### Neutrino Oscillations with DeepCore and PINGU





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Cross-section

 Original IceCube design focused on neutrinos with energies above a few hundred GeV

IceCube Lab

IceTop 81 Stations, each with DeepCore provides 50 m i 2 IceTop Cherenkov detector tanks 2 optical sensors per tank 324 optical sensors reduced volume with IceCube Array 86 strings including 8 DeepCore strings lower energy threshold 60 optical sensors on each string 5160 optical sensors December, 2010: Project completed, 86 strings 1450 m Amanda II Array (precurser to IceCube) DeepCore 8 strings-spacing optimized for lower energies 360 optical sensors Eiffel Tower 324 m 2450 m 2820 m Bedrock

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- O(10<sup>5</sup>) atmospheric neutrino triggers per year



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  - Eight special strings plus 12 nearest standard strings
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#### Oscillations with Atmospheric Neutrinos

- Neutrinos oscillating over one Earth diameter have a  $v_{\mu}$  survival minimum at ~25 GeV
  - Corresponding maximum in  $v_{\tau}$  appearance probability
- Neutrinos from all terrestrial baselines are available for free
  - Compare observations from different baselines and energies to mitigate impact of systematics
- Hierarchy-dependent matter effects below ~10-20 GeV



#### May 13, 2013

### Muon Disappearance

- As a first step, compare zenith-dependent response of standard IceCube muon analysis (high energy) to a modified version for DeepCore
  - Look for oscillation signature in event rate suppression at low energies
  - Detector systematics reduced by comparing HE and LE rates
  - Based on traditional muon analysis, no new techniques designed for DeepCore – lower efficiency accepted



## Muon Neutrino Disappearance



Statistically significant angle-dependent suppression at low energy, high energy sample provides constraint on uncertainties in simultaneous fit

 Shaded bands show range of uncorrelated systematic uncertainties; hatched regions show overall normalization uncertainty

## Muon Neutrino Disappearance

- Oscillation parameter allowed regions extracted from zenith distributions
  - Systematics included
- Excellent agreement with world average measurements (with large uncertainties)
  - Potential for significant improvement with inclusion of energy estimators, more advanced reconstructions and event selections



# Ongoing Improvements

- Parallel analysis of first year of data from DeepCore
  - Introduce specialized data analysis and background rejection techniques for DeepCore
  - Low energy event yield improved by almost an order of magnitude



- Also including an energy estimator based on track length of contained neutrino-induced muons, 2 more DeepCore strings
  - Potentially substantial improvements in precision, depending on impact of systematics

### Future Directions

- Preliminary estimates of sensitivity suggest competitive measurements of oscillation parameters will be possible soon
  - Final precision will depend on improvements in energy and angular resolution, understanding of systematics – progress ongoing!
- Also studying possibility of extending low energy reach of IceCube with an even denser infill array – PINGU
  - Possibility of exploiting neutrino/anti-neutrino asymmetries and matter oscillation effects to measure neutrino mass hierarchy, given the large value of  $\theta_{13}$
  - Studies of feasibility and performance requirements now underway

### PINGU

#### One of several candidate geometries under investigation

- Exploring requirements for mass hierarchy measurement additional strings may be added if better angular and energy resolution is needed
   Cross-section
  - Systematics can be addressed with additional in situ calibration devices



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#### Signature of the Mass Hierarchy

- Idealized case with no background, perfect flavor ID, 100% signal efficiency
- Different assumed resolutions smear the signature but do not eliminate it
  - NB: angular resolution is for muon – kinematic effects *are* included
  - Expected efficiencies and resolutions under investigation now



## Sensitivity vs. Performance

- Numerically evaluate confidence of hierarchy determination after 1 year as a function of assumed energy and muon angular resolution
  - For now, require 20 DOMs hit in PINGU as a proxy for analysis efficiency
  - Need to fold in systematics and physics degeneracies (e.g. Δm<sub>31</sub><sup>2</sup>)
  - Details of analysis technique still being tuned for power, robustness
  - Sensitivity to maximality, octant, etc. under study



### Advantages of PINGU

- Well-established detector and construction technology
- Relatively low cost: ~\$10M design/startup plus ~\$1.25M per string (depending on number of sensors, cost of fuel, etc.)
- Rapid schedule: deployment could be complete by 2017-18, depending on final scope
  - Quick accumulation of statistics once complete
- Provides a platform for more detailed calibration systems to reduce detector systematics
  - Enhance physics at PINGU energies e.g. hierarchy,  $v_{\tau}$  appearance
  - Opportunity for R&D toward other future ice/water Cherenkov detectors
- Working toward a Letter of Intent now