

# A parameterization of the radio signal as measured on the ground

or

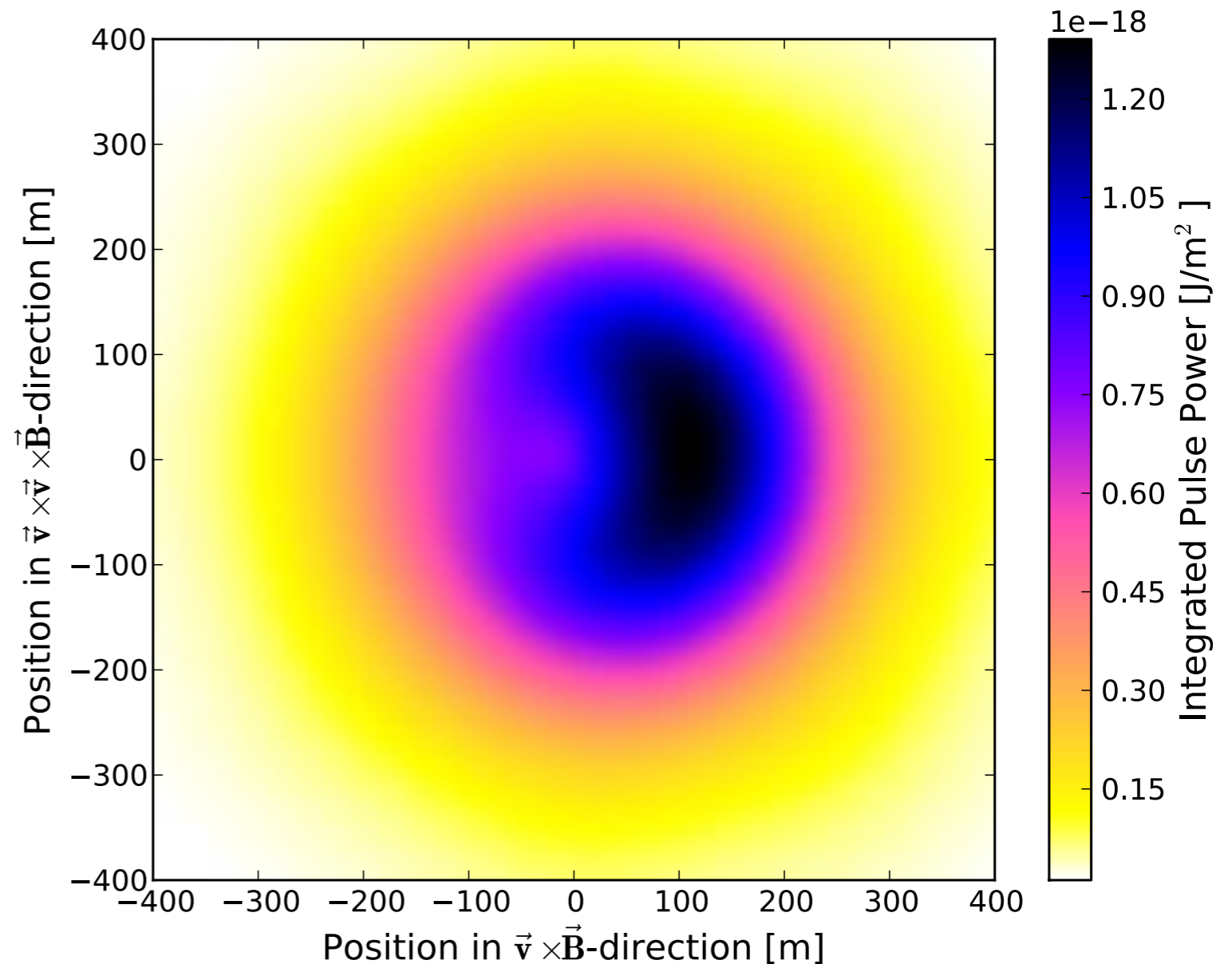
“A radio LDF”

Anna Nelles

Radboud University Nijmegen



and the LOFAR  
Key Science Project  
Cosmic Rays



**CORSIKA + CoREAS simulation of air shower**

zenith angle: 45 degrees, azimuth angle: 13 degrees  
energy:  $2 \times 10^{18}$  eV

# Idea: Model the “bean shape”

All simulations show “bean shape” in shower plane, when rotated in  $\mathbf{v} \times \mathbf{B}$  direction

$\mathbf{v}$  : shower axis

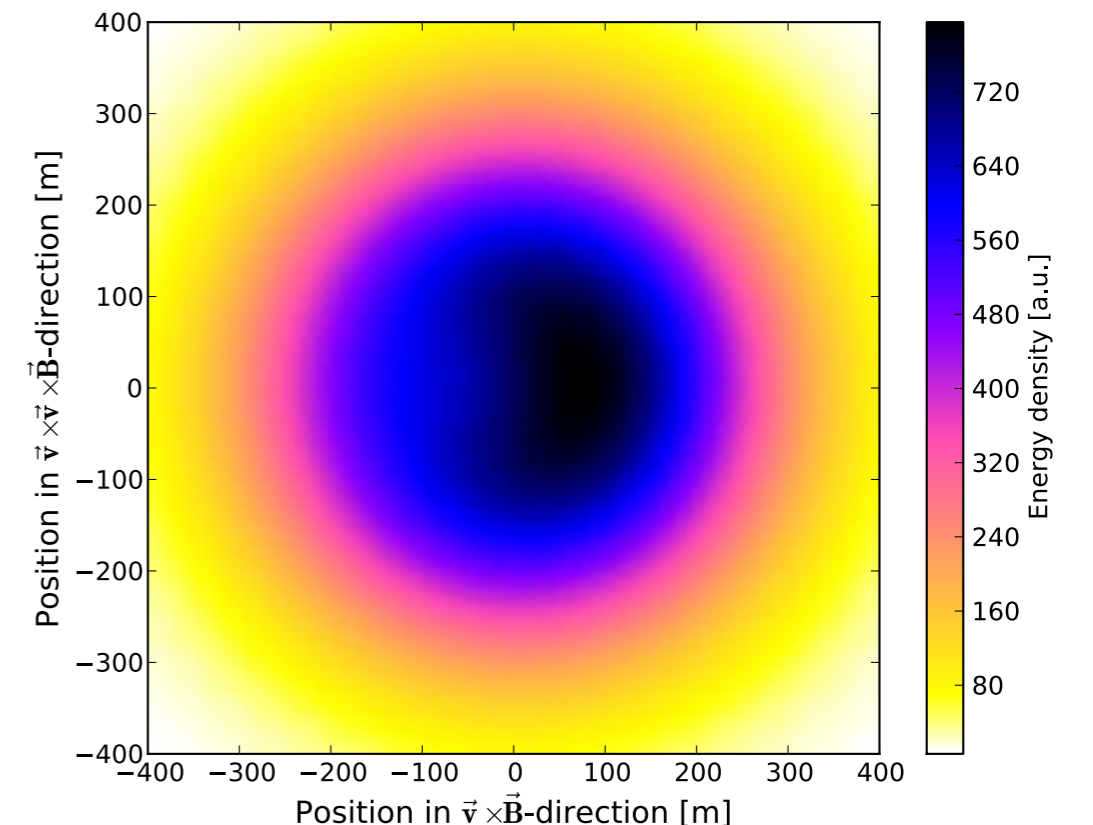
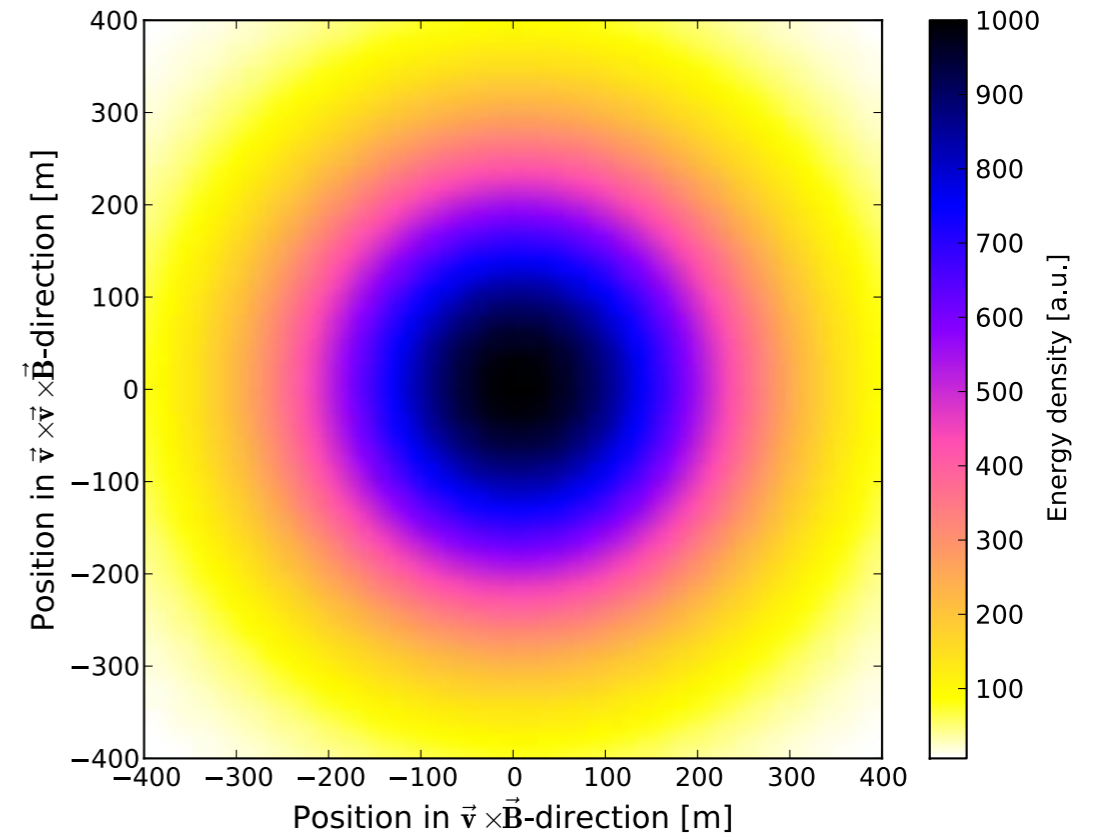
$\mathbf{B}$ : magnetic field vector

Basic distribution: two-dimensional Gaussian

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_+)^2 + (y' - Y_+)^2]}{\sigma_+^2}\right)$$

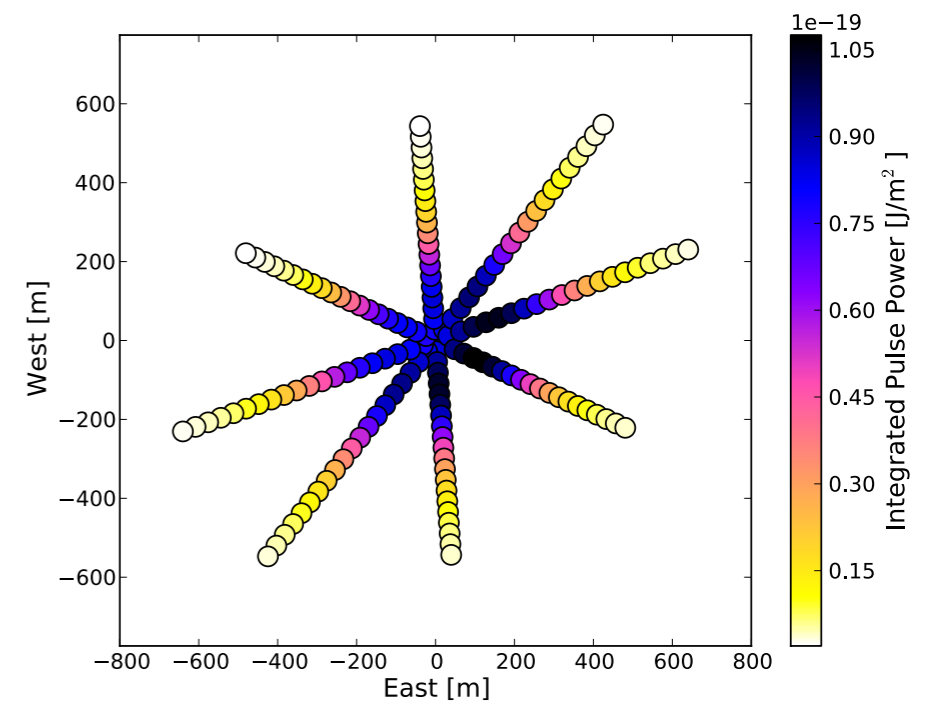
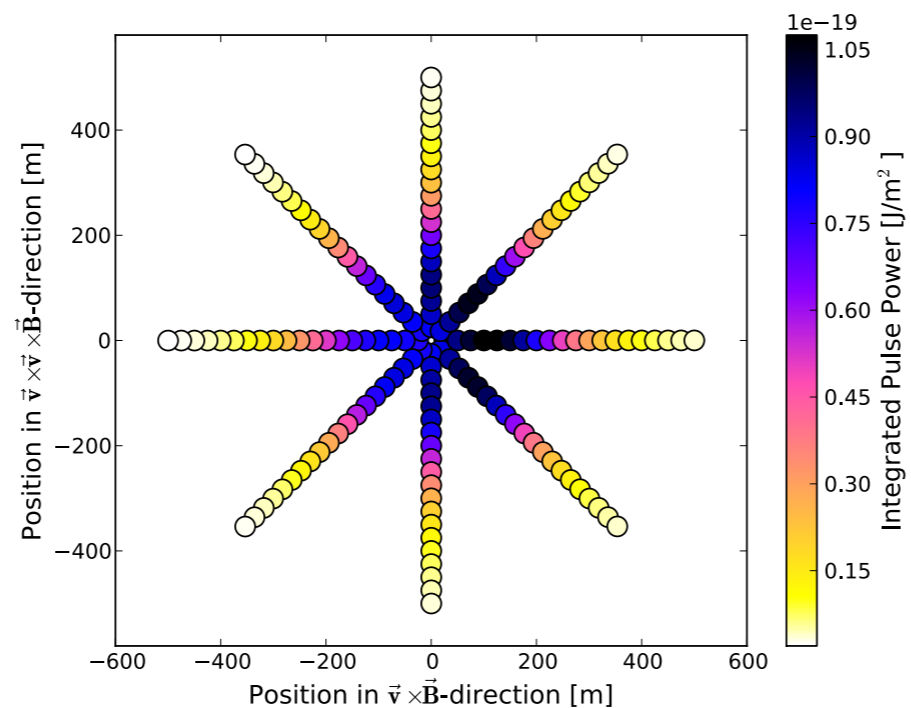
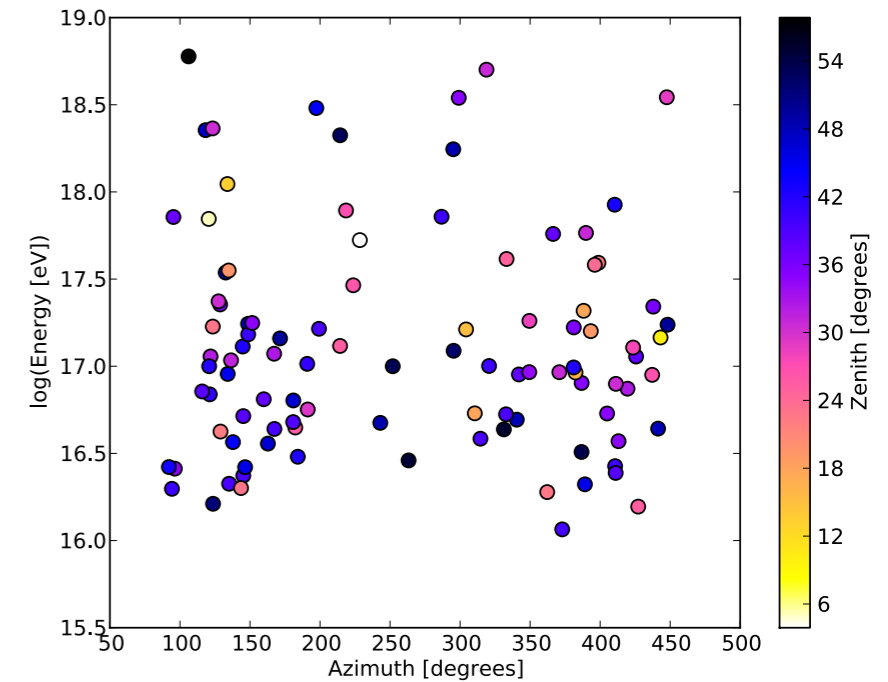
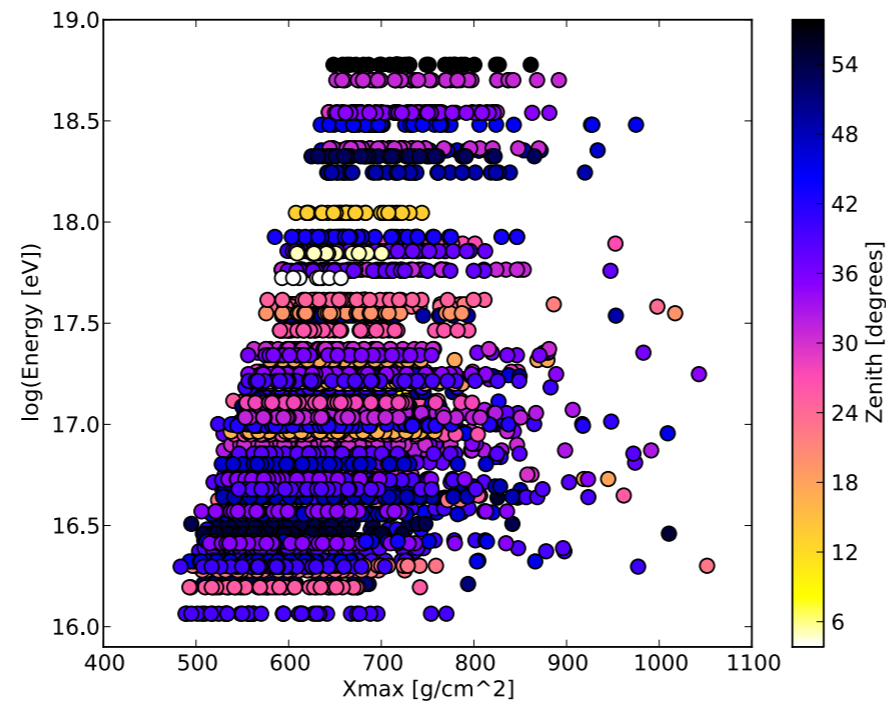
Create “bean shape” by subtracting second Gaussian

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_+)^2 + (y' - Y_+)^2]}{\sigma_+^2}\right) - A_- \cdot \exp\left(\frac{-[(x' - X_-)^2 + (y' - Y_-)^2]}{\sigma_-^2}\right)$$



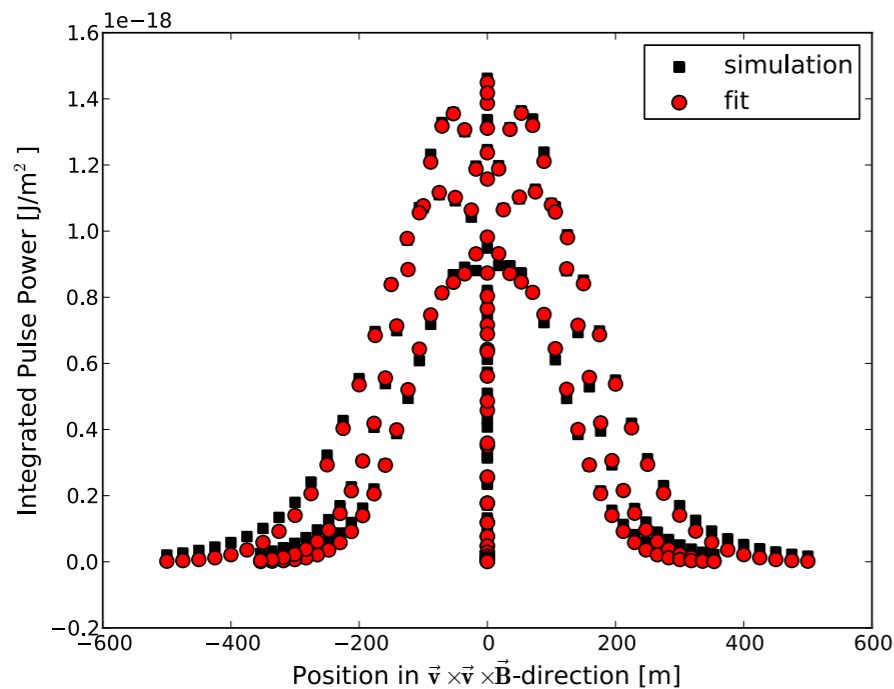
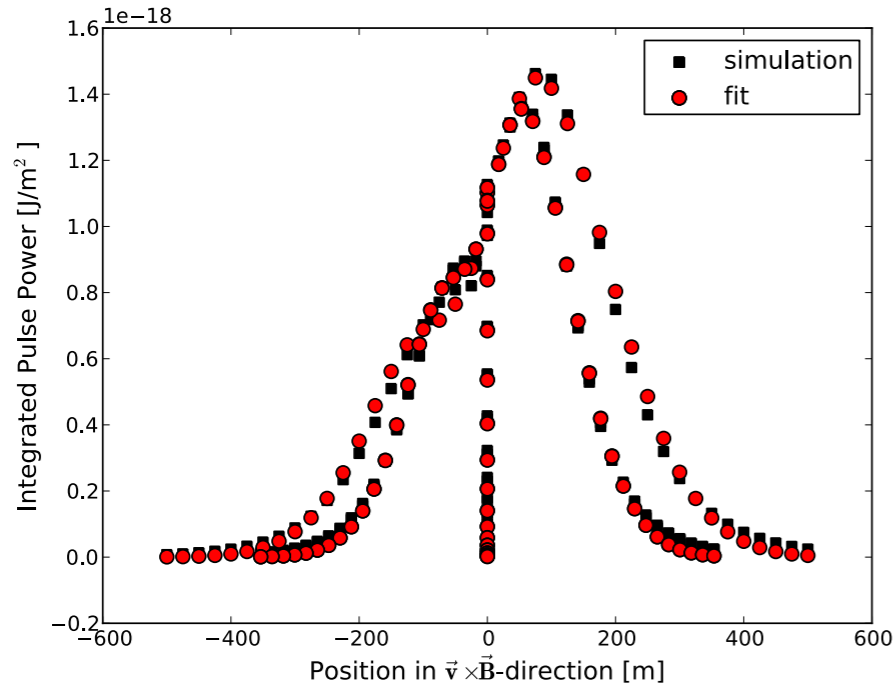
# Simulation Set-up

- about 1250 proton and 750 iron air showers
- CORSIKA 7.400  
FLUKA 2011.2b  
QGSJETII.04 95  
CoREAS
- LOFAR specific height and magnetic field
- Antennas on star-shaped pattern in shower plane, allows good interpolation of signals
- Mis-aligned star-pattern show same results

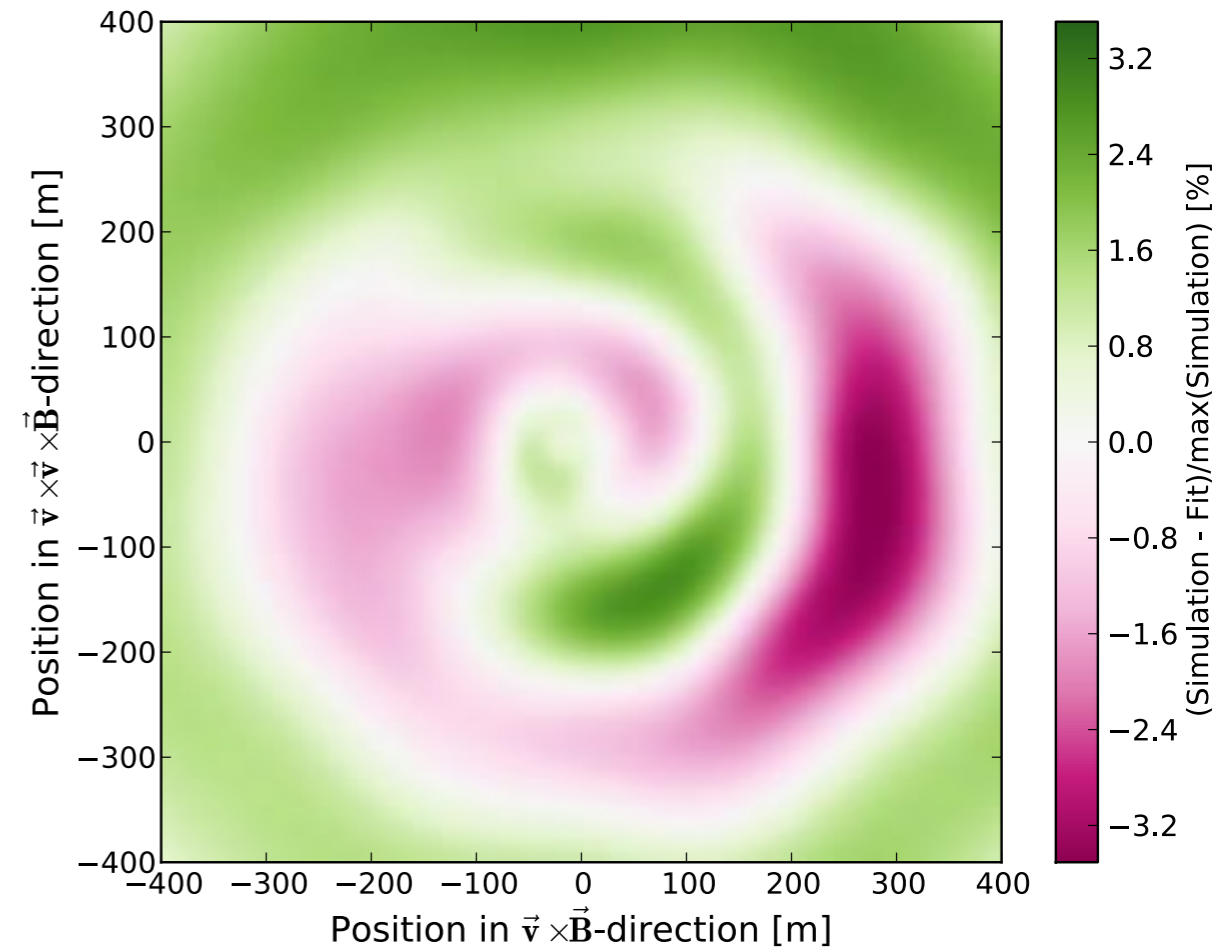


# Fitting the model

## Two-dimensional Fit



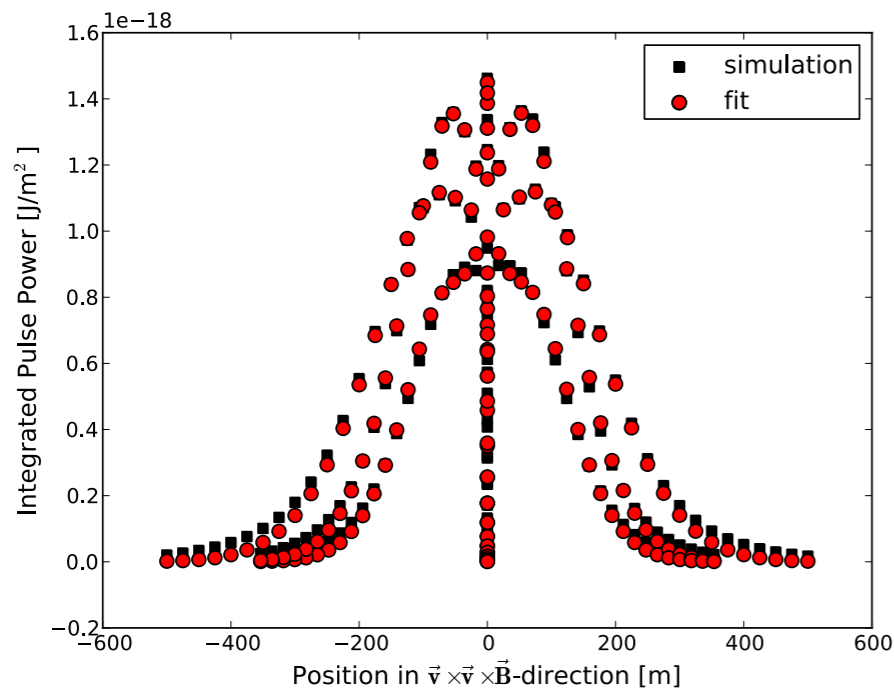
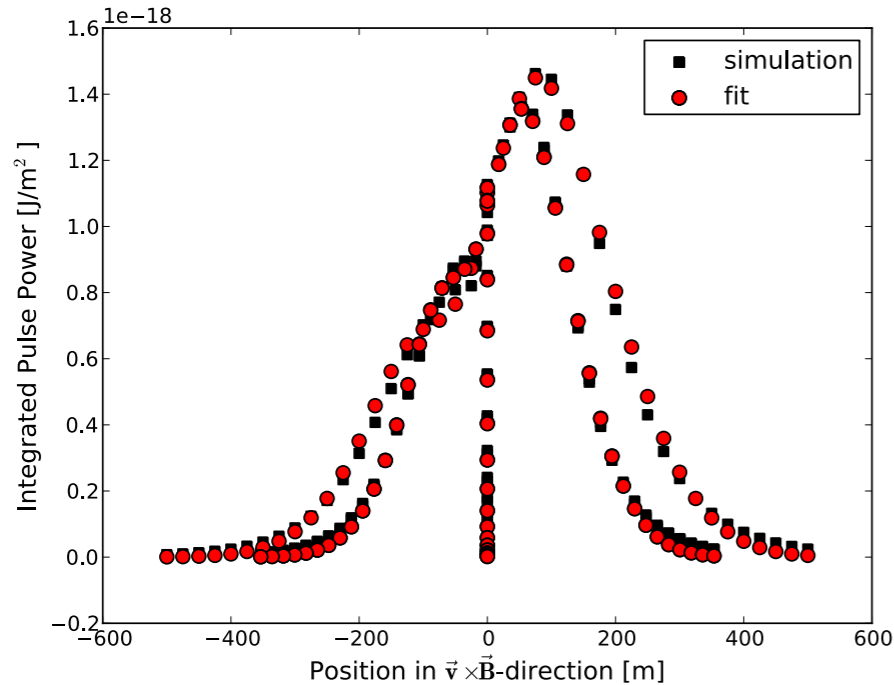
## Relative Differences



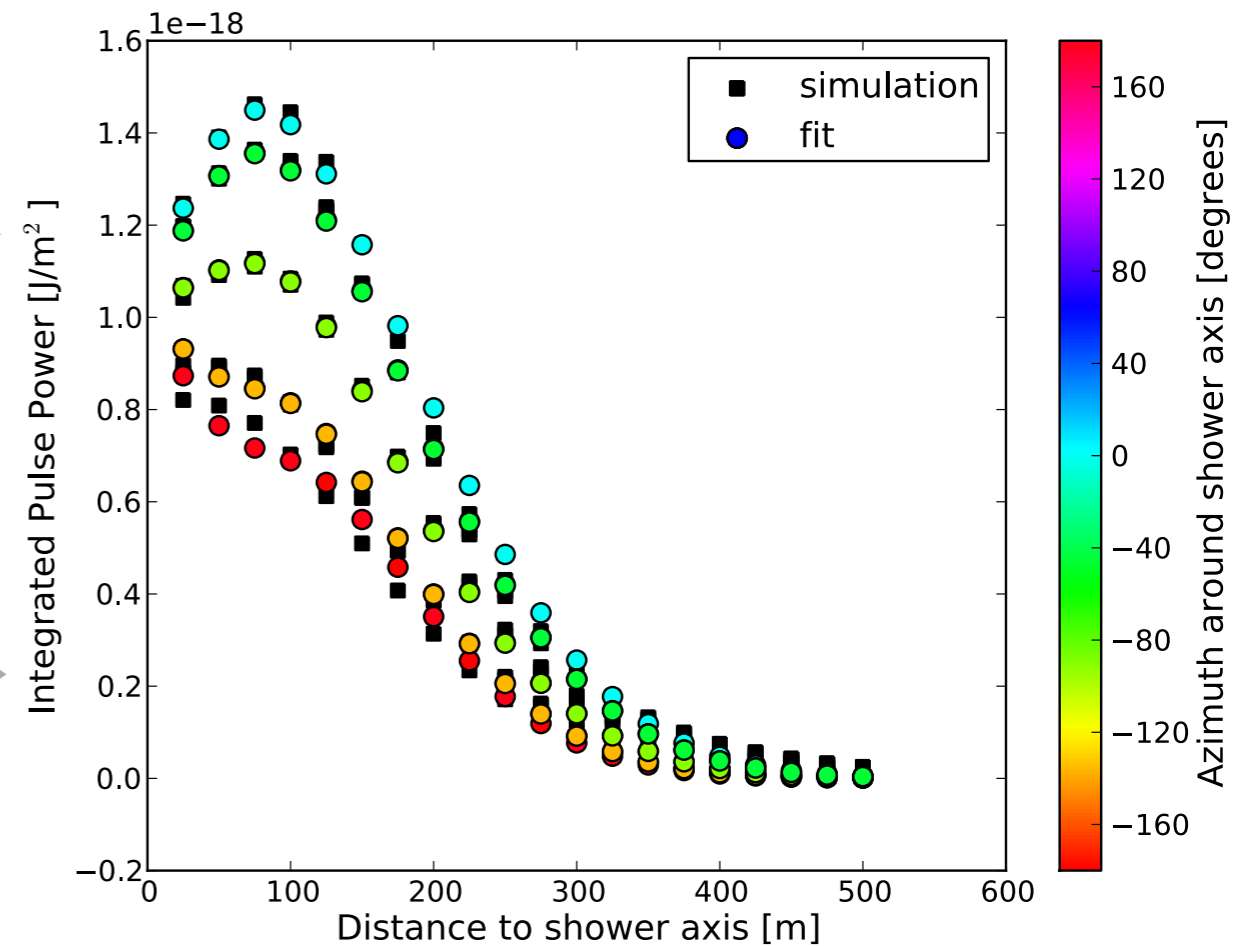
- Deviations less than 10% for most events

# Fitting the model

## Two-dimensional Fit

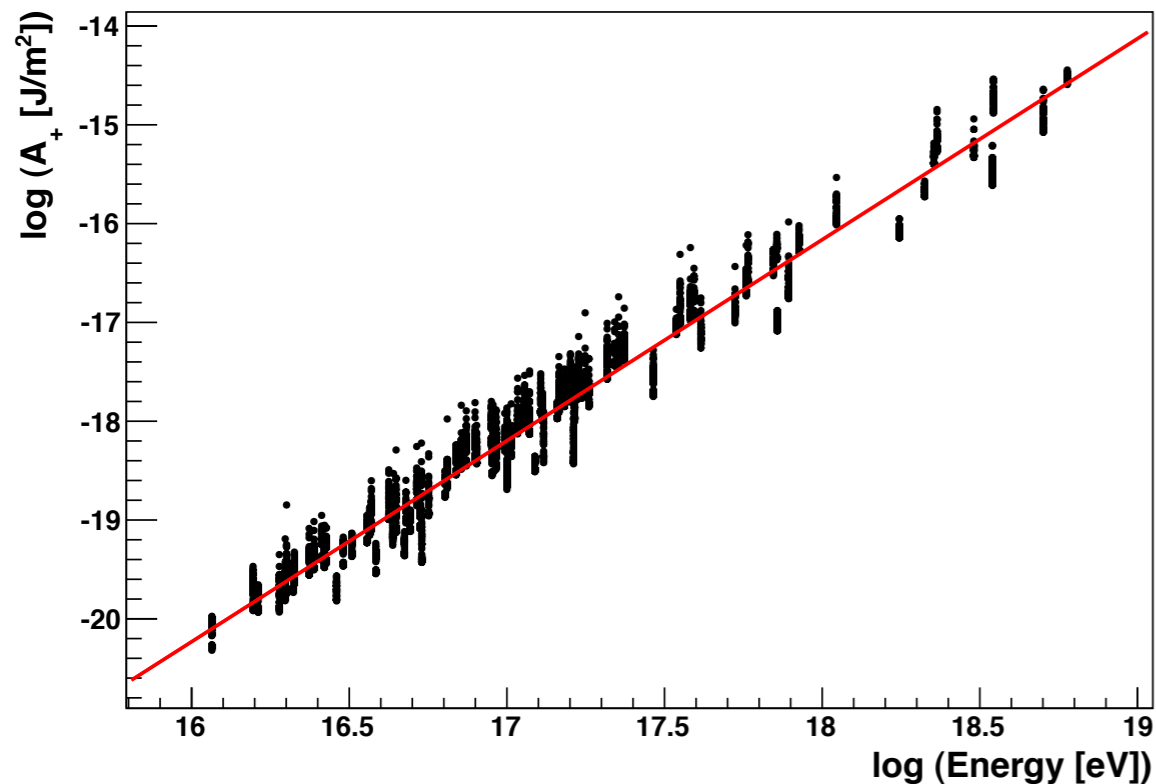


## “Classical LDF view”



- Fit captures all features of the shape
- Azimuthal asymmetry
- 1-d LDF cannot work without correction

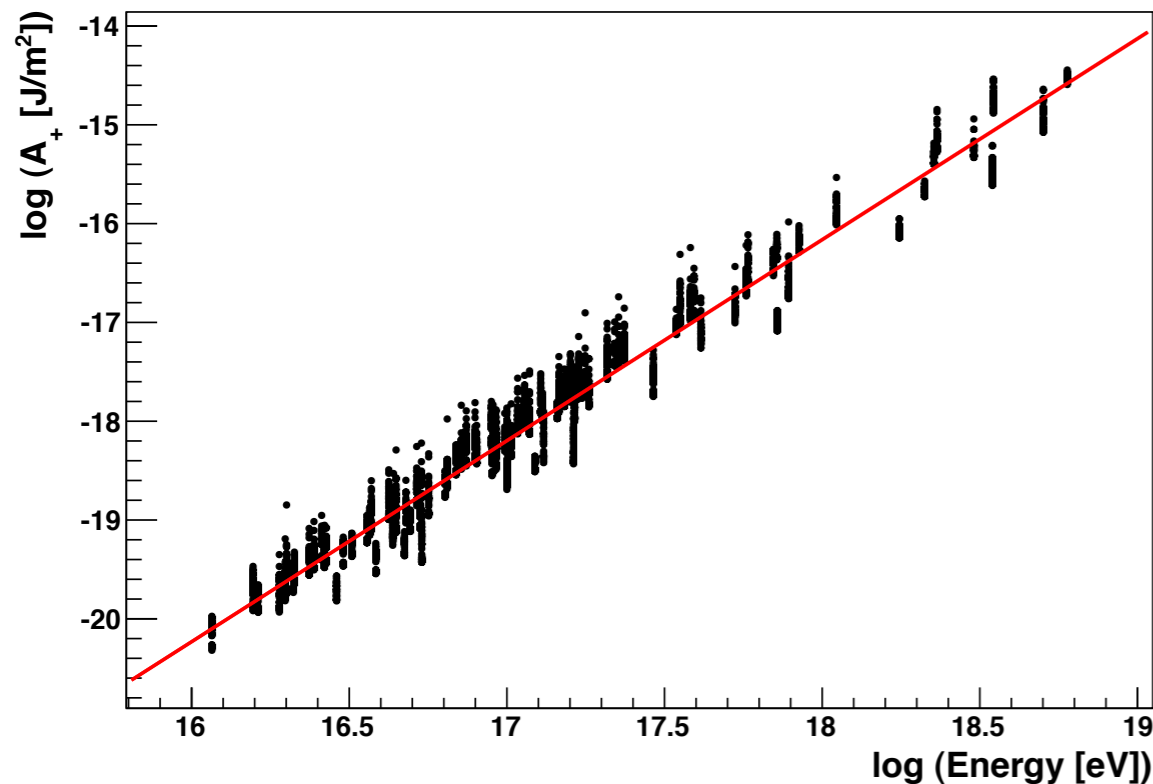
# Study of air shower characteristics



- Scaling parameter  $A_+$  is a function of the energy
- Slope in log-log depiction  $\sim 2.0$ , i.e. power depends quadratic on energy
- Remaining scatter is a function of angle with magnetic field,  $\sin(\alpha)$  and distance to the shower maximum



# Study of air shower characteristics

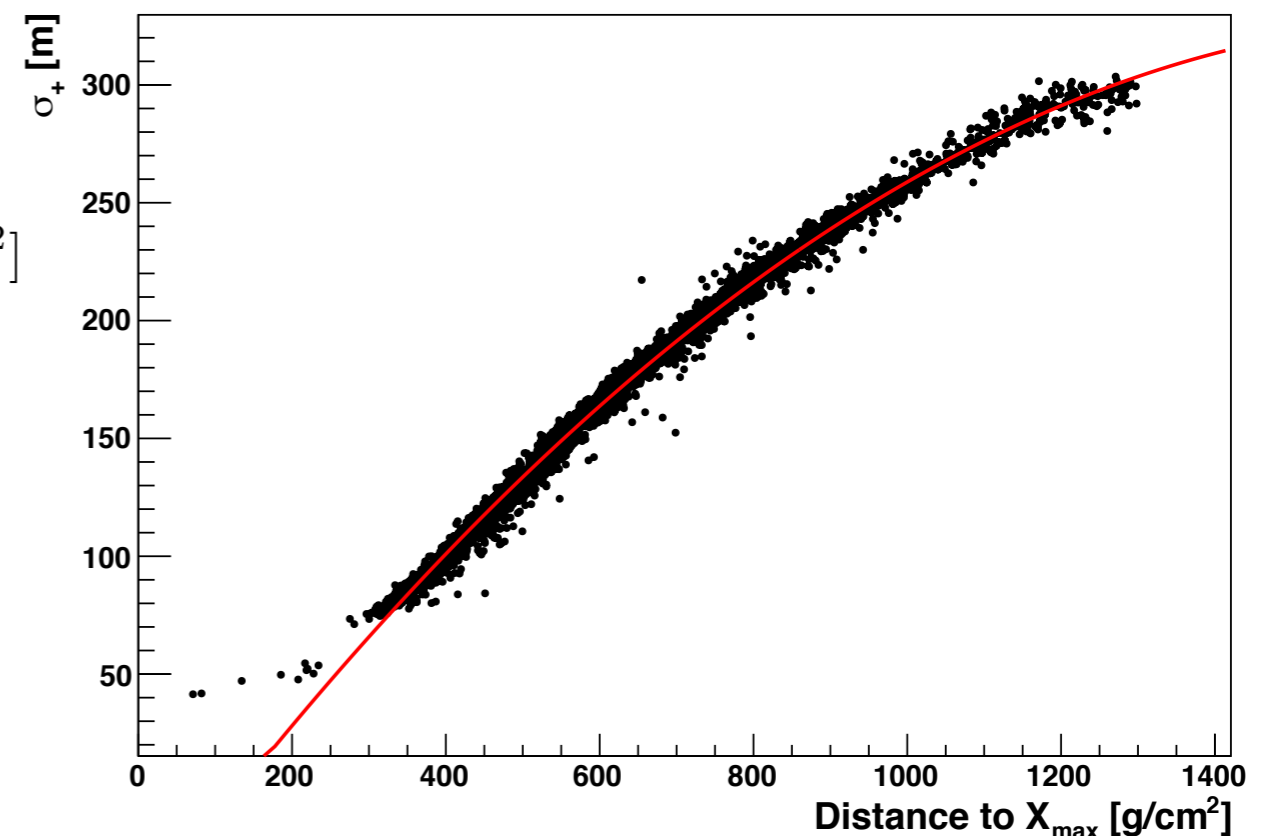


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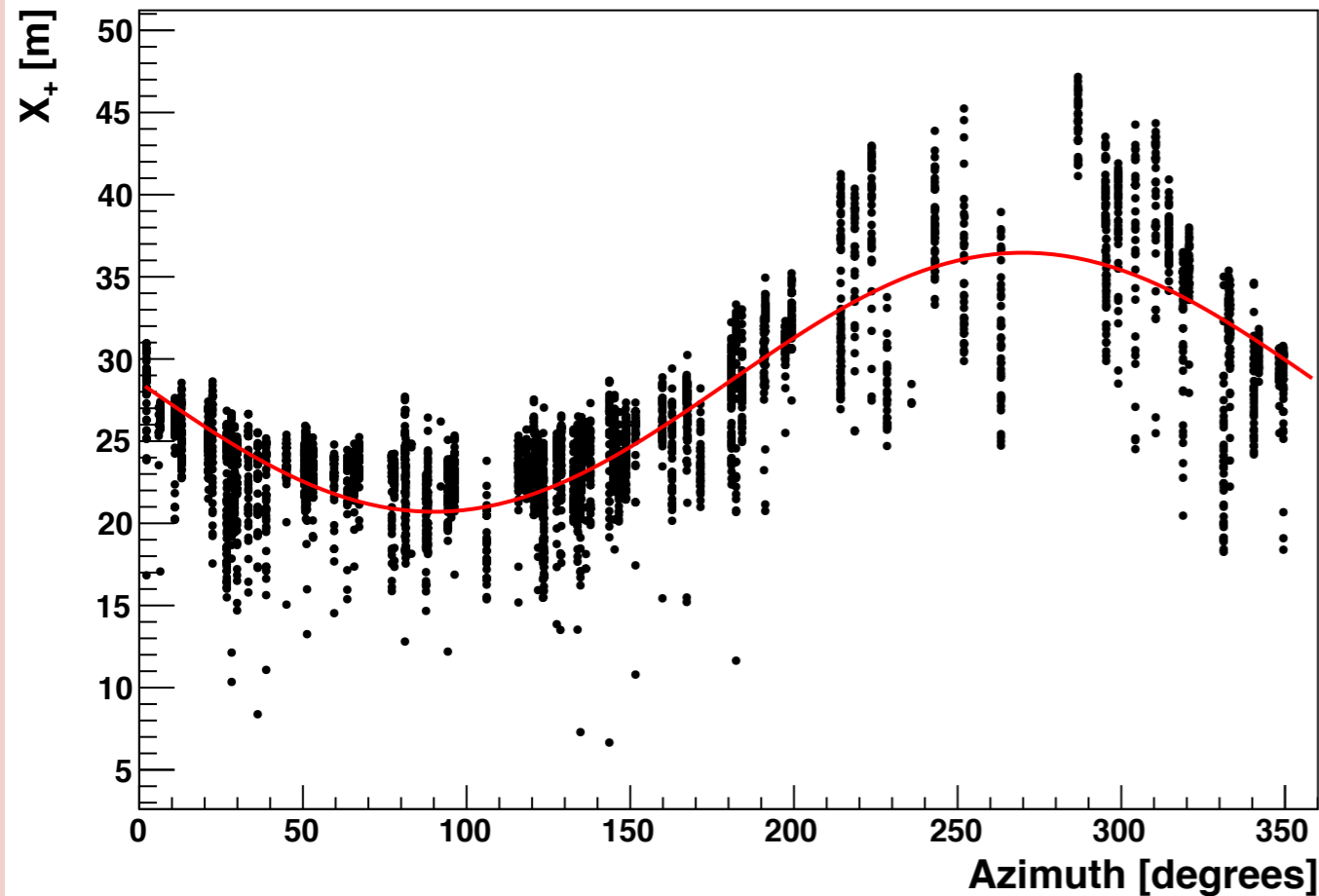
- Width parameter  $\sigma_+$  is a function of the distance to the shower maximum

$$D(X_{\max})[\text{g}/\text{cm}^2] = X_{\text{atm}}[\text{g}/\text{cm}^2] / \cos(\theta) - X_{\max}[\text{g}/\text{cm}^2]$$

- $\sigma_+$  scales similarly
- Remaining scatter is no function of primary air shower parameters



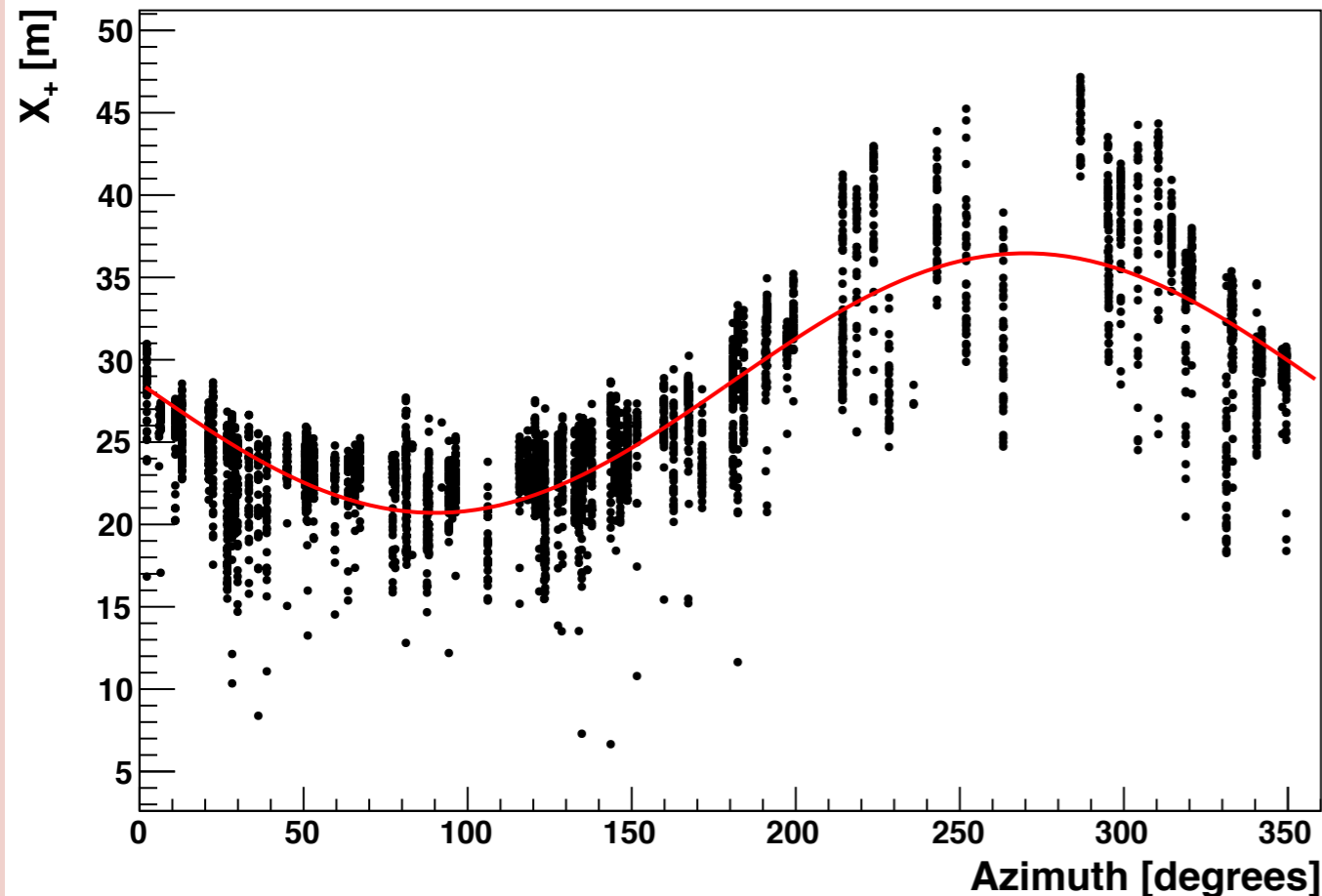
# Study of air shower characteristics



- **Shift parameter  $X_+$**  is a function of the azimuth angle of the shower
- Shift represents interplay between emission mechanisms
- Largest shift for azimuth parallel to magnetic field
- Small but not measurable **shift in  $Y_+$**  direction

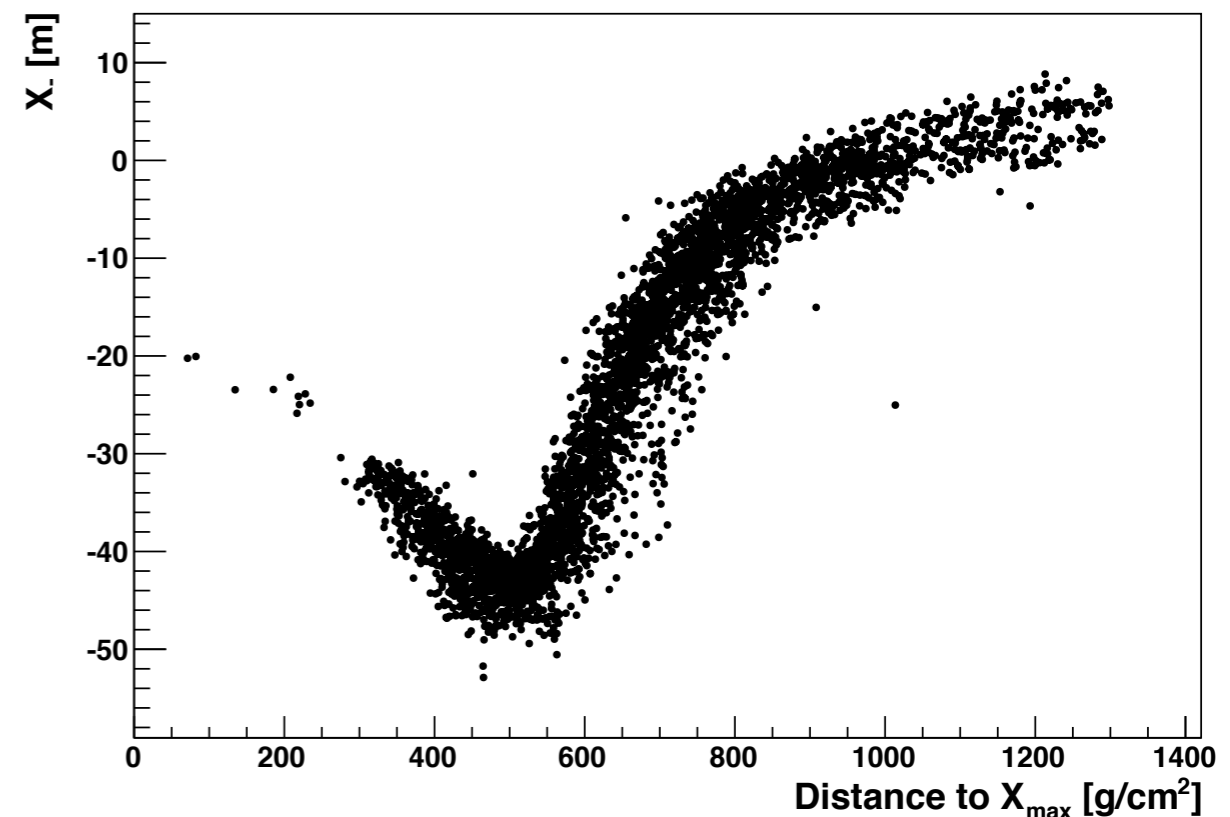


# Study of air shower characteristics



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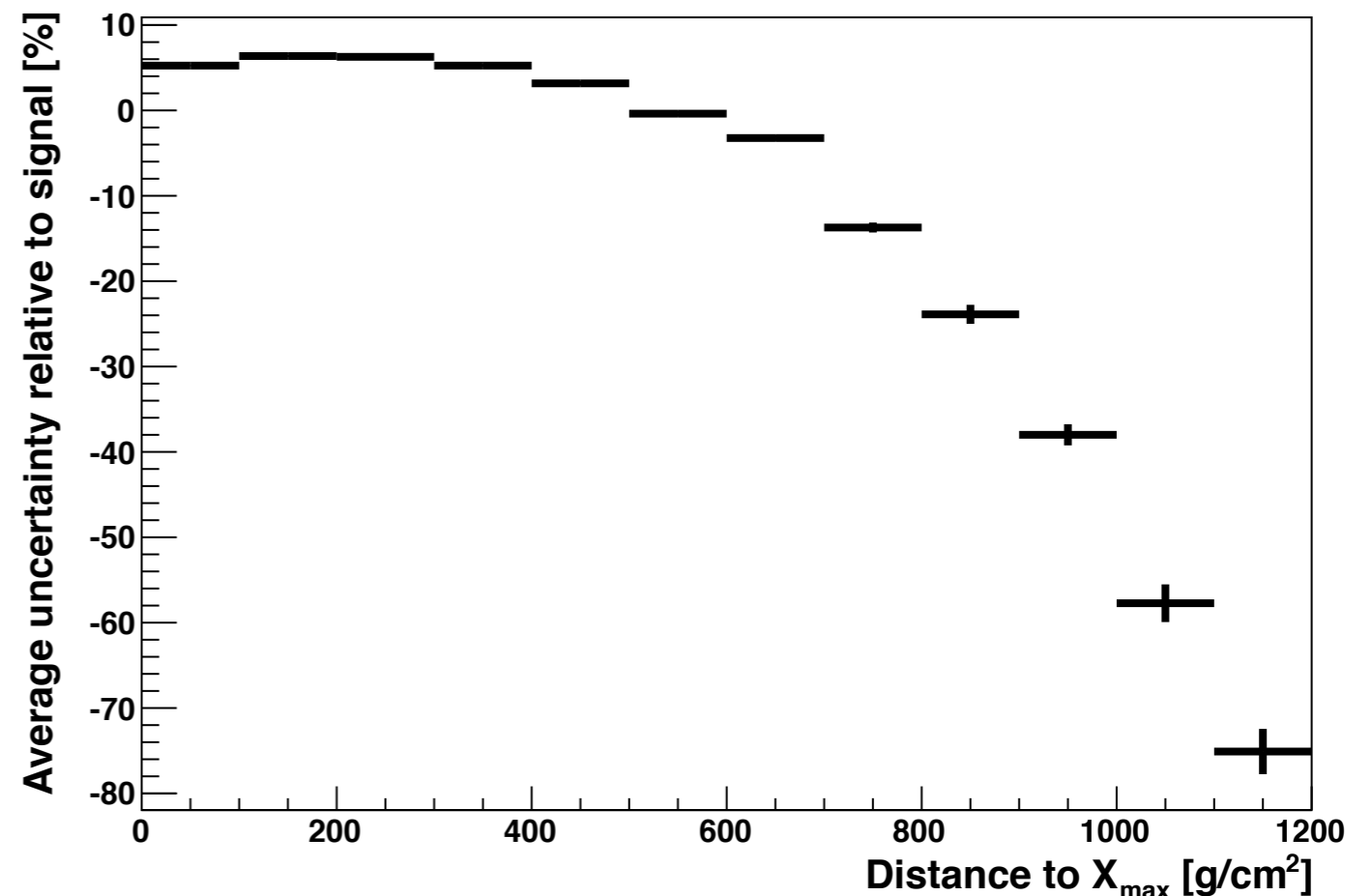
- **Offset of small Gaussian** to shower core ( $X_-$ ) is function of distance to the shower maximum
- Behavior changes direction and medium distances
- Shift needed for Cherenkov and charge excess, dominance of contribution



# Using function as prediction

$$P(x', y') = f_1(E) \cdot \exp\left(\frac{-f_2(\phi, X, Y, x', y')}{f_3(\theta, X_{\max})}\right) \\ - C_0 \cdot f_1(E) \cdot \exp\left(\frac{-f_4(\theta, X_{\max}, X, Y, x', y')}{f_5(\theta, X_{\max})}\right)$$

- Use functions of physical air shower parameters instead of fit parameters
- Very fast and simple prediction for studies of efficiencies, detection thresholds etc.

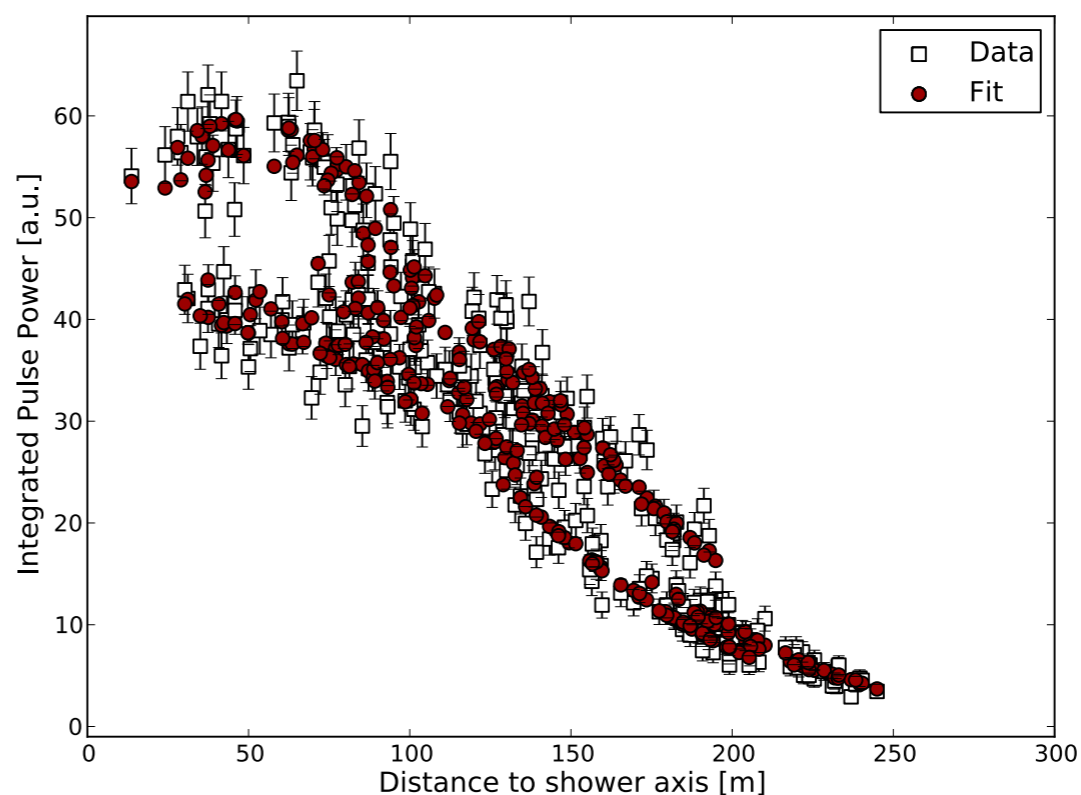
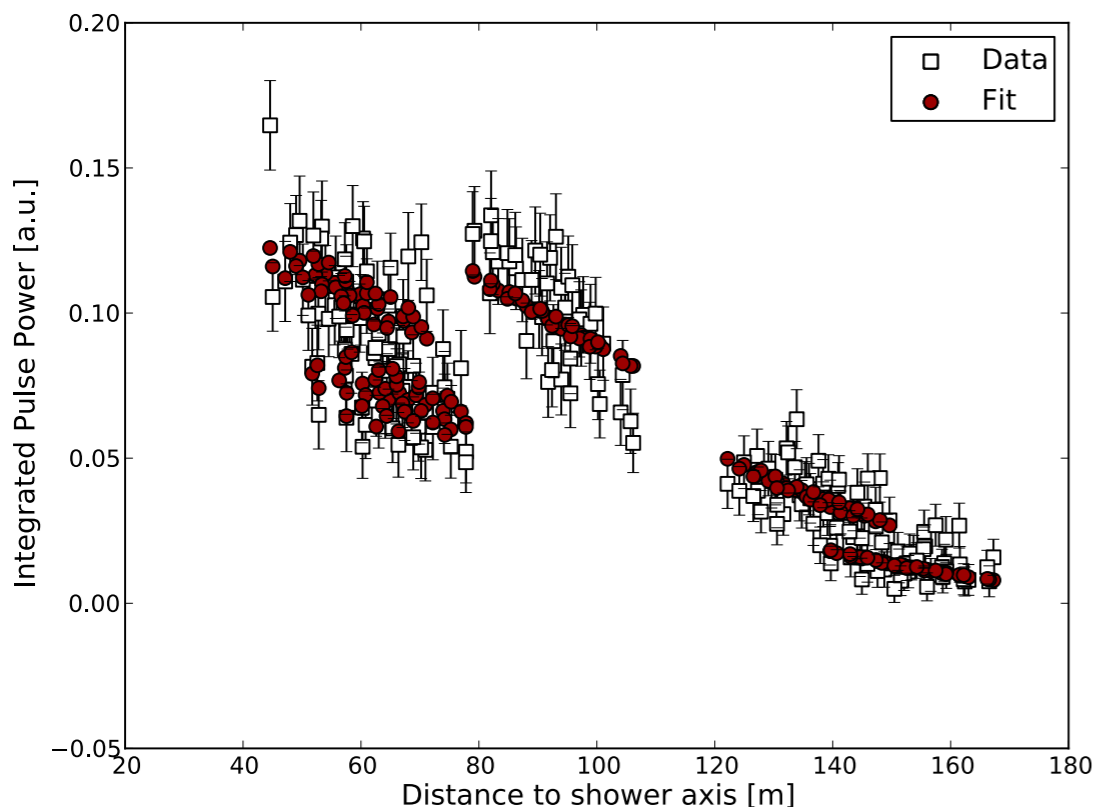
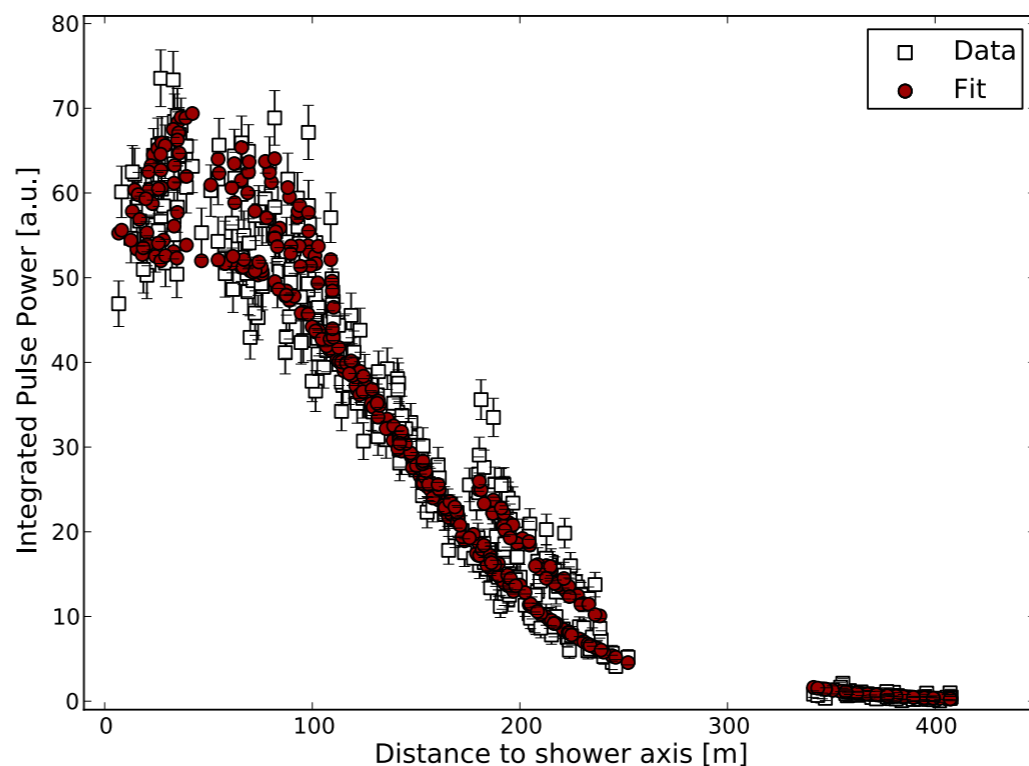
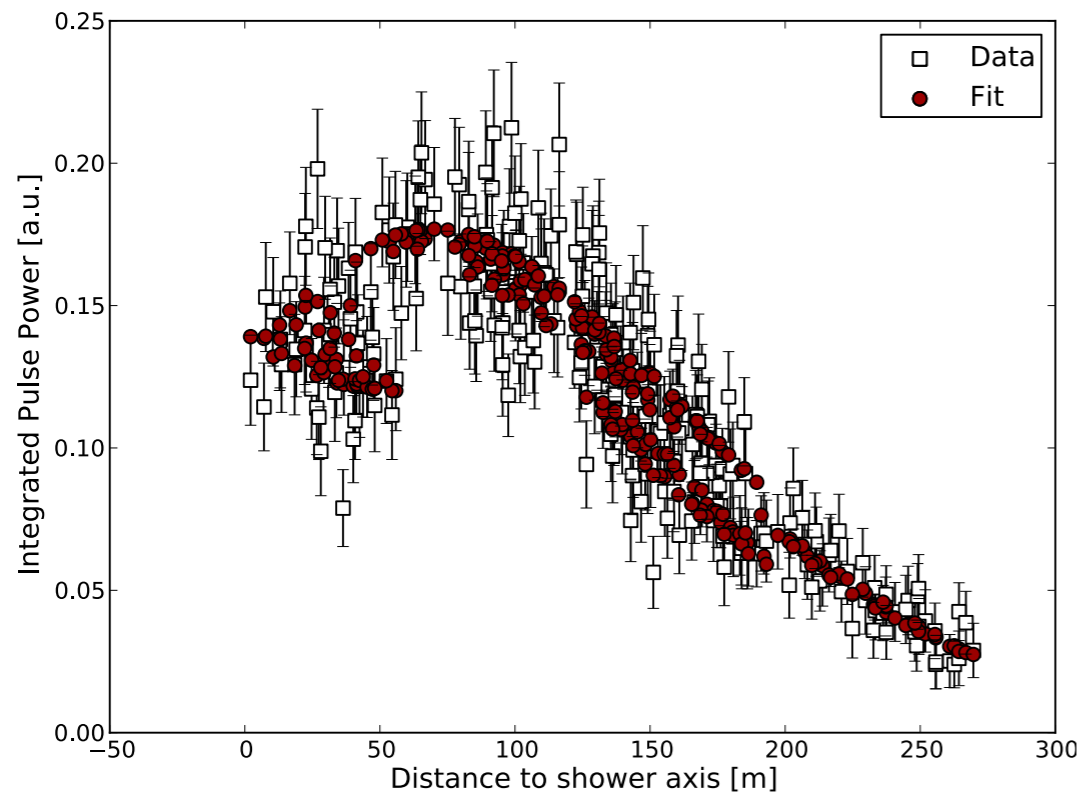


# Reduction of parametrization

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_c)^2 + (y' - Y_c)^2]}{\sigma_+^2}\right) - C_0 \cdot A_+ \cdot \exp\left(\frac{-[(x' - (X_c - C_3))^2 + (y' - Y_c)^2]}{(e^{C_1 + C_2 \cdot \sigma_+})^2}\right)$$

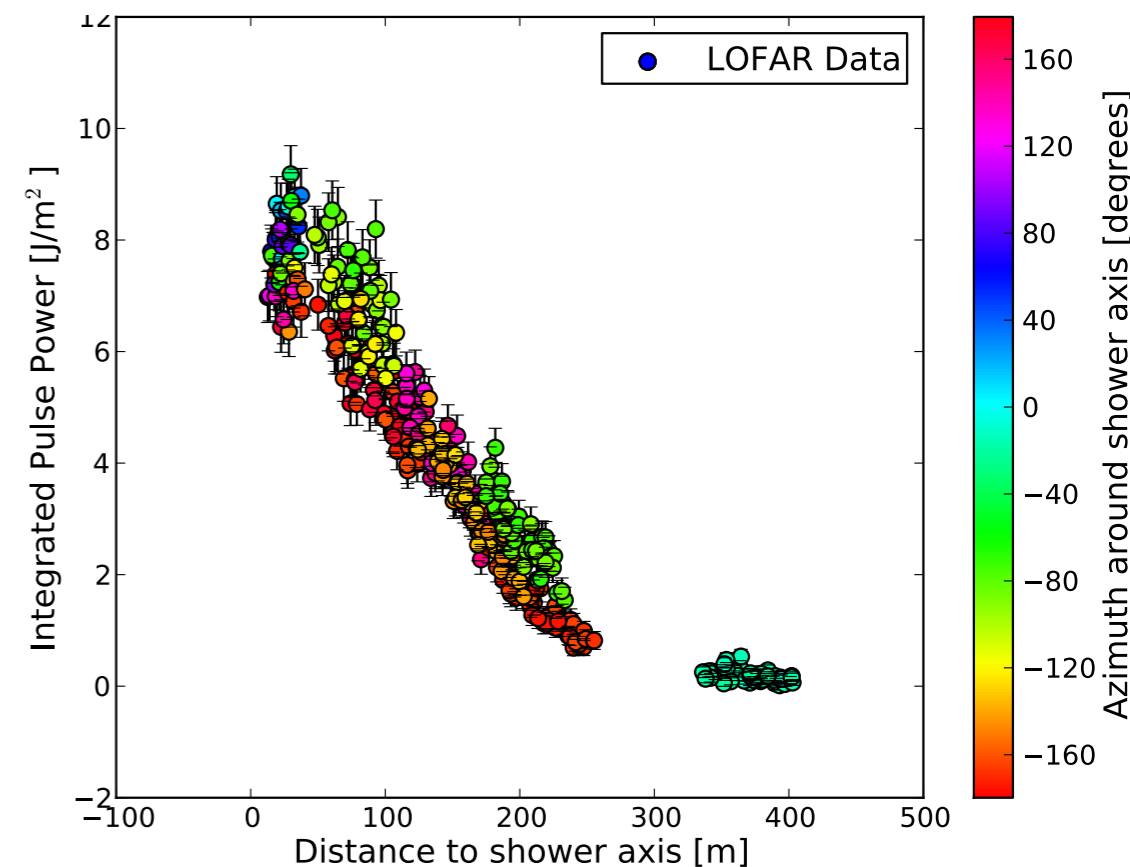
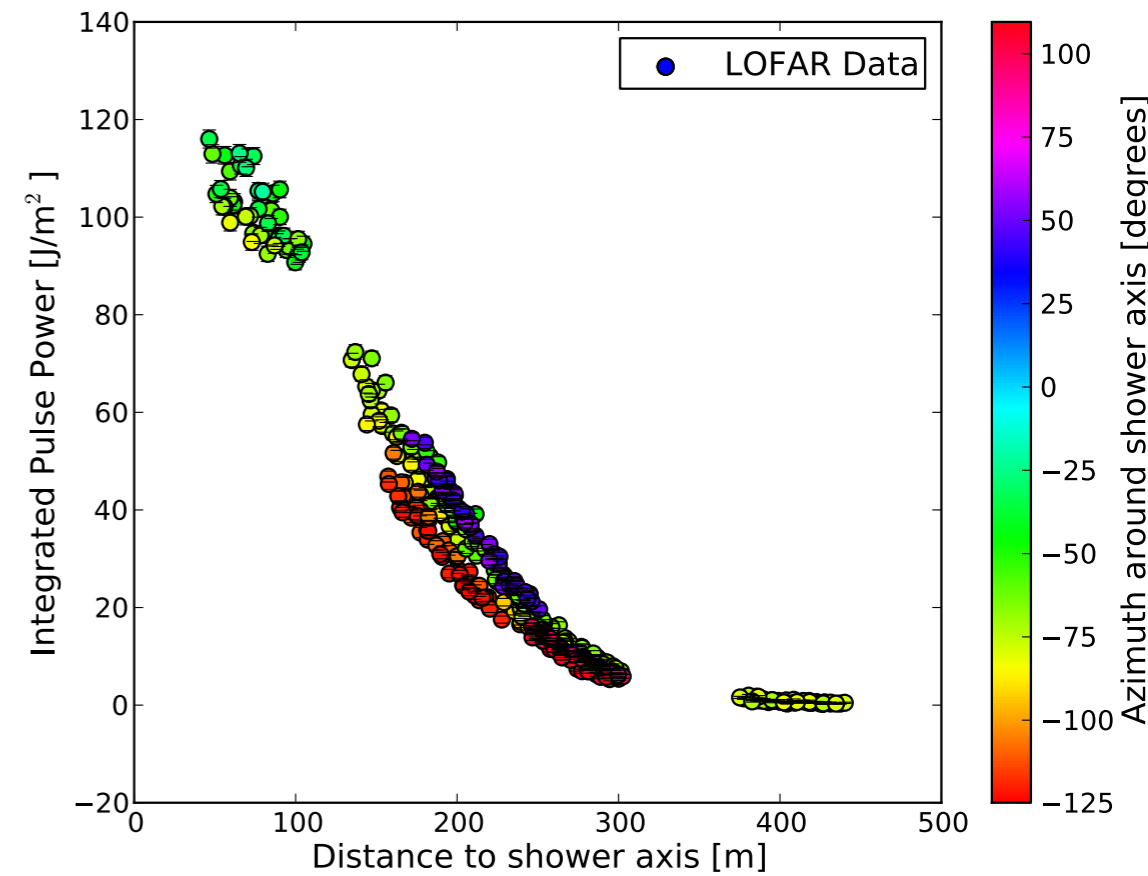
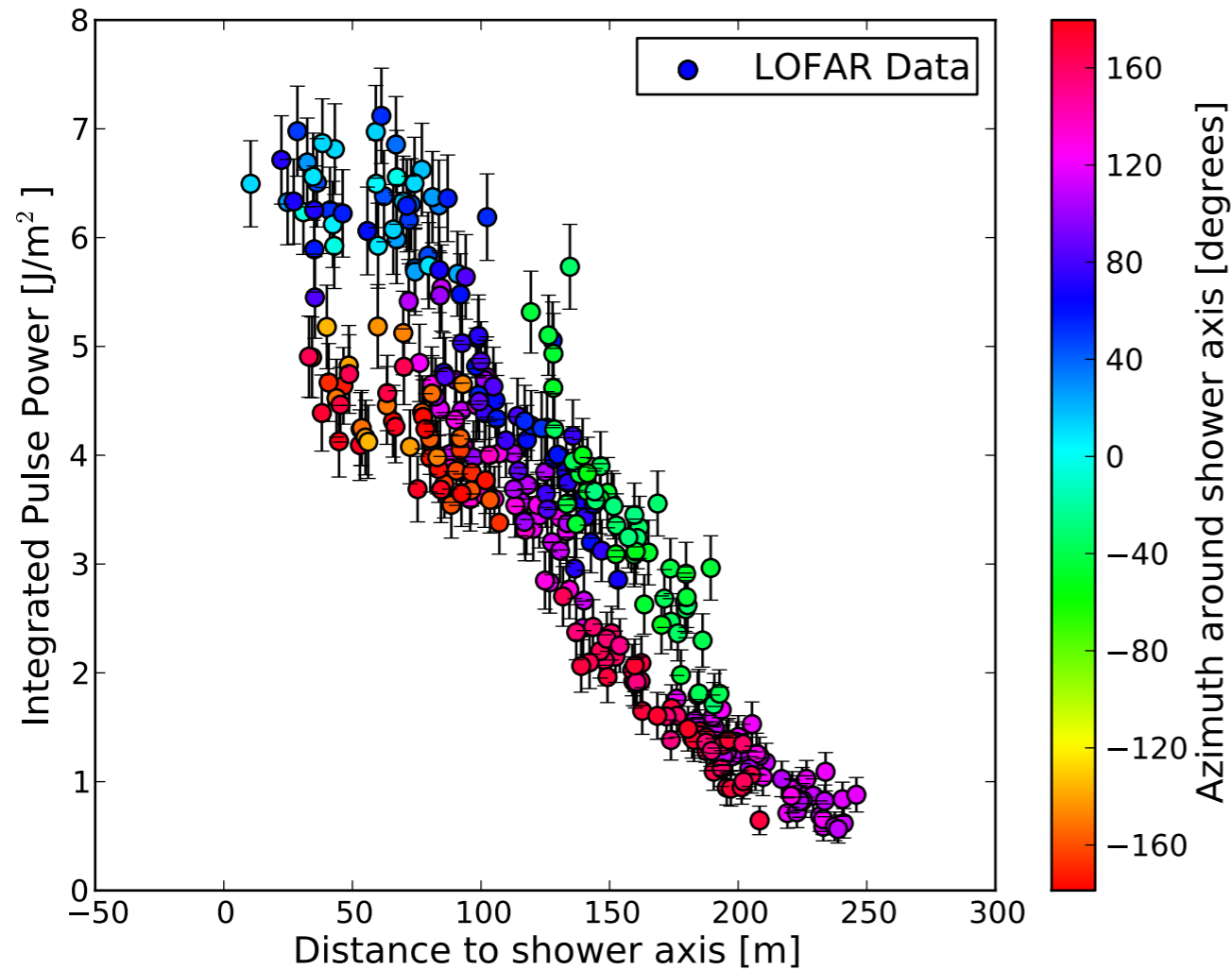
- Reduction in several ways possible, exploiting correlations between parameters
- Maximum reduction: 4 free parameters  $A_+$ ,  $\sigma_+$ ,  $X_c$  and  $Y_c$ 
  - $C_0$ ,  $C_1$  and  $C_2$  constant
  - $C_3$  binned for zenith angle
- At LOFAR:
  - $C_3$  free parameter as sufficient number of antennas
  - $C_0$  can vary in restricted range

# Test on Data



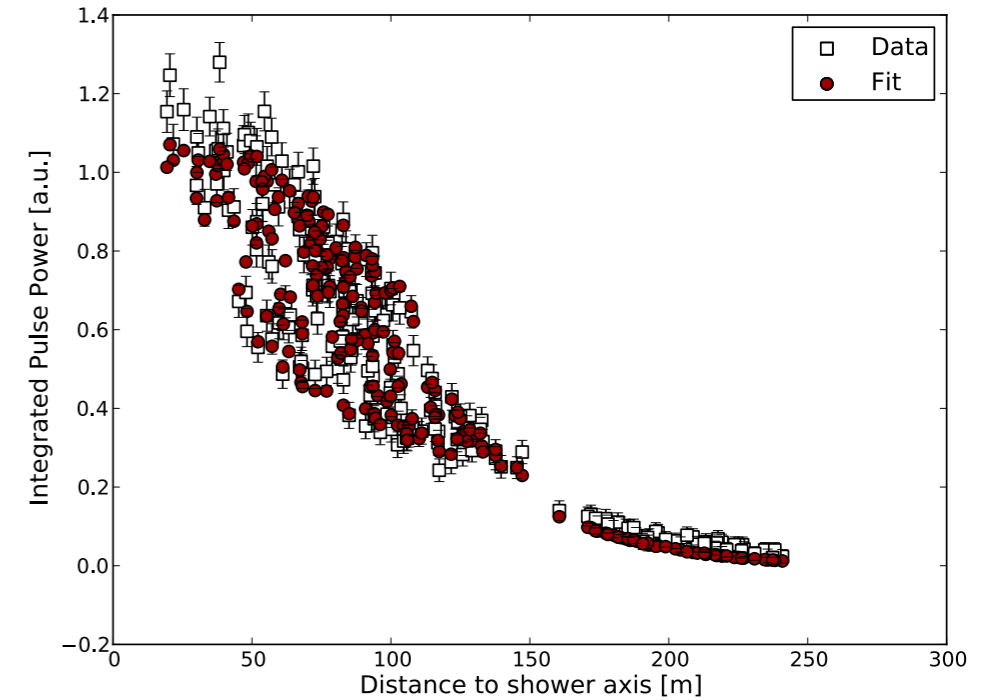
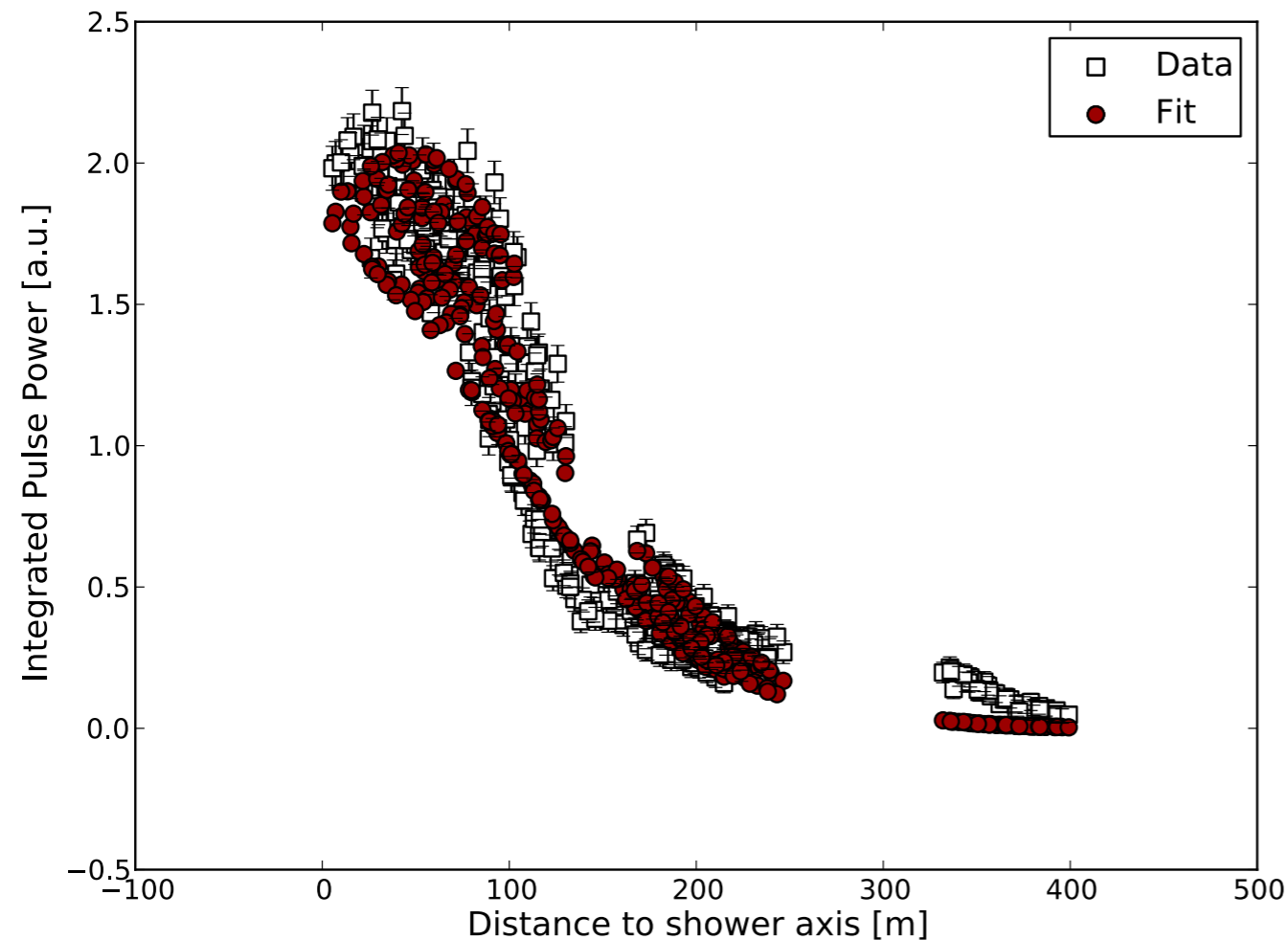
- All data from LOFAR can be fitted with this function
- If sufficient “structure” is measured
- $\chi^2/\text{ndof}$  distribution centered around 1, only some exceptions
- ...

# Azimuthal asymmetry in data



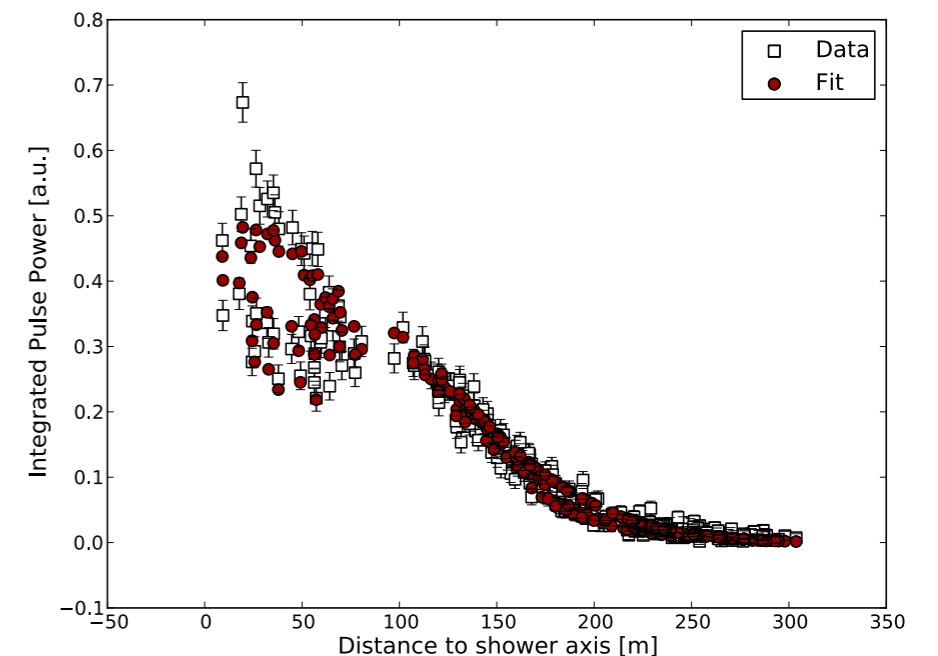
# Test on Data

## Thunderstorm



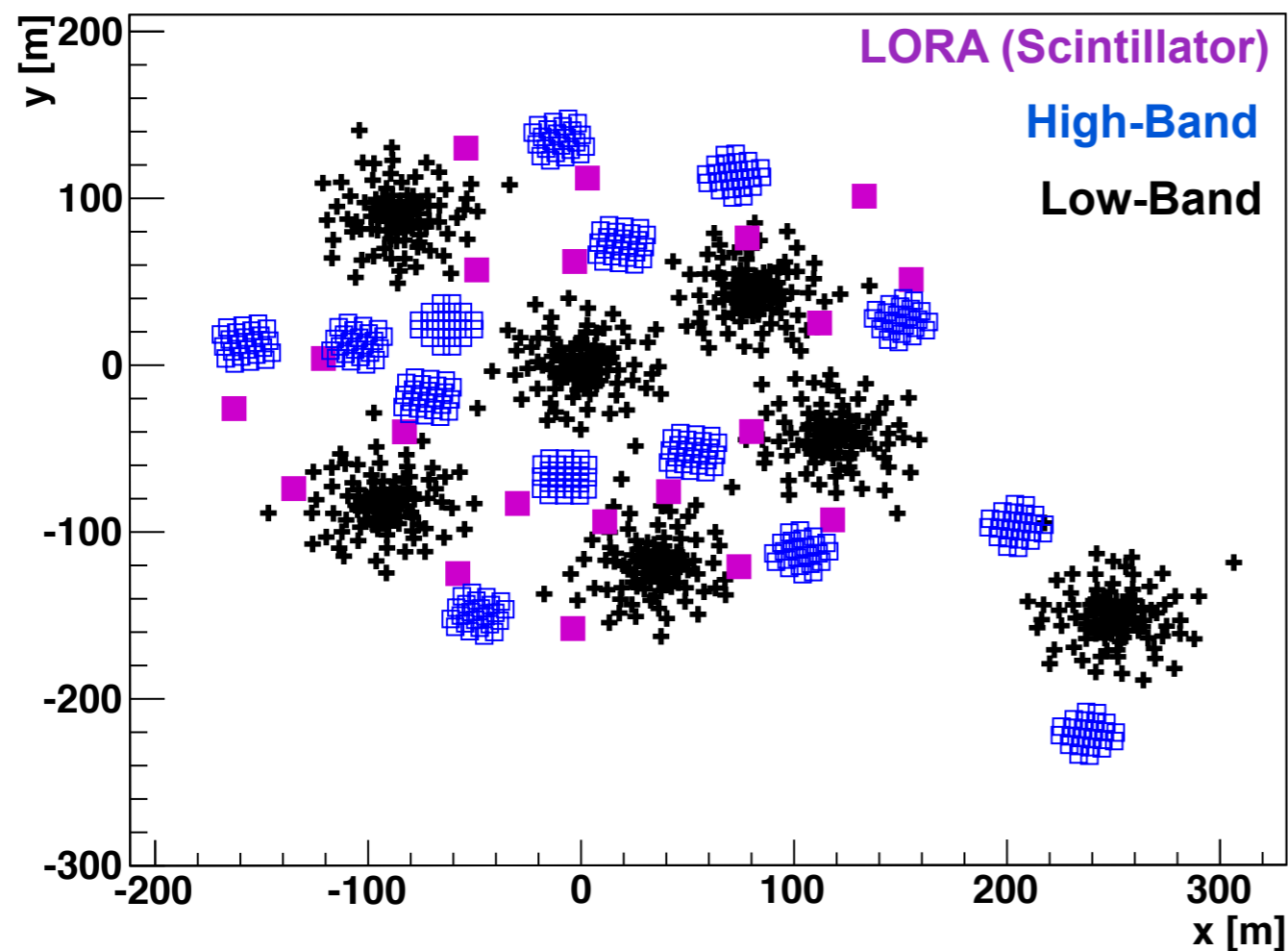
“Vertical events”  $< 15^\circ$

- (Most) events recorded during **strong electric field** cannot be fitted with good quality
- Vertical events show slightly less good fit



# Reconstruction of shower parameters

Testing the predictions of the simulations:



- Independent measurement from **particle detectors LORA**:
- instrumented area smaller than LOFAR area
- strong cuts on reconstruction limits sample (cuts on NKG parameters and distance of core position)
- test against full sample contains mis-reconstructions
- (Partly)-Independent reconstruction from **Full Monte Carlo method** (see Talk Stijn Buitink)
- Both methods based on CoREAS
- Both methods use radio only
- Only 50 air showers



# Conclusions

- We presented a “Double Gaussian” function with minimal 4 free parameters
- Scaling parameter directly sensitive to energy of air shower
- Width parameter directly sensitive to distance to the shower maximum
- All LOFAR data can be fitted with suggested function
- Correlations of parameters experimentally confirmed
- Next up: cross-check with experiment that has different way of measuring  $X_{\max}$  as for example AERA at Pierre Auger Observatory

