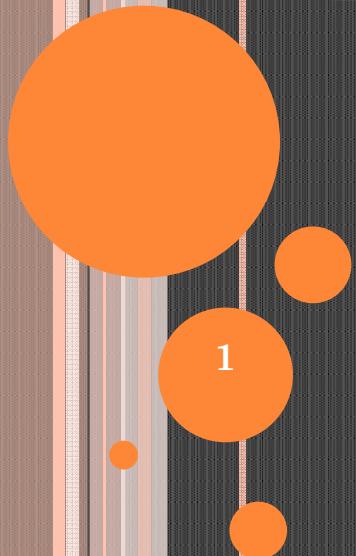


LIDAR USE FOR IACT (CTA/HESS)

G.Vasileiadis LPTA/CNRS/Un.Montpellier II



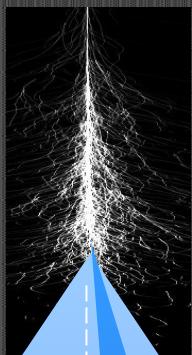
- Imaging Atmospheric Cherenkov Technique?!
- Use of Lidars for IACT
- LPTA HESS Lidar
- Towards a more adapted solution

Gamma-ray

Detection of TeV gamma rays
using Cherenkov telescopes

Particle shower

Cherenkov light

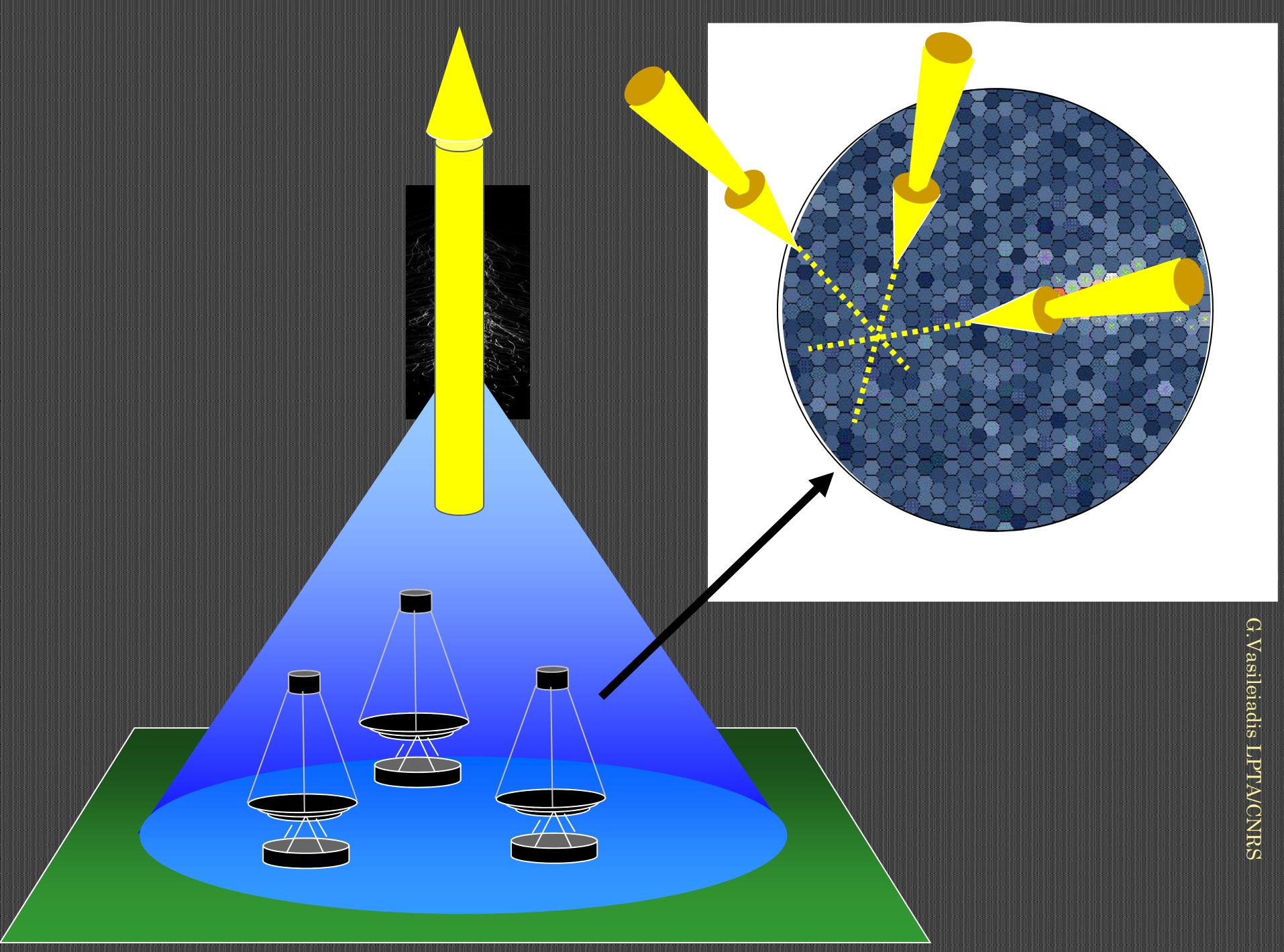


$\sim 10 \text{ km}$

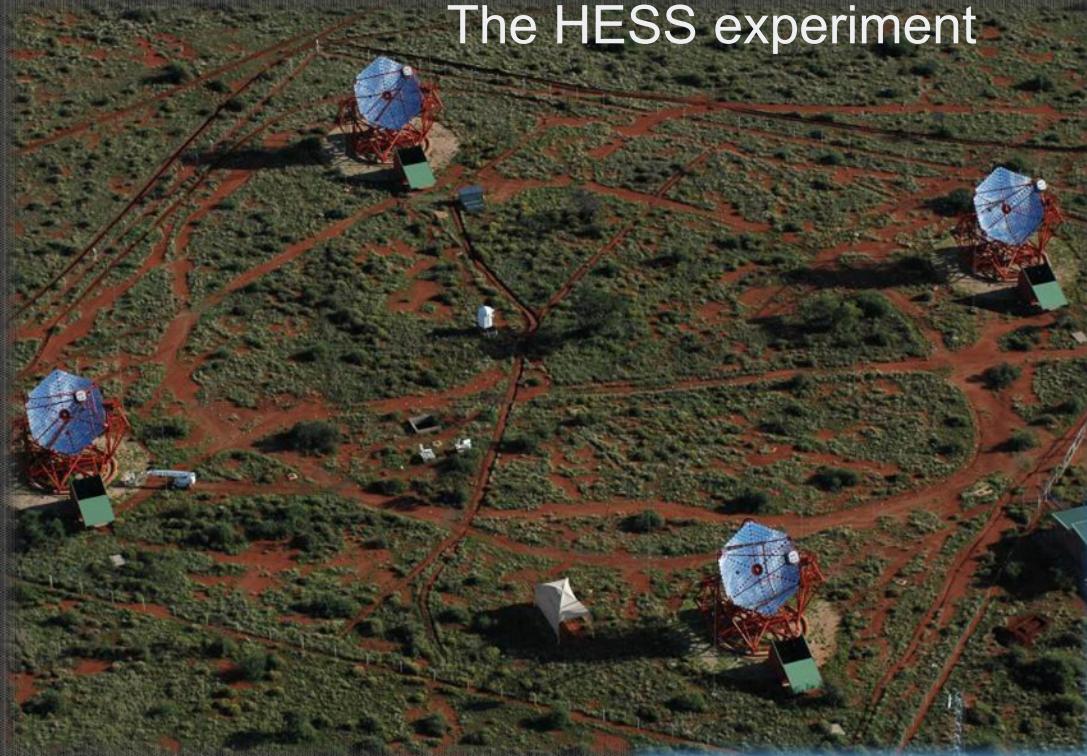
1°

$\sim 120 \text{ m}$

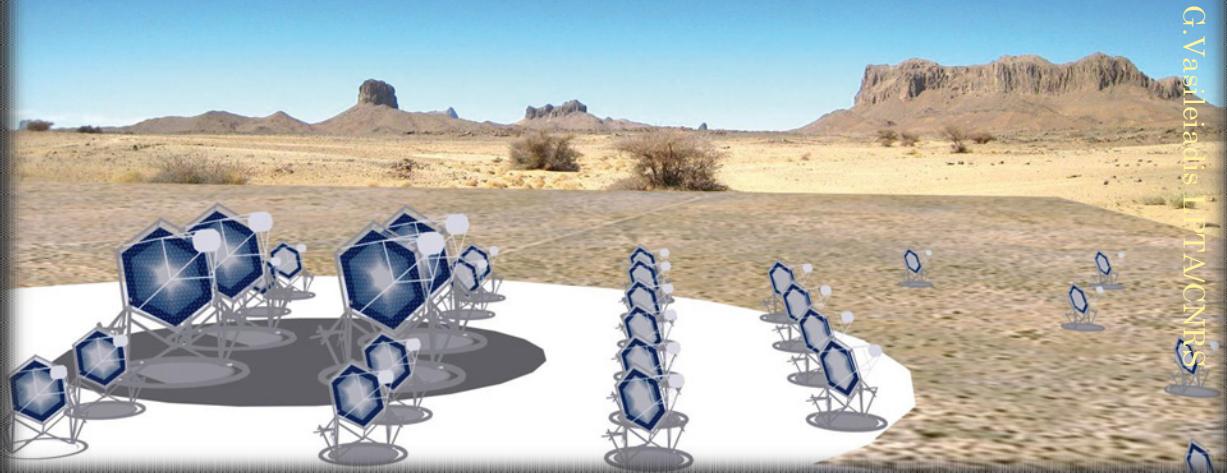




The HESS experiment



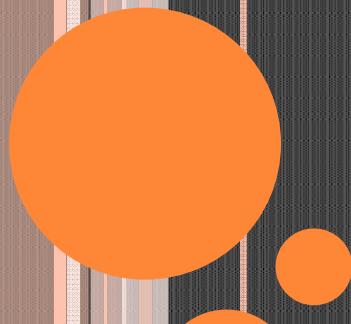
The CTA Project



WHY LIDARS AND FOR WHAT USE !

- Cherenkov showers develops in the atmosphere
- Photon absorption critical
 - Aerosols
 - Clouds
 - Dust
 - ...
- Atmospheric profile need (function of absorption coeff α) to be known.
 - Input to MC studies
 - Effective Area calculations
 - Energy Spectrum & Detected Source Photon Yield
 - *see also S.Nolan's & C.Rulten talks*

So from a LIDAR point of view we need this α



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FIRST APPROACH

Elastic Type LIDAR

(Leosphere,LPTA -Lidar at the HESS site)

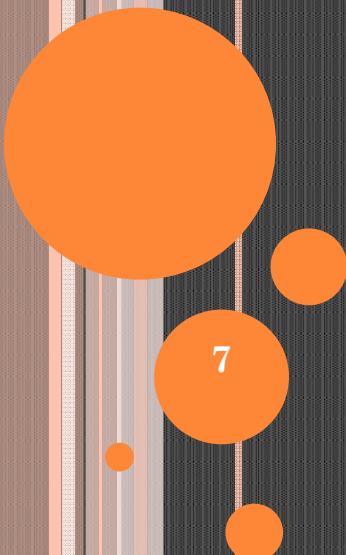
• Pros

- Double WL (532&355nm)
 - Covers better detected Cherenkov photon spectrum
- Easier to construct
- Comfortable Photon Yield
- Easier to Operate.

• and Cons

- Parameters (α, β) more difficult to extract
- Big systematic error ($\sim 15\text{-}20\%$)
- No information available after Clouds & Aerosols

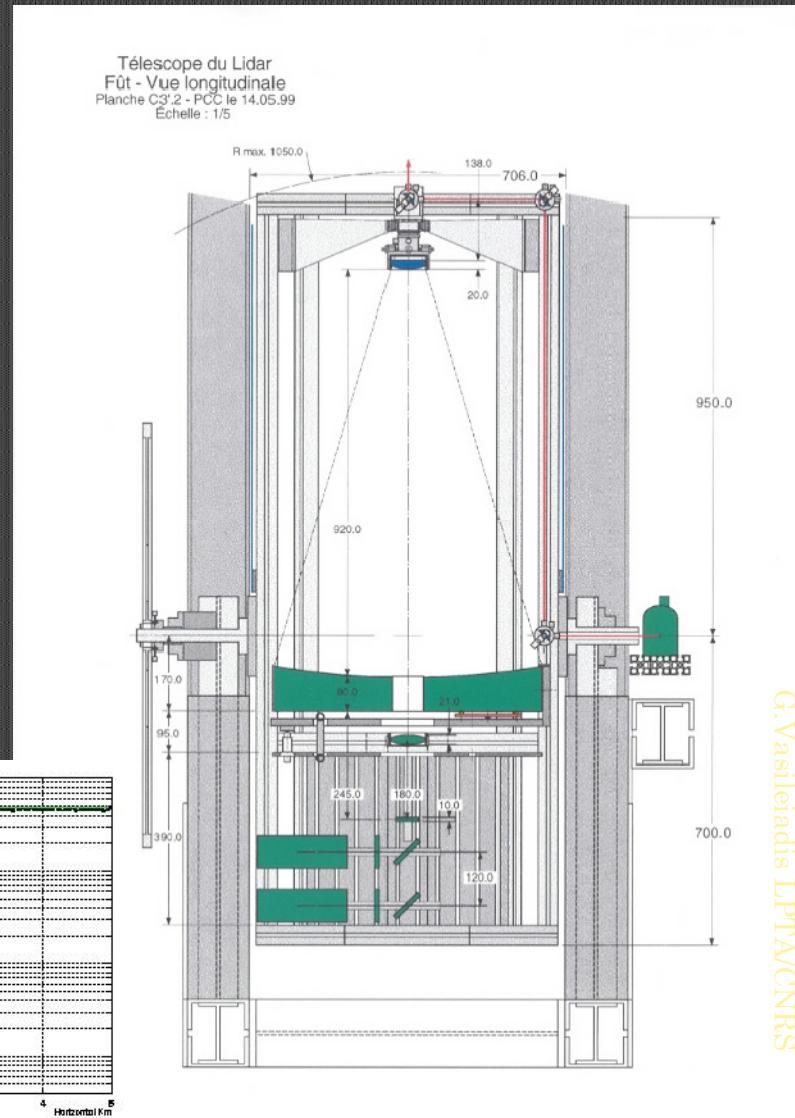
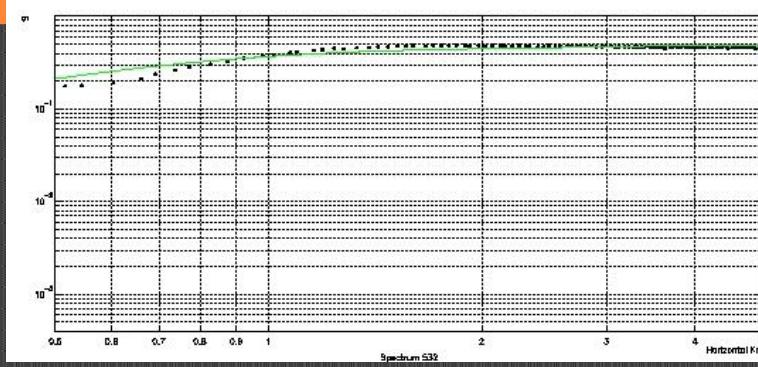
$$P(R) = \frac{E_0 \eta_L}{R^2} O(R) \beta(R) \exp \left[-2 \int_0^R \alpha(r) dr \right]$$



LPTA LIDAR AT HESS-I

In a nutshell :

- Double WL (355,532 nm)
- Overlap at 800m-1km
- CG Type Telescope
 - 60cm Ø, f=102cm
- Quantel Brilliant 20 Nd:YAG
 - 10Hz
 - 180mJ (65mJ)
 - 2.5m resolution
 - Octopus CAGE 8265
- PMTs Readout (XP2012B)
- Linux/LabView SW



LPTA LIDAR AT HESS-II

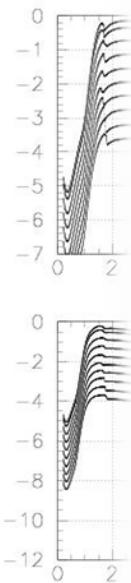
In a nutshell :

- Double WL
- CG Type Telescope
- Useful Range 1-20 km
- Completely Automated
- Remote Operated from FR
- Operational (..or so..) for the last 2y on site.
- Simultaneous LIDAR&Physics data available.
- Data available via Web (Hess)
- First analysis shows limitations
- Mechanical and other issues
 - Namibian weather conditions
 - Not really conceived for harsh enviroment
 - Still one of the few currently working Cherenkov Lidars.



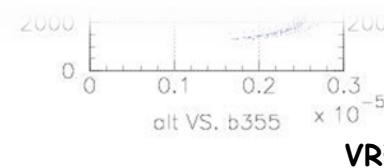
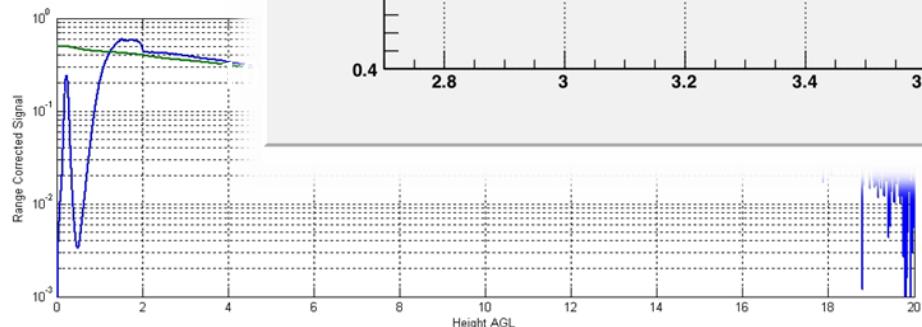
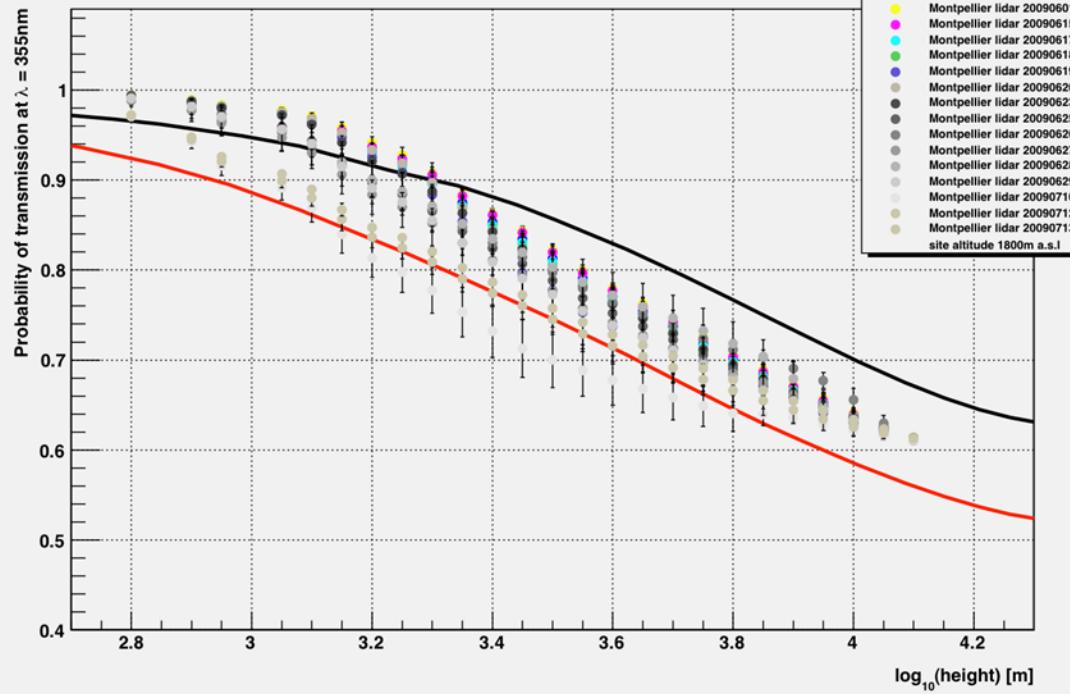
ACTUAL DATA ANALYSIS SAMPLE

Raw Data

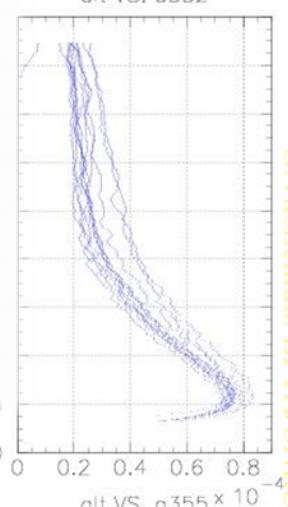
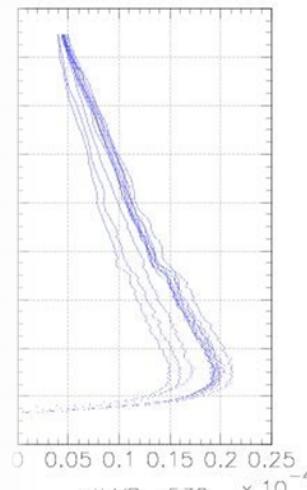


C.Rulten

Comparing lidar measurement to models



Simulated data (provided by P.Ristori) in green overlapping



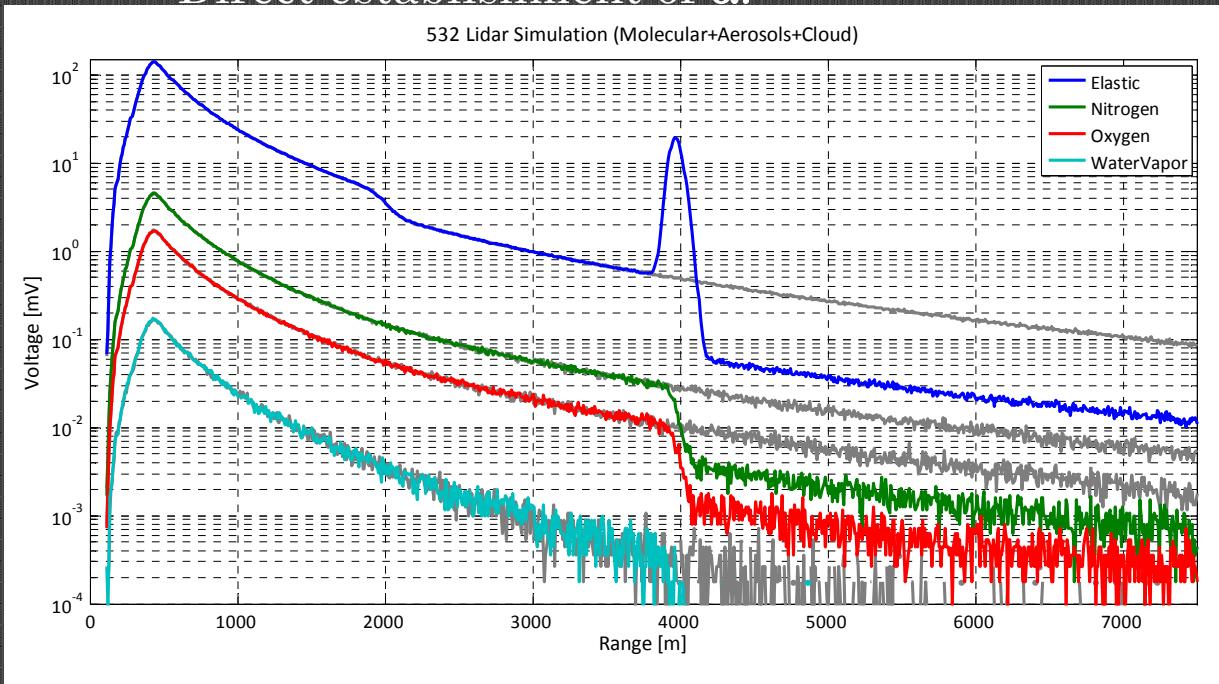
G.Vasileiadis LPTA/CNRS

SO WHAT'S NEXT ?

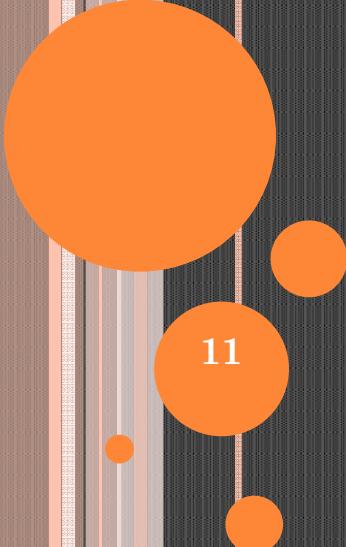
As a results of divers discussions but mainly as an outcome of the 1st workshop in Prague we have orient ourselves on developing a RAMAN Lidar.

Benefits :

- Direct establishment of α .



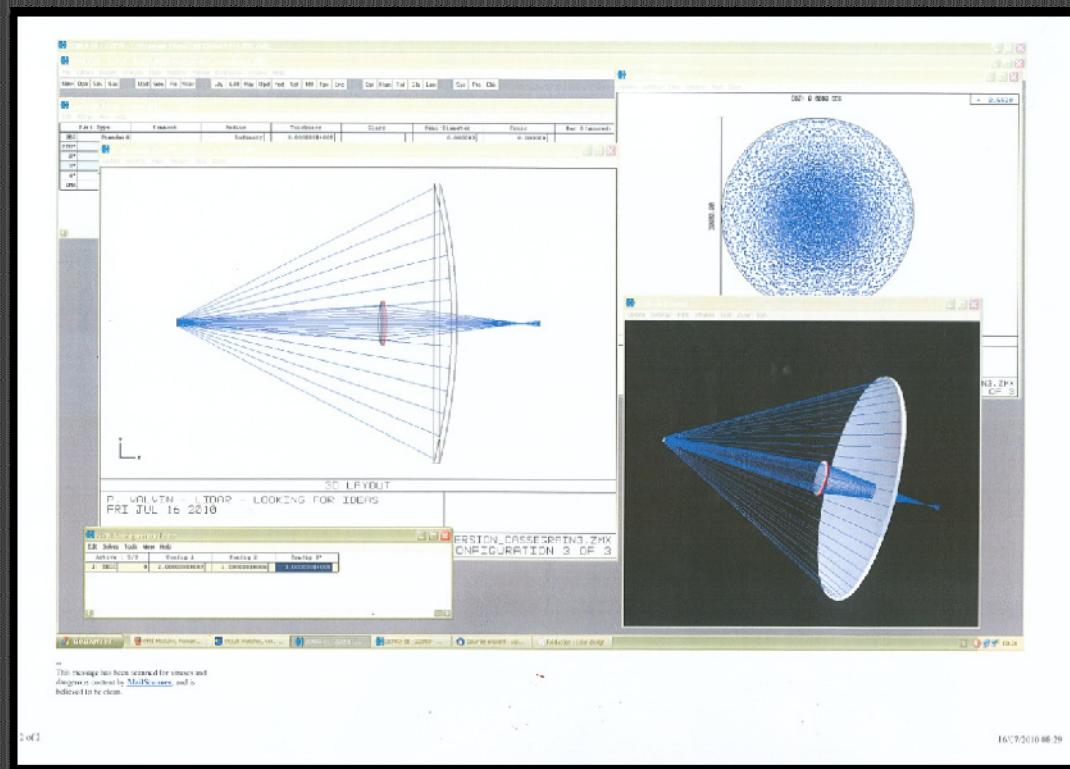
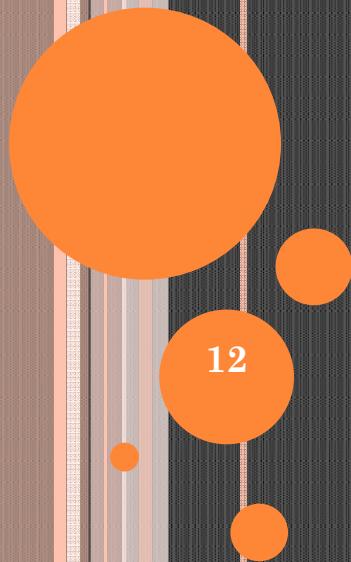
Another IACT peculiarity : *We need to obtain a reasonable Raman spectrum within 120sec.*



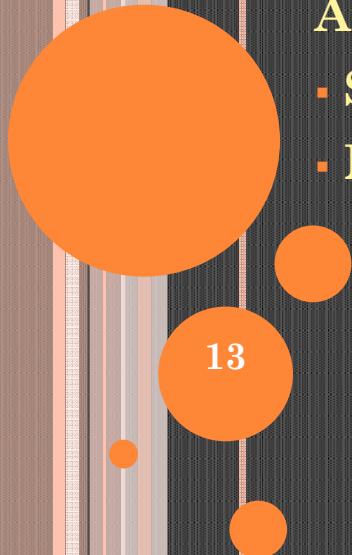
OPTICAL SIMULATION STUDIES

Based on the CLUE mirror specs :

- Diameter 1.8m/f1
- Simulate performance
- Compare Cassegrain vs Optical fiber readout
- Optimize performance and Photon Detection
- Zemax Optical SW tool.

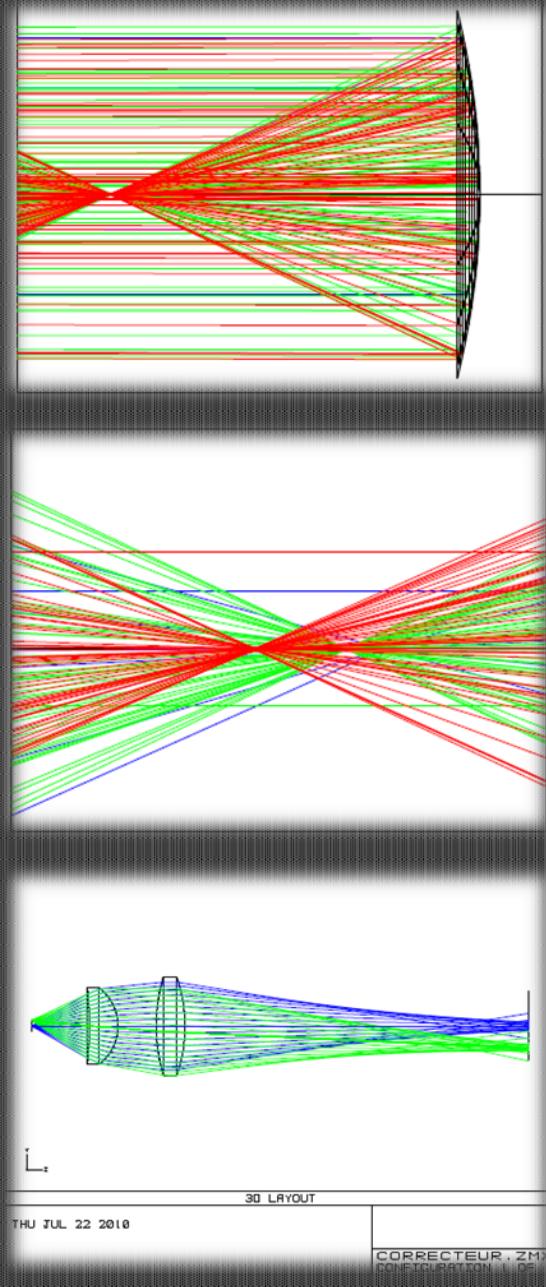


OPTION ONE



Fiber Newton Approach

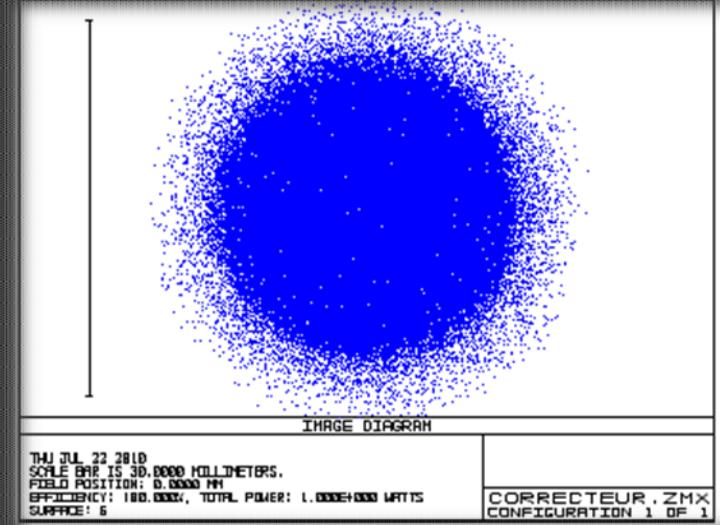
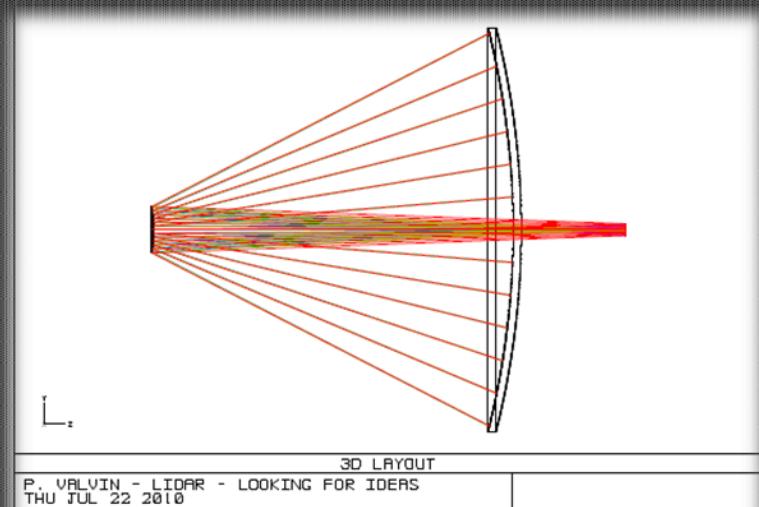
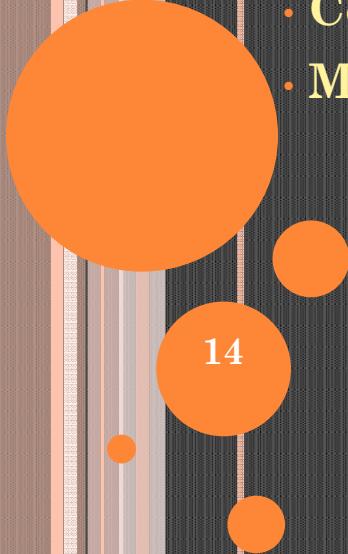
- Altitude range 300-20000m
- Liquid Fiber 6mm Ø f/1
- $Q_e \sim 15\%$
- Collimation using 2-lens system
- Numerical Output 0.6
- Difficult to collimate for both Altitude extremes
- Still durable
- More analysis ongoing



OPTION TWO

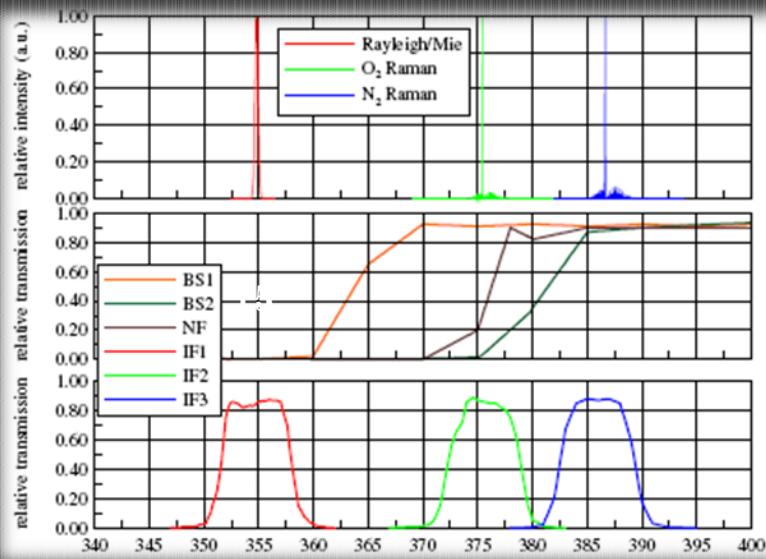
Cassegrain

- Secondary Optics
 - Parabolic 20-25cm Ø
- Output Trace within 50mm
- Durable
- High Qe
- Costly
- More difficult to optimise



WORKING BENCH

- Baseline CLUE Container
- LASER : Inlite II or Quantel CFR200 Mil spec.
- Register O₂&N₂ Raman Lines
- Barr Co. Raman optical Bench
- LICEL
- Photomultiplier R/O
- ACS Control SW
- *See also Alicia's talk tomorrow for another approach*





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CONCLUSIONS

After some years of experimenting we concluded that we have to opt for a Raman LIDAR capable of extracting the absorption coefficient α within a 2' run.

Three groups actually are pursuing this path and we eagerly wait (~2011) the outcome.