

**Neutrino reconstruction for PINGU** 

#### A. Kappes, R. Shanidze

#### for the IceCube collaboration & PINGU

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## IceCube/DeepCore/PINGU

		At	tmospher	ic-v ·	flux			
	Neutrino osc.[1,2]		Astrophysical-v sources					
	MSW effect		Point sources - good angular resolution			Diffuse f - good e	Diffuse flux - good energy resolution	
							log (E/GeV)	
( 1	) 1 GeV <b>PINGU</b>	2 DeepCore	1	3 TeV	4 IceCub	5 e	6 1 PeV	



- \* Cherenkov threshold  $E_{th} > [n/(n^2-1)^{\frac{1}{2}}] \times m$ for  $n_{ref} \sim 1.3$  (water/ice):  $E_e > \sim 0.8$  MeV,  $E_{\mu} > \sim 160$  MeV
  - [1] Antares collaboration, Phys.Lett. B714 (2012) 224 [1206.0645]
    [2] IceCube Collaboration, Phys. Rev. Lett. 111(2013), 081801 [1305.3909]



#### **PINGU** configurations



Closest distance between 2 strings: 20 m. (same for all strings in this configuration)

IceCube string #36 is indicated by a blue square. Red squares: 8 DeepCore strings Radius of the Circle, around IceCube String #36 is R=75 m.  $\rho V_{PINGU} \sim 5 M ton$ 



#### **Track reconstruction in IceCube**

- ν<sub>µ</sub>-CC and atmospheric-µ
   reconstruction in IceCube:
   "infinite track" approximation.
- PDF based reconstruction[1] includes best knowledge of calibrated detector and medium(ice) optical properties.
- The PDF is calculated with respect to "seed track" wich is obtained with first guess pattern recognition[2].



- [1] AMANDA Collaboration (J. Ahrens et al.), NIM. A524 (2004) 169 Muon track reconstruction and data selection techniques in AMANDA
- [2] IceCube Collaboration (M.G. Aartsen et al), arXiv:1308.5501 Improvement in fast particle track reconstruction with robust statistics



- Level 2 (L2) reconstruction: Standard IceCube/DeepCore reconstruction of the triggered/filtered events:
  - Hit cleaning algorithms
  - Various IceCube algorithms for track, vertex and energy reconstruction: SPE, MPE, FiniteReco, . . ., Monopod
- L2 reconstruction for DeepCore events where modified and updated for the PINGU MC data.
- Level2 reconstruction for PINGU includes additional algorithms:
  - SANTA [DESY]
  - IgelFit [Bonn U.]
  - HybridReco [PSU]



# SANTA



 Used in current analysis of PINGU sensitivity [talk by A.Gross, MANTS-2013, 14/10/2013]



## SANTA

 Energy and angular resolution obtained for PINGU MC events reconstructed with SANTA (direction) amd Monopod(energy)



A.Gross, presentations at ICRC-2013 (Rio de Janeiro) and VLVnT-2013 (Stokholm).



# **IgelFit**

- Divides sphere around a vertex in N-segments/tracks (N=128 by default).
- In each segment IceCube PDF-based reconstruction is applied.
- Requires vertex coordinates as a "seed"

Currently studies: M. Day (UW) - for DeepCore data







- muon neutrino-CC event reconstruction with 8-parameters:
  - neutrino interaction vertex and time ( $x_0, y_0, z_0, t_0$ )
  - $\mu(v)$  direction:  $\cos\Theta$ ,  $\phi$
  - $\mu$ -energy (defined from length  $I_{\mu}$ ) and cascade energy (E<sub>x</sub>)
- Requires PDF / depends on the knowledge of ice properties.
  - Current PINGU MC simulations use SPICEMie.
  - PDFs are calculated from SPICE 1.
- Relatively slow, different minimization and LH-scan algorithms tested:
  - MINUIT algorithms
  - Markov Chain(MC): Vanilla Metropolis-Hastings MCalgorithm
  - MultiNest algorithm[1]
- [11 F. Feroz et al, arXiv< 0809,3437



#### **MultiNest algorithm**

#### João Pedro Athayde Marcondes de André @IceCube meeting, Munchen 2013



- Zenith angle resolution of less than 10° for most events
- Inner Gaussian contains about 57% of the events

Energy resolution of neutrino events reconstructed with HybridReco/Multinest (bottom left) and Monopod (bottom right)



#### Atmospheric v-flux and event selection

#### Atmospheric neutrino flux ("neutrino beam"):

#### M. Honda et al. Phys. Rev. D 83, 123001 (2011) [arXiv:1102.2688]

"Improvement of low energy atmospheric neutrino flux calculation using the JAM nuclear interaction model" HKKM fluxes: http://www.icrr.u-tokyo.ac.jp/~mhonda/nflx2011/index.html 6 sites: Frejus, Gran Sasso. INO, Kamioka, Soudan, Sudbury



- ν<sub>µ</sub>-CC events ("track like") at low energy (E < ~20 GeV) will be contaminated with v-NC and electron and tau neutrino CC ("cascade") like events. PID-methods could help to separate these events.</li>
- Method based on the "superluminal" hits for the separation of track and cascade type neutrino events studied in PSU. (talk by Ty. DeYoung)
- Another method is based on the variables from collider experiments for PID in the DeepCore/PINGU.
   Currently under development in NBI by M.Jørgensen.



## **Current PINGU MC data**

- All PINGU MC data sets produced with GENIE.-GEANT-CLSIM\* (/data/sim/PINGU/2012/triggered/GENIE-in-ice) are copied and stored at DESY. (\* Ken Clark, GPU/WestGrid system)
- Largest statistics of simulated data: v15 (40 str., 60 PDOM) neutrino energy range: 1-80 GeV, E<sup>-2</sup>

	Ve	$ u_{\mu}$	$v_{\tau}$
Simulated(x10 <sup>8</sup> )	3.75	8.0	3.75
Triggered (x10 <sup>6</sup> )	0.428	1.250	0.102



 All triggered events from PINGU v15 are reconstructed with "standard" reconstruction ("PINGU L2" including SANTA, Monopod)

Simulation volume for PINGU events



## **Summary and outlook**

- Level2 reconstruction scripts for PINGU MC events based on standard IceCube/DeepCore algorithms and few new methods were set and tested. The PINGU reconstruction is working at different sites in USA/Canada and Europe.
- Current algorithms used in the reconstruction of v-direction (SANTA) and  $E_v$  (Monopod) provide sufficient angular and energy resolution for the NHM determination .
- New promising methods for the neutrino direction and energy reconstruction with improved resolutions were presented at the IceCube meeting last week.
- Work is in progress to adapt IceCube tables and algorithms developed for high energy neutrinos evens to lower energies..



# The IceCube Collaboration & PINGU

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