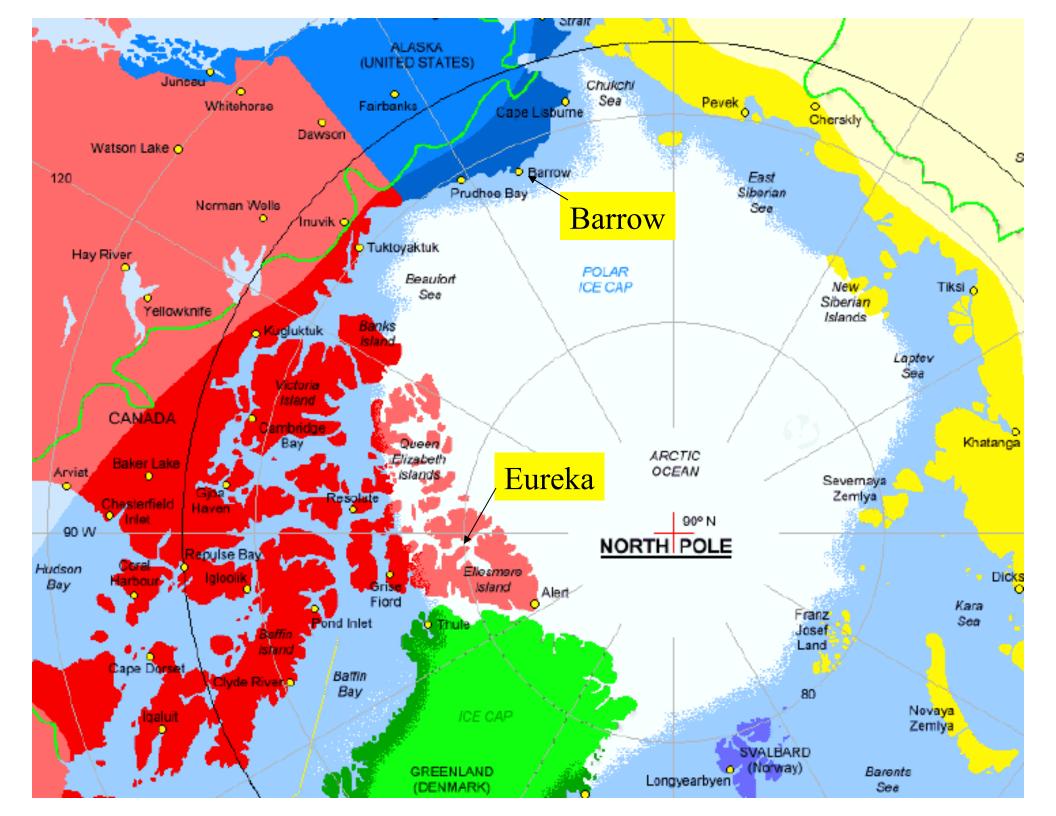
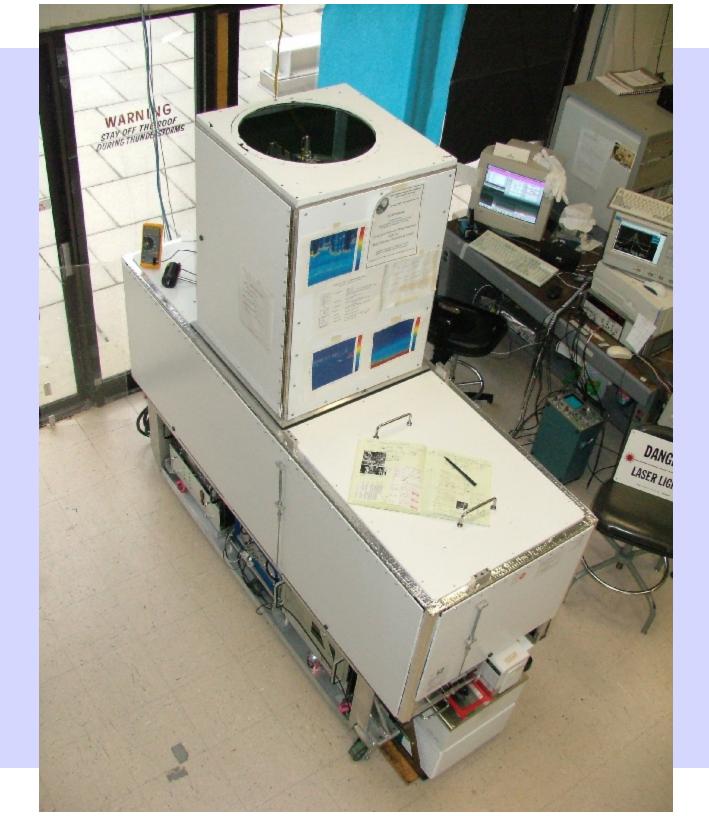
High Spectral Resolution Lidar Ed Eloranta http://lidar.ssec.wisc.edu

Designation in the

**MEM** 

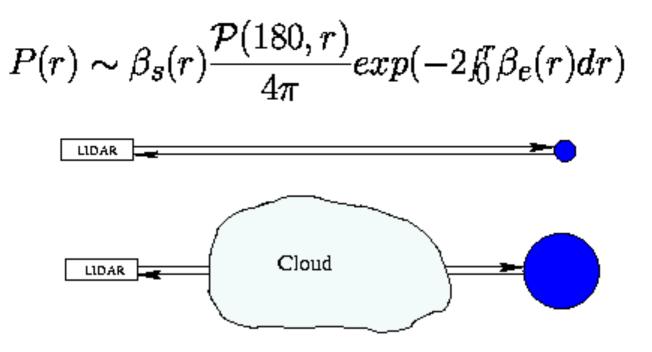
Alaska, Barrow, 2004 by Igor Razen







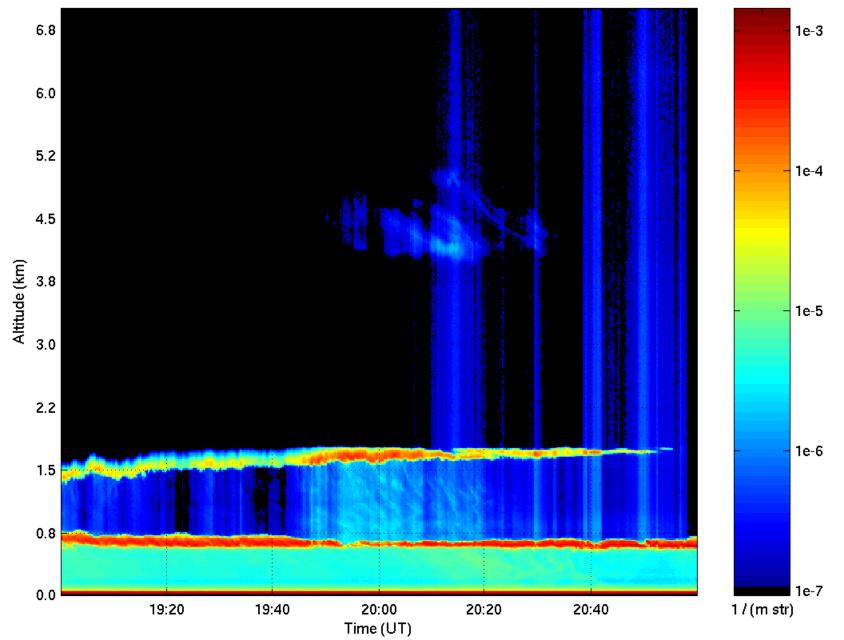


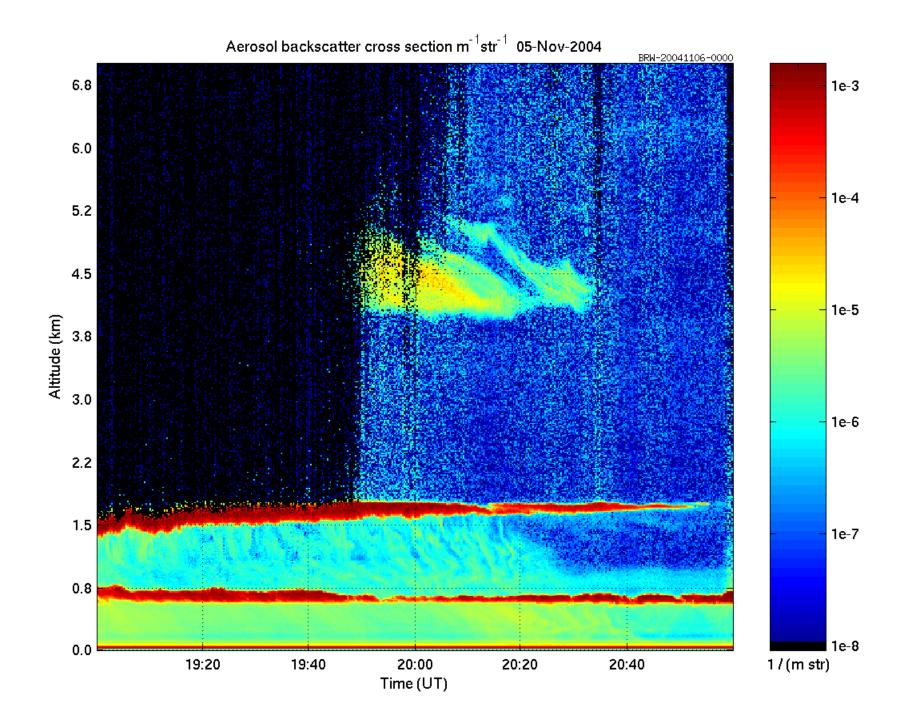


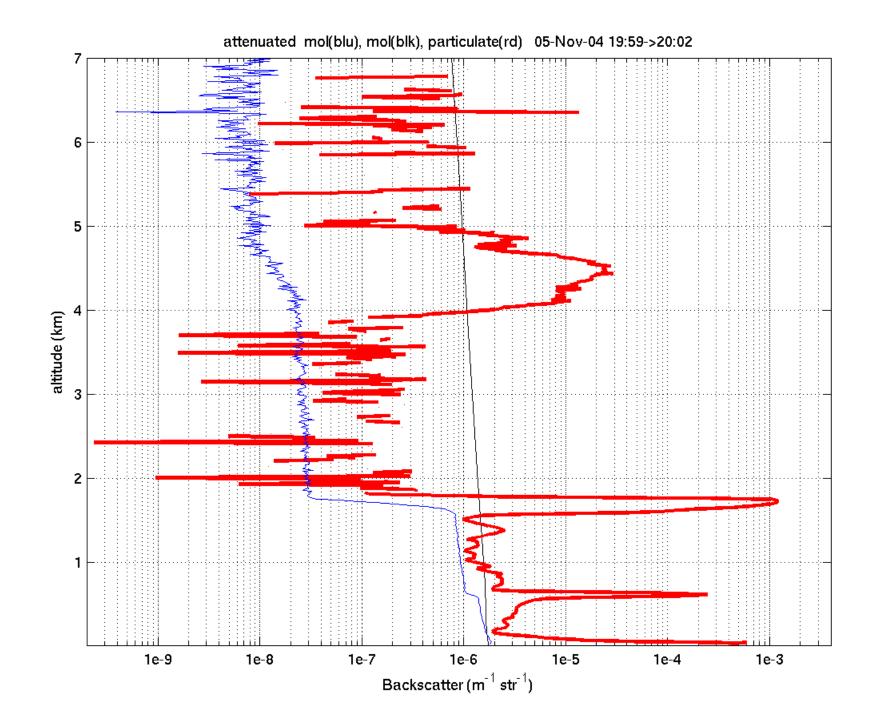
Traditional aerosol lidar can not distinguish between changes in target reflectivity and attenuation between the lidar and the target

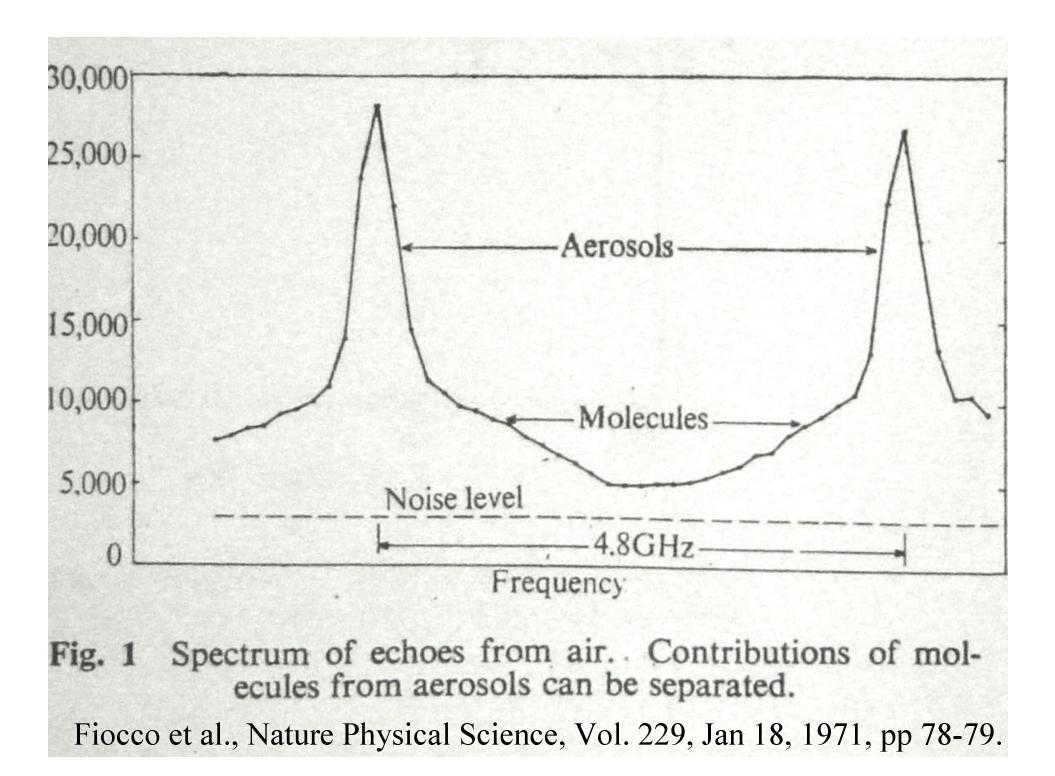
The power,  $p_c(r)$ , received by a standard aerosol lidar where the scattering cross sections of aerosols,  $\beta_a(r)$ , and molecules,  $\beta_m(r)$ , both contribute:  $p_c(r) \sim \frac{1}{r^2} \cdot \left(\frac{P(180,r)}{4\pi}\beta(r) + \frac{3}{8\pi}\beta_m(r)\right) \cdot exp(-2\int (\beta_a(r) + \beta_m(r)) \cdot dr)$ If we can separate returns from aerosols,  $p_a(r)$ , and molecules,  $p_m(r)$ :  $p_a(r) \sim \frac{1}{r^2} \cdot \frac{P(180,r)}{4\pi} \beta_a(r) \cdot exp(-2\int (\beta_a(r) + \beta_m(r)) \cdot dr) - \text{aerosol return},$  $p_m(r) \sim \frac{1}{r^2} \cdot \frac{3}{8\pi} \beta_m(r) \cdot exp(-2\int (\beta_a(r) + \beta_m(r)) \cdot dr) - \text{molecular return}$ where  $\frac{P(180,r)}{4\pi}$  is the backscatter phase function. The backscatter ratio can be obtained from the ratio of these returns:  $R(r) = \frac{p_a(r)}{p_m(r)} = \frac{\frac{P(180,r)}{4\pi} \cdot \beta_a(r)}{\frac{3}{2} \cdot \beta_m(r)}$  where  $\beta_m(r)$  ~atmospheric density profile,  $\rho(r)$ .  $\beta_a'(r) = \frac{P(180,r)}{4\pi} \cdot \beta_a(r) = \frac{3}{8\pi} \cdot \beta_m(r) \cdot \frac{p_a(r)}{n_m(r)}$ The optical depth between  $r_1$  and  $r_2$  is derived by comparing the molecular return to that expected from a purely molecular atmosphere:  $\tau(r_1, r_2) = \frac{1}{2} \cdot log(\frac{r_1^2 \rho(r_2) \cdot p_m(r_1)}{r_2^2 \rho(r_1) \cdot p_m(r_2)})$ 

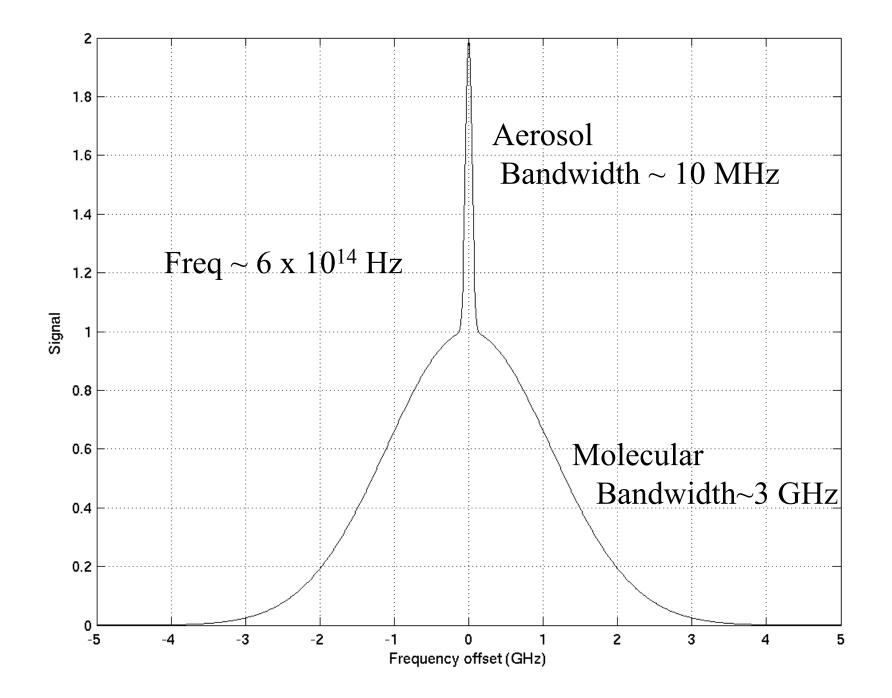
Attenuated backscatter (m<sup>-1</sup>str<sup>-1</sup>) 05-Nov-2004

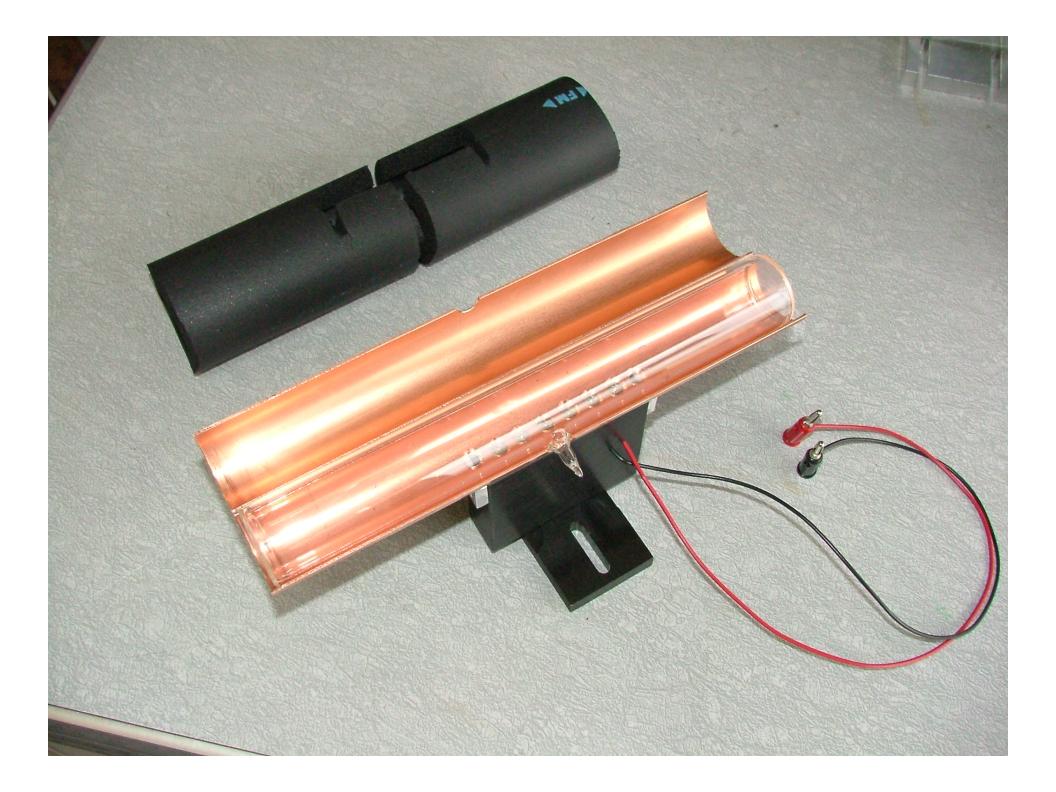


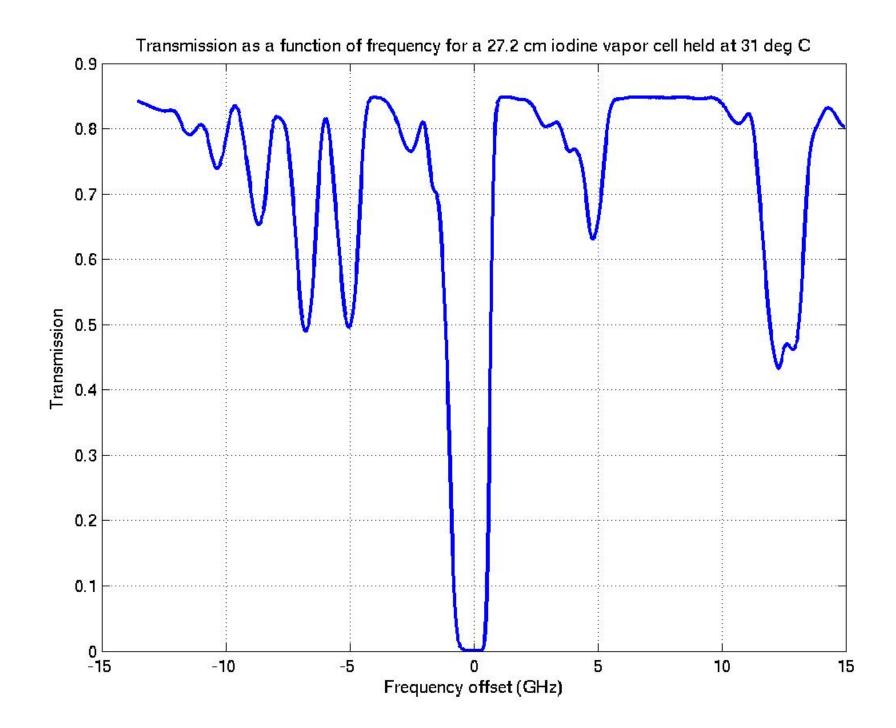


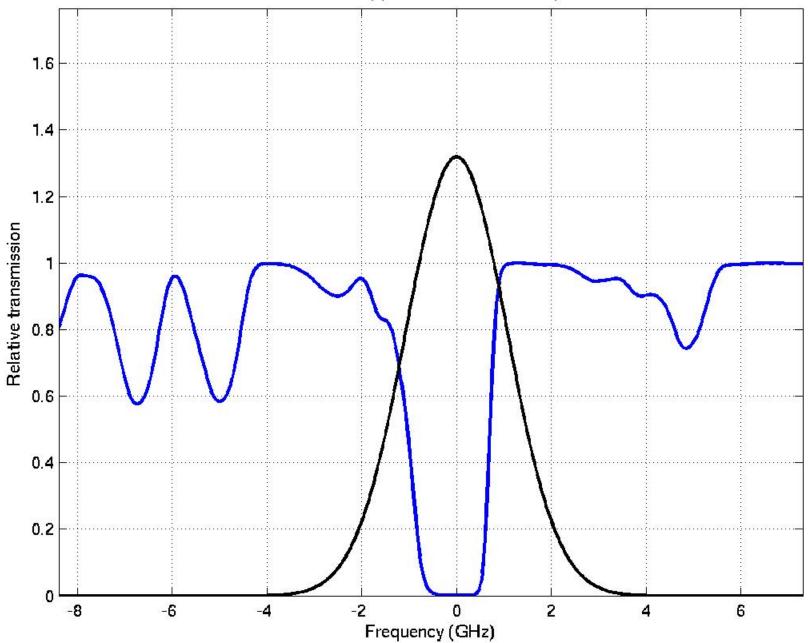




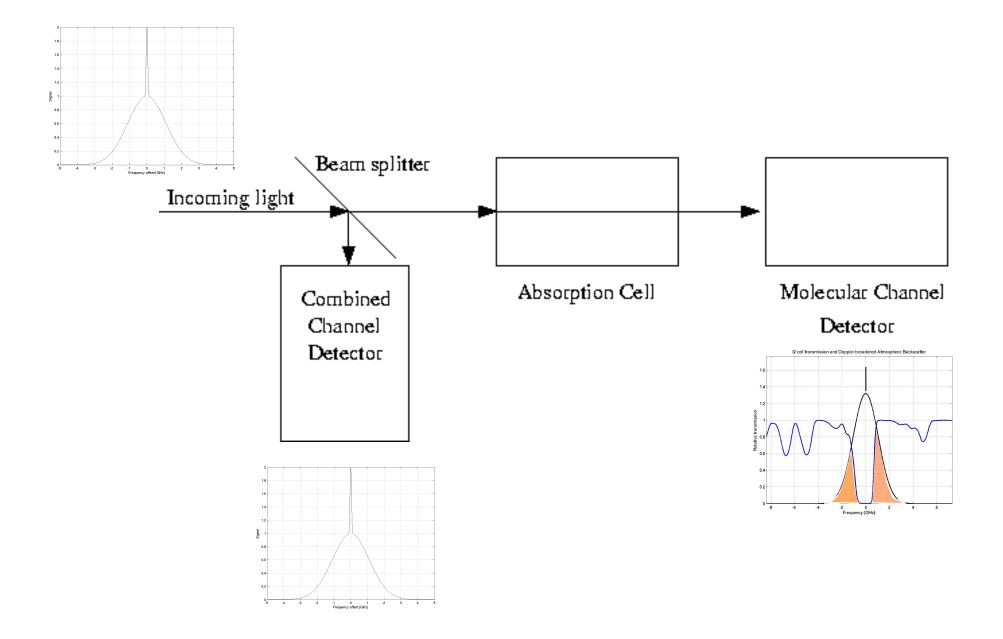








I2 cell transmission and Doppler broadened Atmospheric Backscatter



### **Basic HSRL Equations**

 $S_c = G_{ac}N_a + G_{mc}N_m$ ; eq 1—Signal in the combined channel

 $S_m = G_{am}N_a + G_{mm}N_m$ ; eq 2—Signal in the molecular channel

Where  $G_{ik}$  are gains of the two channels when exposed to  $N_a$  aerosol and  $N_m$  molecular photons. Solving for  $N_m$  and  $N_a$  yields:

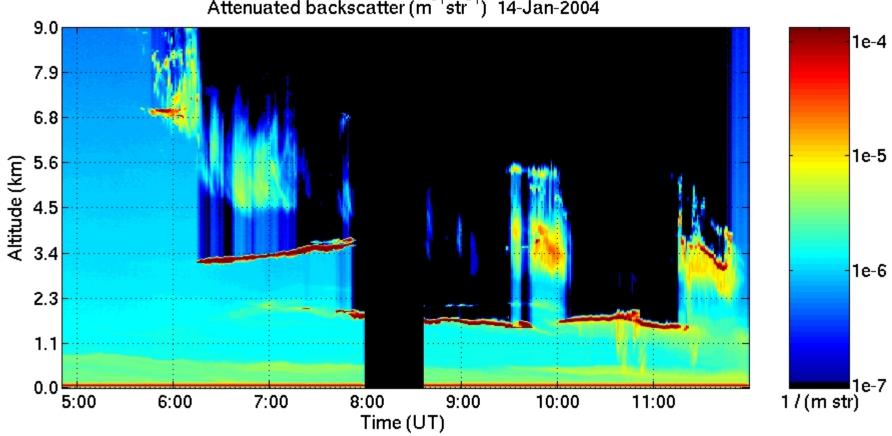
 $N_m = \frac{S_m/G_{am} - S_c/G_{ac}}{(G_{mm}/G_{am}) - (G_{mc}/G_{ac})}$ ; eq 3—Number of molecular photons incident as function of signals  $N_a = \frac{S_c/G_{mc} - S_m/G_{mm}}{(G_{ac}/G_{mc}) - (G_{am}/G_{mm})}$ ; eq 4–Number of aerosol photons incident as function of signals With  $G_{ac}$  =gain of the combined channel when exposed to aerosol photons Define other gains relative to  $G_{ac}$ :

$$\begin{aligned} G_{mc} &= C_{mc} \cdot G_{ac}, \ G_{am} = C_{am} \cdot G_{ac}, \ G_{mm} = C_{mm} \cdot G_{ac} \\ N_m &= (1/G_{ac}) \cdot \frac{S_m/C_{am}-S_c}{(C_{mm}/C_{am})-C_{mc}} = (1/G_{ac}) \cdot \frac{S_m-C_{am}S_c}{C_{mm}-C_{mc}C_{am}} \\ N_a &= (1/G_{ac}) \cdot \frac{S_c/C_{mc}-S_m/C_{mm}}{(1/C_{mc})-(C_{am}/C_{mm})} = (1/G_{ac}) \cdot \frac{S_c/C_{mc}-S_m/C_{mm}}{(1/C_{mc})-(C_{am}/C_{mm})} = (1/G_{ac}) \cdot \frac{C_{mm}S_c-C_{mc}S_m}{C_{mm}-C_{mc}C_{am}} \\ \end{aligned}$$
The scattering ratio is then:
$$\frac{N_a}{N_m} = \frac{C_{mm}S_c-C_{mc}S_m}{S_m-C_{am}S_c} \end{aligned}$$

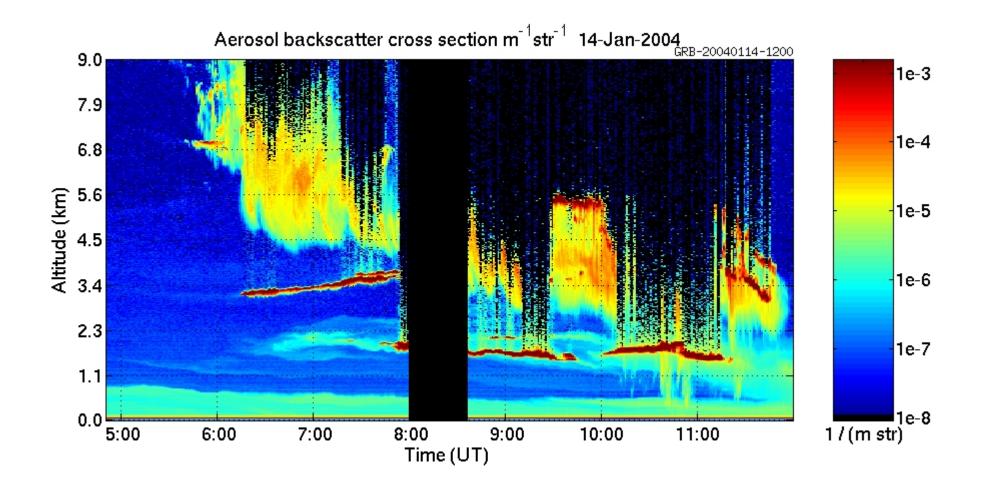
The backscatter cross section,  $\beta'_a$ , is:

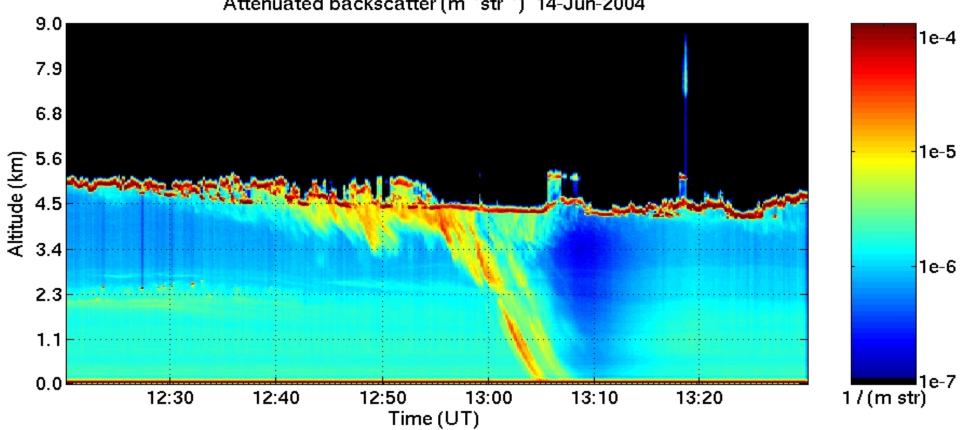
 $\beta'_{a}(r) = \beta_{a}(r) \cdot \frac{P(180,r)}{4\pi} = \frac{N_{a}(r)}{N_{m}(r)} \cdot \beta_{m}(r), \text{ where } \beta_{a} = \text{scattering cross section}, \frac{P(180,r)}{4\pi} = \text{backscatter phase function.}$ the optical depth, $\tau$ , between two points  $r_{1}$  and  $r_{2}$  is:  $\tau(r_{2} - r_{1}) = \frac{1}{2} \cdot \log(\frac{r_{1}^{2}\rho(r_{2})\cdot N_{m}(r_{1})}{r_{1}})$  where  $\rho(r) =$  the atmospheric density profile

$$r(r_2 - r_1) = \frac{1}{2} \cdot log(\frac{r_1^2 \rho(r_2) \cdot N_m(r_1)}{r_2^2 \rho(r_1) \cdot N_m(r_2)})$$
, where  $\rho(r)$  = the atmospheric density profile

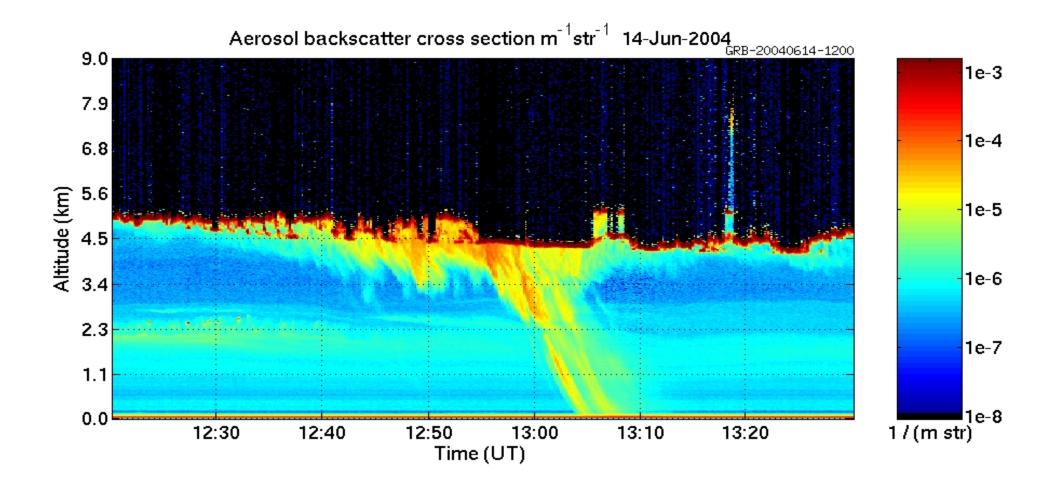


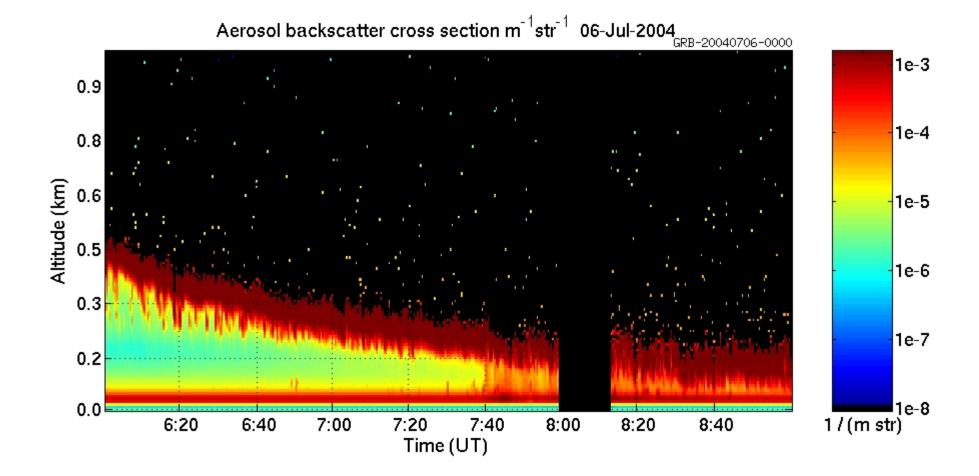
Attenuated backscatter (m<sup>-1</sup>str<sup>-1</sup>) 14-Jan-2004

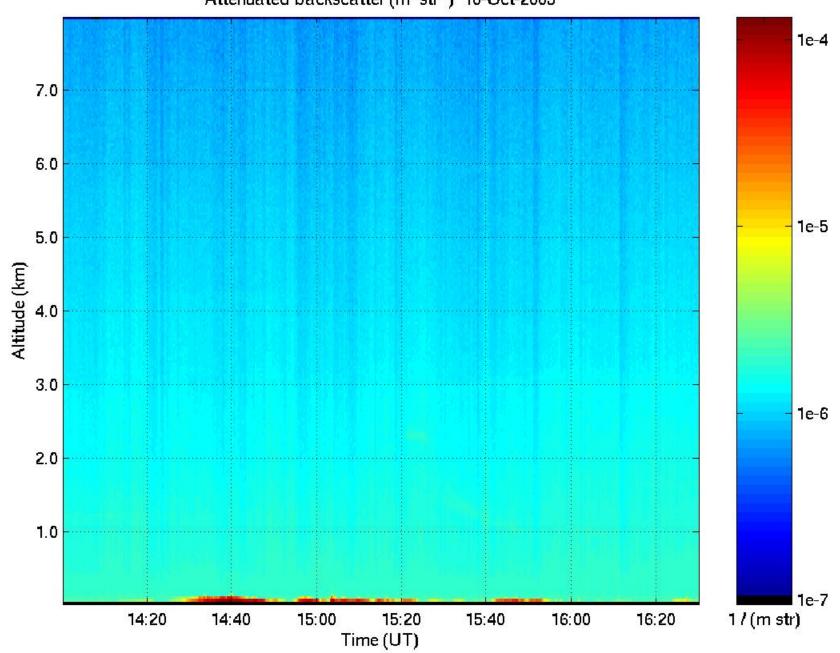




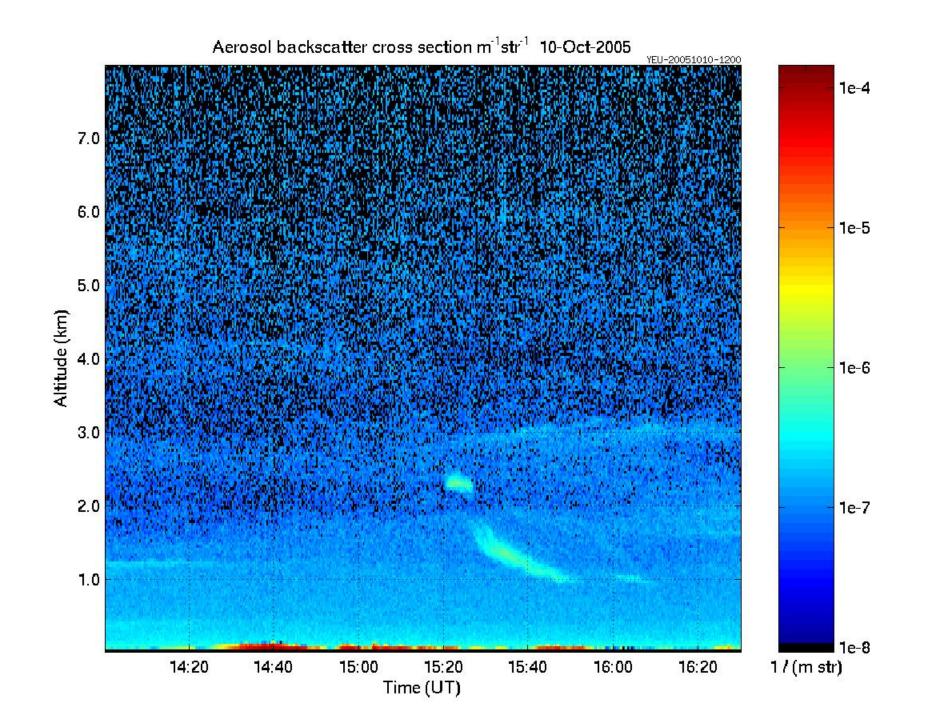
Attenuated backscatter (m<sup>-1</sup>str<sup>-1</sup>) 14-Jun-2004

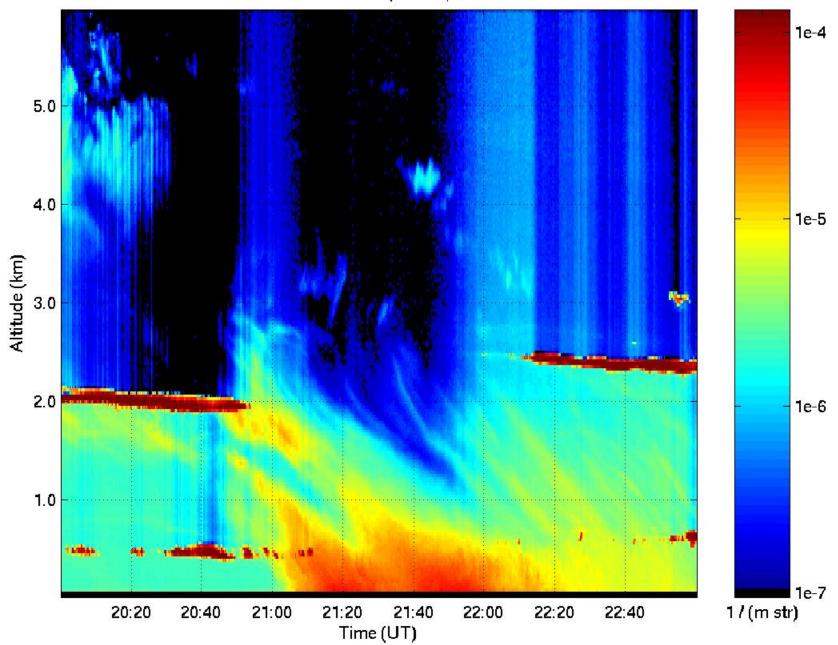




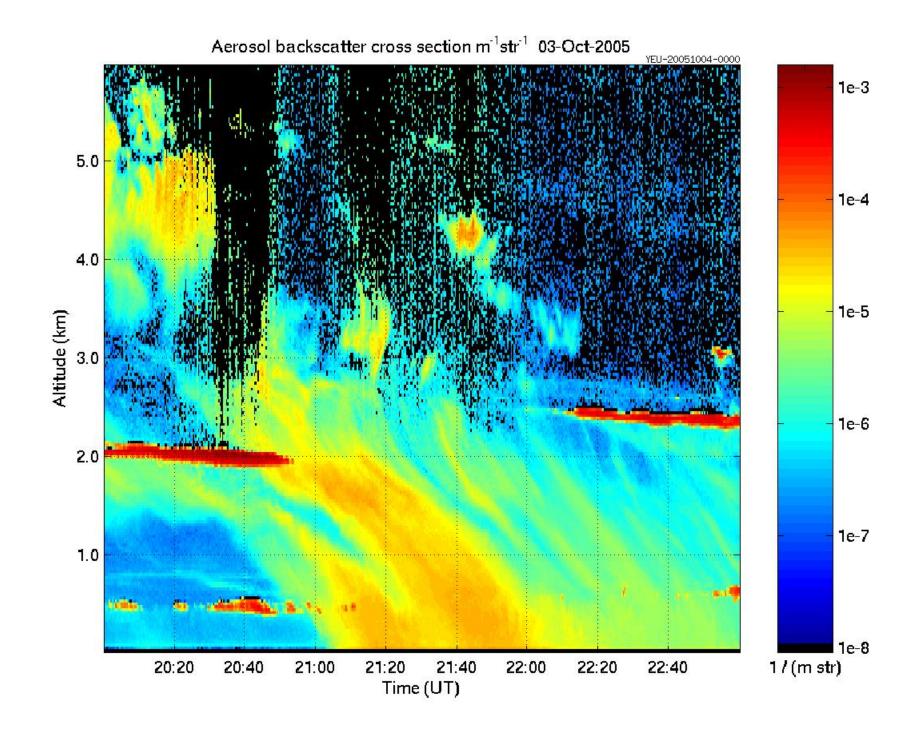


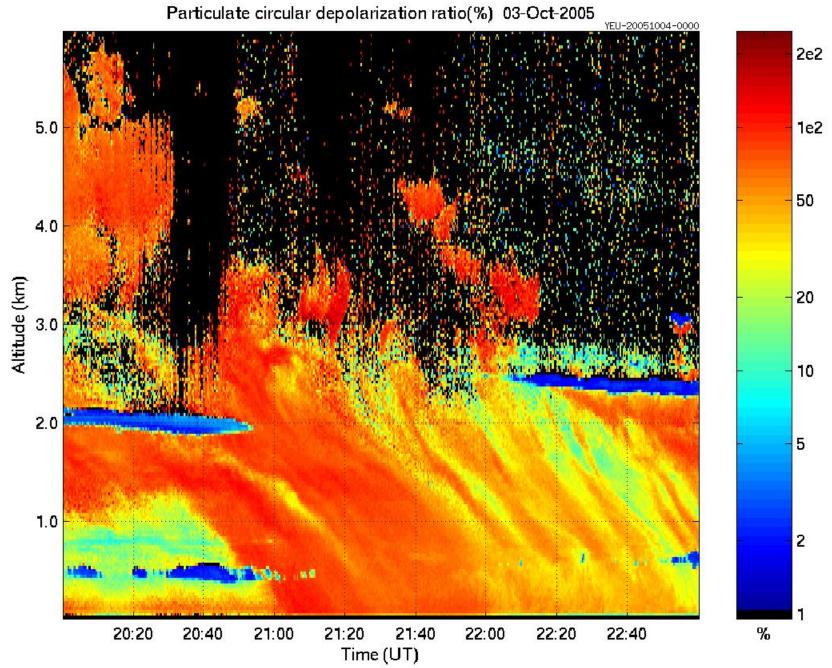
Attenuated backscatter (m<sup>-1</sup>str<sup>-1</sup>) 10-Oct-2005



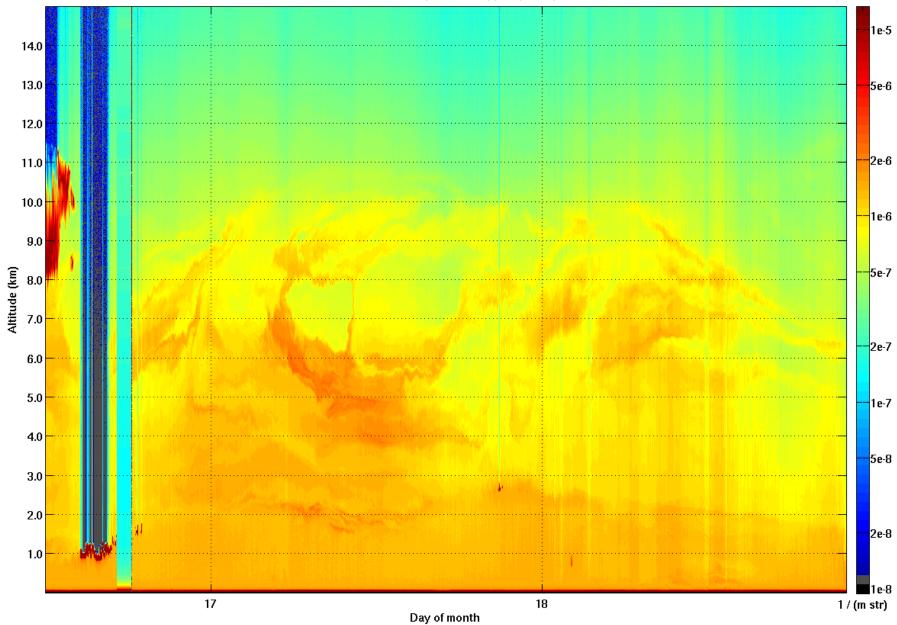


Attenuated backscatter (m<sup>-1</sup>str<sup>-1</sup>) 03-Oct-2005

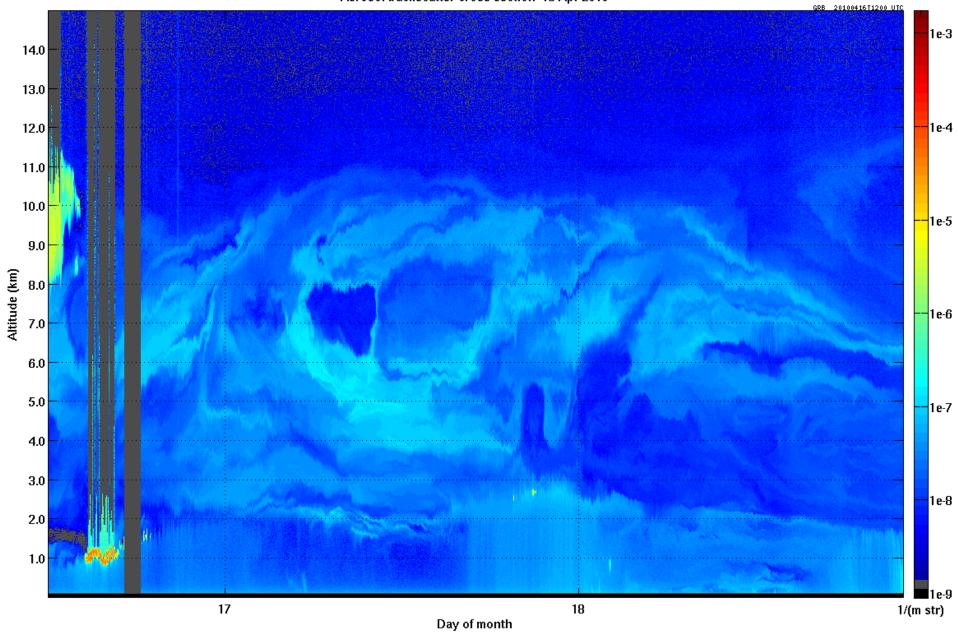


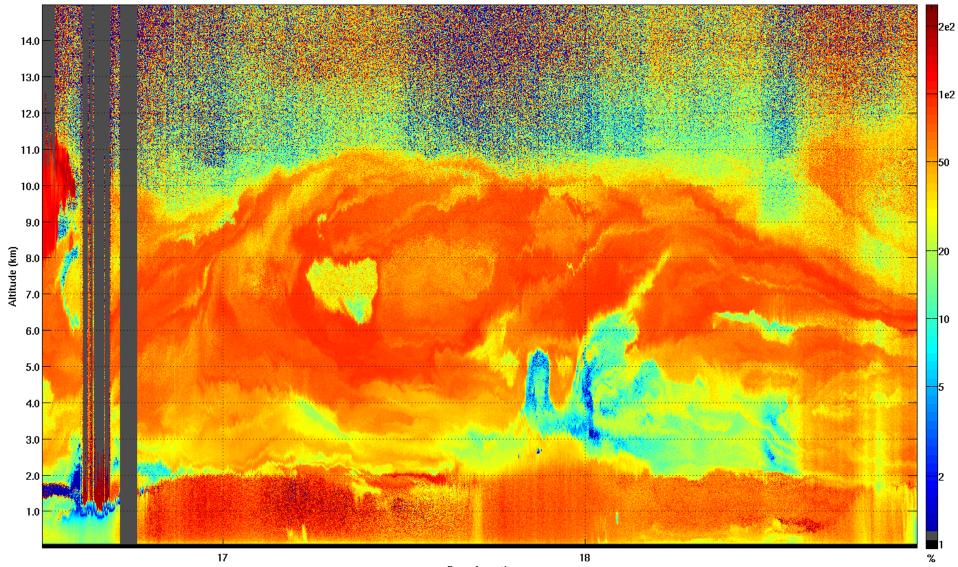


Attenuated backscatter (no masks applied) 16-Apr-2010



Aerosol backscatter cross section 16-Apr-2010





Day of month

# Advantages of 532 nm operation

- --Iodine adsorption line for filtering
- --Important wavelength for radiative transfer
- --Allows use of doubled Nd:YAG laser
- --Strong molecular scattering

# Problem with 532 nm—eye safety

--Wavelength region with smallest permitted exposure max single pulse exposure =  $5e-7 \text{ J/cm}^2$ 

### Eye safety

532 nm wavelength has smallest permitted exposure ANSI safe exposure  $\leq 5e-7$  (PRF/4)-1/4 J/cm<sup>2</sup> Where PRF = the pulse repetition frequency This forces high repetition rates and large apertures Range ambiguity limits R < ~4kHz, i.e.  $r_{max} < ~40 \text{ km}$ Cost, complexity, turbulence limit aperture to  $\sim 0.5$  m. Thus max transmitted energy laser pulse is limited to:  $p25^{2*}5e-7^{1000^{-1/4}}=0.174 \text{ mJ/pulse}$ 

and the maximum transmitted power is:

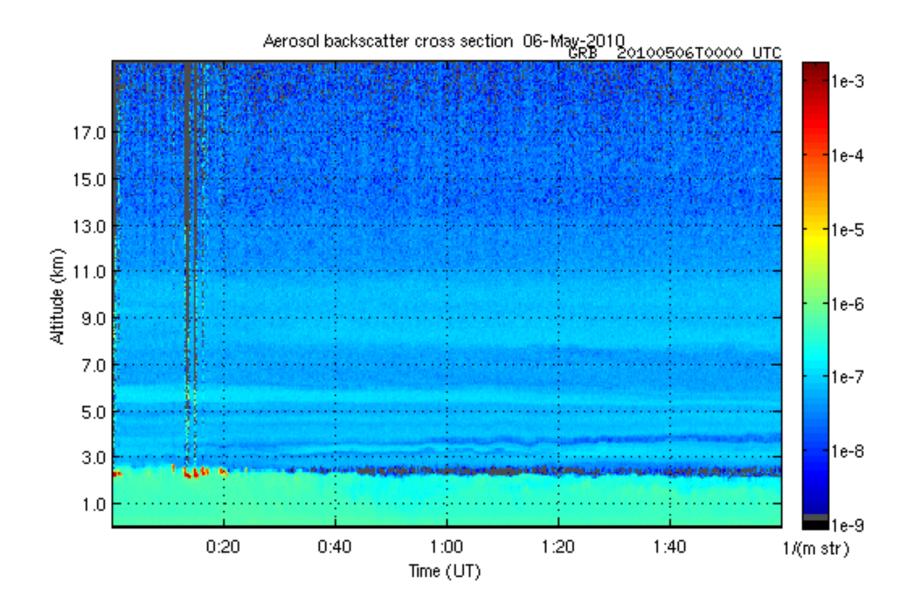
0.174e-3\*4000 Hz = 0.7 Watt

-- Low energy laser pulses reduce the signal to noise ratio of the lidar small lidar returns must compete with scattered sunlight

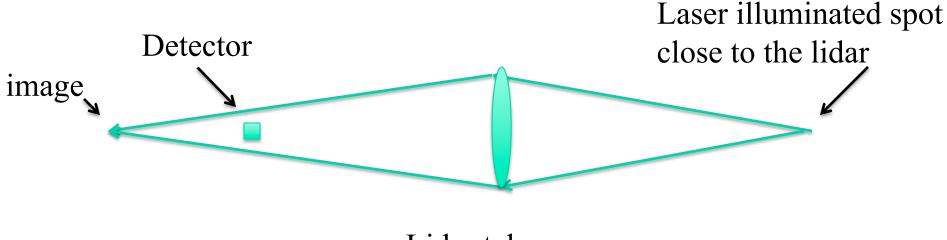
--this forces a small receiver acceptance angles (40-100 microradians) a narrow spectral bandpass ~8 GHz

# Arctic HSRL Specifications

•	Altitude coverage	~75m>30 km
•	Altitude resolution	7.5 m
•	Time resolution :	
•	-Backscatter, depolarization	profiles 0.5 sec
•	-Optical depth profiles	>20 sec
•	Eye safe at output	
٠	Wavelength	532 nm
•	Power	200 → 600 mW
•	Repetition rate	4 kHz
٠	Field of view	45 microradians
•	Sky noise filter bandwidth	8 GHz
•	Typical background noise/bin	>1 photon/1000 laser pulses
•	Receiver diameter	0.4 m
•	I2 filter bandwidth	1.8 GHz

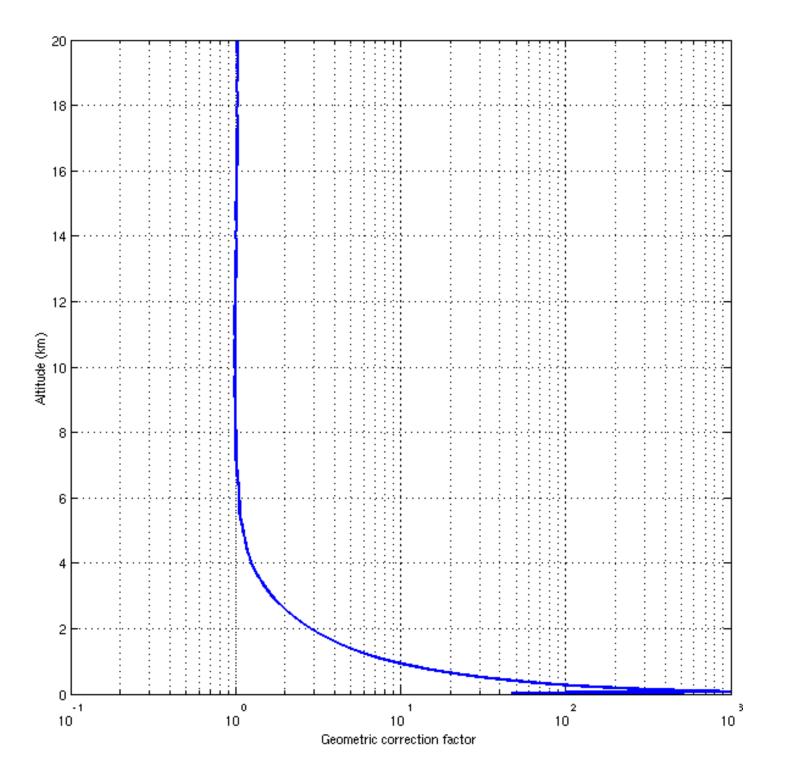


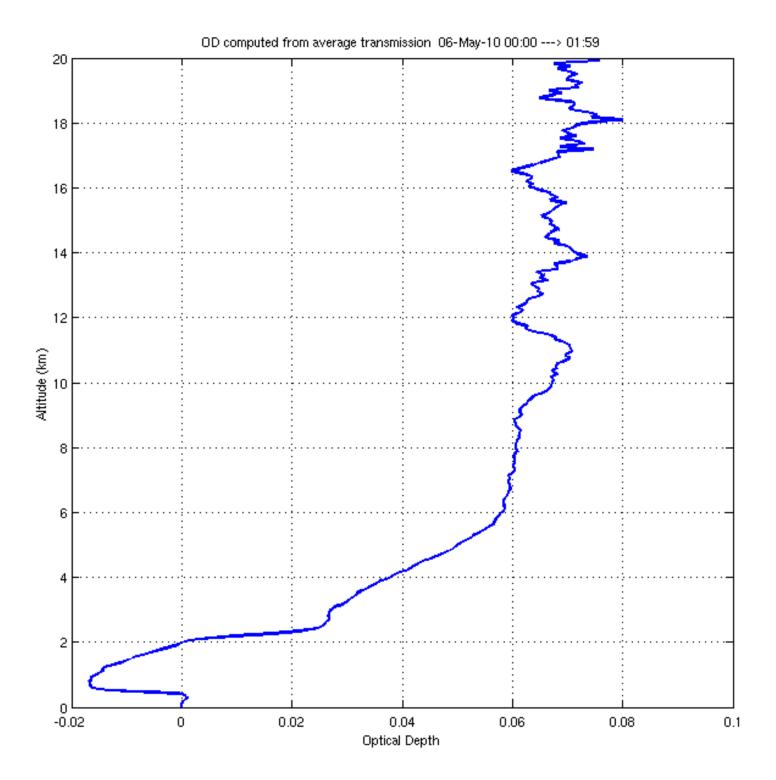
### Geometric correction --- overlap correction



Lidar telescope

A lidar with a small field-of-view will require a large geometric correction at close range because the out of focus image becomes much larger than the detector.





Comparison between potential systems

-- Raman

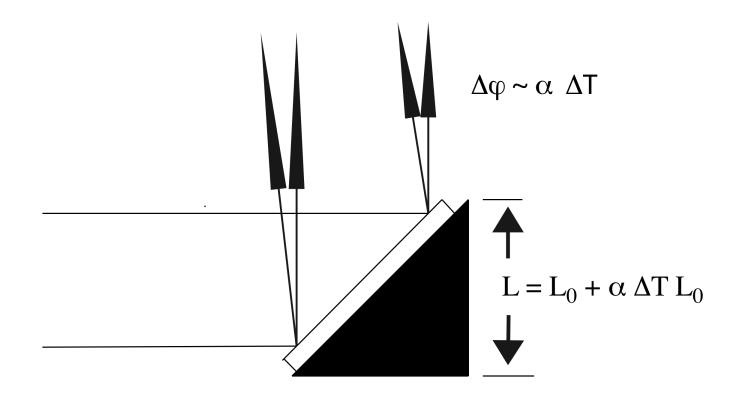
scattering cross section~ Rayleigh/1000 requires high average power ~10's W large diameter receiver, typically ~0.6 m can accommodate larger FOV's relatively simple

--I2 HSRL 532nm

scattering cross section ~Rayleigh/4 lower average power ~0.5 W somewhat smaller receiver ~0.4 m can accommodate larger FOV's relatively complex

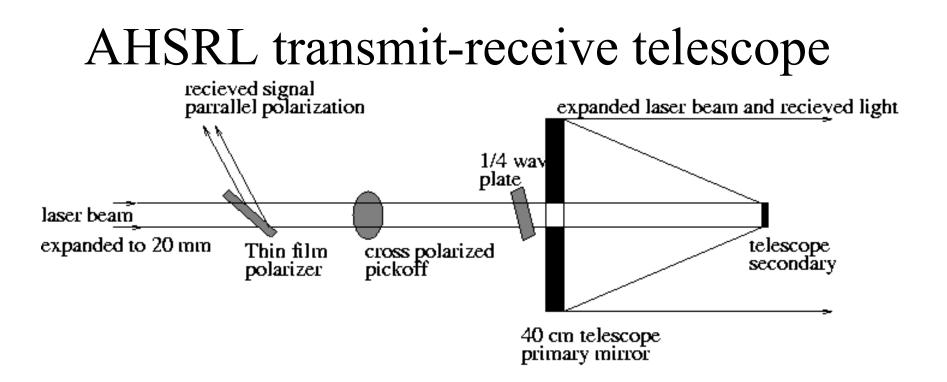
--Fabry-Perot HSRL 355 nm larger cross section that 532nm more complex than I2 field-of-view due to Fabry-Perot HSRL data can be found at: http://lidar.ssec.wisc.edu

Fureka Anaust 2005 by Jaor Razenkov



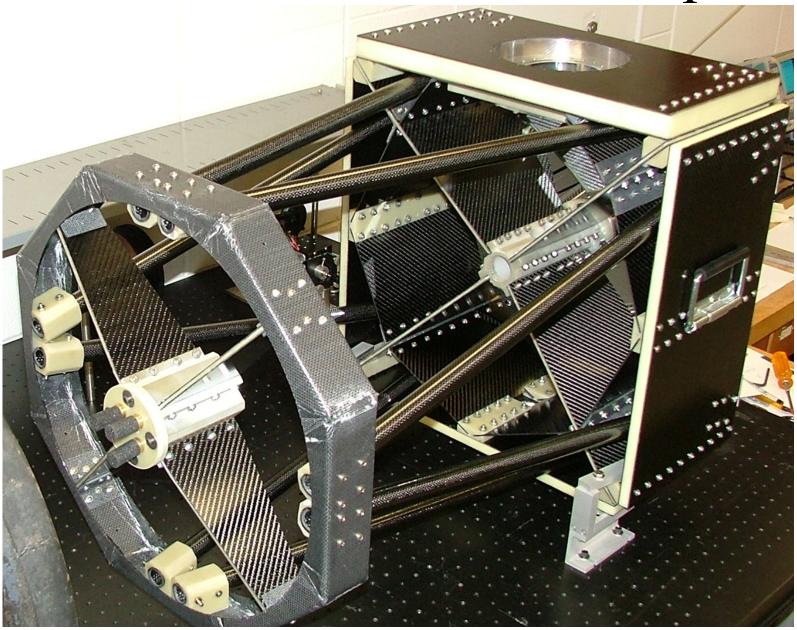
Thermal expansion of components effect the alignment of transmitter with the receiver. Here we consider the example of an 45 deg aluminum mountin block for a beam turning mirror.

Angle shift due to 10 deg C temperature change:  $\Delta \phi \sim \alpha \Delta T \sim 2.5 * 10^{-5} * 10 \Delta \phi \sim 250$  microradian



- --The 20 mm diameter linearly-polarized laser beam is converted to circular polarization by <sup>1</sup>/<sub>4</sub> wave plate before expansion 40 cm.
- --The received signal is converted to linear polarization on return through the ¼ wave plate. Approx. 10% of the signal is separated to measure the cross-polarized component. The parallel-polarized component is separated from the transmit beam by the thin-film polarizer.

# Transmit-Receive Telescope



I was asked to comment on following:

--Cost? ~\$ 1M

--Laser power, possible interference with Cerenkov measurements? 532 nm, maximum eyesafe power ~0.5 W

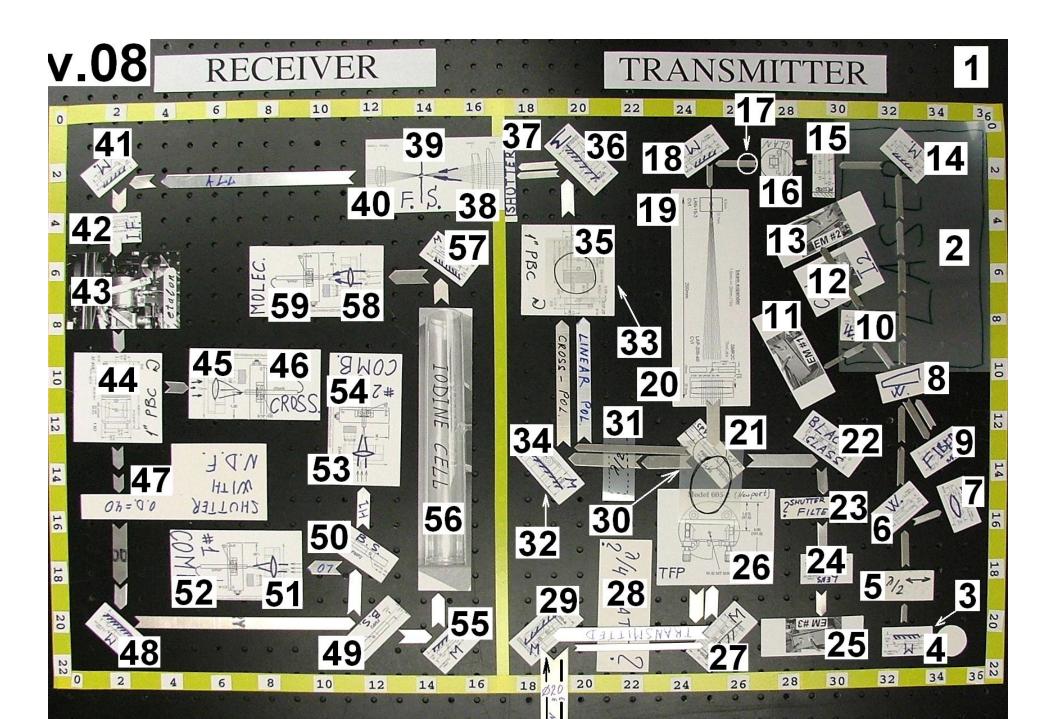
--Maintenance cost?

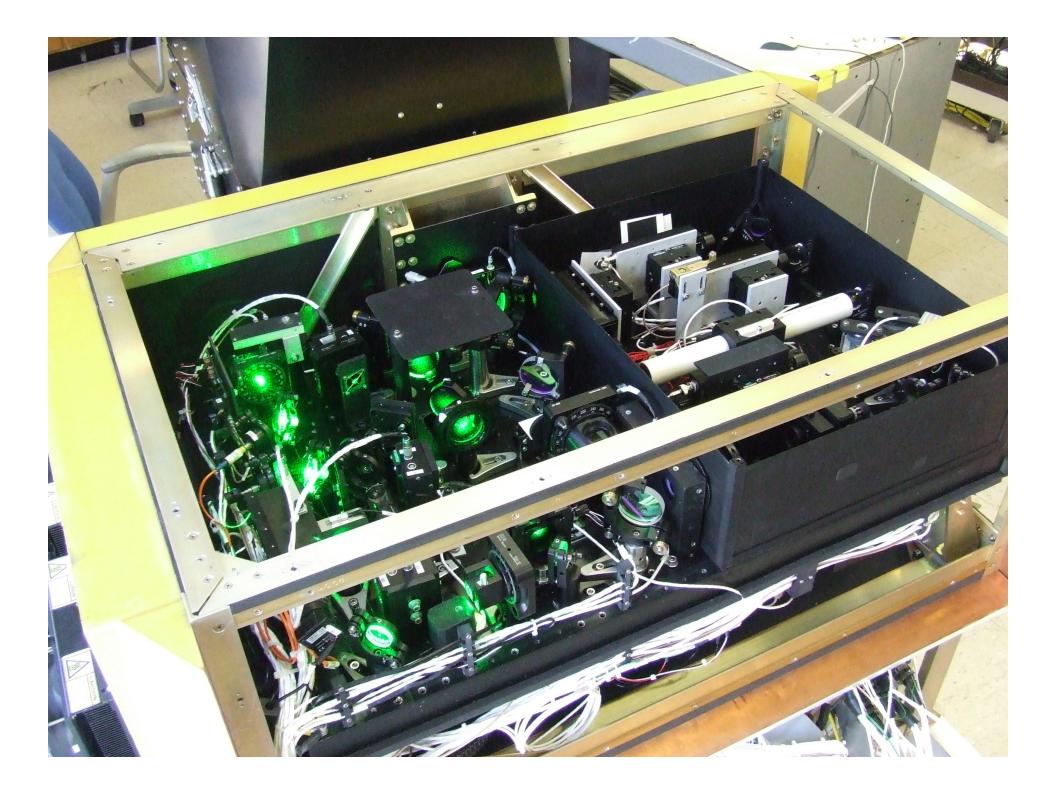
Relatively low, visit Arctic site ~ 2-3 times/year

--External data requirements?

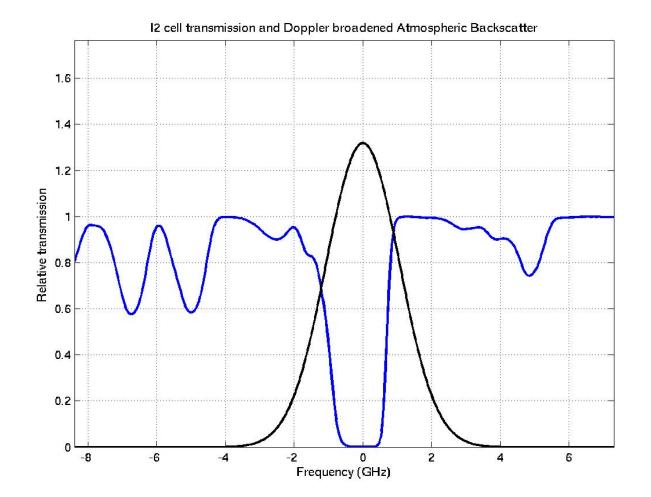
Radiosonde profiles from national weather service.



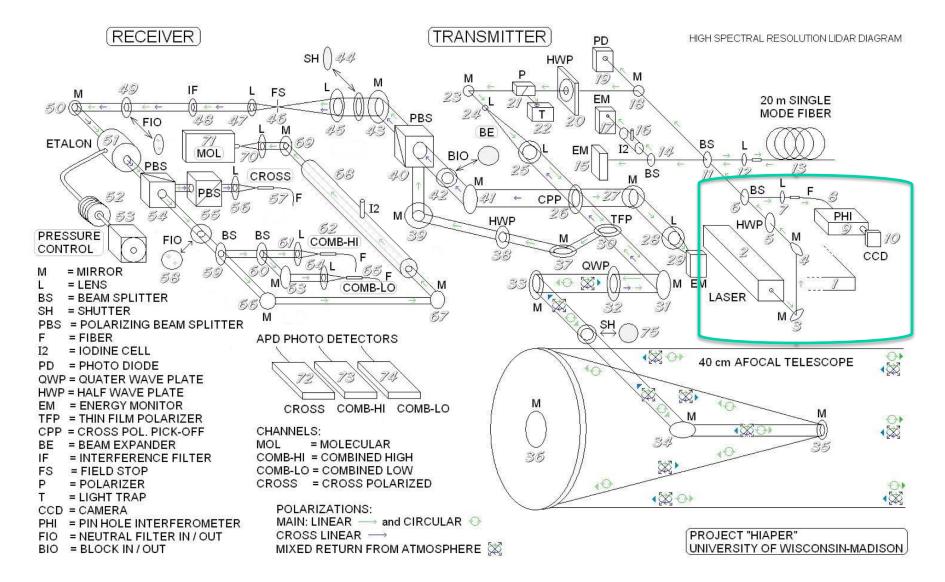




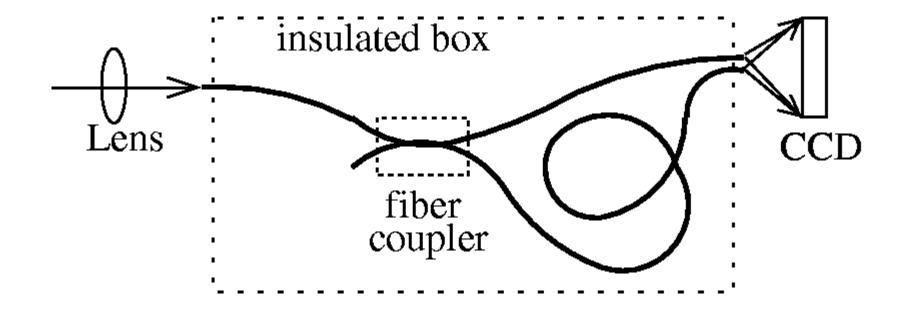
# HSRL calibration requires a spectral scan to determine bandpass of filters



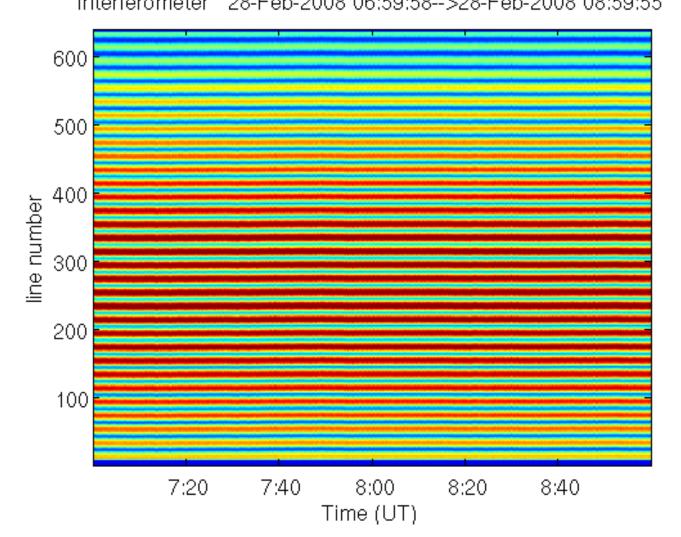
### Laser and interferometer



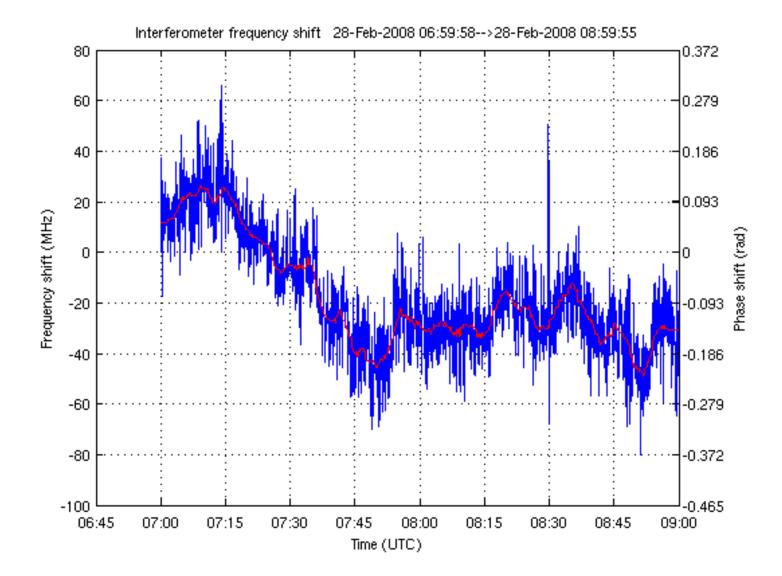
## Young's pinhole interferometer

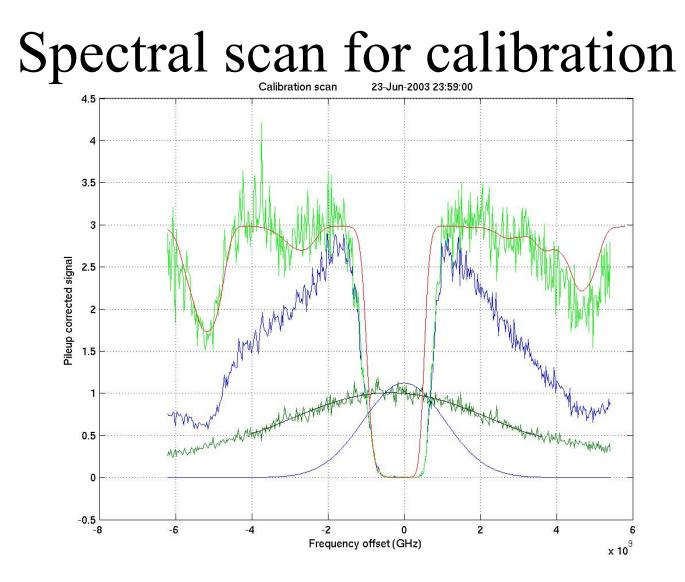


## One column from interferometer Interferometer 28-Feb-2008 06:59:58-->28-Feb-2008 08:59:55

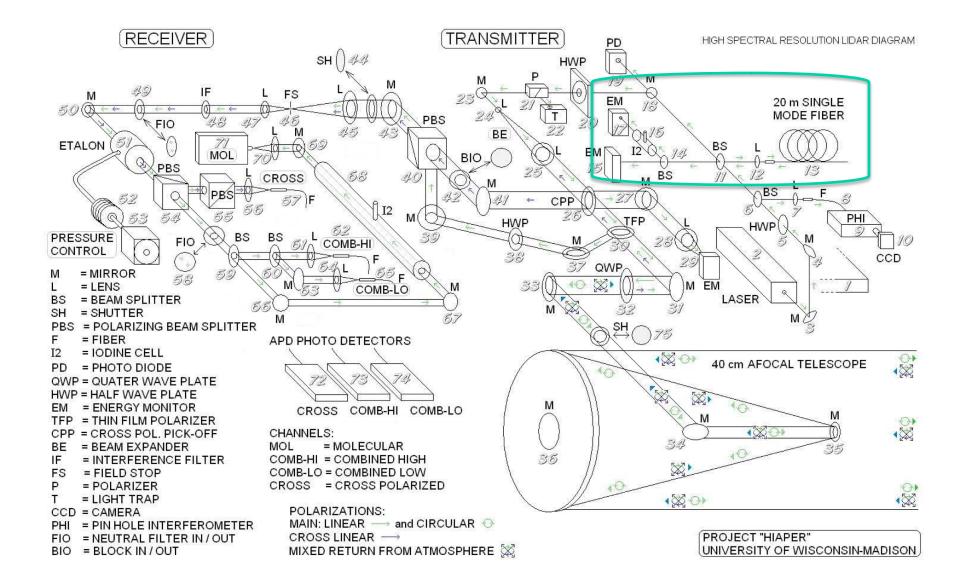


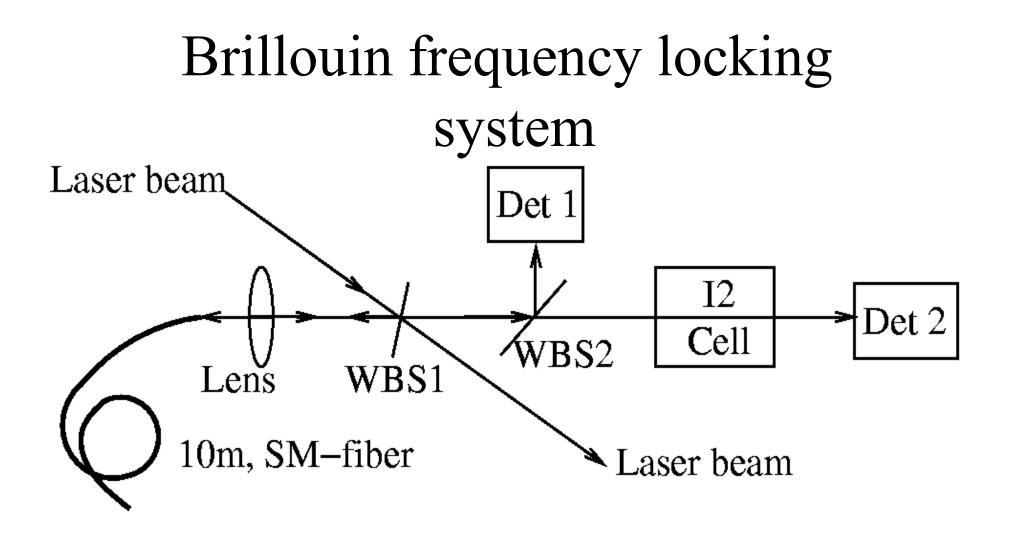
# Frequency vs time from interferometer



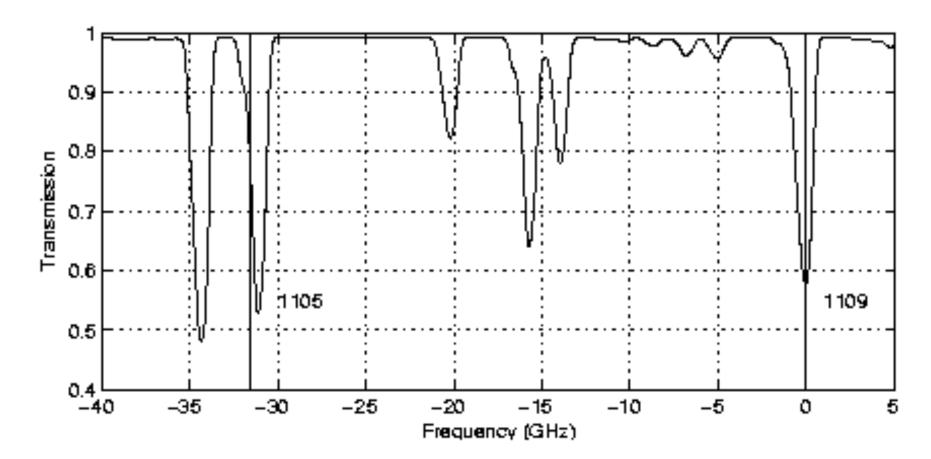


## Brillouin frequency locking

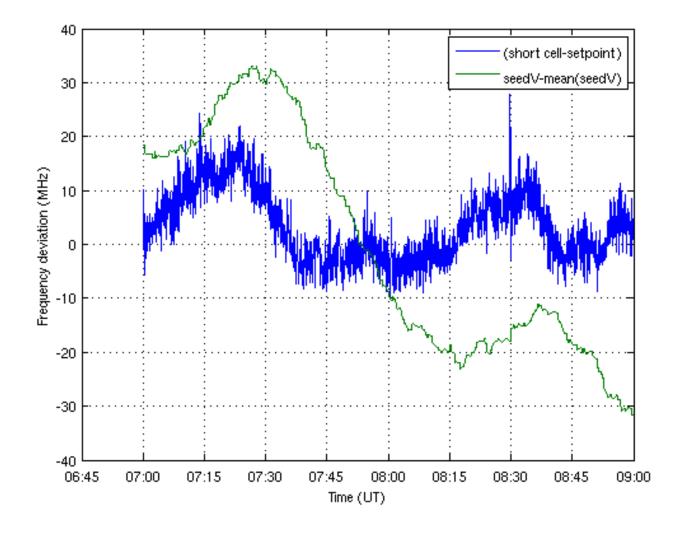


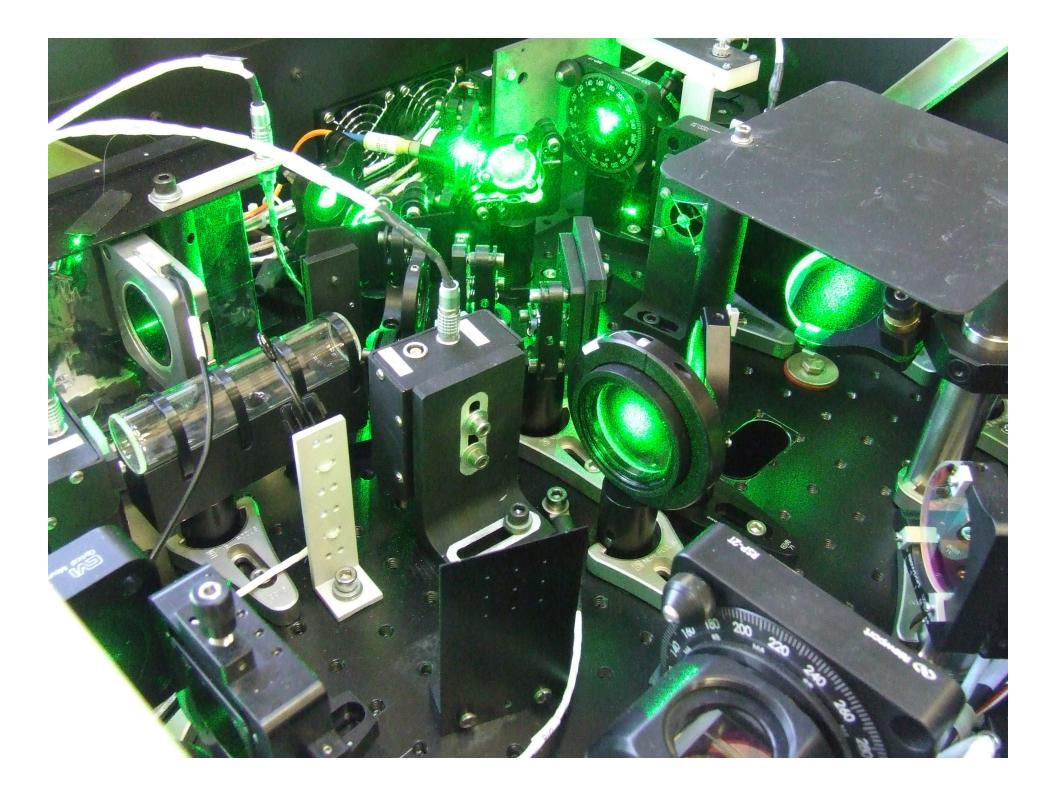


## Transmission of 2-cm iodine cell

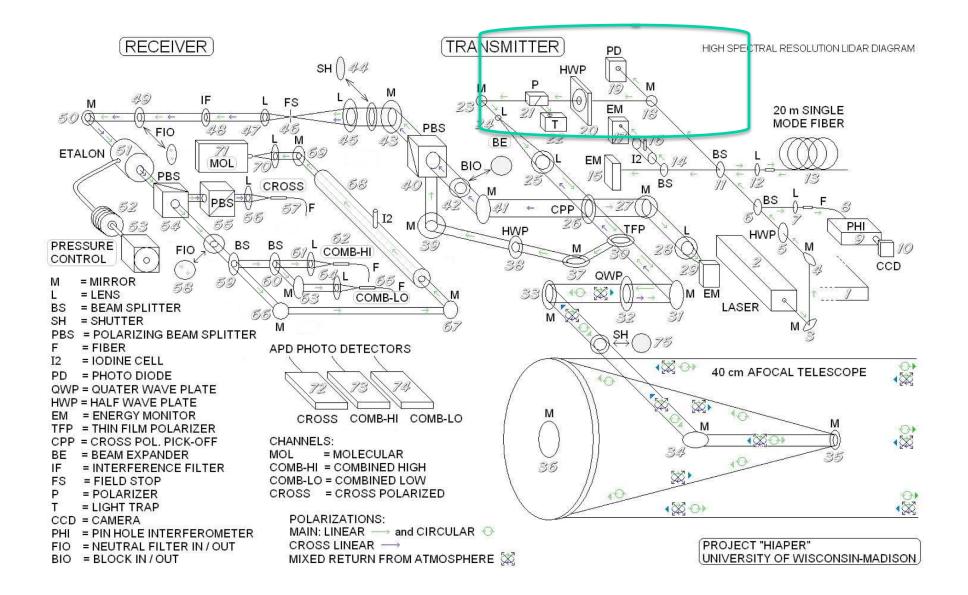


# Example of frequency locking

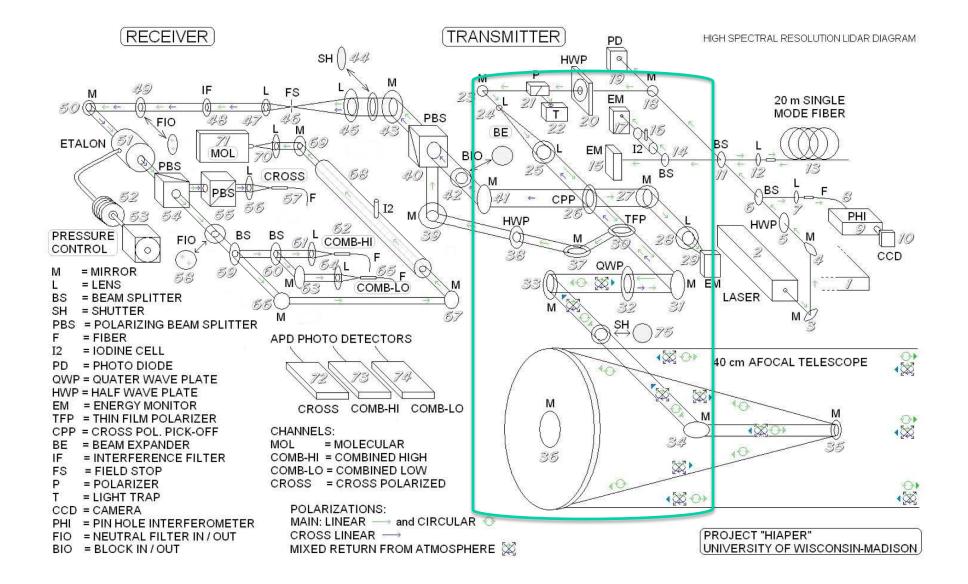




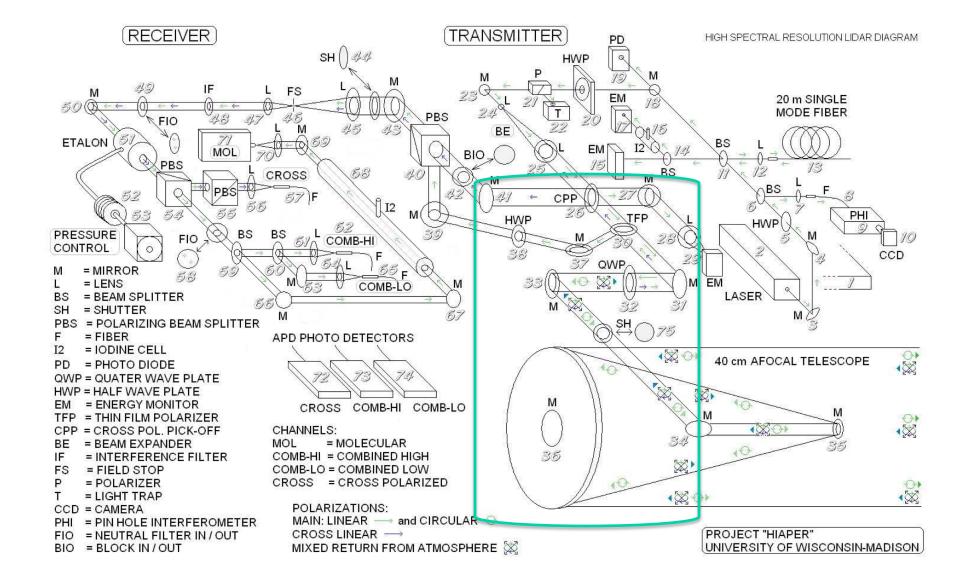
#### Power control and photo diode timing detector

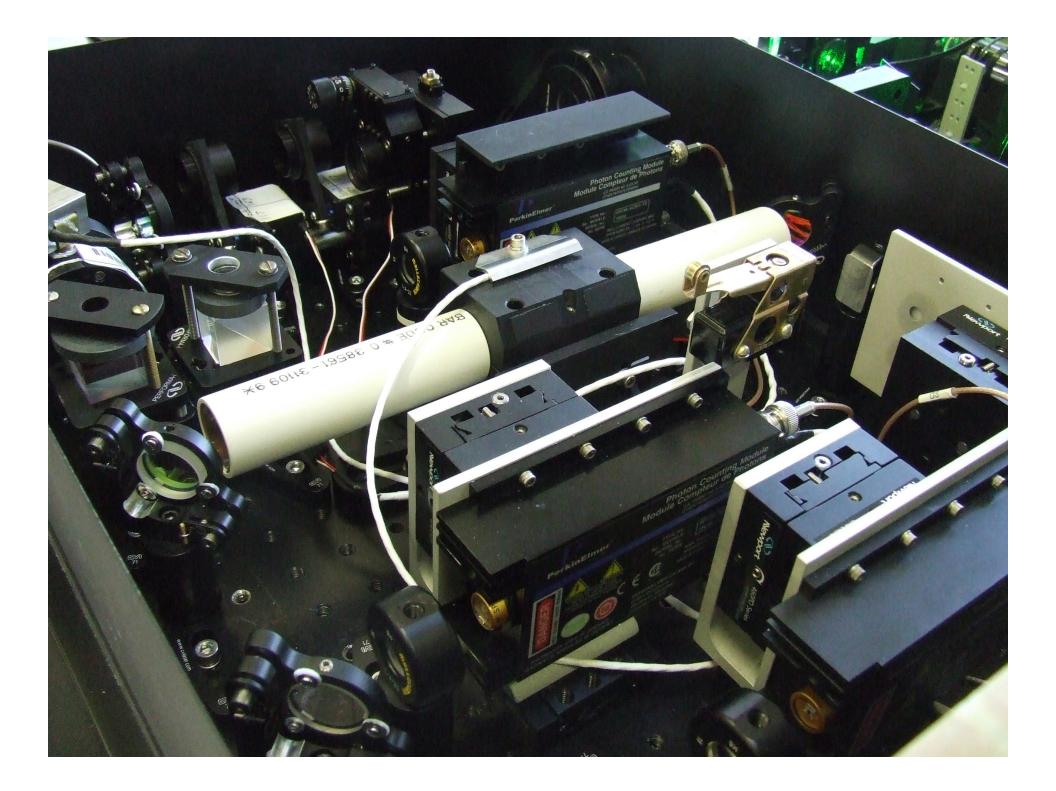


#### Beam expansion and e-monitor



### Received light passes through 1/4 plate, cross-pol and parallel pickups





Specifications Transmitter:	GVHSRL	Langley HSRL
Repetition rate	4000 Hz	200 Hz
Wavelength	532 nm	532 nm
Energy	82 uJ	2.5 mJ
Ave power	339 mW	500 mW
Receiver: Aperture Bandwidth Quantum Eff Field of View Optical trans	40 cm 8 GHz 55% 100 mrad ~34%	40 cm 60 GHz 10% (?) 250-1000 mrad 57%
Signal strength ~ Sky Noise ~		0.27 (Area*Pwr*QE*h) 3.4 (Area*BW*W*QE*h)