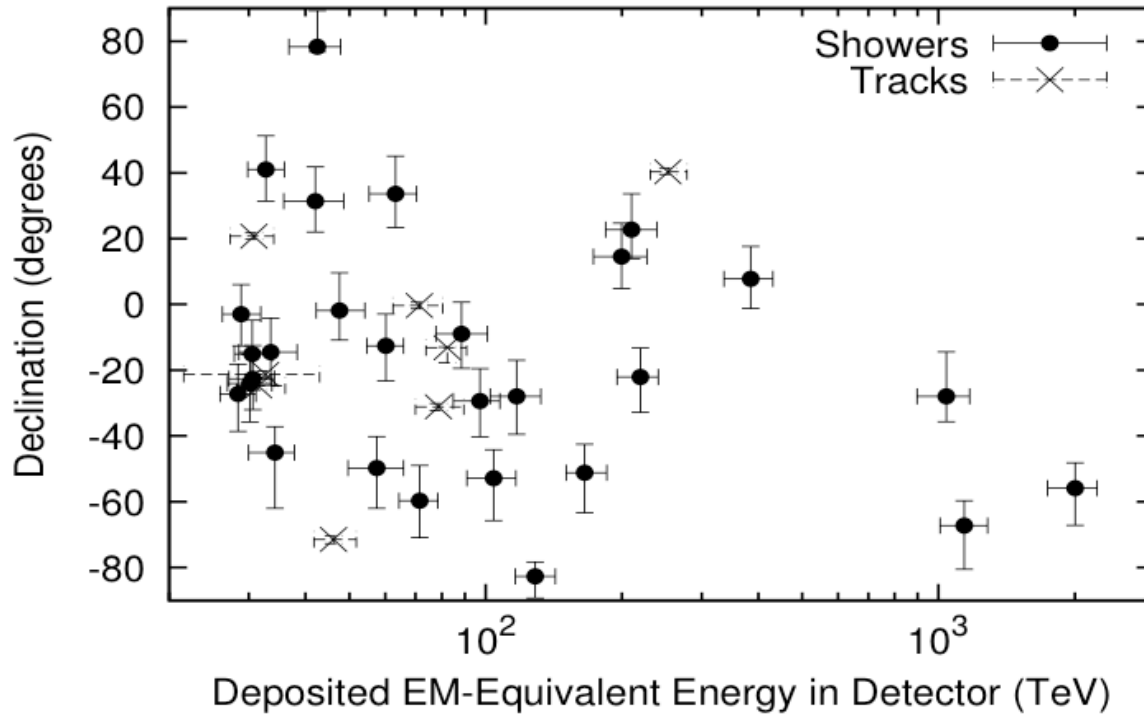




Where do the neutrinos come from?



Phys. Rev. Lett. 113, 01-22, (2014).

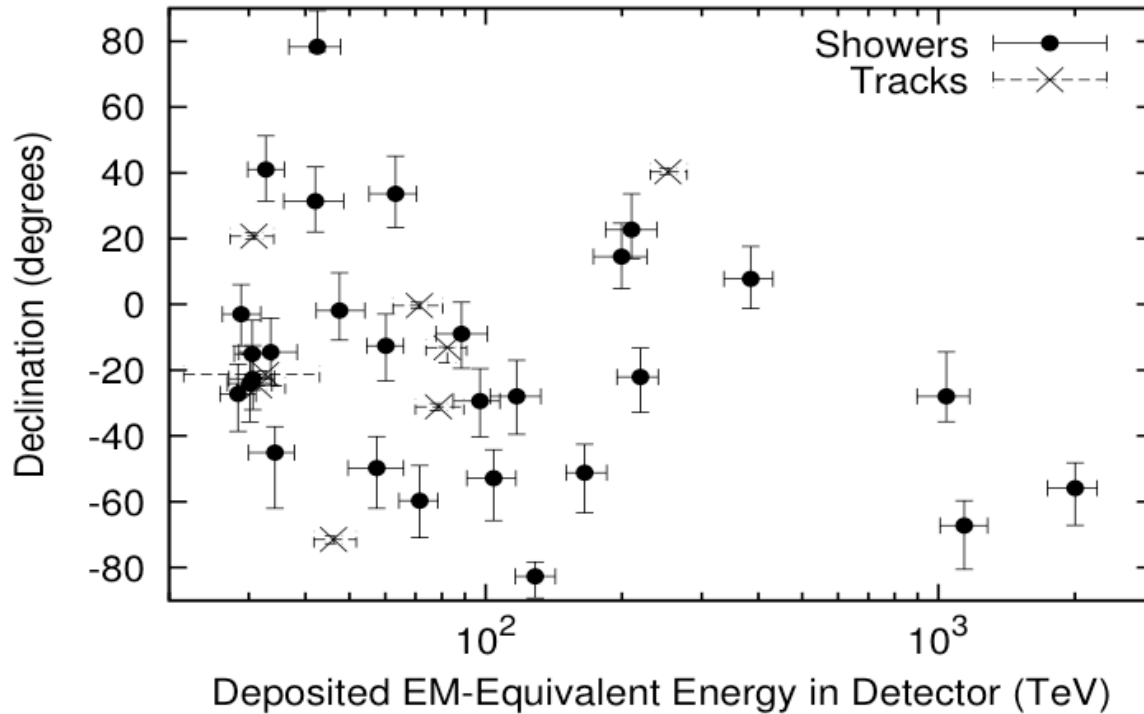


- only few tracks
- showers don't point

Where do the neutrinos come from?



Phys. Rev. Lett. 113, 01-22, (2014).

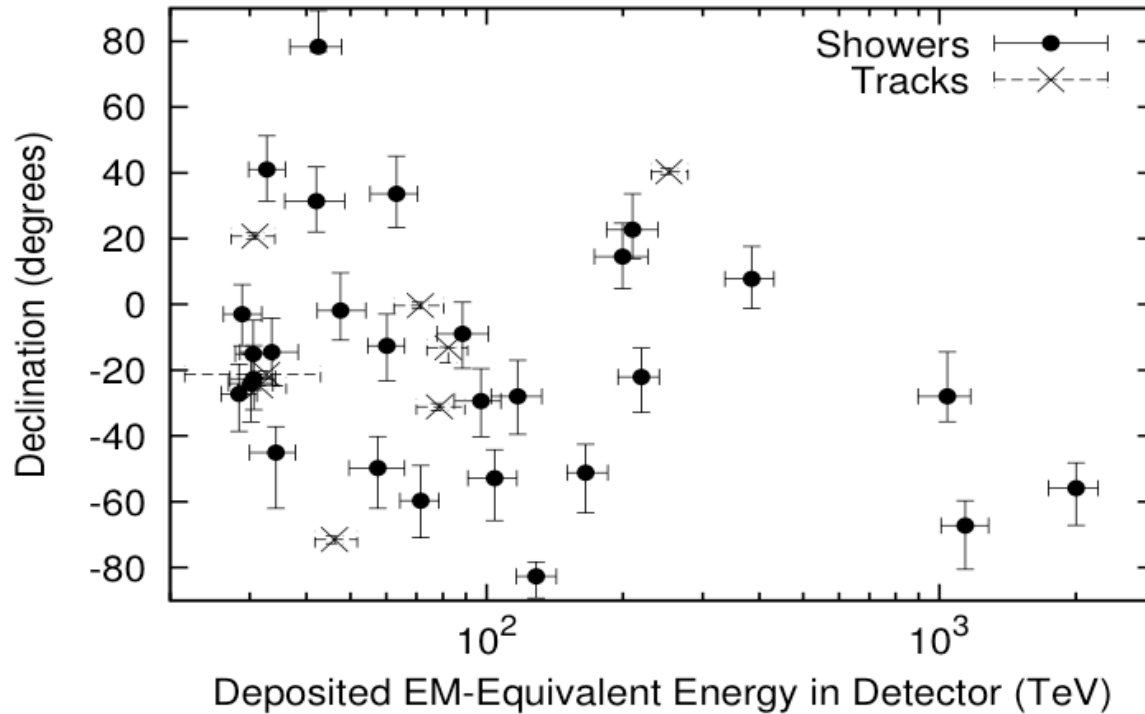


- only few tracks
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- PeV neutrinos only from the southern sky

Where do the neutrinos come from?



Phys. Rev. Lett. 113, 01-22, (2014).



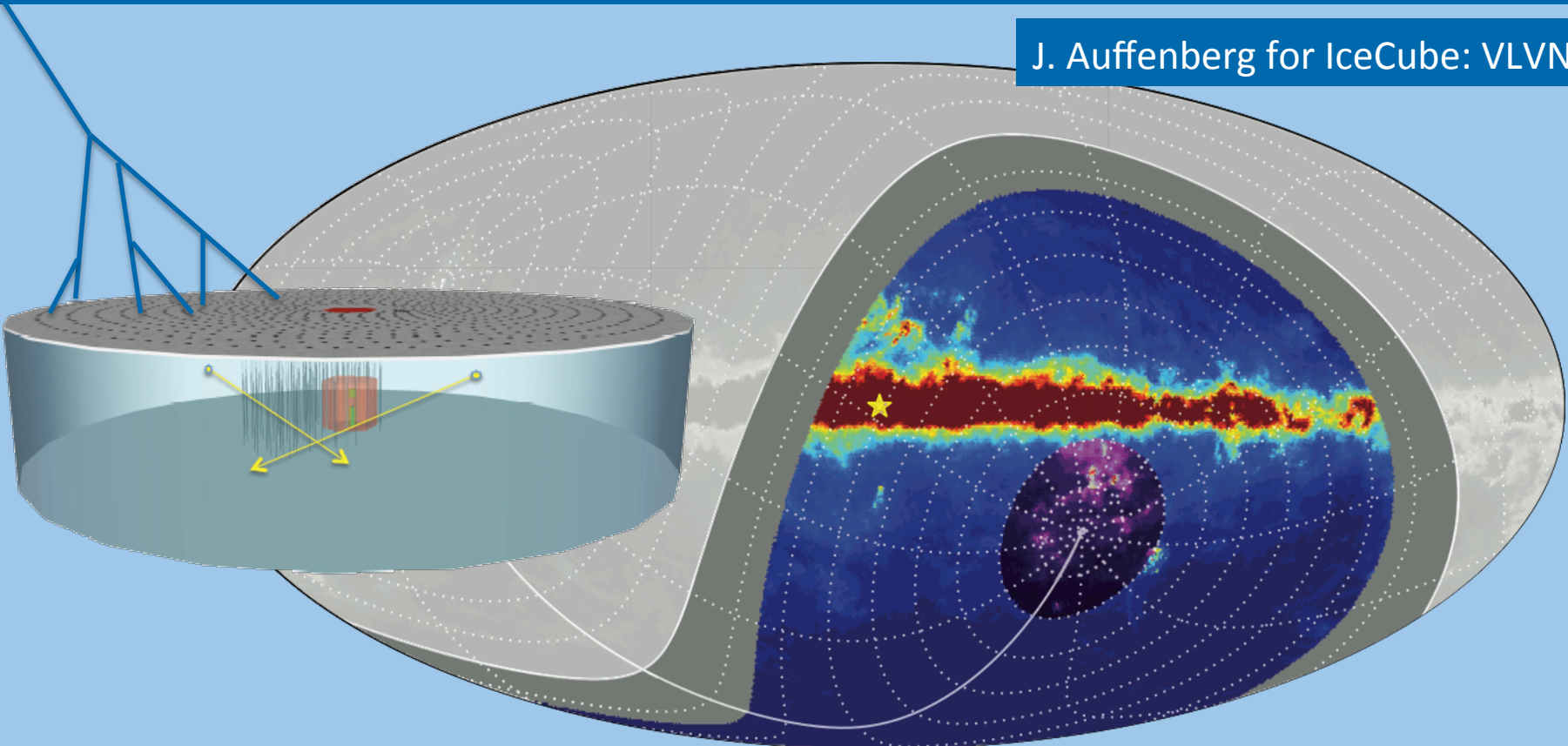
- only few tracks
- showers don't point
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We need more high energy tracks e.g. from the southern sky!

The future extension: IceVeto

We need more high energy tracks e.g. from the southern sky!

J. Auffenberg for IceCube: VLVNT13



Open the southern sky for $E < 100$ TeV Neutrino induced muon tracks by vetoing signals with coincident air showers

IceTop Veto

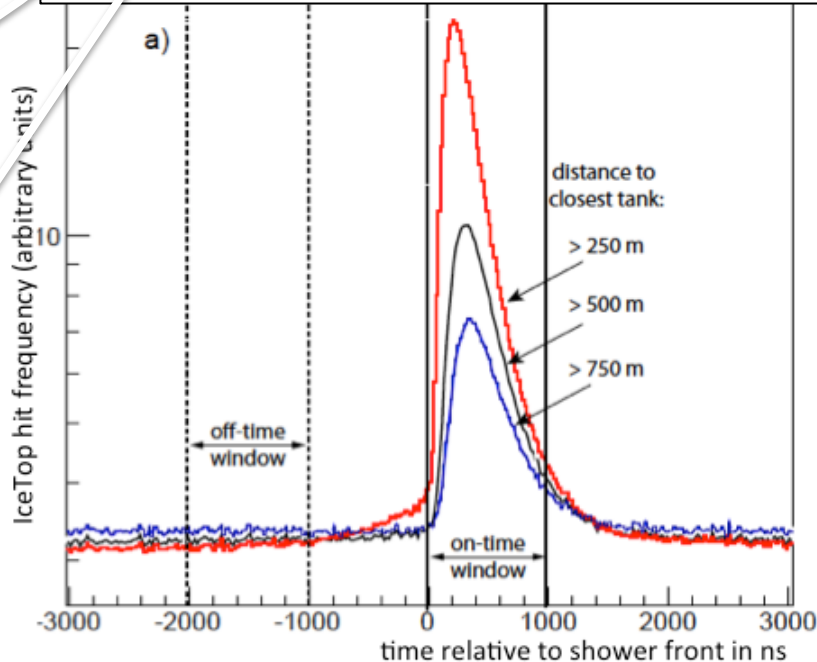


← Air-Shower Front

Track

IceCube
(InIce)

R. Abbasi, J. Auffenberg et al. Nucl. Inst. Meth. A700, 188-220 (2013).



J. Auffenberg for IceCube: ICRC13 ID 0373

IceTop Veto

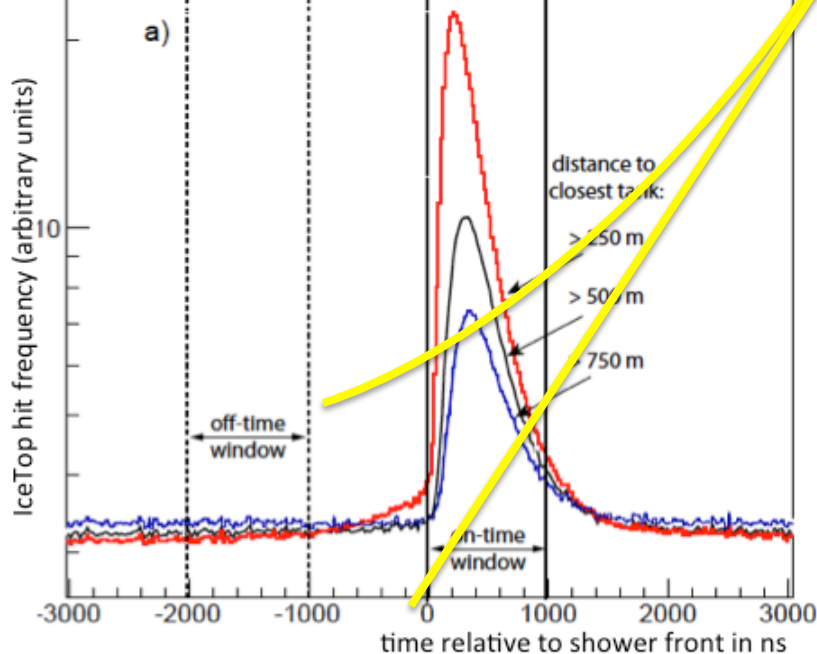


Air-Shower Front

IceCube
(InIce)

Coincidence

R. Abbasi, J. Auffenberg et al. Nucl. Inst. Meth. A700, 188-220 (2013).



J. Auffenberg for IceCube: ICRC13 ID 0373

IceTop Veto

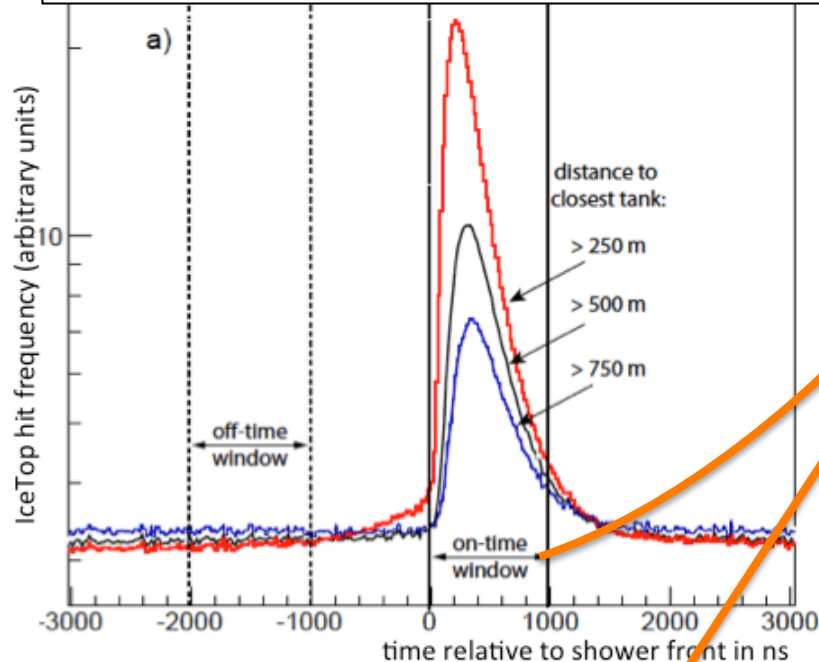


Air-Shower Front

IceCube
(InIce)

Coincidence

R. Abbasi, J. Auffenberg et al. Nucl. Inst. Meth. A700, 188-220 (2013).

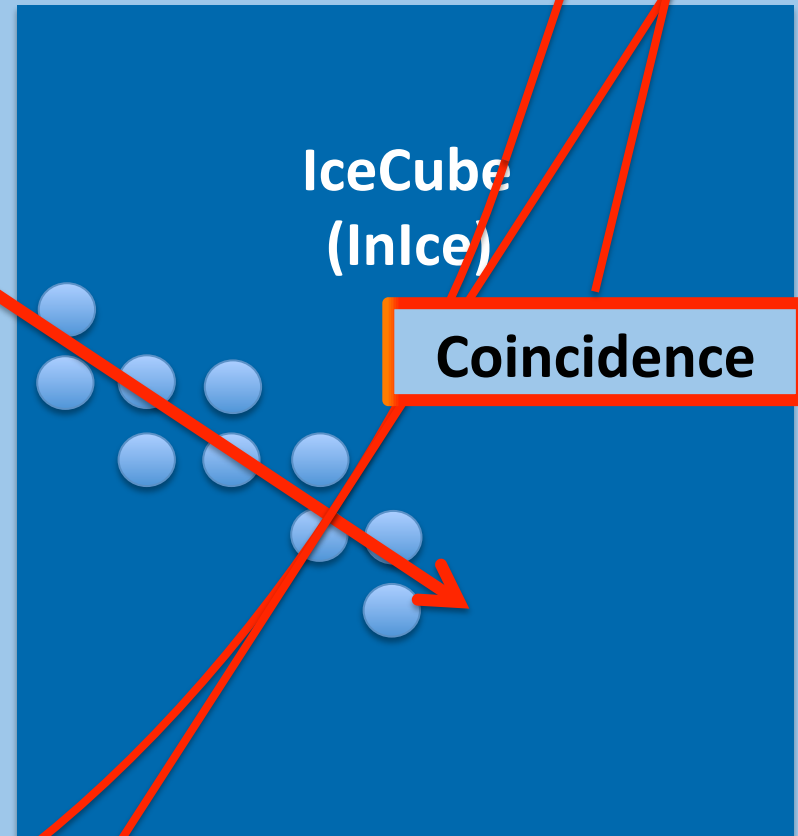


J. Auffenberg for IceCube: ICRC13 ID 0373

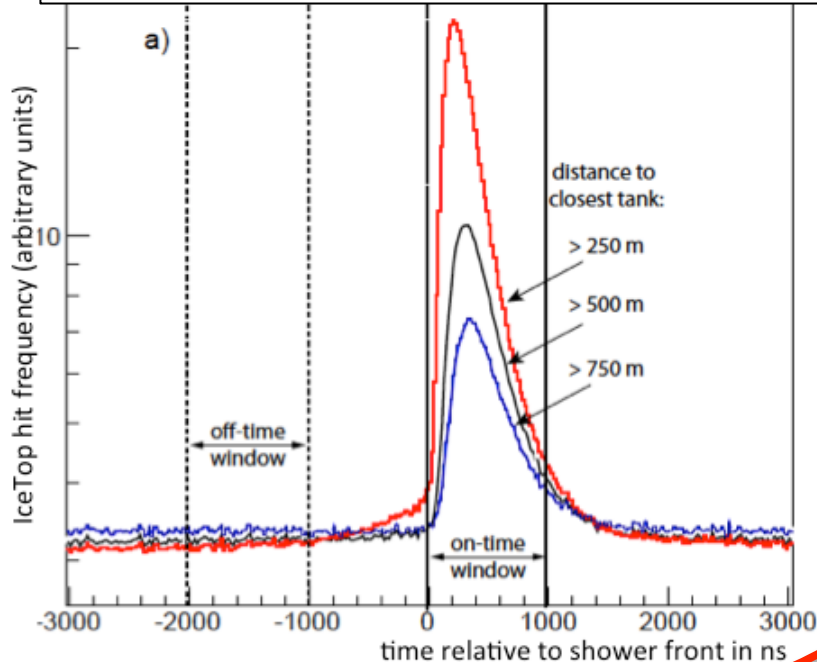
IceTop Veto



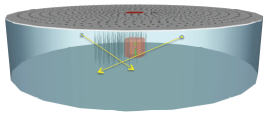
Air-Shower Front



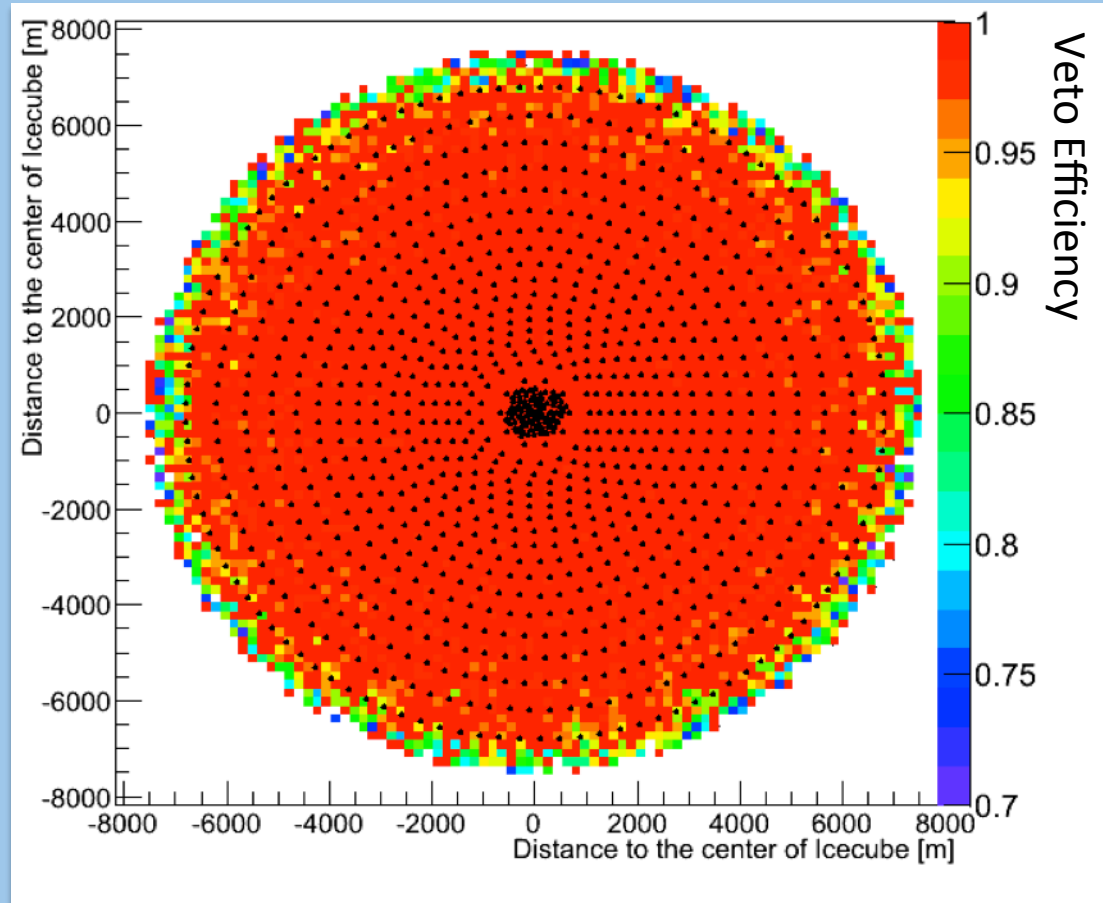
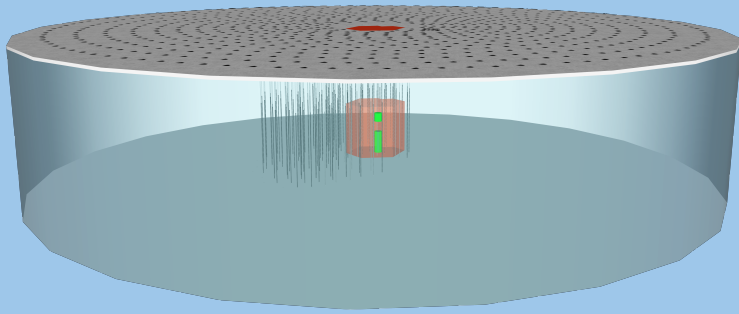
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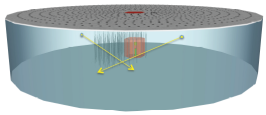


J. Auffenberg for IceCube: ICRC13 ID 0373

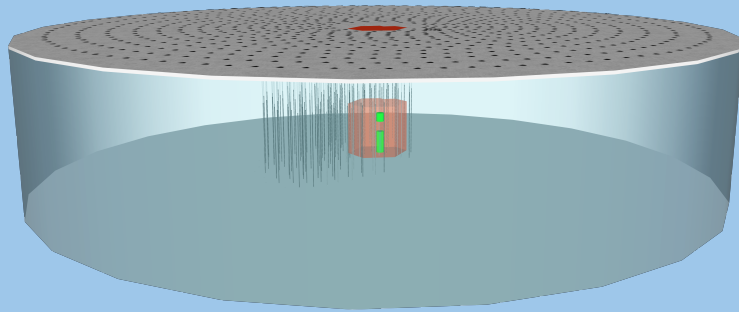


IceVeto Simulation

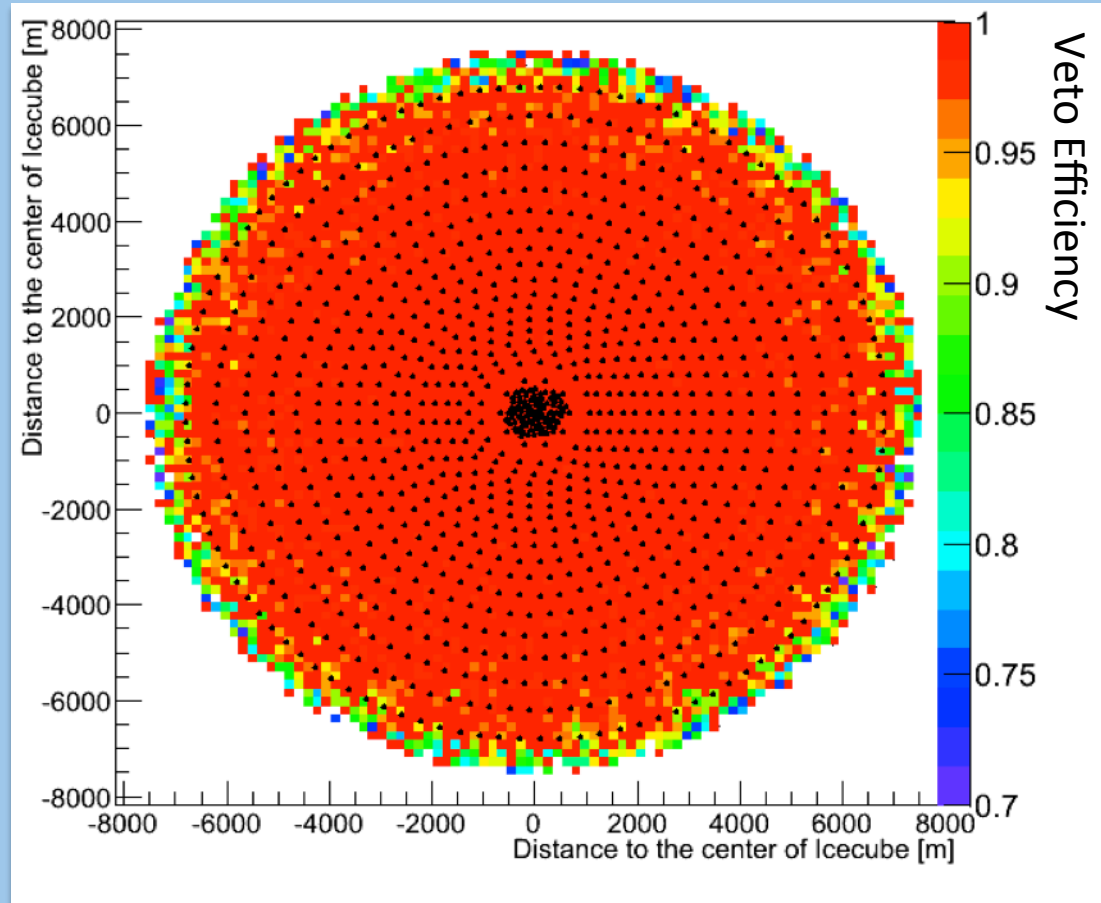


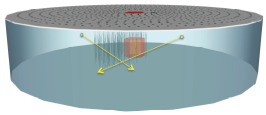


IceVeto Simulation

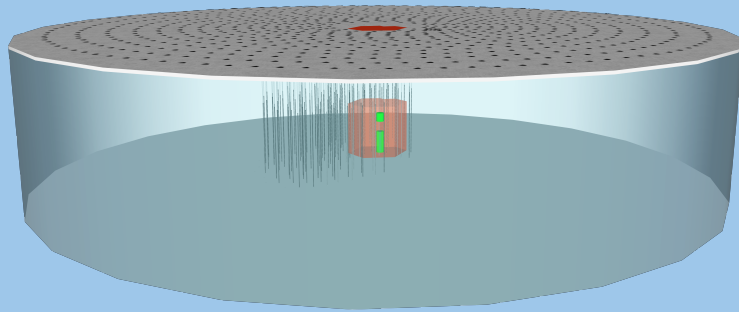


- **943 additional modules on surface**

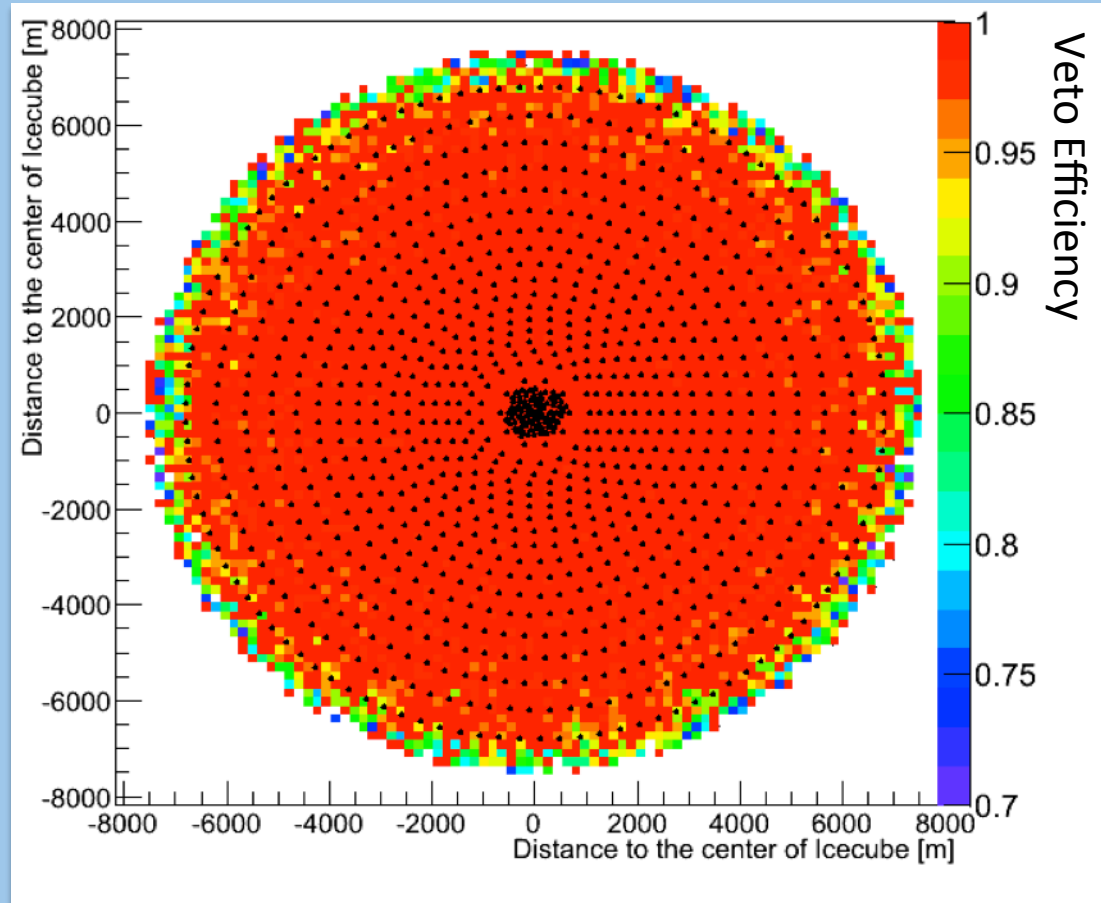


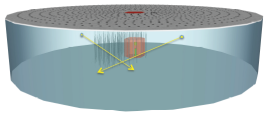


IceVeto Simulation

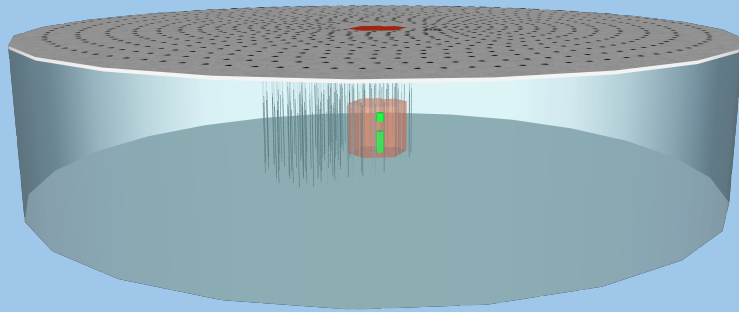


- 943 additional modules on surface
- 99.999% Veto efficiency For PE > 4000

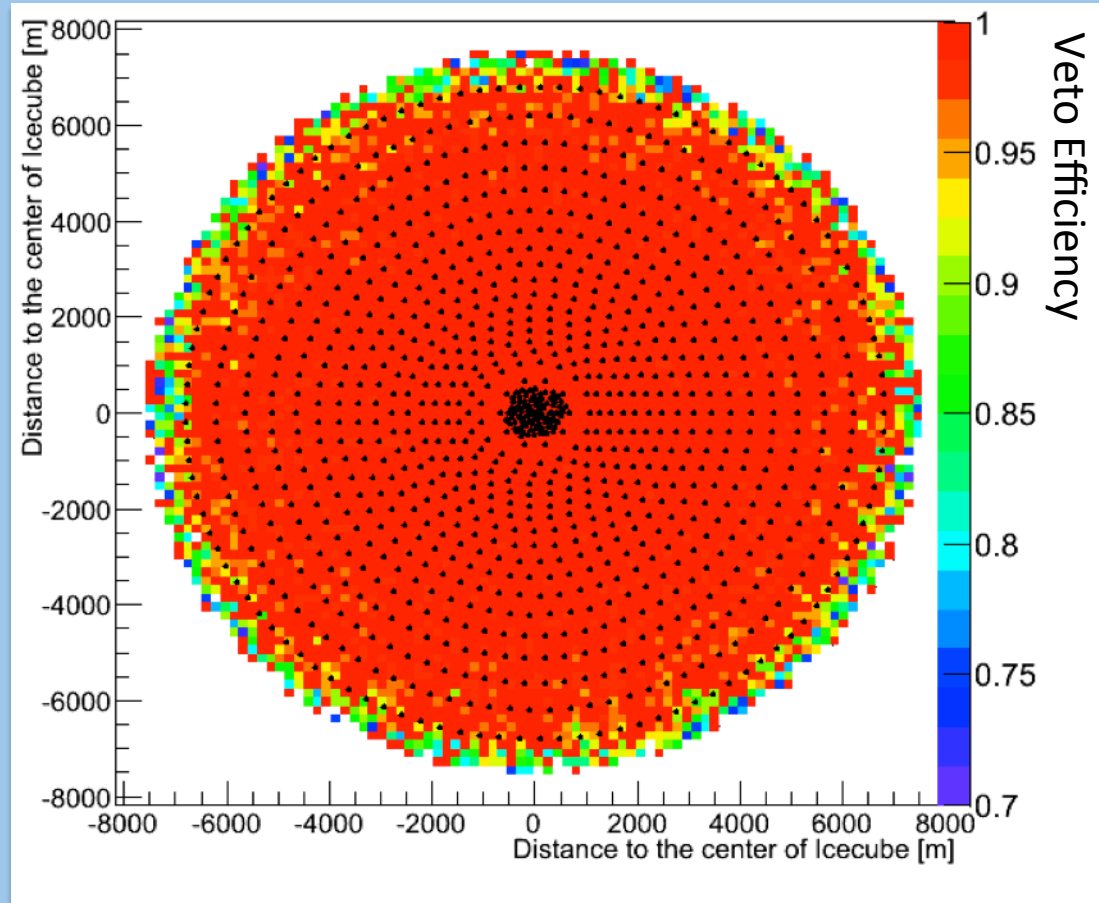




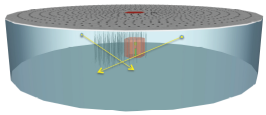
IceVeto Simulation



- **943 additional modules on surface**
- **99.999% Veto efficiency**
For PE > 4000



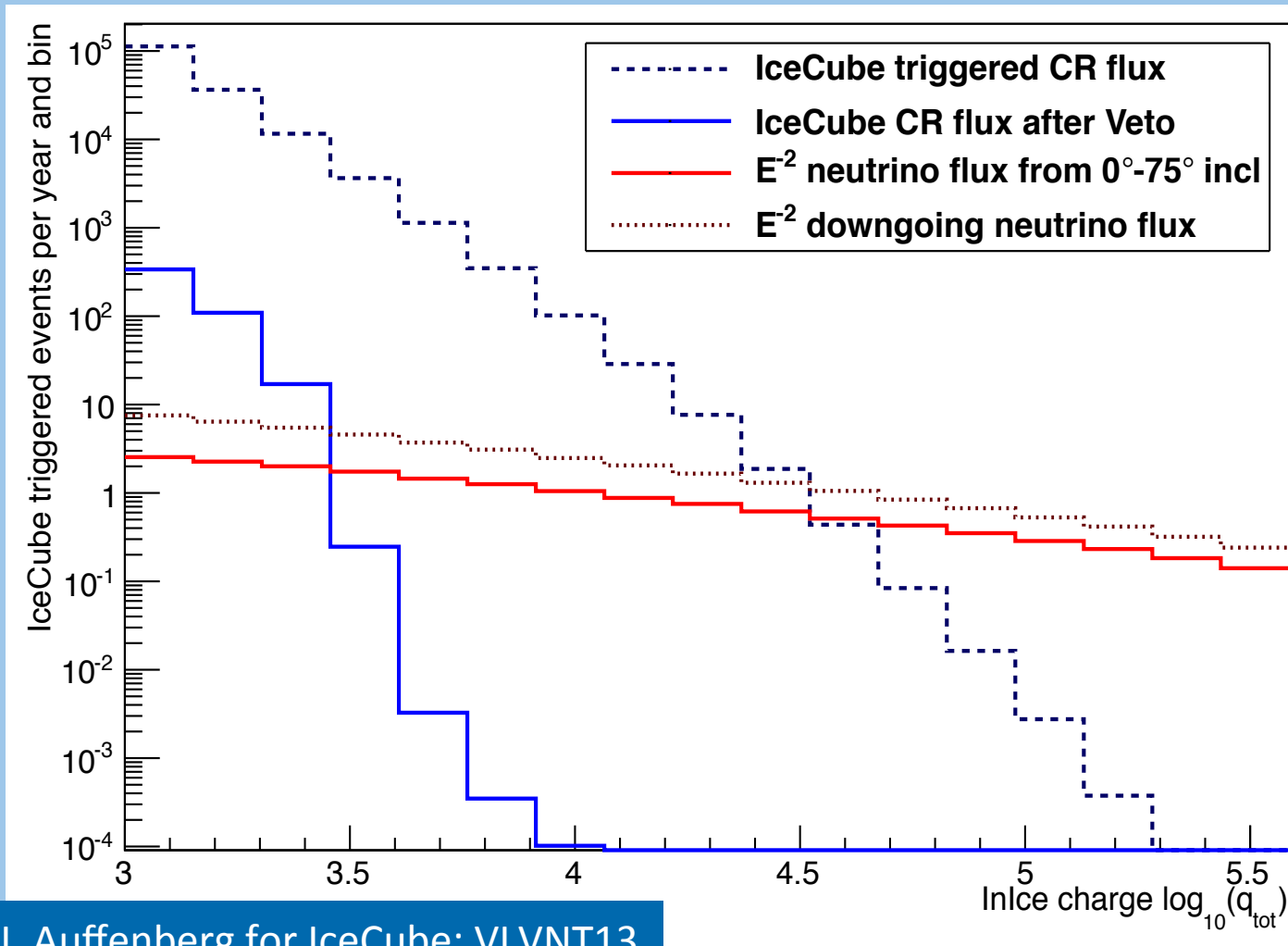
IceVeto is a sub-PeV cosmic-ray energy veto
with 10^{-4} rejection power!



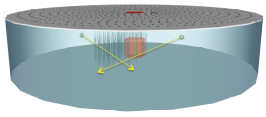
IceVeto Performance



Veto efficiency and neutrino flux calculated based on real data.



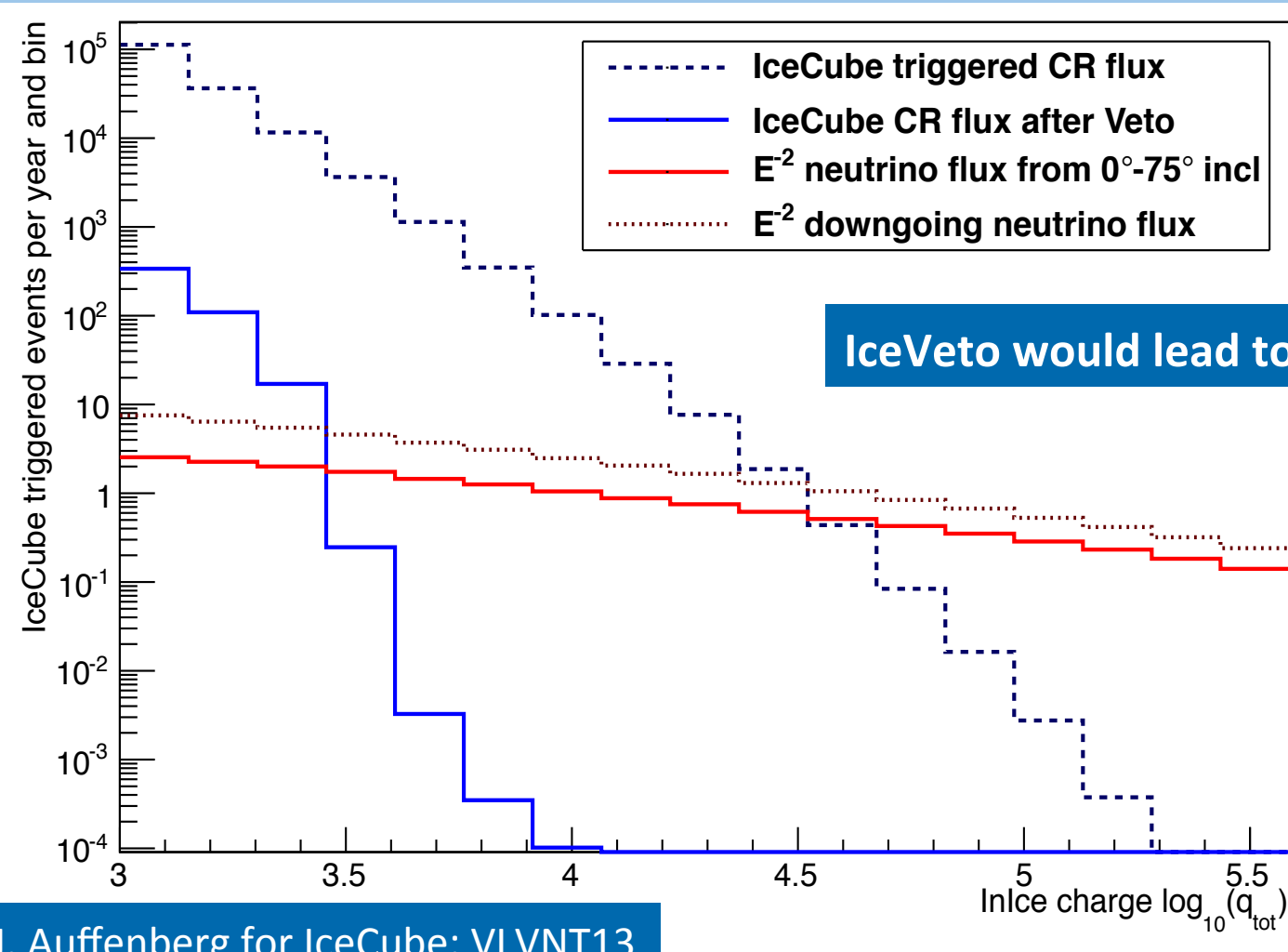
J. Auffenberg for IceCube: VLVNT13



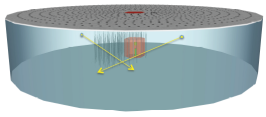
IceVeto Performance



Veto efficiency and neutrino flux calculated based on real data.



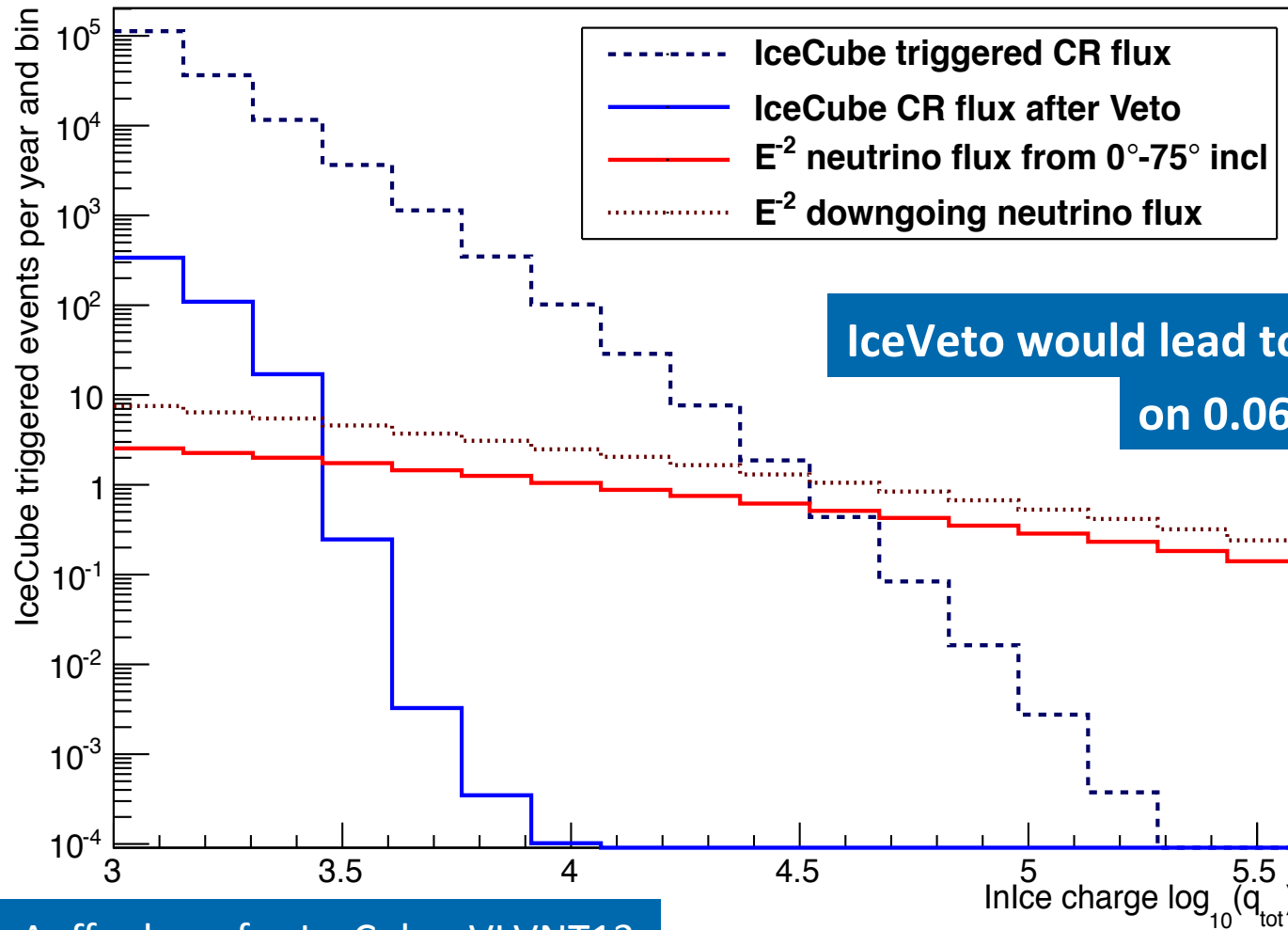
J. Auffenberg for IceCube: VLVNT13



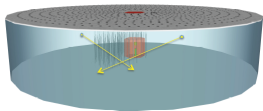
IceVeto Performance



Veto efficiency and neutrino flux calculated based on real data.



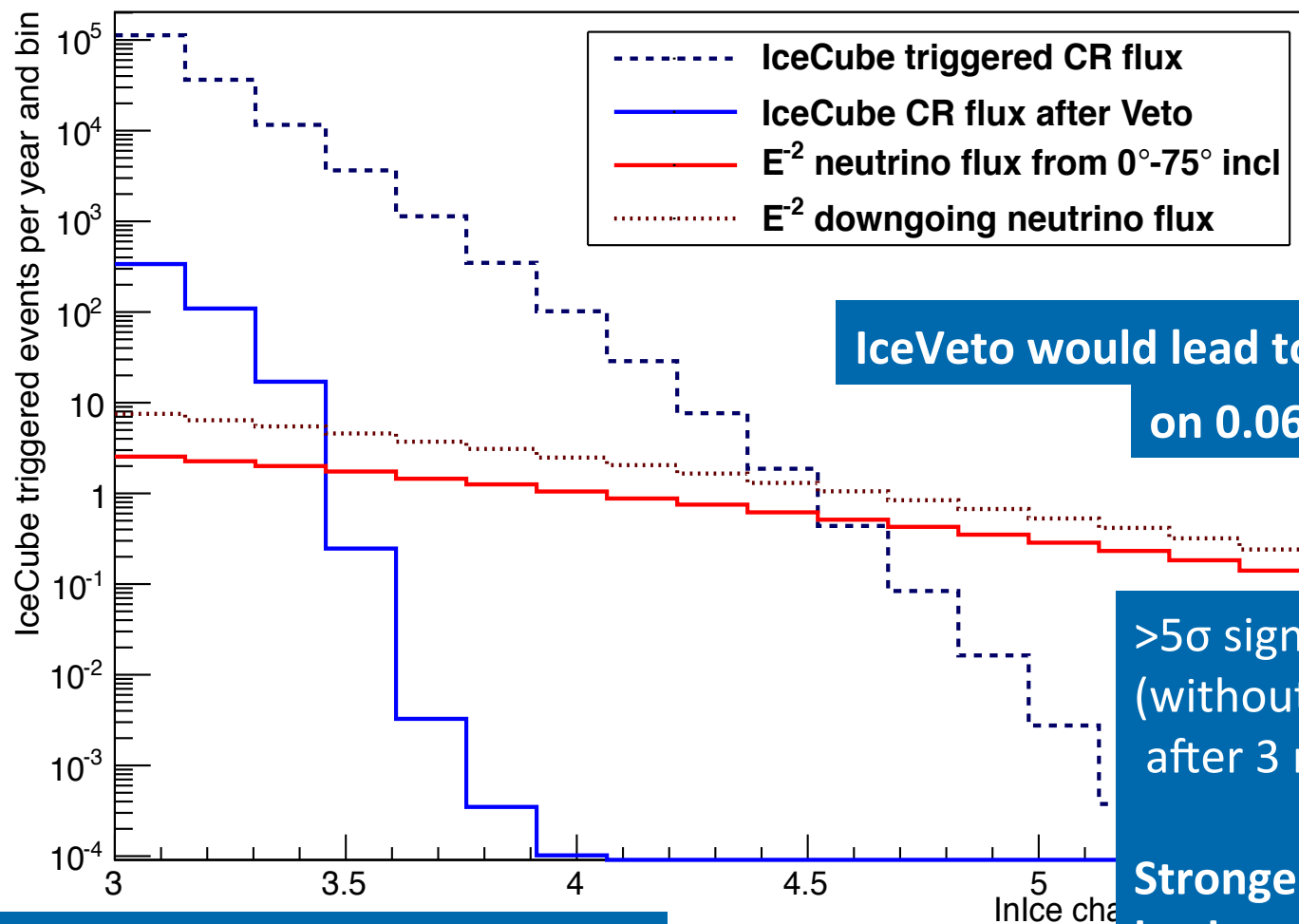
J. Auffenberg for IceCube: VLVNT13



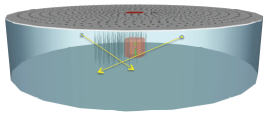
IceVeto Performance



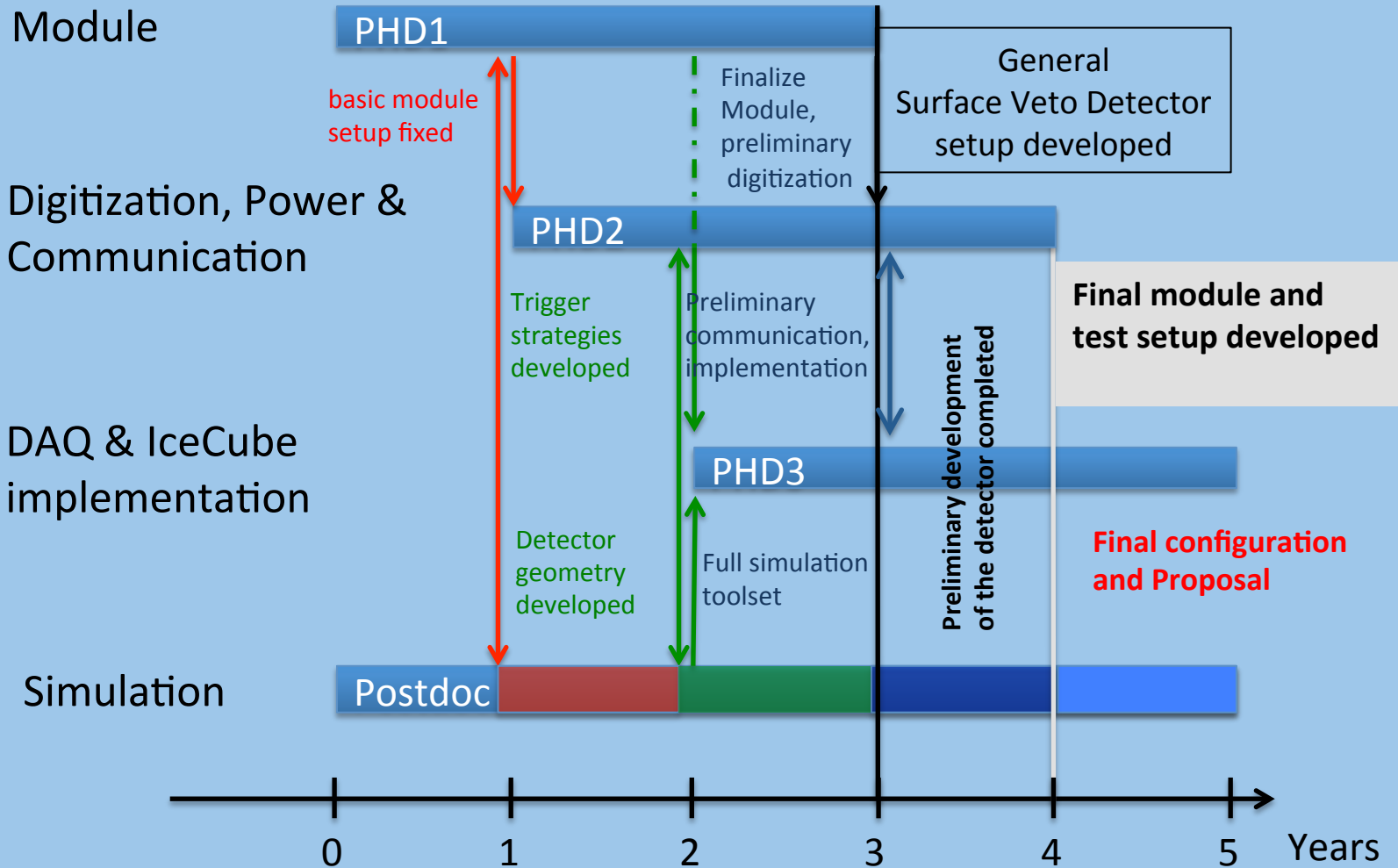
Veto efficiency and neutrino flux calculated based on real data.

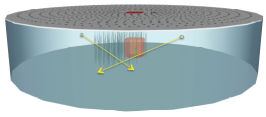


J. Auffenberg for IceCube: VLVNT13

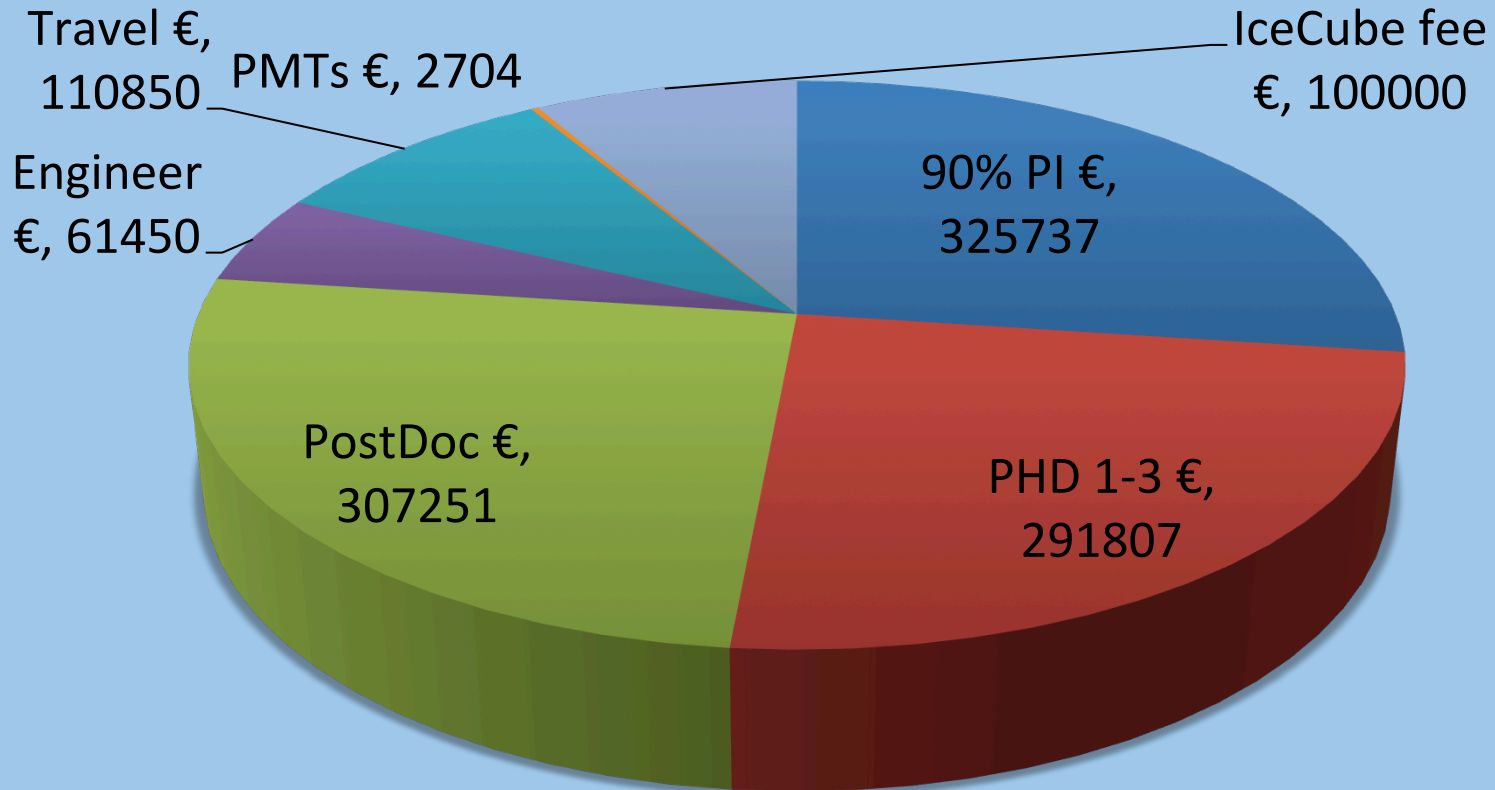


The IceVeto Project

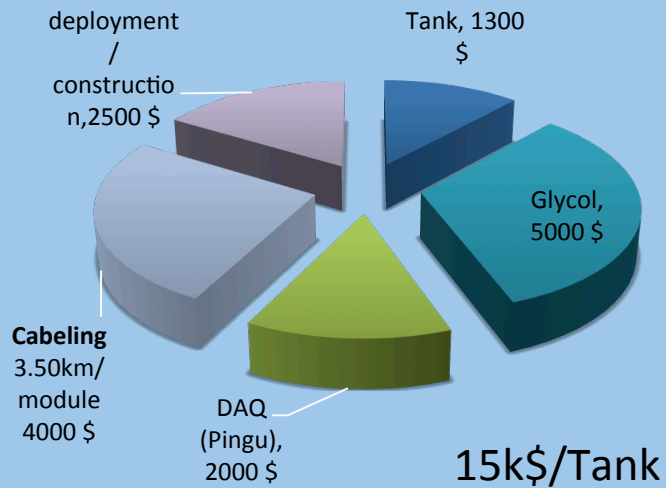




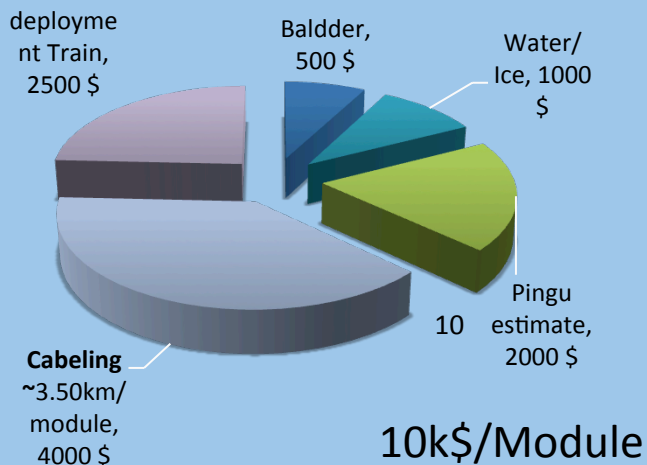
The IceVeto Project



A- Total Direct cost (i + ii) €	1199799
B- indirect cost (overheads)) €	299950
Total Est. Cost & Requested EU contribution) €	1499749



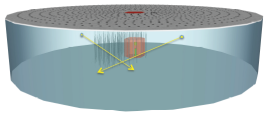
- Tank: \$1300 (2007: \$1135) or just a bladder (\$500)
- Tyvec Liner: \$330 (2007: \$300)
- Maybe Glycol instead of Water: \$5000 (price at the south pole vs. just south pole water \$1000)
- DAQ + PMT: \$2000 (PINGU estimate very likely less)
- Cabling (~3.50 km per module): ~\$4 000 000
- Deployment \$2500 per module or less?



Total: \$8400 - \$13500 per module

\$ 10 – 20 Million for IceVeto

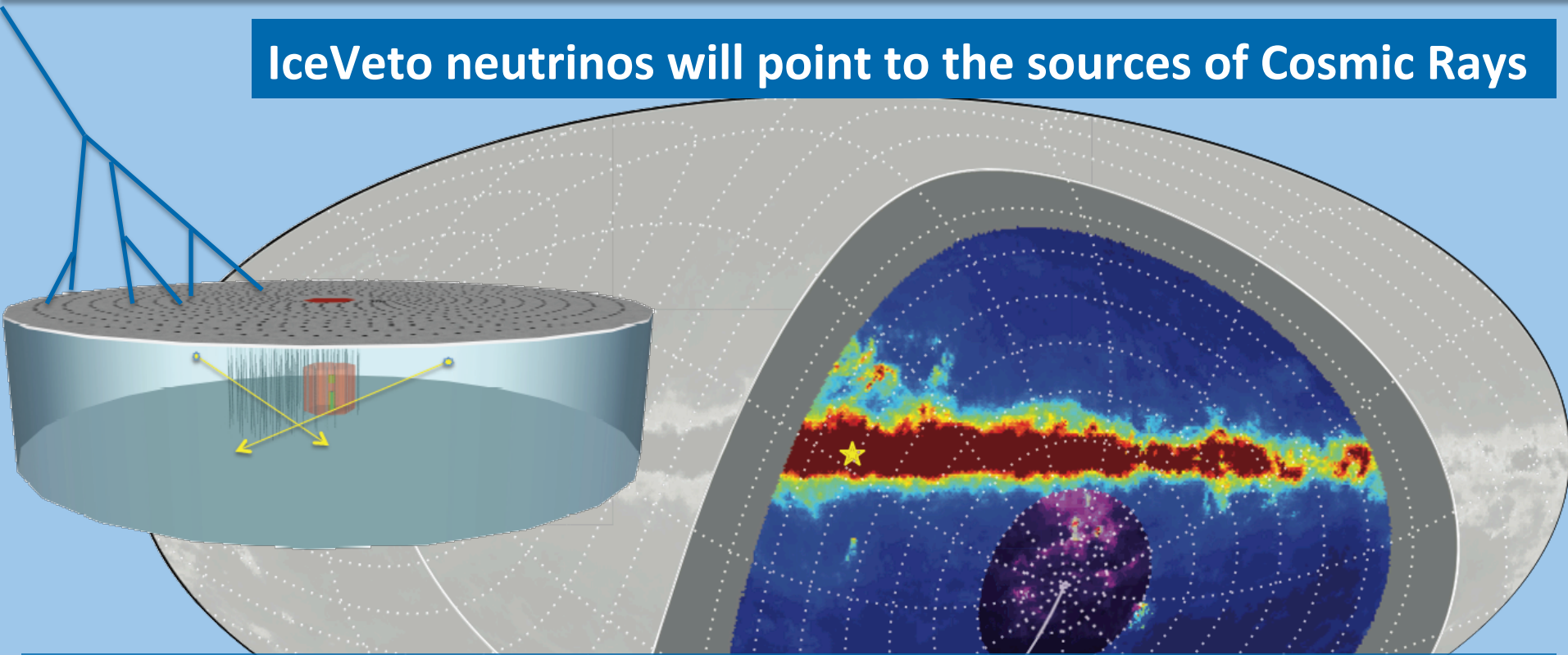
< 1/10 IceCube



IceVeto summary



IceVeto neutrinos will point to the sources of Cosmic Rays



I propose:

- Develop IceVeto, an IceCube add on till 2020
- Use Ice Cherenkov tanks for cosmic-ray detection on the surface
- Move IceCube with IceVeto towards a multi component astroparticle detector
- Open the Southern Sky for PeV muon neutrinos

Helmholtz Alliance for Astroparticle Physics



Detector Design and Technology
for Next Generation Neutrino Observatories

**HAP Workshop Topic 4:
Advanced Technologies**

Program

- Neutrino detection from MeV to EeV energies
- Air shower physics with surface detectors
- Veto strategies
- Optical sensor development
- Radio and acoustic detection technology
- Design studies of future detectors
- New ideas

December 08-10, 2014
at RWTH Aachen

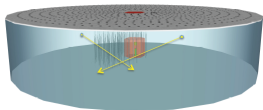
Local Organisation:
Jan Auffenberg, Christopher Wiebusch

Program Committee:
Gisela Anton (Uni Erlangen),
Klaus Helbing (Uni Wuppertal),
Timo Karg, Marek Kowalski (DESY)

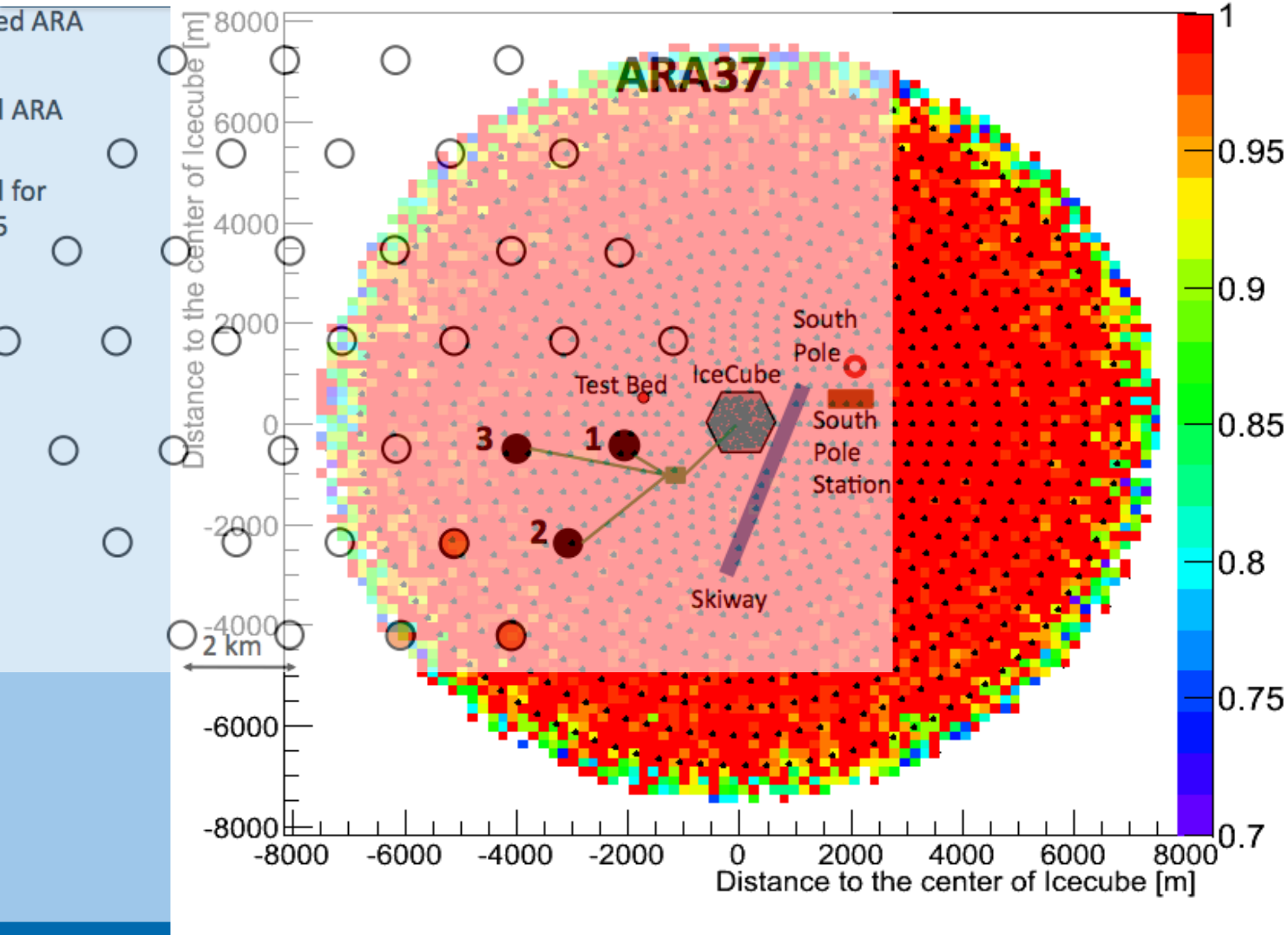
hap2014@physik.rwth-aachen.de
<http://hap2014.physik.rwth-aachen.de>

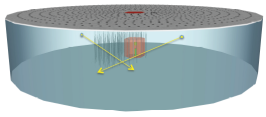
- Neutrino detection from MeV to EeV
- Air-shower physics with surface detectors
- Veto strategies
- Sensor Development and strategies
- Detector Design
- New Ideas

December 8th – 10th 2014
at RWTH Aachen University

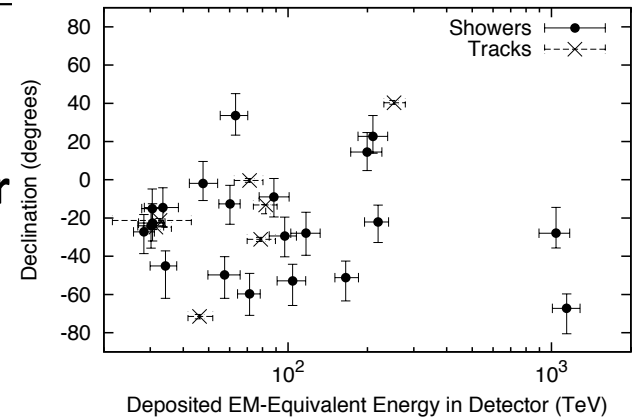
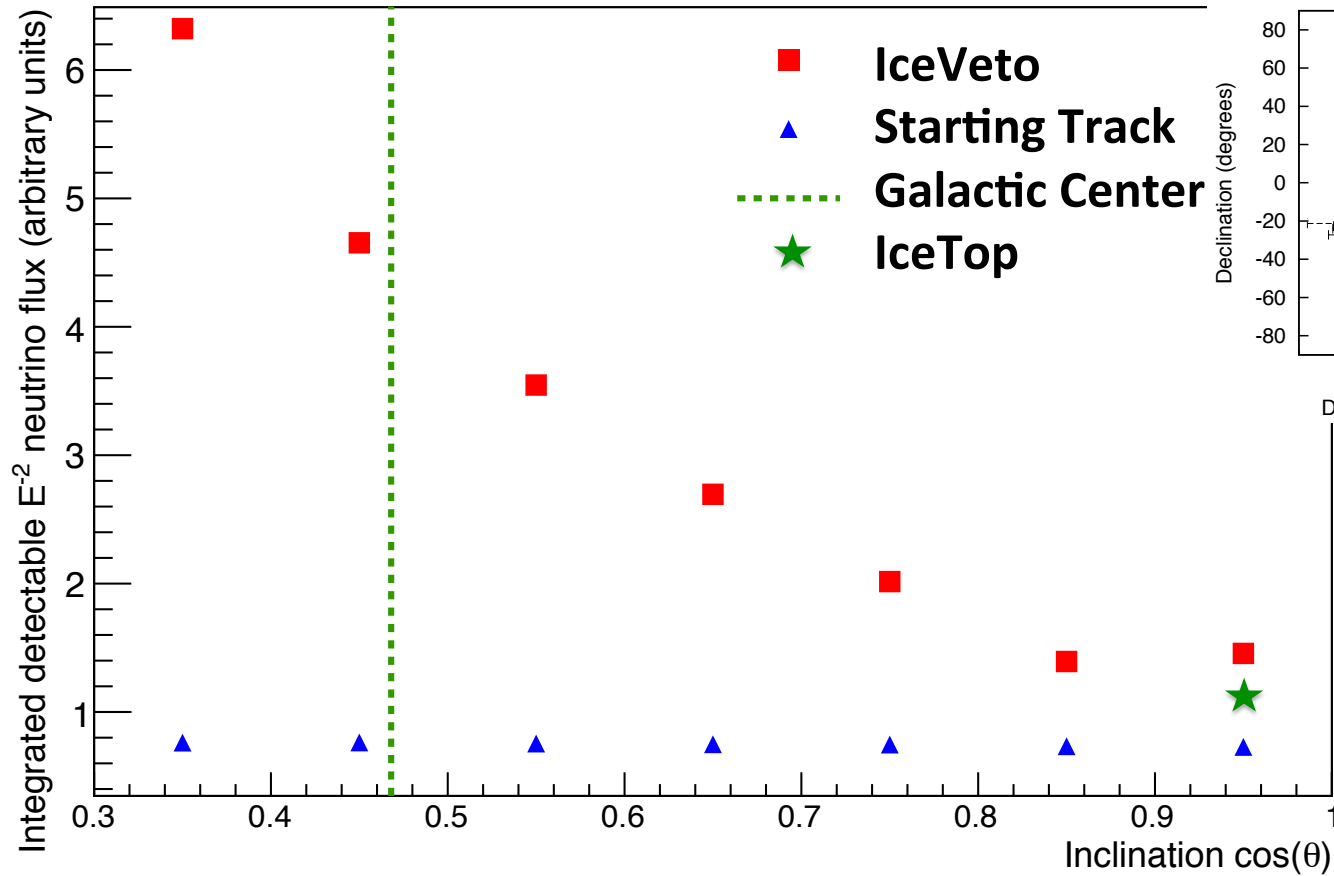


Simulation results

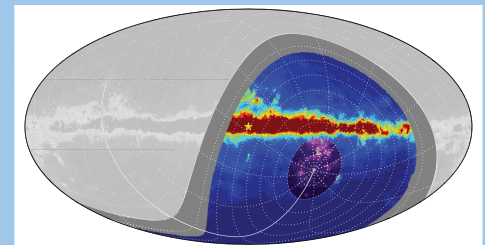




Many more tracks compared to starting track analyses

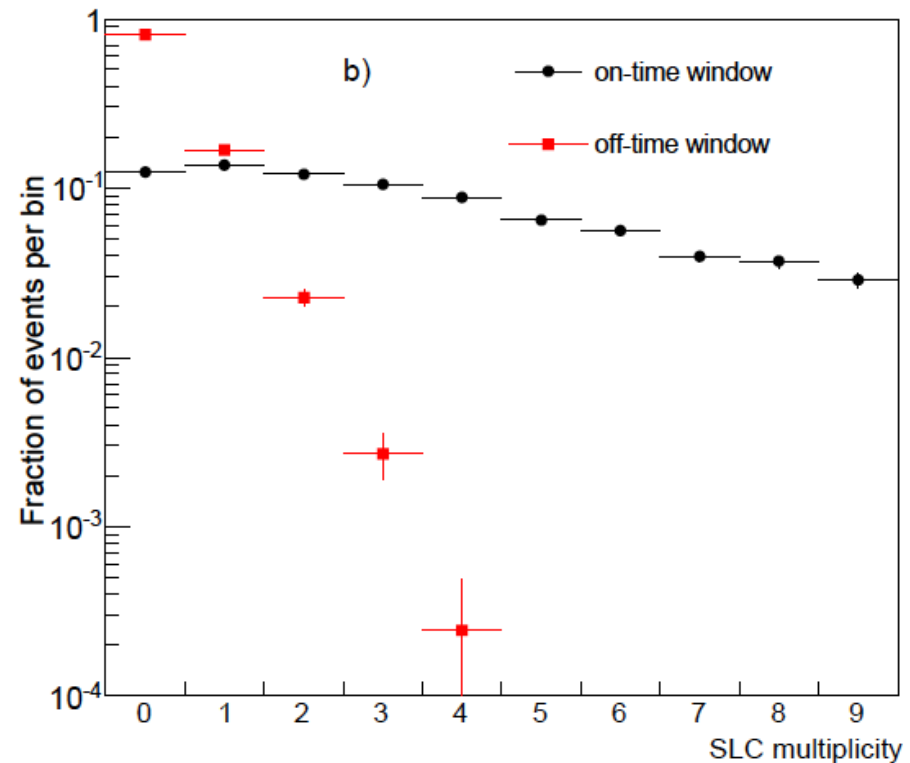
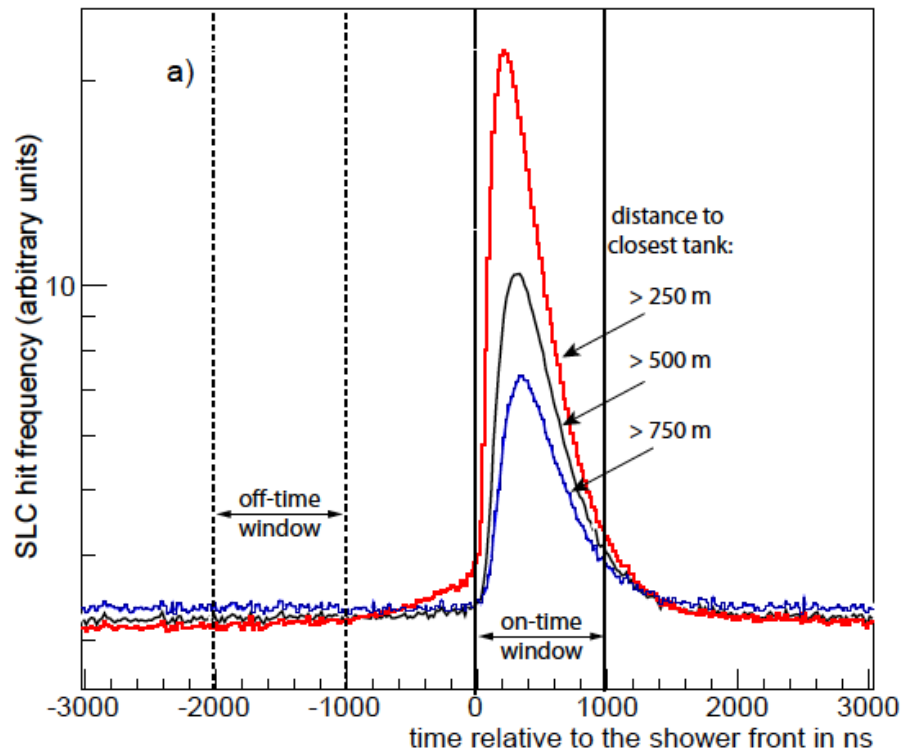


Only the crosses are tracks!



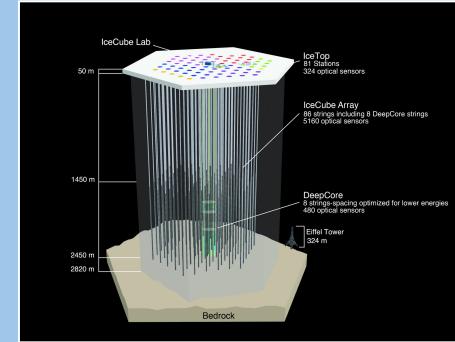
5.5 times more neutrino induced muon tracks from the Galactic Center (62° inclination) between 30 TeV -5 PeV based on an E^{-2} neutrino flux.

Signal Loss due to an IceTop Veto

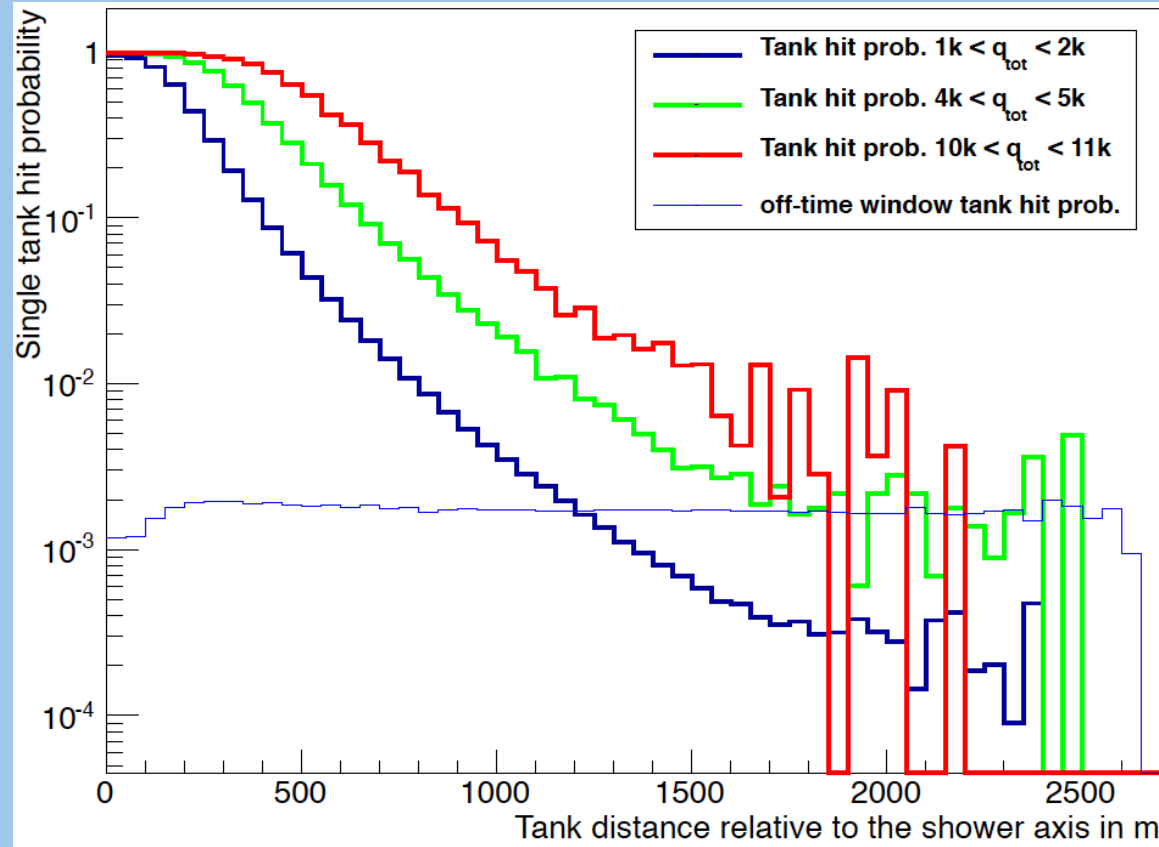


Signal loss below 2% for a >1 Hit cut !

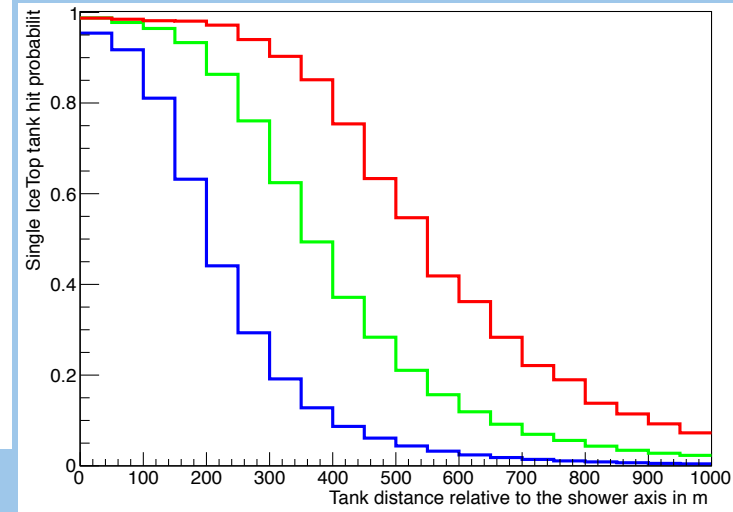
Single IceTop tank hit probability

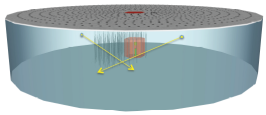


Tank hit probability increases with increasing NPE in the deep detector



The background hit probability is at 2×10^{-3}





What about KM3NeT?

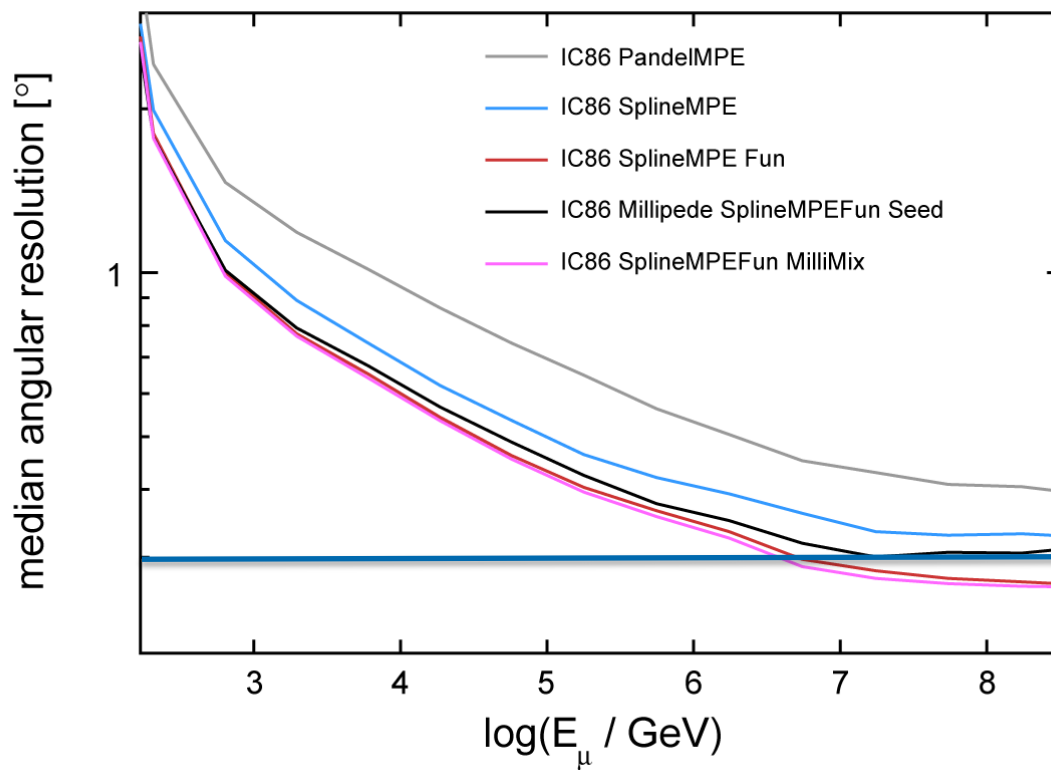


- KM3NeT has the same potential for a surface Veto. (due to good cascade direction reconstructions less important?)
- No high precision positioning for surface Veto modules necessary. (a floating buoy grid? Engineering is not trivial.)
- IceVeto for IceCube is a high energy extension for the observation space of KM3NeT. (Can't see PeV neutrinos from the south).



Slide and work credit to Kai Schatto

Averaging SplineMPE Fun and Millipede

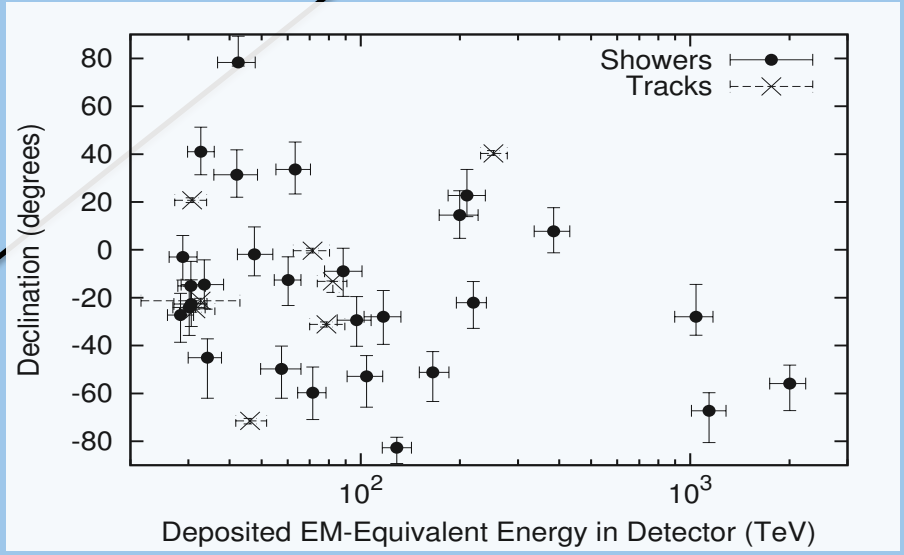
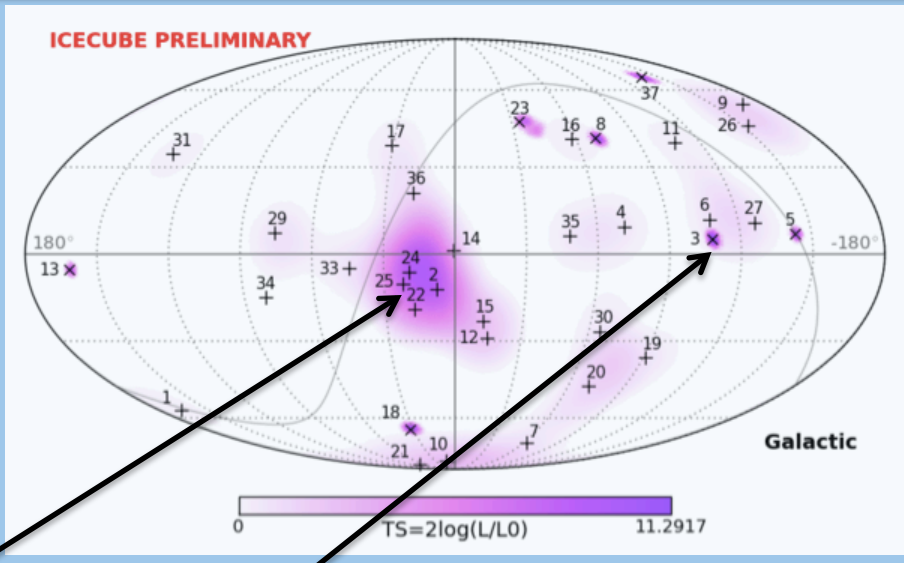


0.1° resolution

Where do the neutrinos come from?



ID	Dep. Energy (TeV)	Observation Time (MJD)	Decl. (deg.)	R.A. (deg.)	Med. Angular Error (deg.)	Event Topology
1	47.6 ^{+6.5} _{-8.4}	55351.3222143	-1.8	35.2	16.3	Shower
2	117 ⁺¹⁵ ₋₁₅	55351.4659661	-28.0	282.6	25.4	Shower
3	78.7 ^{+10.8} _{-3.7}	55451.0707482	-31.2	127.9	$\lesssim 1.4$	Track
4	165 ⁺²⁰ ₋₁₅	55477.3930984	-51.2	169.5	7.1	Shower
5	71.4 ^{+9.0} _{-9.0}	55512.5516311	-0.4	110.6	$\lesssim 1.2$	Track
6	28.4 ^{+2.7} _{-2.5}	55567.6388127	-27.2	133.9	9.8	Shower
7	34.3 ^{+3.5} _{-4.3}	55571.2585362	-45.1	15.6	24.1	Shower
8	32.6 ^{+10.3} _{-4.3}	55608.8201315	-21.2	182.4	$\lesssim 1.3$	Track
9	63.2 ^{+11.1} _{-8.0}	55685.6629713	33.6	151.3	16.5	Shower
10	97.2 ^{+10.4} _{-12.4}	55695.2730461	-29.4	5.0	8.1	Shower
11	88.4 ^{+12.5} _{-10.7}	55714.5909345	-8.9	155.3	16.7	Shower
12	104 ⁺¹³ ₋₁₃	55739.4411232	-52.8	296.1	9.8	Shower
13	253 ⁺²⁶ ₋₂₉	55756.1129844	40.3	67.9	$\lesssim 1.2$	Track
14	1041 ⁺¹³² ₋₁₄₄	55782.5161911	-27.9	265.6	13.2	Shower
15	57.5 ^{+8.3} _{-7.8}	55783.1854223	-49.7	287.3	19.7	Shower
16	30.6 ^{+3.6} _{-3.5}	55798.6271285	-22.6	192.1	19.4	Shower
17	200 ⁺²⁷ ₋₂₇	55800.3755483	14.5	247.4	11.6	Shower
18	31.5 ^{+4.6} _{-3.3}	55923.5318204	-24.8	345.6	$\lesssim 1.3$	Track
19	71.5 ^{+7.0} _{-7.2}	55925.7958619	-59.7	76.9	9.7	Shower
20	1141 ⁺¹⁴³ ₋₁₃₃	55929.3986279	-67.2	38.3	10.7	Shower
21	30.2 ^{+3.5} _{-3.3}	55936.5416484	-24.0	9.0	20.9	Shower
22	220 ⁺²¹ ₋₂₄	55941.9757813	-22.1	293.7	12.1	Shower
23	82.2 ^{+8.6} _{-5.4}	55949.5693228	-13.2	208.7	$\lesssim 1.9$	Track
24	30.5 ^{+3.4} _{-2.6}	55950.8474912	-15.1	282.2	15.5	Shower
25	33.5 ^{+4.9} _{-5.0}	55966.7422488	-14.5	286.0	46.3	Shower
26	210 ⁺²⁹ ₋₂₆	55979.2551750	22.7	143.4	11.8	Shower
27	60.2 ^{+5.6} _{-5.6}	56008.6845644	-12.6	121.7	6.6	Shower
28	46.1 ^{+5.7} _{-4.4}	56048.5704209	-71.5	164.8	$\lesssim 1.3$	Track
29	32.7 ^{+2.9} _{-2.9}	56108.2572046	41.0	298.1	7.4	Shower
30	129 ⁺¹⁴ ₋₁₂	56115.7283574	-82.7	103.2	8.0	Shower
31	42.5 ^{+5.4} _{-5.7}	56176.3914143	78.3	146.1	26.0	Shower
32	—	56211.7401231	—	—	—	Coincident
33	385 ⁺⁴⁶ ₋₄₉	56221.3424023	7.8	292.5	13.5	Shower
34	42.1 ^{+6.5} _{-6.3}	56228.6055226	31.3	323.4	42.7	Shower
35	2004 ⁺²³⁶ ₋₂₆₂	56265.1338677	-55.8	208.4	15.9	Shower
36	28.9 ^{+3.0} _{-2.6}	56308.1642740	-3.0	257.7	11.7	Shower
37	30.8 ^{+3.3} _{-3.5}	56390.1887627	20.7	167.3	$\lesssim 1.2$	Track



Hotspot of cascades only
p-value: 7%

All event Hotspot
p-value: 87%

MC simulation(based on real data) ICECUBE

Assumptions:

IceVeto detection module acts as a CR Veto similar to an IceTop tank.

Input parameters:

- Veto efficiency for events with reconstruction between 0-75° inclination with 1000 PE light deposit in the detector with >99.9% .
- IceVeto tanks are forced on rings around IceTop.

Input from real data:

- IceTop tank hit probability as function of PE.
- Geometrical event distribution.

