

Zenith Angle Reconstructions in AURA Data

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Outline

- 1 Method
- 2 Results
- 3 Conclusion

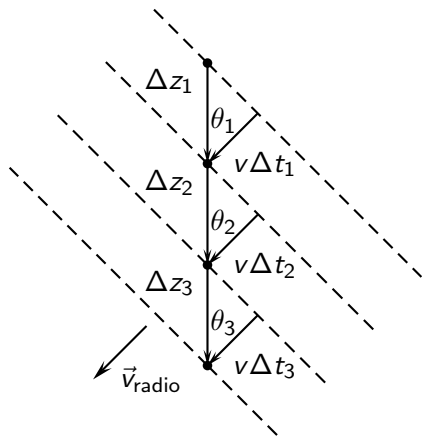
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2 Results

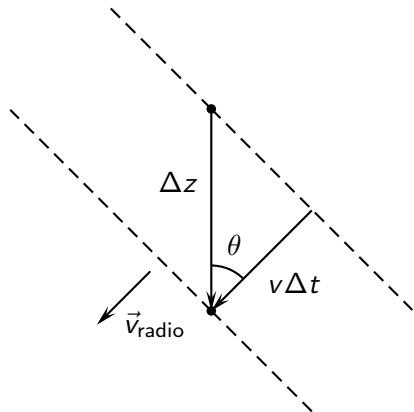
3 Conclusion

Method I



- Use single-cluster data
- Assume plane waves
- Consecutive antennas are Δz apart
- Wavefront travels $v\Delta t$ between hitting consecutive antennas

Method II



- Basic trig gives

$$\theta = \cos^{-1} \left(\frac{v\Delta t}{\Delta z} \right)$$

- Define

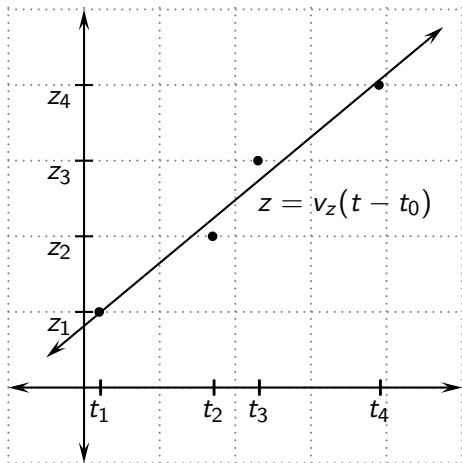
$$v_z := \frac{\Delta z}{\Delta t}$$

so that

$$\theta = \cos^{-1} \left(\frac{v}{v_z} \right)$$

- **Important:** $v_z > v$ and is not a physical speed

Method III



- Require $n_{\text{hit}} \geq 2$
- For $n_{\text{hit}} > 2$, use a least squares fit to get v_z
- If $v_z > v$, compute θ
- Place θ values in 1° bins

Outline

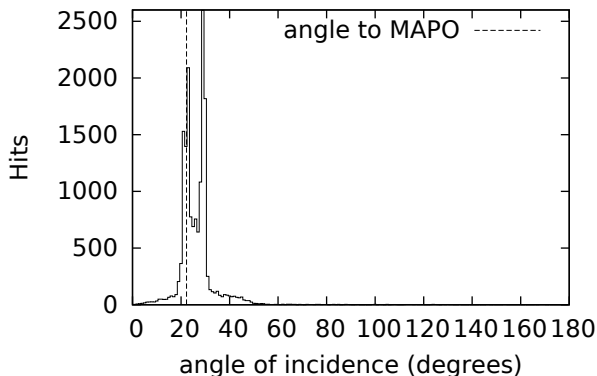
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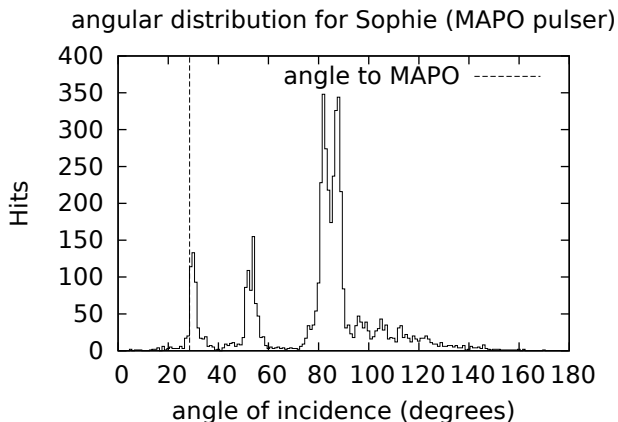
Pulsar data for single clusters I

angular distribution for Danielle (MAPO pulsar)



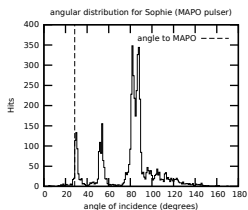
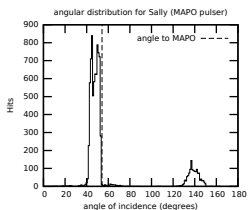
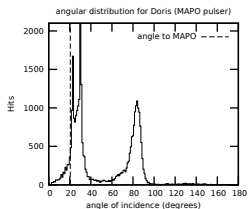
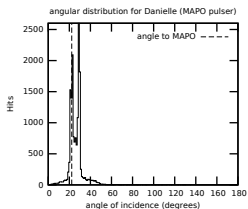
- Why this double peak?
- Feature exists in both individual runs and combined many-run data.

Pulsar data for single clusters II

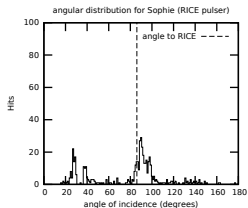
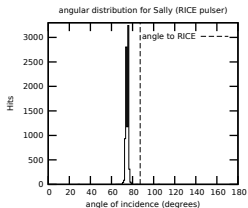
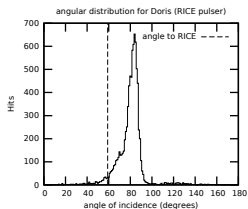
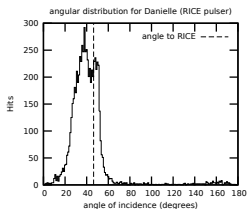


- Why discrete peaks here?
- **Note:** Vertical line is straight-line zenith angle (no ray tracing)

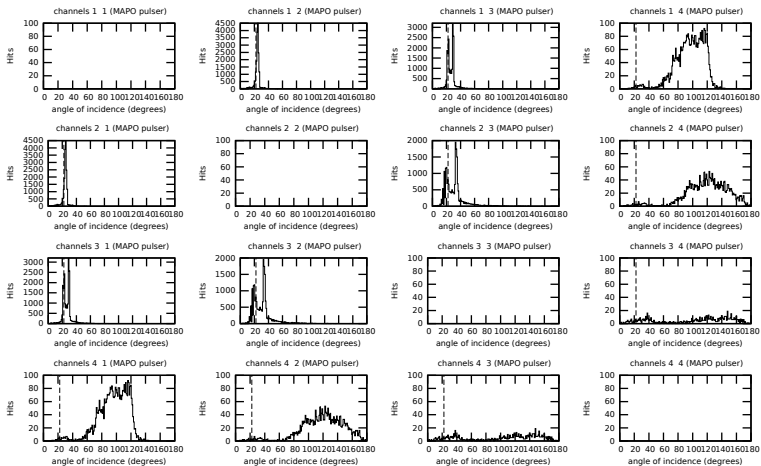
Pulsar data for all clusters — MAPO surface pulsar



Pulsar data for all clusters — RICE in-ice pulsar



MAPO surface pulser data per antenna pair in Danielle

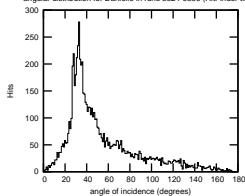


■ Low frequency channels may skew some combinations

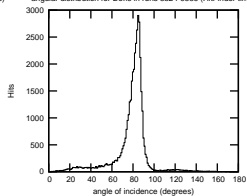
▶ Other clusters & plots for RICE in-ice pulser

Regular data for single clusters

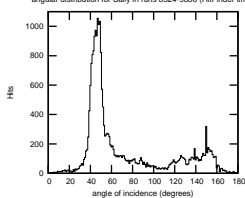
angular distribution for Danielle in runs 8524-9886 (HitFinder time)



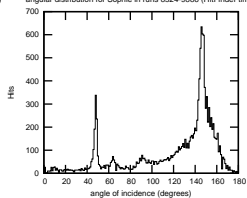
angular distribution for Doris in runs 8524-9886 (HitFinder time)



angular distribution for Sally in runs 8524-9886 (HitFinder time)



angular distribution for Sophie in runs 8524-9886 (HitFinder time)



More plots available online

- Also tried using the first 5σ time and the last 5σ time
- For normal runs, also tried 2D angle \times time histograms
- Additional plots are available online:
 - ... /[~mrichman/angular-distributions](#)
 - ... /[~mrichman/angular-distributions-first-5-sigma-times](#)
 - ... /[~mrichman/angular-distributions-last-5-sigma-times](#)
 - ... /[~mrichman/pulser-ang-dists-by-cluster](#)
 - ... /[~mrichman/pulser-ang-dists-by-pulser](#)
 - ... /[~mrichman/pulser-pairwise-ang-dists-by-pulser](#)

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Caveats

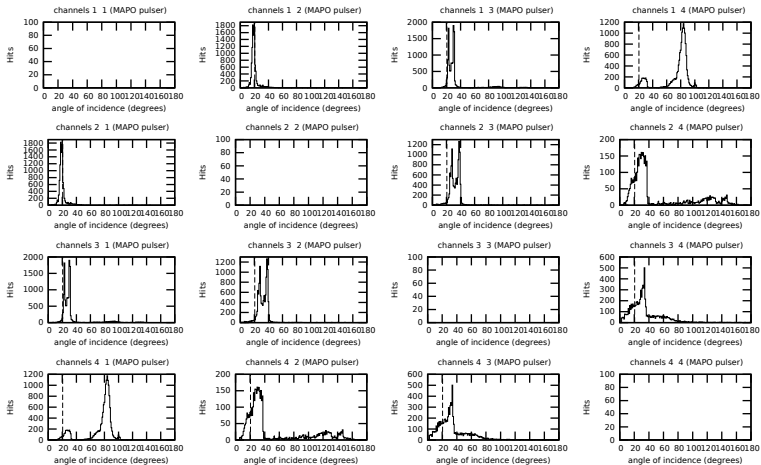
- Haven't filtered events with wrong periodicity from pulsar hits
- Haven't accounted for varying $n(z)$ (is this needed?)

Next steps

- Address those caveats
- Try plotting Δt distributions directly
- Check for hardware crosstalk
- Try scatterplots of zenith angle, antenna pair vs. antenna pair
- More stuff TBD

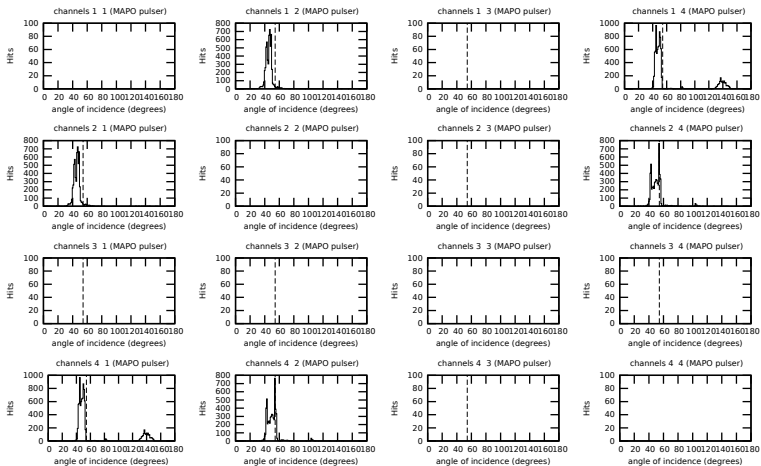
Backup Slides

MAPO surface pulser pulser data per antenna pair in Doris



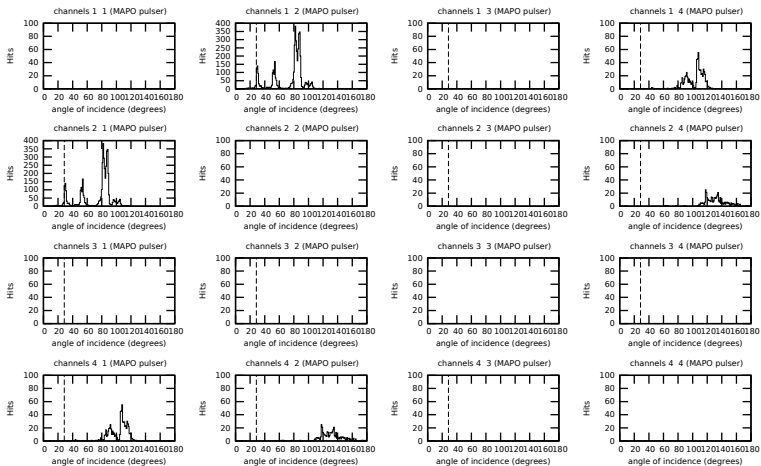
◀ Back to Danielle's MAPO surface pulser data

MAPO surface pulser pulser data per antenna pair in Sally



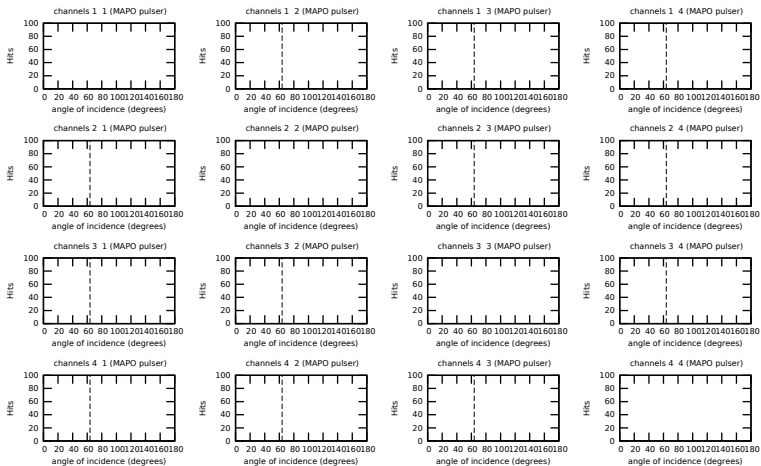
◀ Back to Danielle's MAPO surface pulser data

MAPO surface pulser data per antenna pair in Sophie



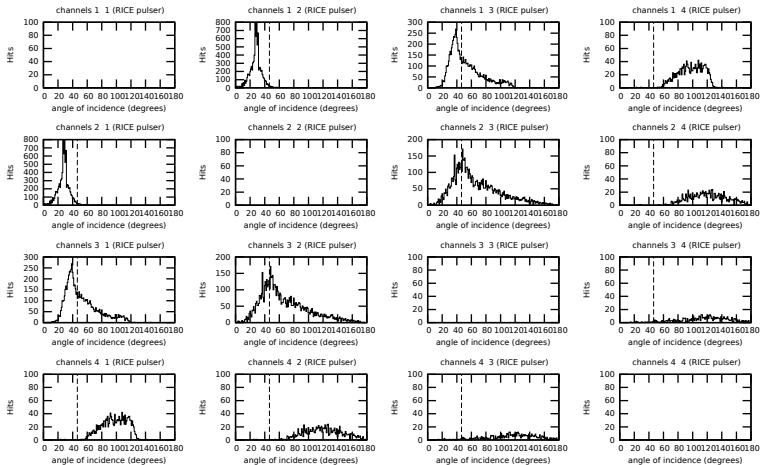
[← Back to Danielle's MAPO surface pulser data](#)

MAPO surface pulser pulser data per antenna pair in Susan



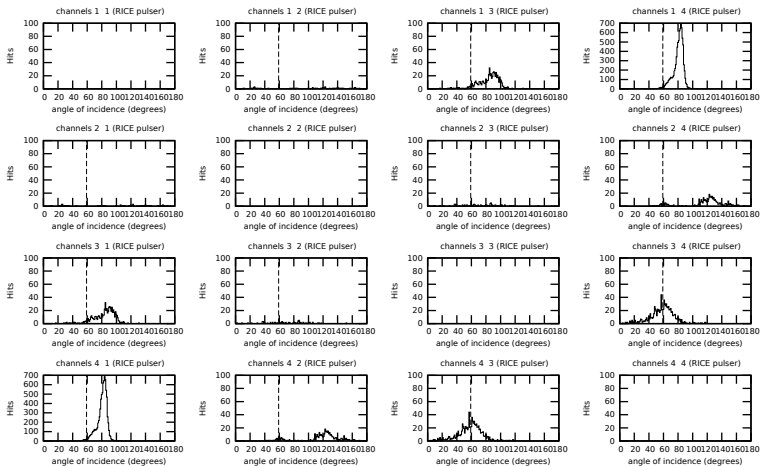
◀ Back to Danielle's MAPO surface pulser data

RICE in-ice pulser data per antenna pair in Danielle



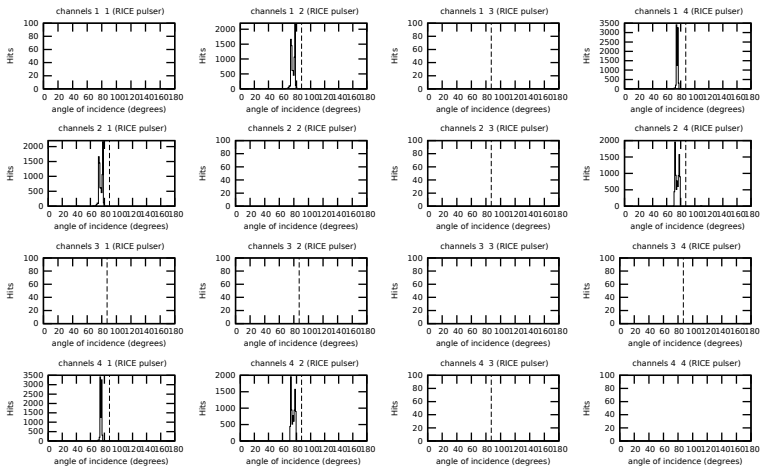
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RICE in-ice pulser data per antenna pair in Doris



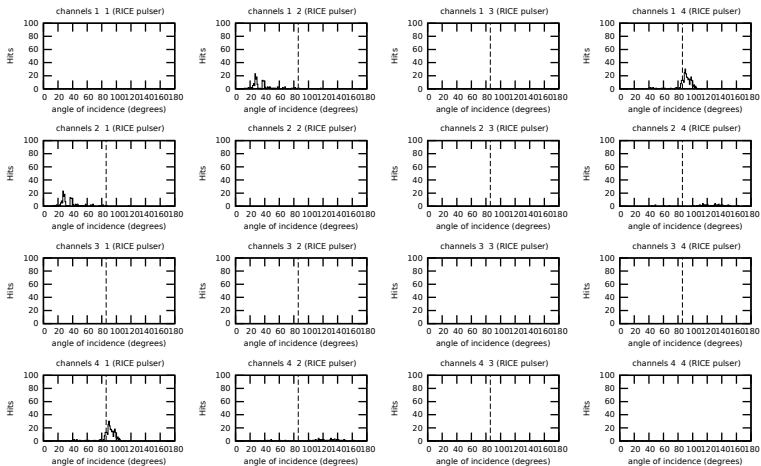
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RICE in-ice pulser data per antenna pair in Sally



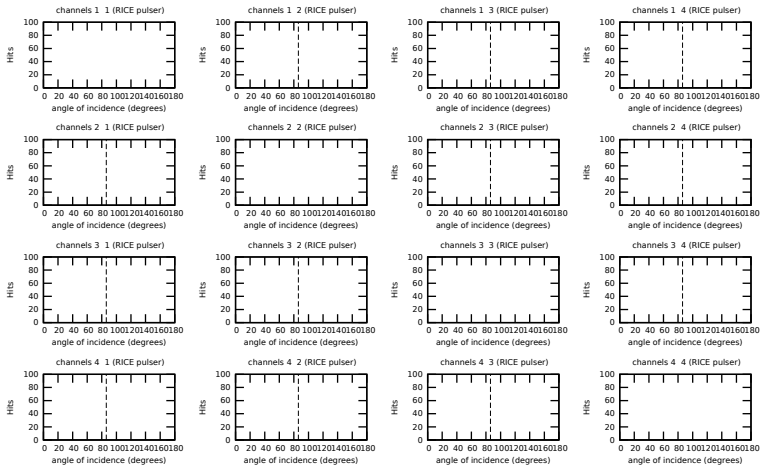
[← Back to Danielle's MAPO surface pulser data](#)

RICE in-ice pulser data per antenna pair in Sophie



◀ Back to Danielle's MAPO surface pulser data

RICE in-ice pulser data per antenna pair in Susan



◀ Back to Danielle's MAPO surface pulser data