



Intro to IceCube Simulation

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IceProd/SimProd Workshop

Madison, WI USA

June 14, 2024

IceCube Laboratory

Data from every sensor is collected here and sent by satellite to the IceCube data warehouse at UW-Madison

1450 m

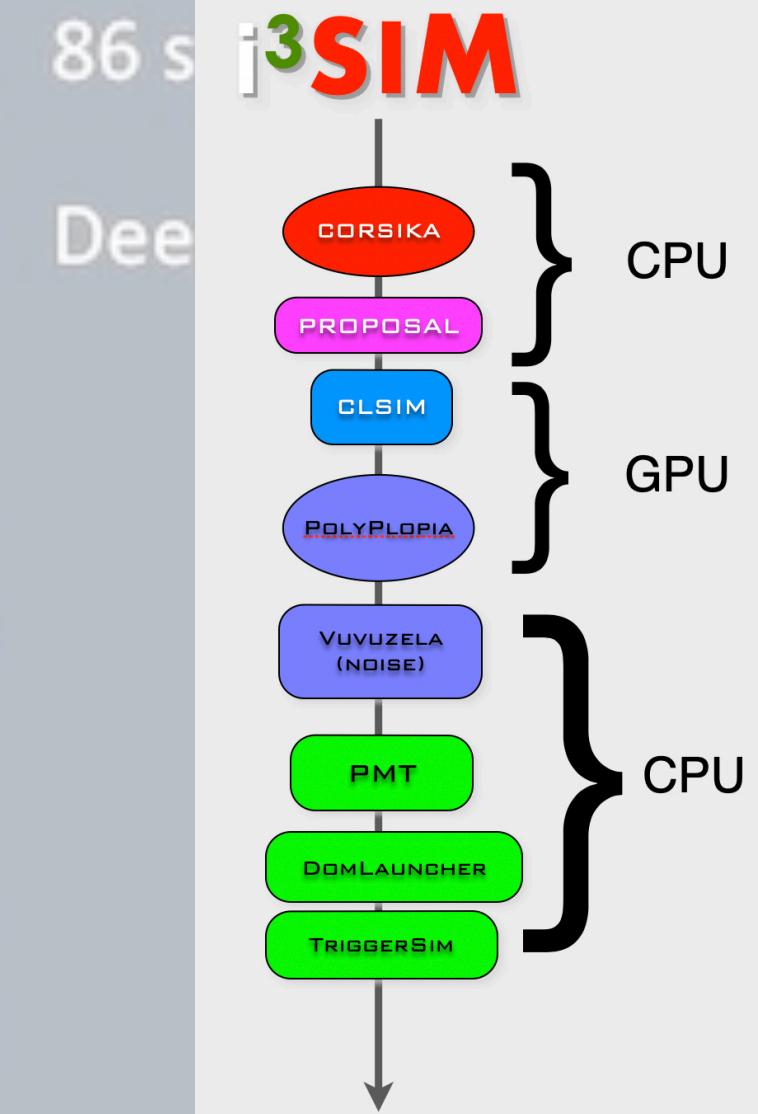
Digital Optical
Module (DOM)
5,160 DOMs deployed
in the ice

50 m

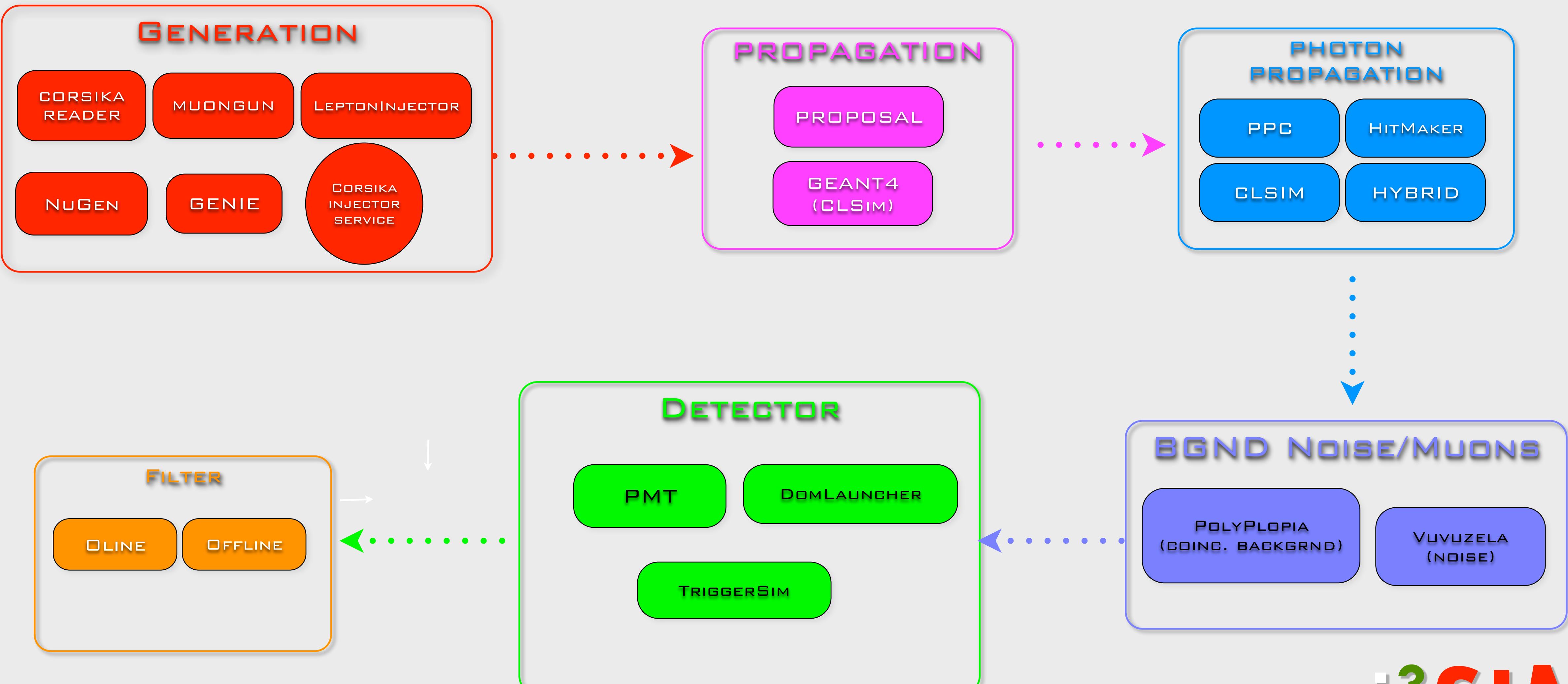
IceCube

1

Amundsen–Scott South Pole Station, Antarctica
A National Science Foundation-managed research facility



Simulation Chain



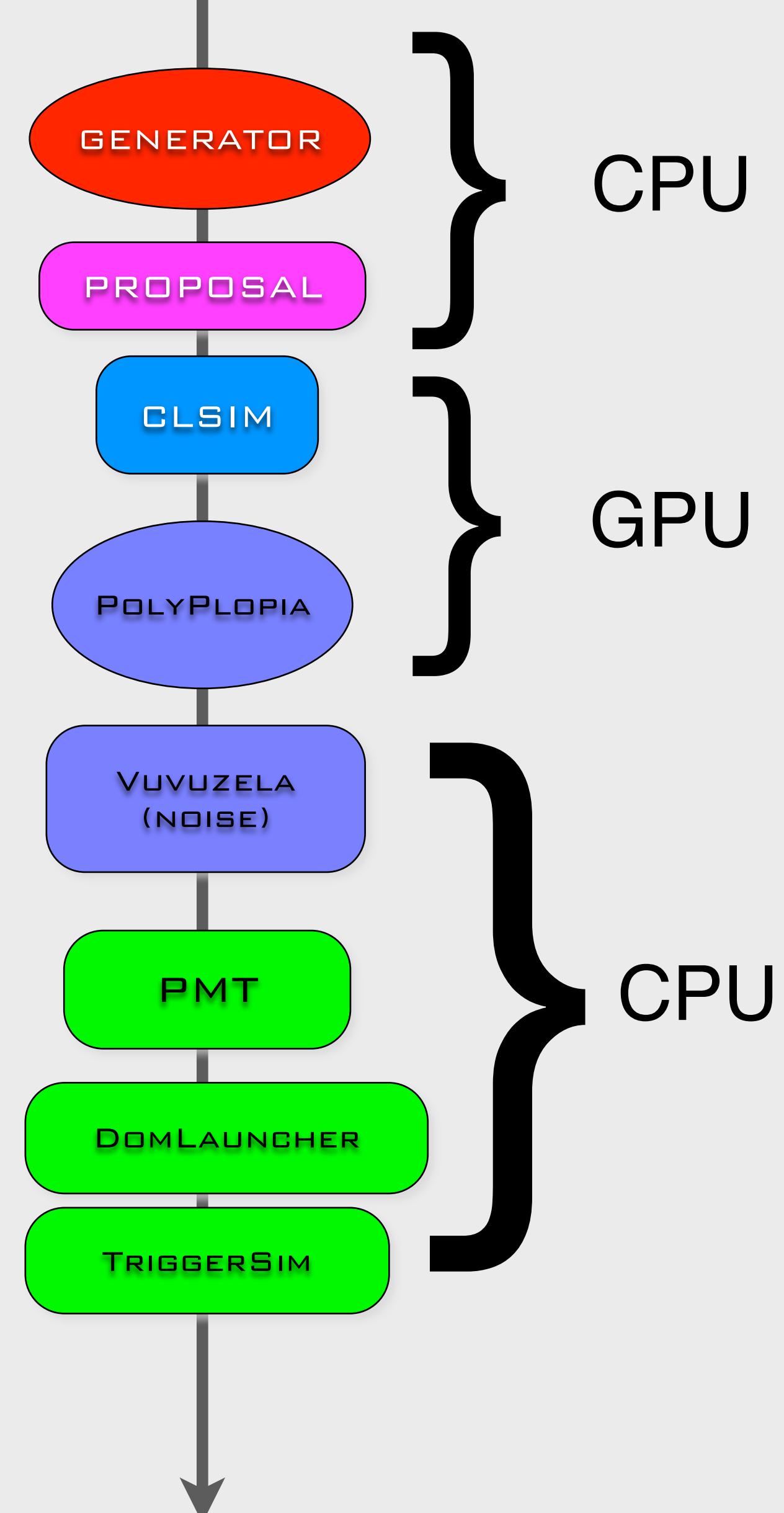
i³SIM

The Simulation Shish Kabob

(Computing Resource Optimization)

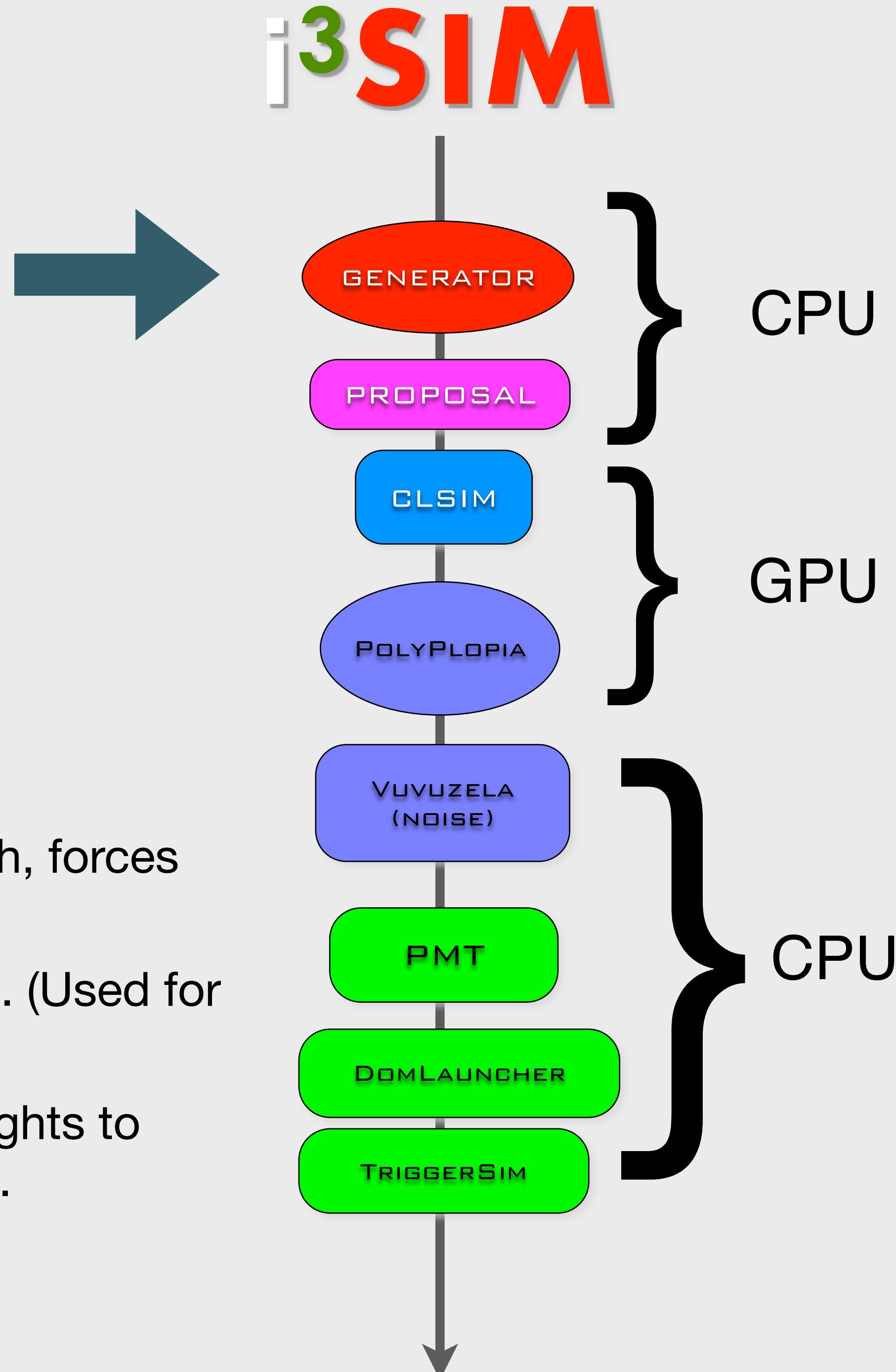
i3SIM

- Optimizing the shish kabob:
 - Different parts of the simulation chain have different resource requirements.
 - CORSIKA is CPU-intensive and requires little RAM
 - Photon propagation run almost exclusively on GPUs
 - Detector simulation is CPU bound and requires more memory.
- Things to keep in mind:
 - Running whole chain on a GPU node will waste GPU resources and limit throughput.
 - Intermediate storage:
 - breaking up chain requires transferring/storing intermediate files.
 - Reduce complexity in workflow



Generators

- ▶ Cosmic-ray Air Showers:
 - ▶ **Triggered-CORSIKA** (parallel corsika client/server w clsim)
 - ▶ **corsika-reader**: IceTray reader for standard format (deprecated)
 - ▶ **CorsikaInjectorService** (IceTop)
- ▶ Muons:
 - ▶ **MuonGun**: parametrization of flux of atm. muons under the ice.
- ▶ Neutrinos:
 - ▶ **neutrino-generator**: injects neutrinos, propagates them through Earth, forces interaction in detector volume.
 - ▶ **genie-reader**: detailed simulation of neutrino interactions with GENIE. (Used for low-energy simulations)
 - ▶ **LeptonInjector / NuFSGen** (not yet available): weighted leptons+weights to account for flux models, interaction models, in-earth propagation, etc.

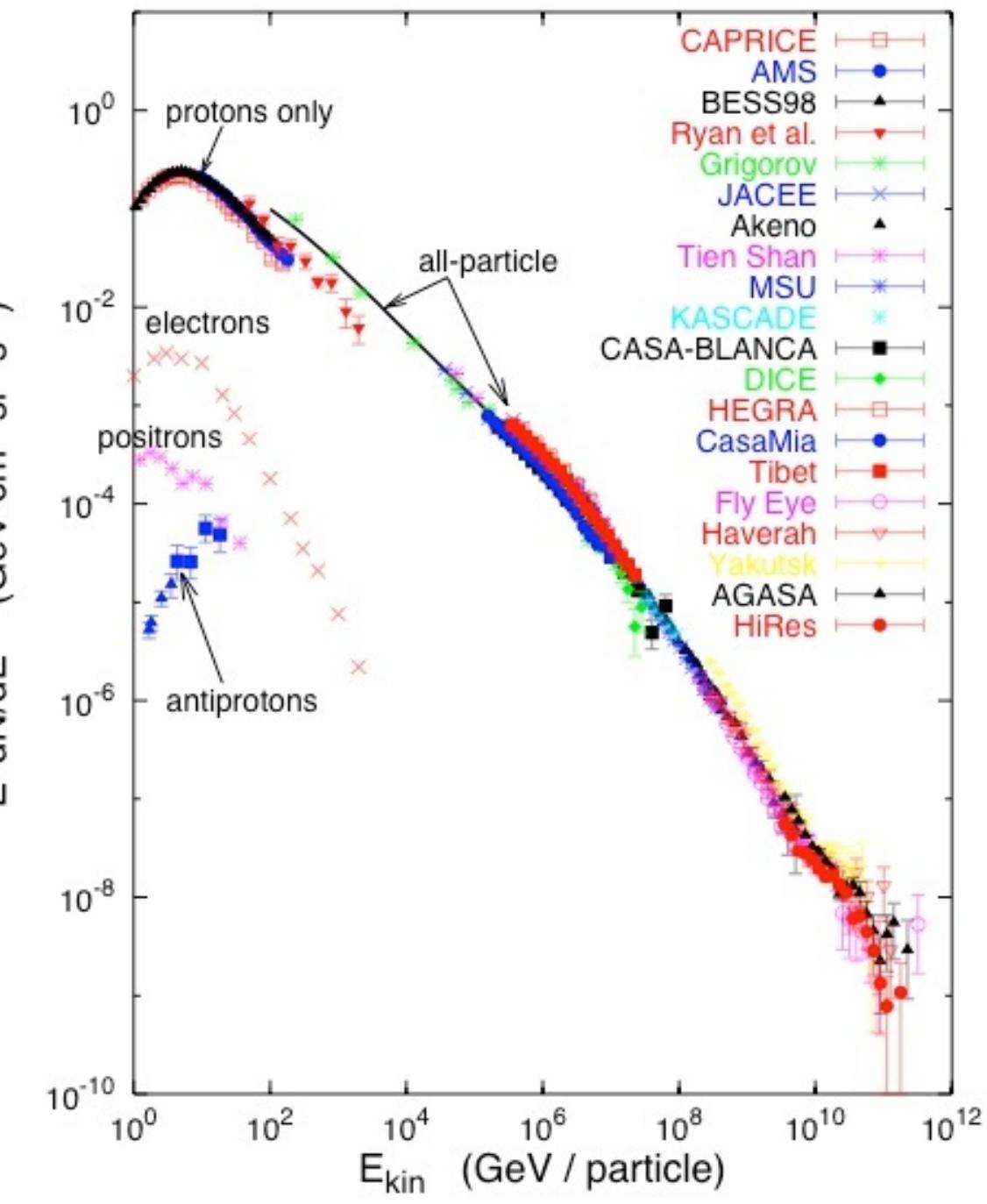


Generators : CORSIKA

(COsmic Ray SImulations for KAscade)

- Particles are tracked through the atmosphere until they undergo reactions with the air nuclei or - in the case of unstable secondaries - decay.
- The hadronic interactions at high energies may be described by several reaction models alternatively:
 - *VENUS*, *QGSJET*, and *DPMJET* (Gribov-Regge theory),
 - *SIBYLL* (minijet model).
 - *neXus*, *EPOS* (combination of *QGSJET* and *VENUS*).
 - *HDPM* (Dual Parton Model).
- Hadronic interactions at lower energies:
 - *GHEISHA*, *FLUKA* , or *UrQMD* models.
- For electromagnetic interactions
 - Tailored version of *EGS4*.
 - Analytical *NKG* formulas.

► weighted events : artificial spectrum $E^{-\gamma}$
 ► 5 representative mass groups: (H,He,CNO,Mg,Fe)



FORSCHUNGSZENTRUM KARLSRUHE
Technik und Umwelt

Extensive Air Shower Simulation
with CORSIKA:
A User's Guide
(Version 5.61 from April 21, 1998)

J. Knapp¹ and D. Heck
Institut für Kernphysik

Forschungszentrum Karlsruhe GmbH, Karlsruhe

¹Institut für Experimentelle Kernphysik, Universität Karlsruhe, D-76021 Karlsruhe, Germany

MuonGun (IceCube implementation of MUPAGE)

arXiv:0907.5563v1 [astro-ph.IM] 31 Jul 2009

- Parametrized muon distributions from many, many CORSIKA showers.
- Produce muons in detector volume by sampling from splinned PDF tables
 - zenith angle
 - depth
 - energy
 - bundle multiplicity
 - radius

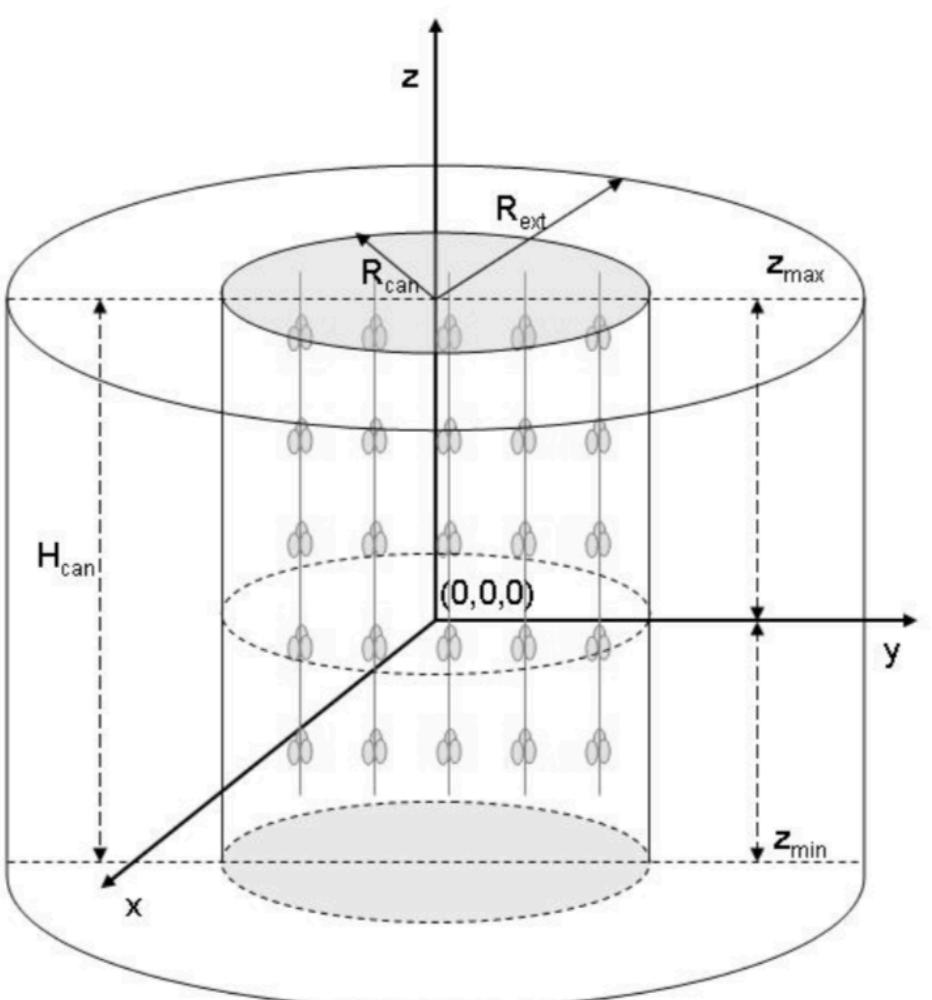
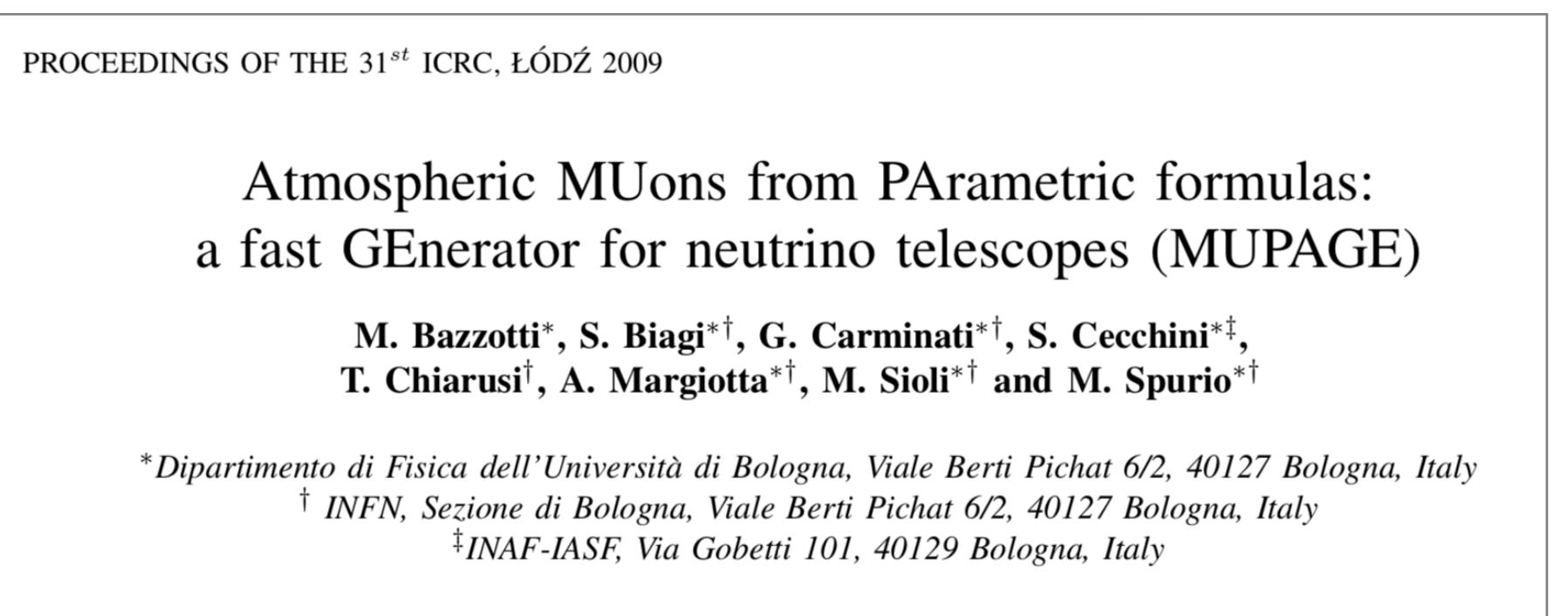
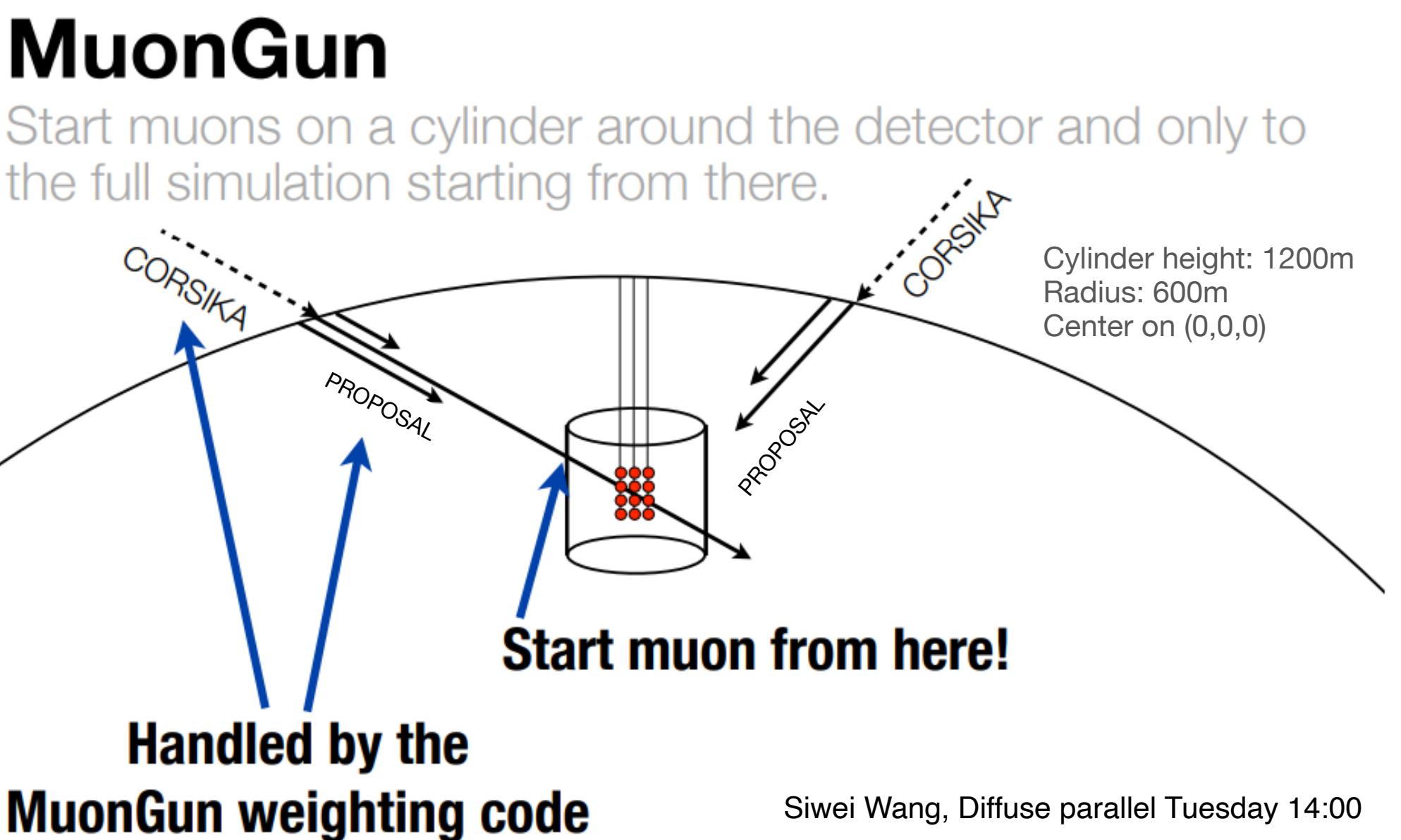
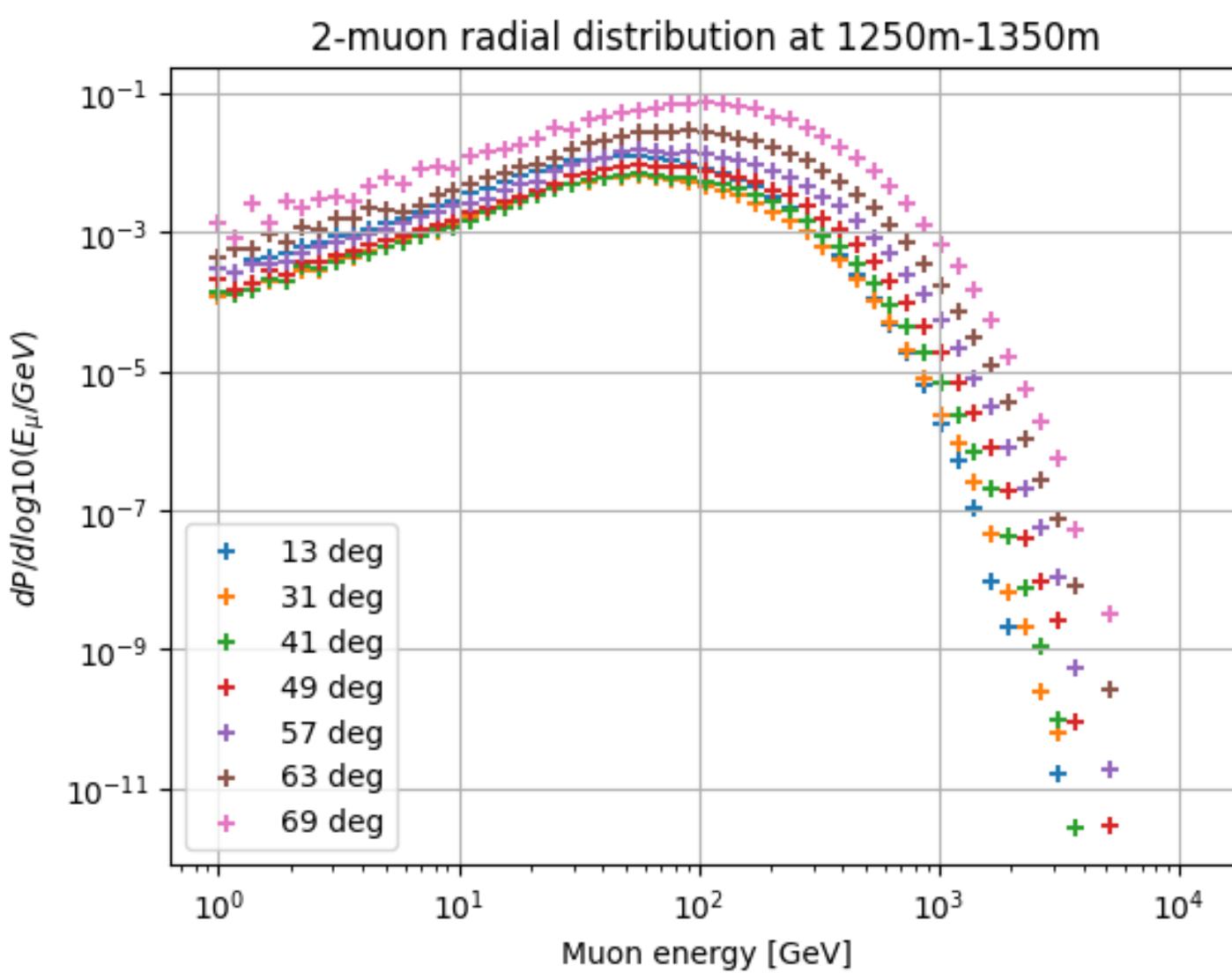
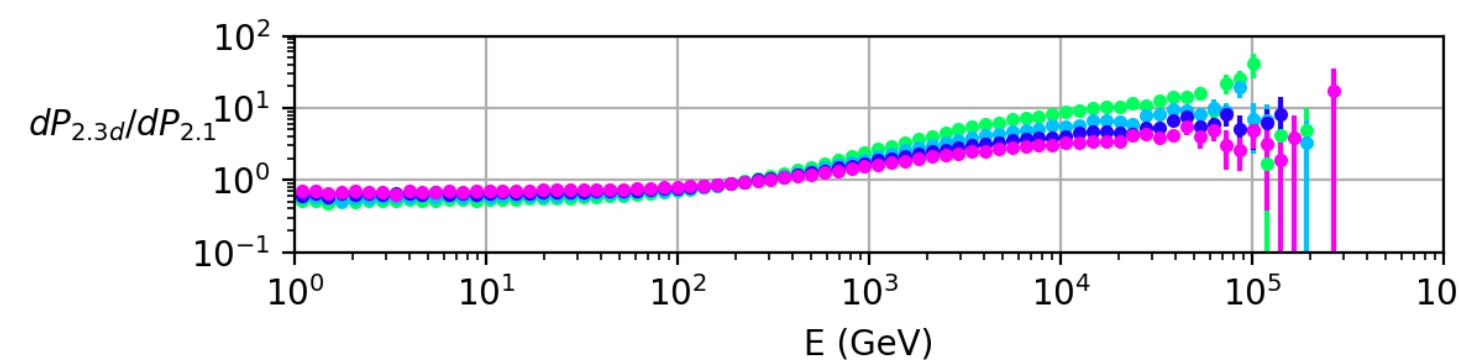
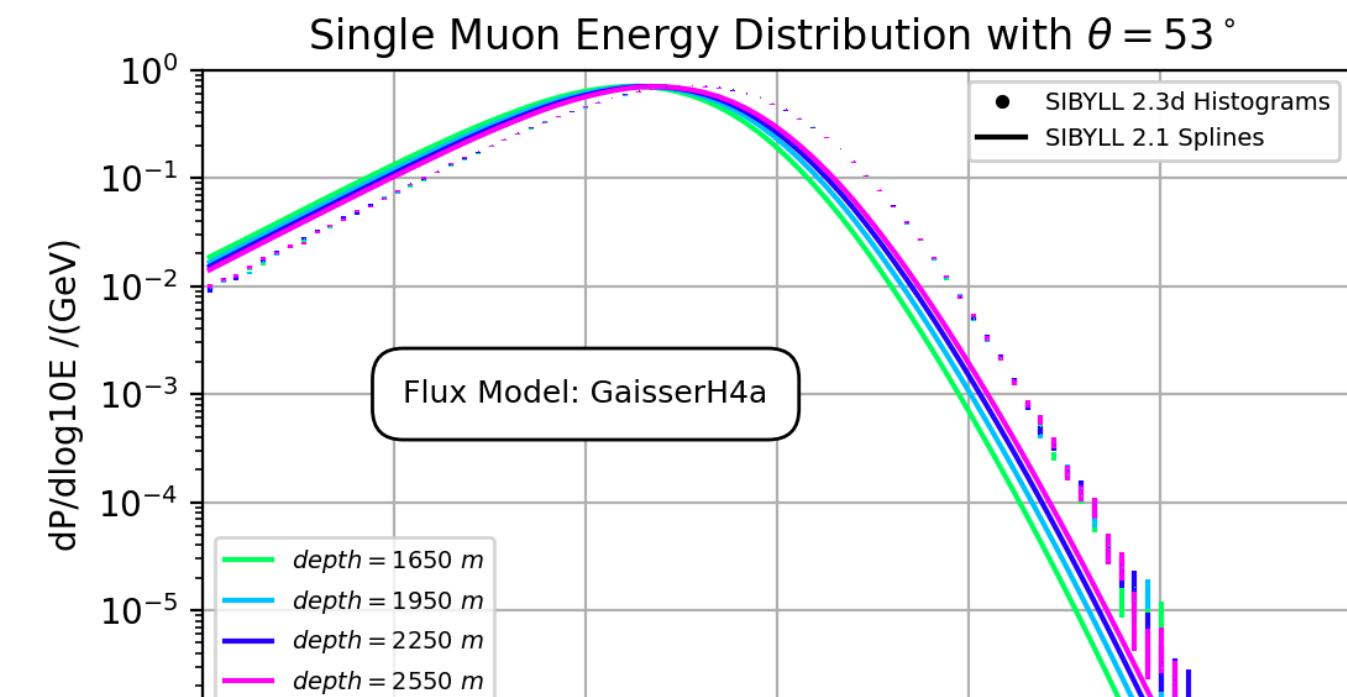


Fig. 1: Sketch of some input parameters. The cylinder surrounding the instrumental volume is the *can*, with radius R_{can} and height H_{can} . The events are generated on an extended can with R_{ext} . The origin of the coordinate system does not have to be located at the center of the detector. The lower disk is at a depth H_{max} with respect to the sea/ice surface.

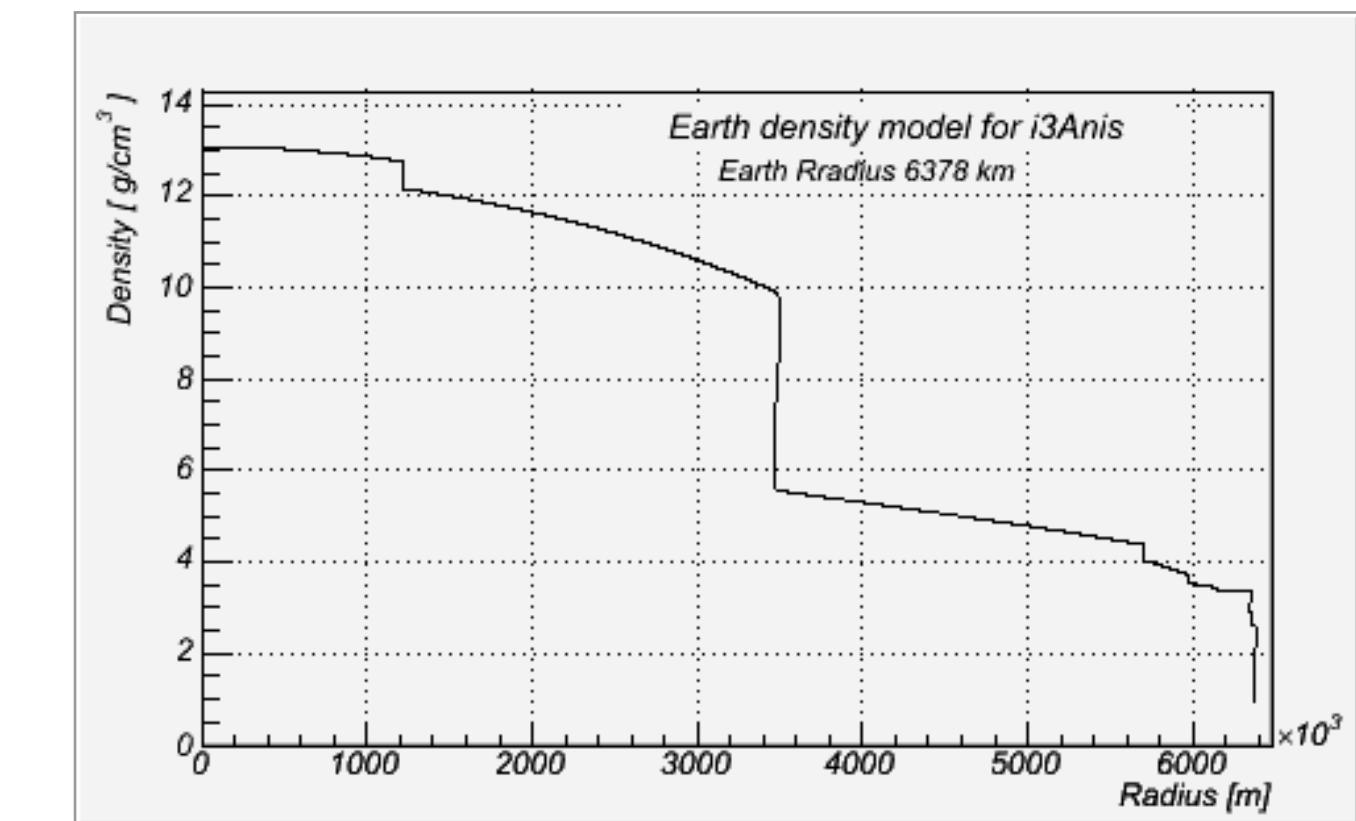
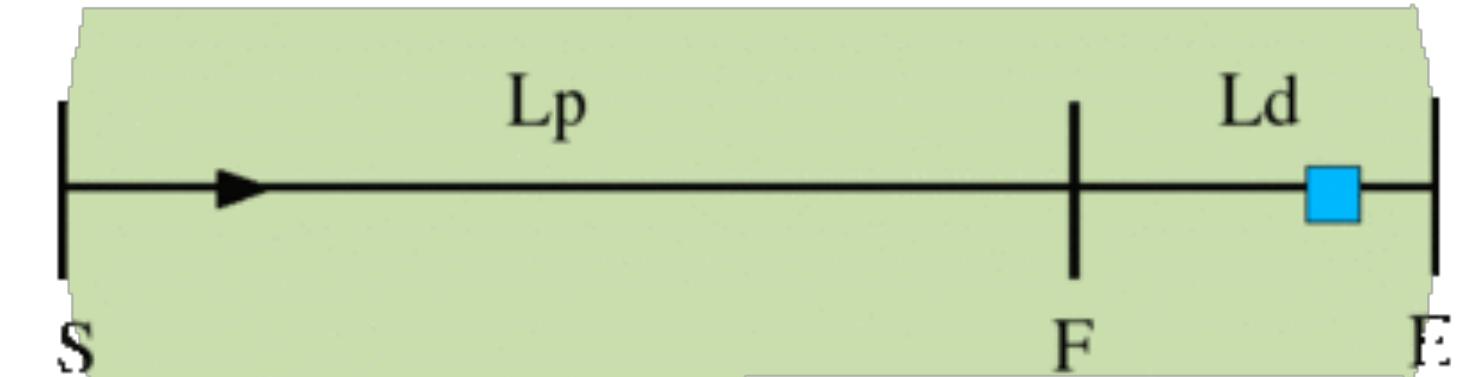


Siwei Wang, Diffuse parallel Tuesday 14:00



neutrino-generator

1. Calculate total path length inside the Earth using injected neutrino geometry.
 - a. Separate the total path length into propagation area (SF) and
 - b. detection volume (FE).
2. Define a step length $dx[m]$ using propagation area and step number.
3. For each step:
 - a. Calculate a column depth and Earth's density at the step point.
 - b. Calculate a total cross section at the step point.
 - c. Calculate a probability that the injected neutrino interacts within the step. Try Monte-Carlo, and decide whether an interaction happened within the step.
 - d. If interaction occurred: choose interaction randomly.
 - i. If CC-interaction is selected with injection particle NuMu or NuE, break (event is killed).
 - ii. else, generate secondaries and continue to next step.
 - e. If nothing happens, continue next step.
4. Finish propagation when injected neutrino + secondaries reach surface of detection volume (point F), then process a weighted interaction.



- produce a $E^{-\gamma}$ ν_μ , ν_e , ν_τ with
 - ▶ PRELIM Earth's density model
- - CC nu + N
 - CC nubar + N
 - NC nu + N
 - NC nubar + N
- parton distribution functions
- prop & interaction of neutrinos into a weight

LeptonInjector/LeptonWeighter

- LeptonInjector and LeptonWeighter are designed for large-volume Cherenkov neutrino telescopes such as IceCube.
- The neutrino event generator allows for quick and flexible simulation of neutrino events within and around the detector volume
- Implements the leading Standard Model neutrino interaction processes relevant for neutrino observatories:
 - neutrino-nucleon deep-inelastic scattering
 - neutrino-electron annihilation.

This is publicly available code.

<https://arxiv.org/abs/2012.10449>

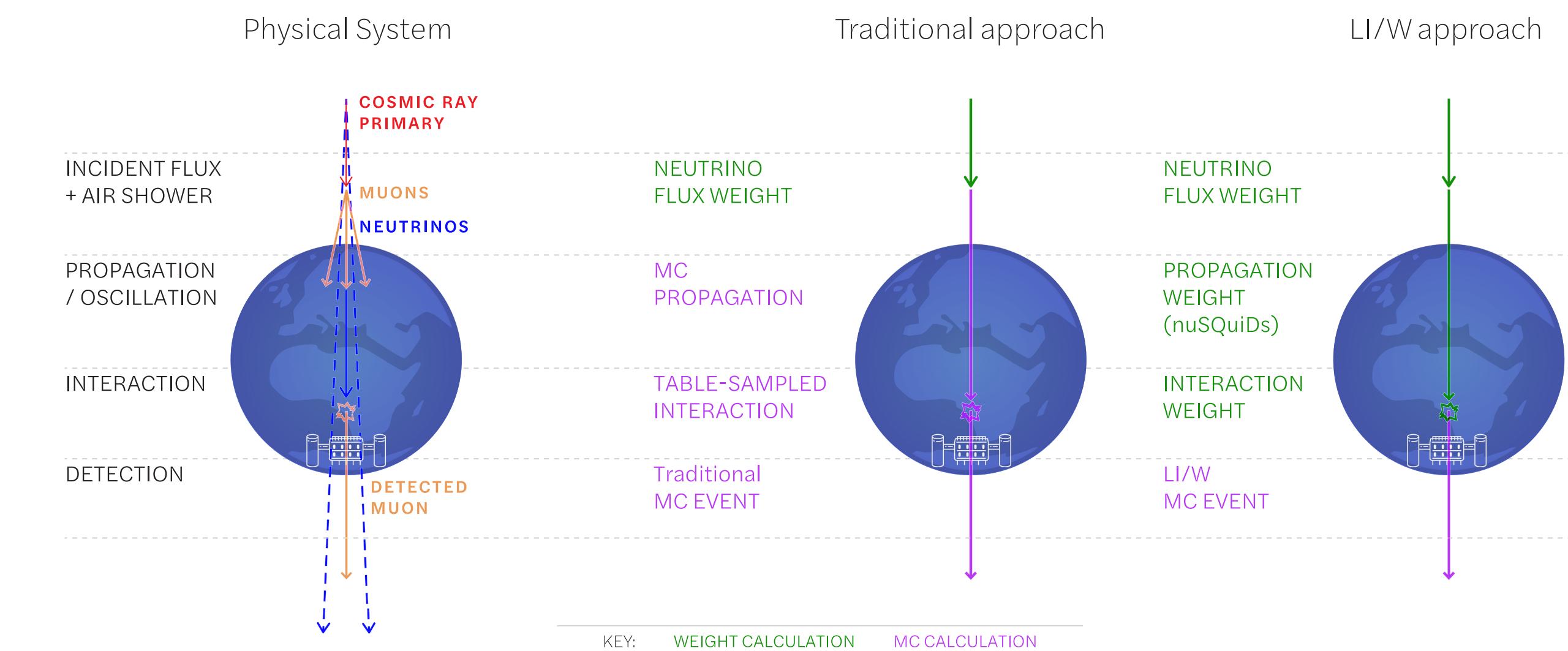


Figure 1.1: A diagram illustrating the different event generation and weighting steps for traditional methods compared with the LeptonInjector and LeptonWeighter philosophy.

Lepton propagation

- ▶ PROPOSAL: parametrized interactions with the medium.

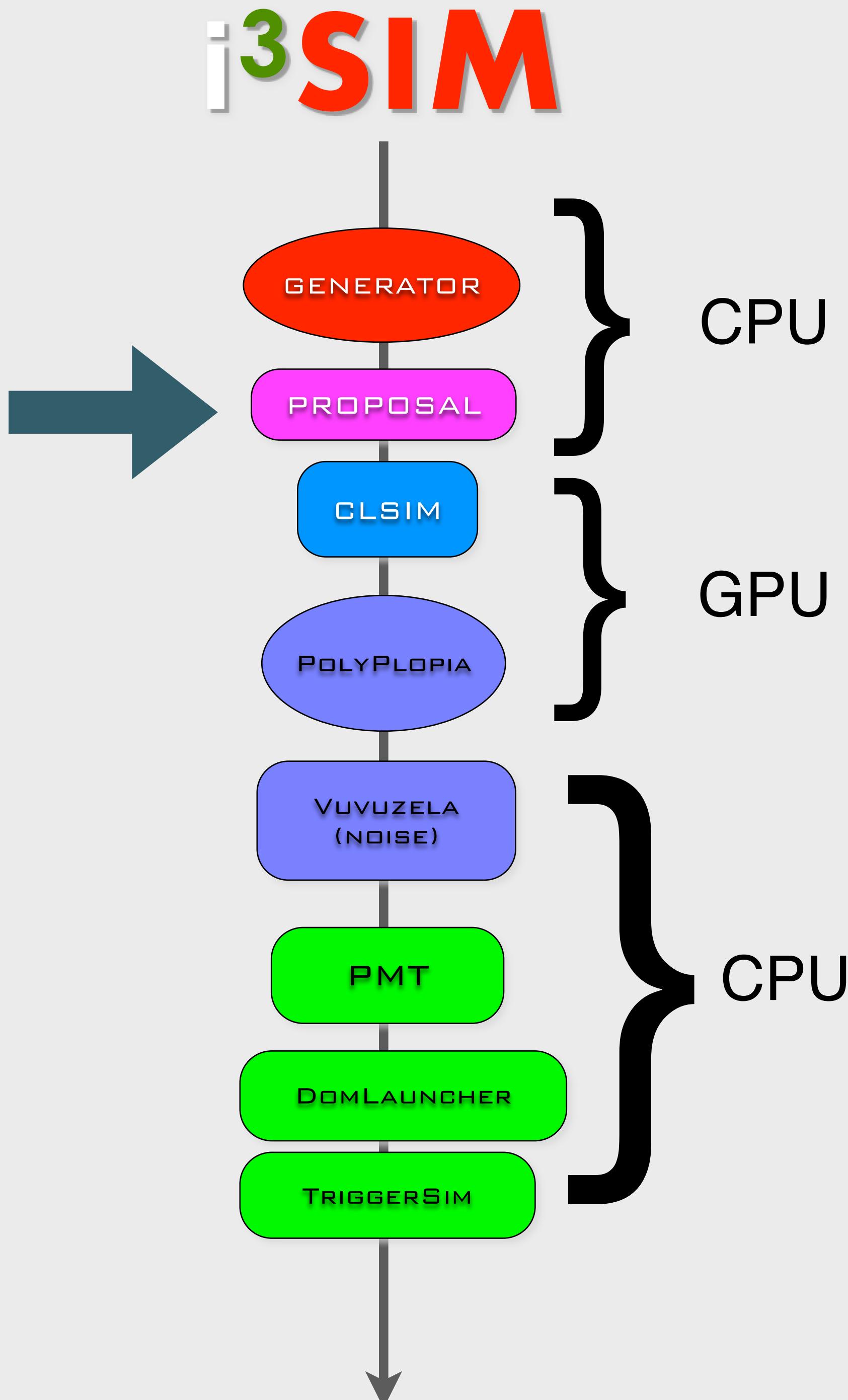
Comp. Phys. Com. 184, 9 (2013), p2070-2090

- ▶ Stochastic energy losses include:
 - ▶ ionization
 - ▶ electron-pair production
 - ▶ bremsstrahlung
 - ▶ photo-nuclear interaction
 - ▶ decay

- ▶ GEANT4: Detailed particle propagation in media.

<https://geant4.web.cern.ch/>

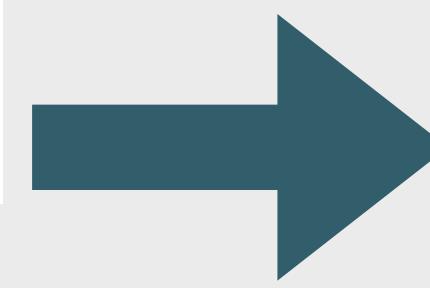
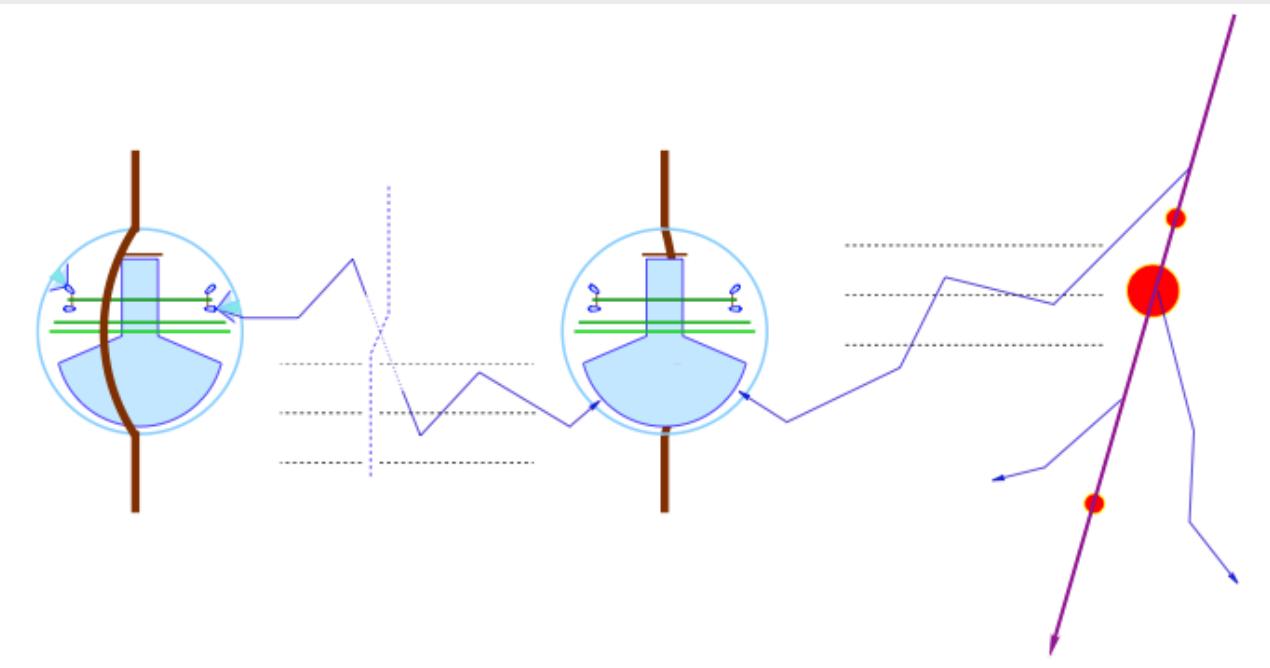
- ▶ 3rd-party G4 library used by CLSim to propagate leptons for low-energy simulations (CPU-intensive).



Photon Propagation (PPC, CLSim)

i3SIM

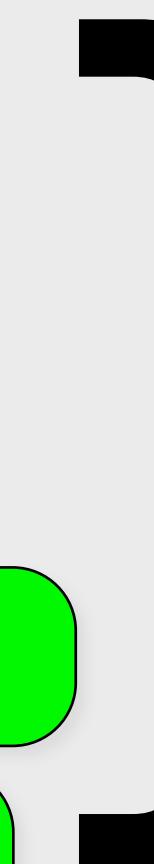
- μ energy loss + cascades \rightarrow photons \rightarrow p.e.
- Photon propagation : ice properties + PMT response + DOM glass/gel
 - Direct photon tracking
 - CLSim, PPC
 - Hybrid photon tracking
 - CLSim + Splined PDF lookup of amplitude and time distribution



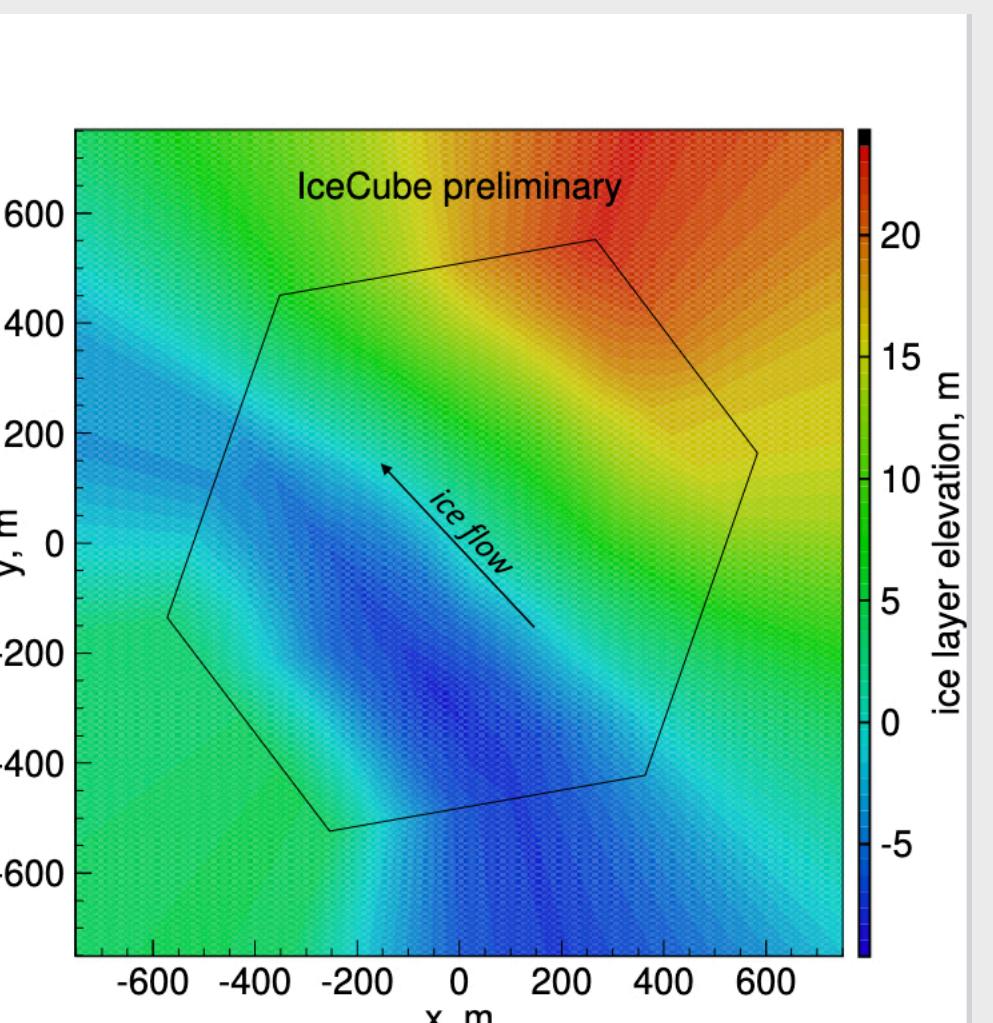
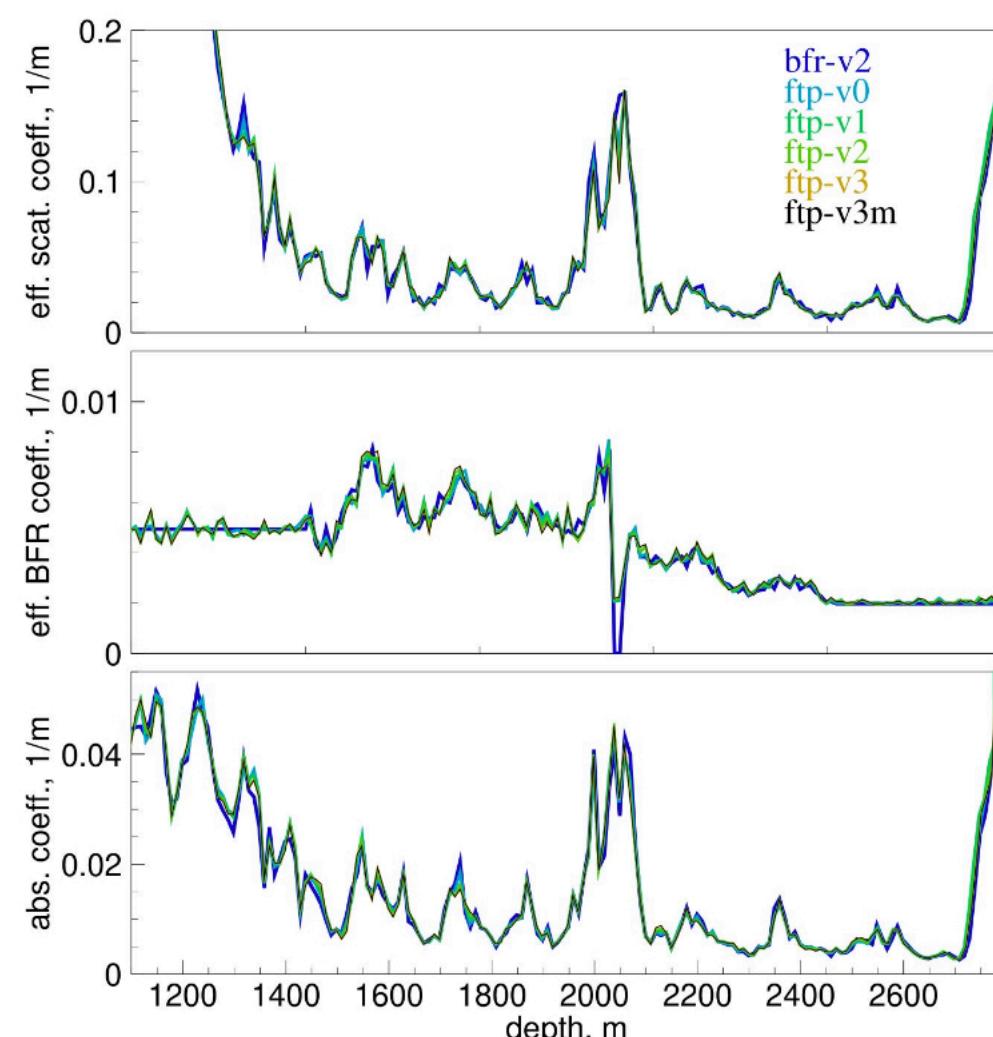
CPU

GPU

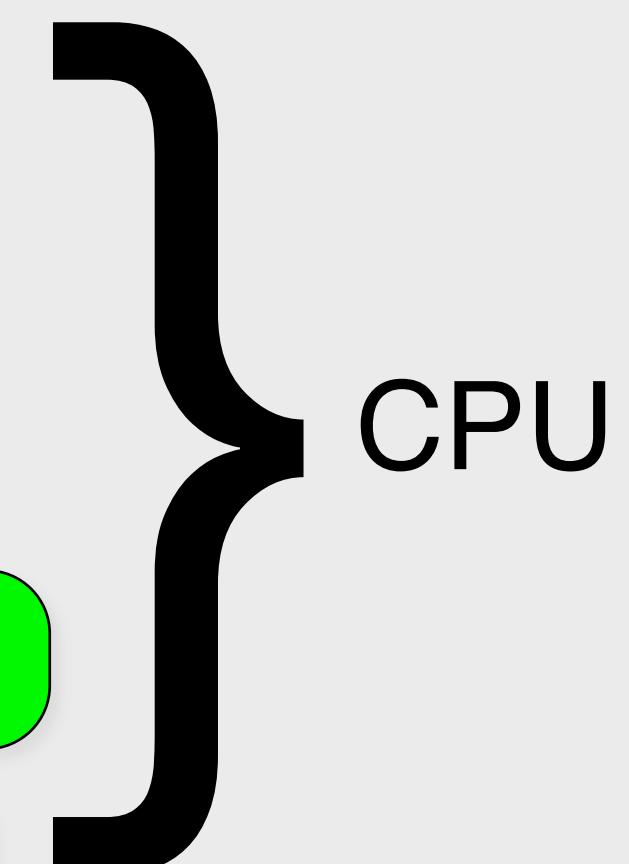
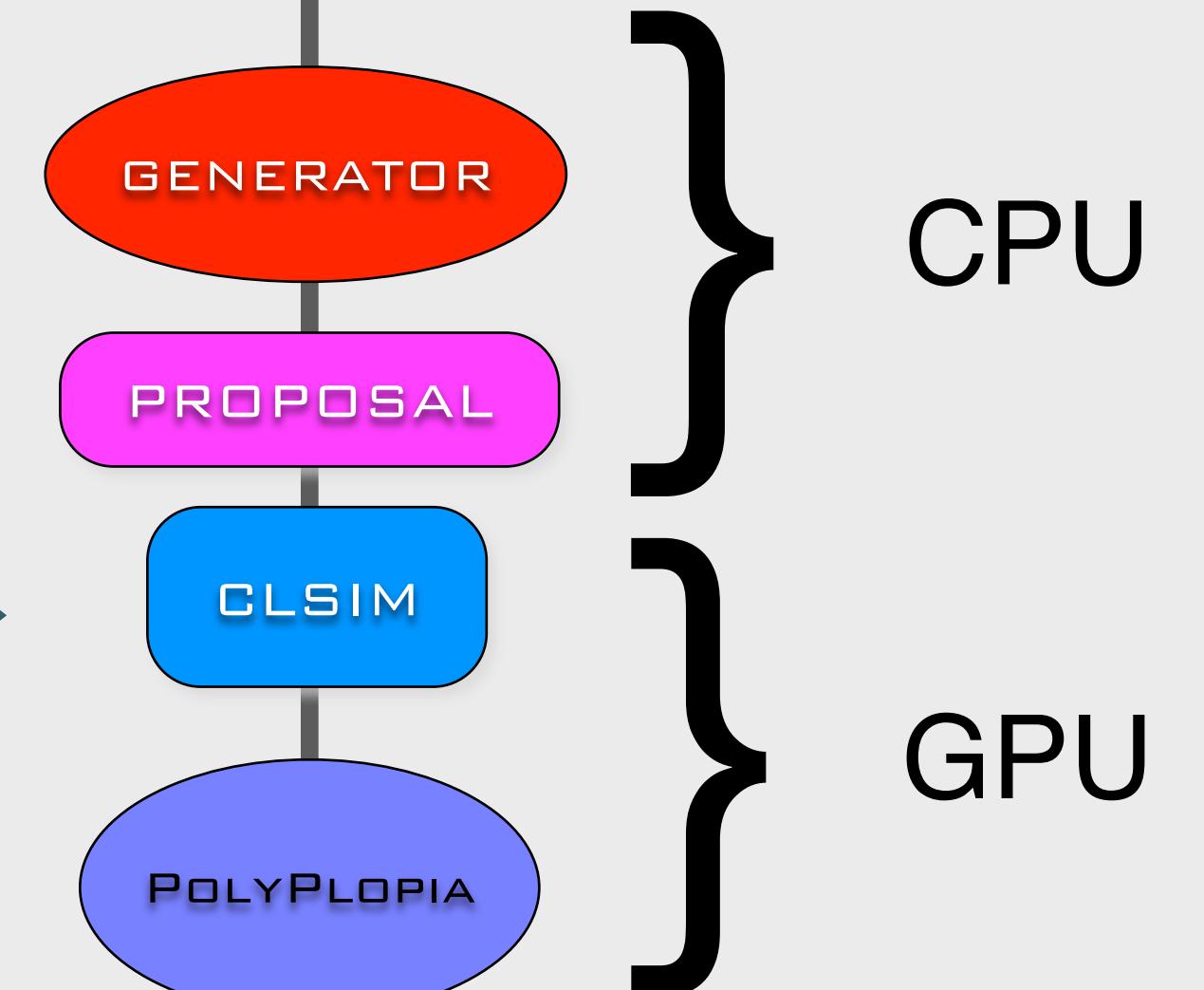
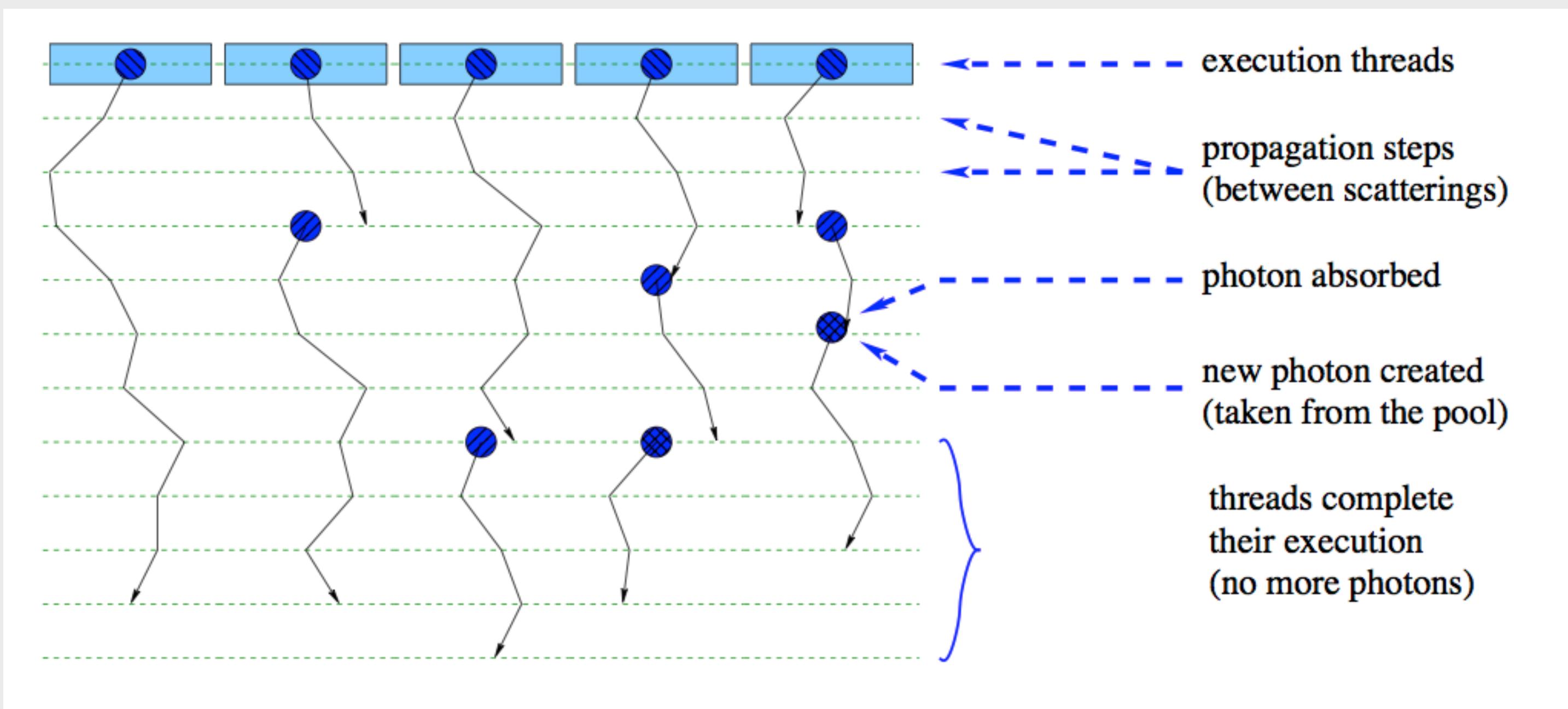
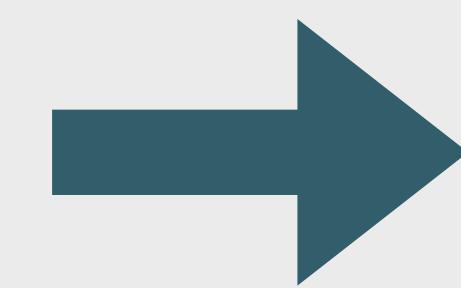
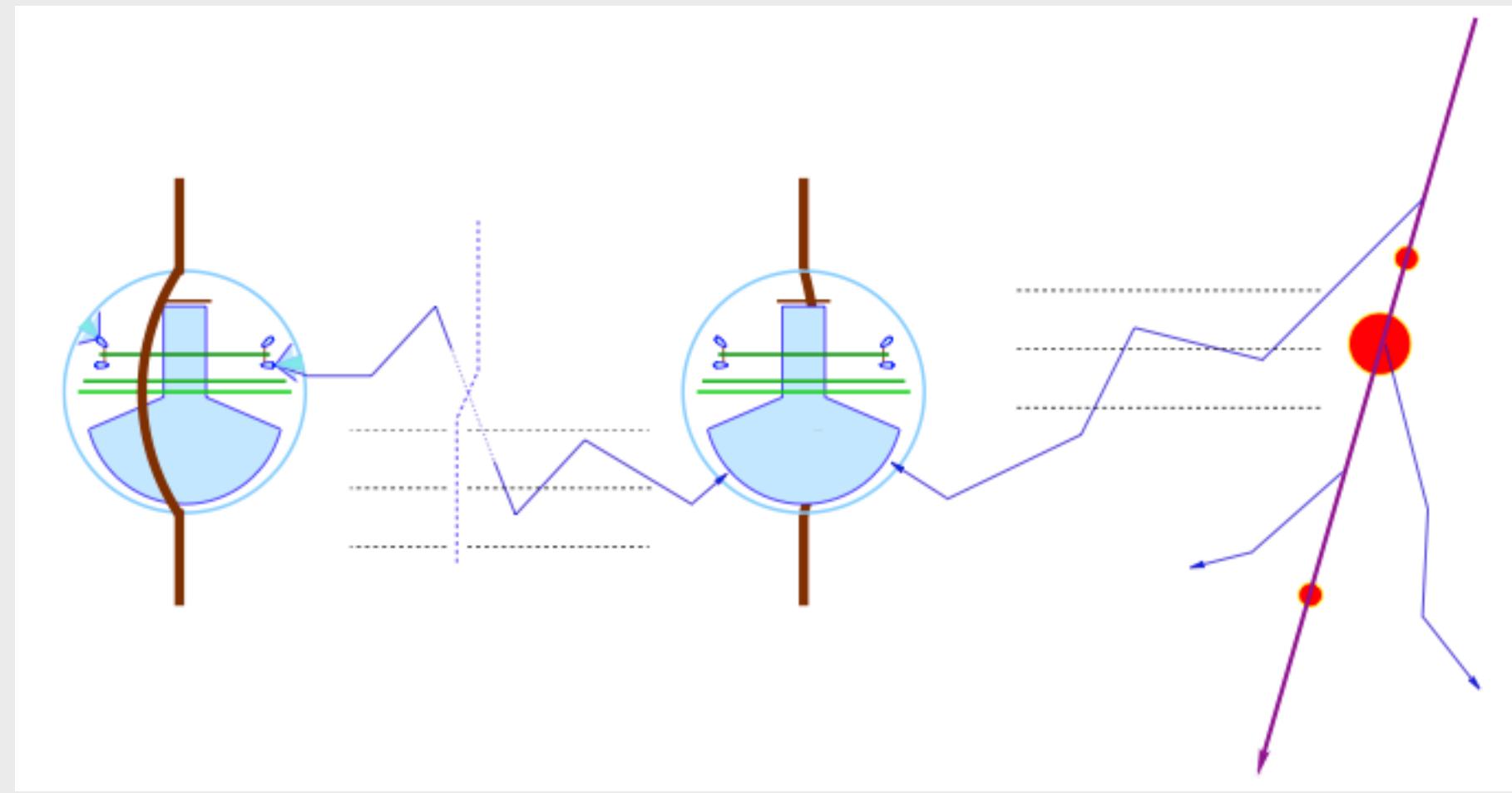
CPU



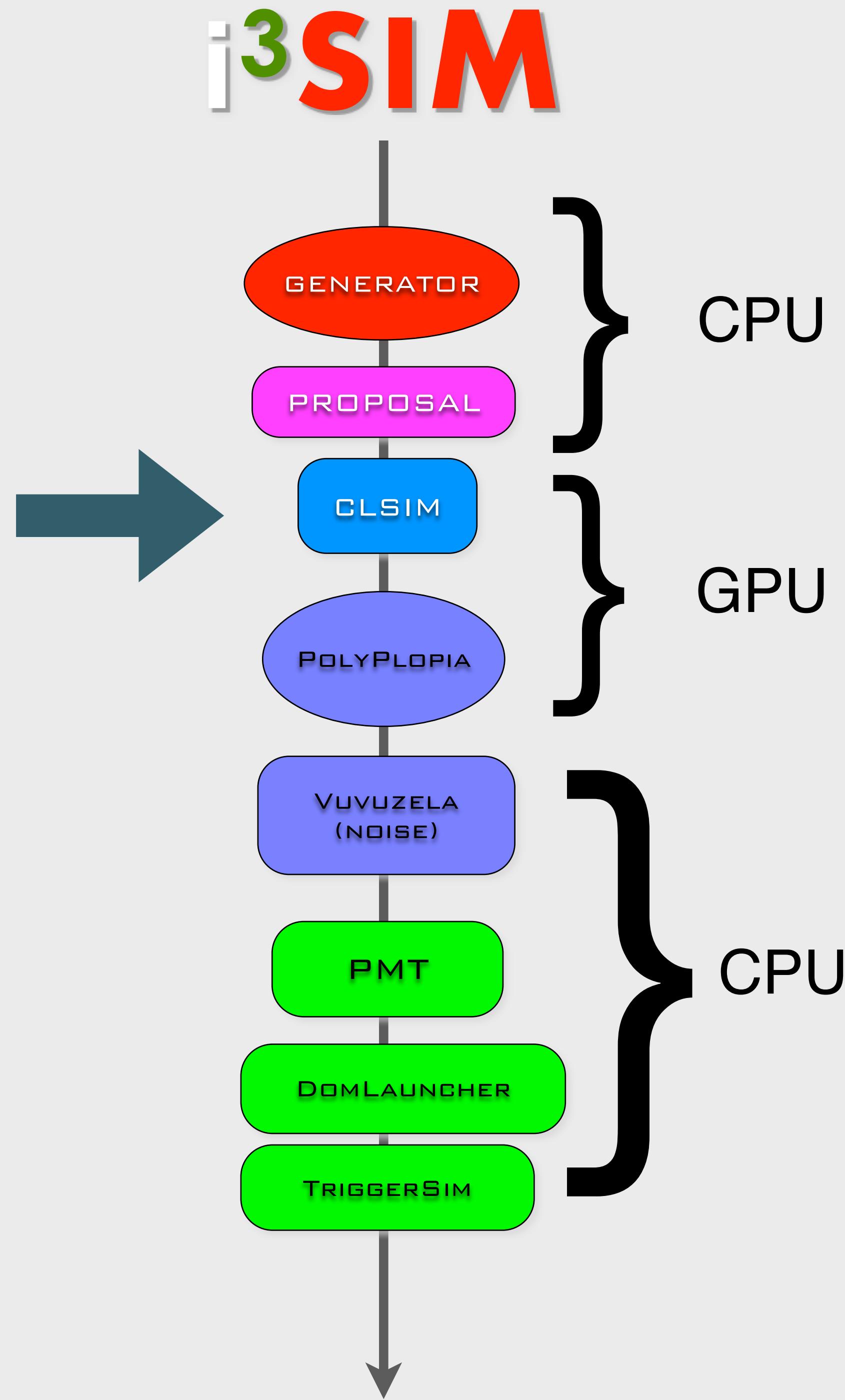
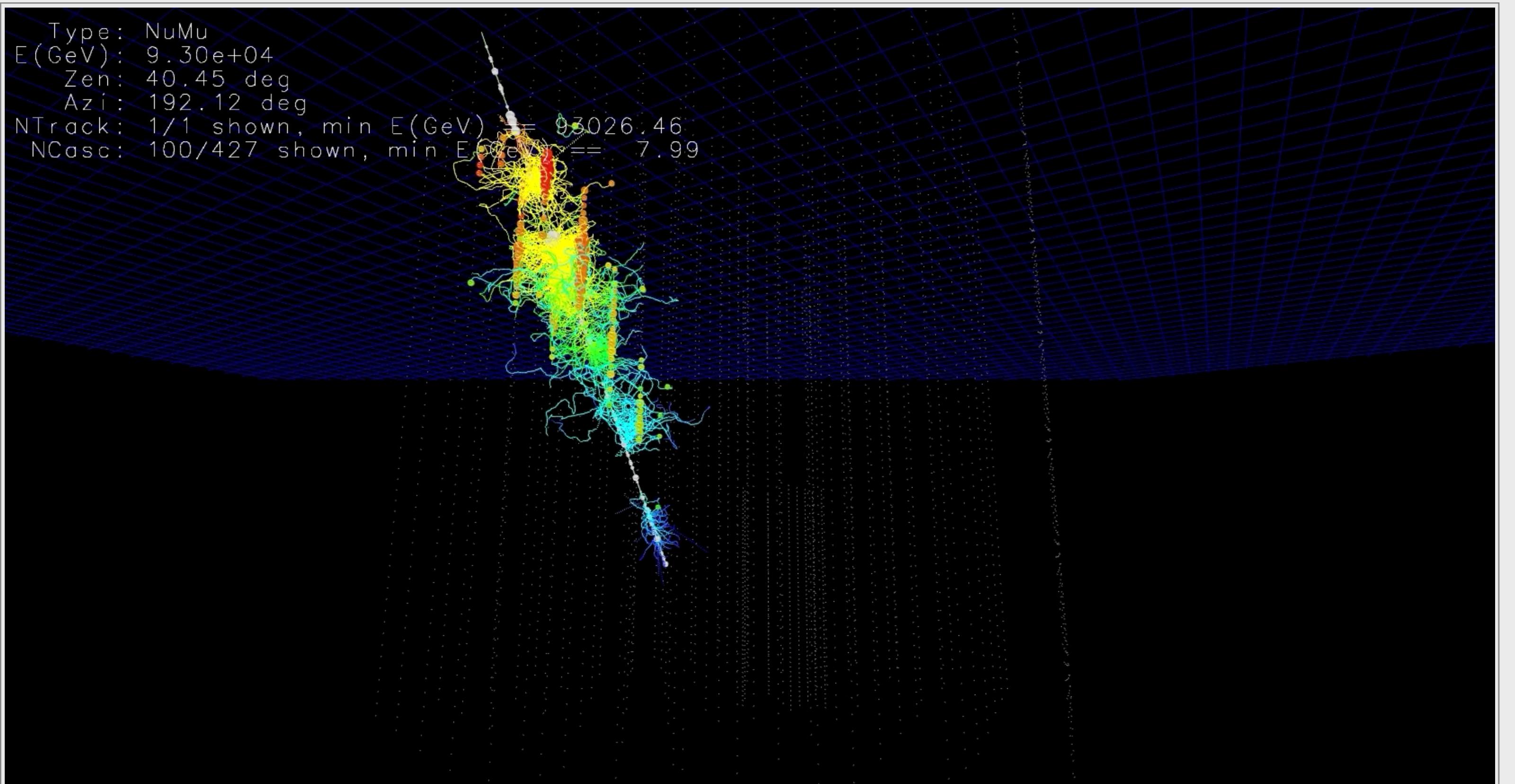
IceModel



Photon Propagation (PPC, CLSim)



Photon Propagation (PPC, CLSim)



PolyPlopia

(from gr. πολύς - polús, "many", and ὄψ-ops , "vision")

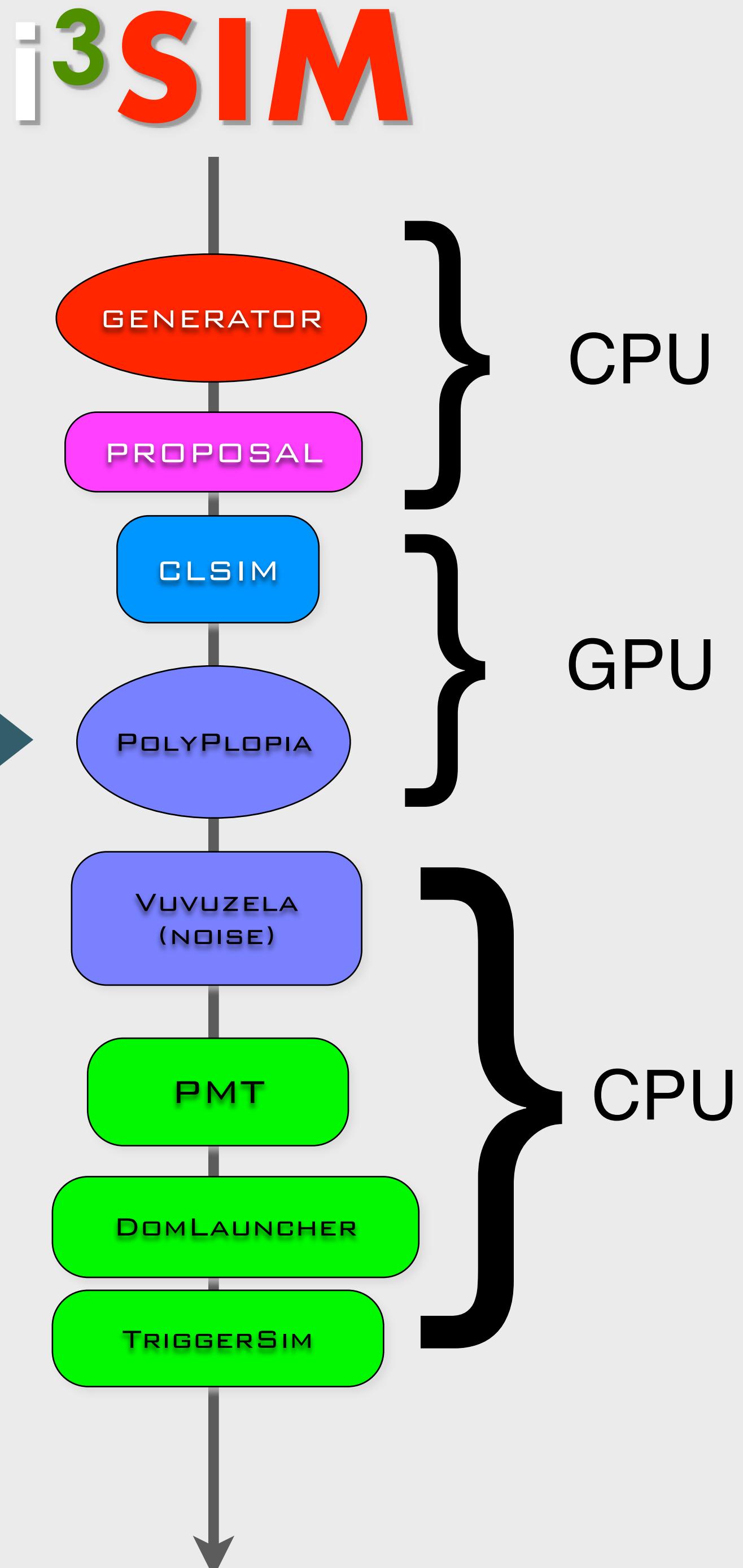
Coincident atmospheric shower events in IceCube

- **PoissonMerger**

- Injects background event read from *CoincidentEventService* on top of primary events in the chain by sampling from a Poisson distribution over a time window Δt .
- Writes a separate I3MCTree with background particles.
- Writes a combined I3MCPE map for signal and background.

- **MPHitFilter**

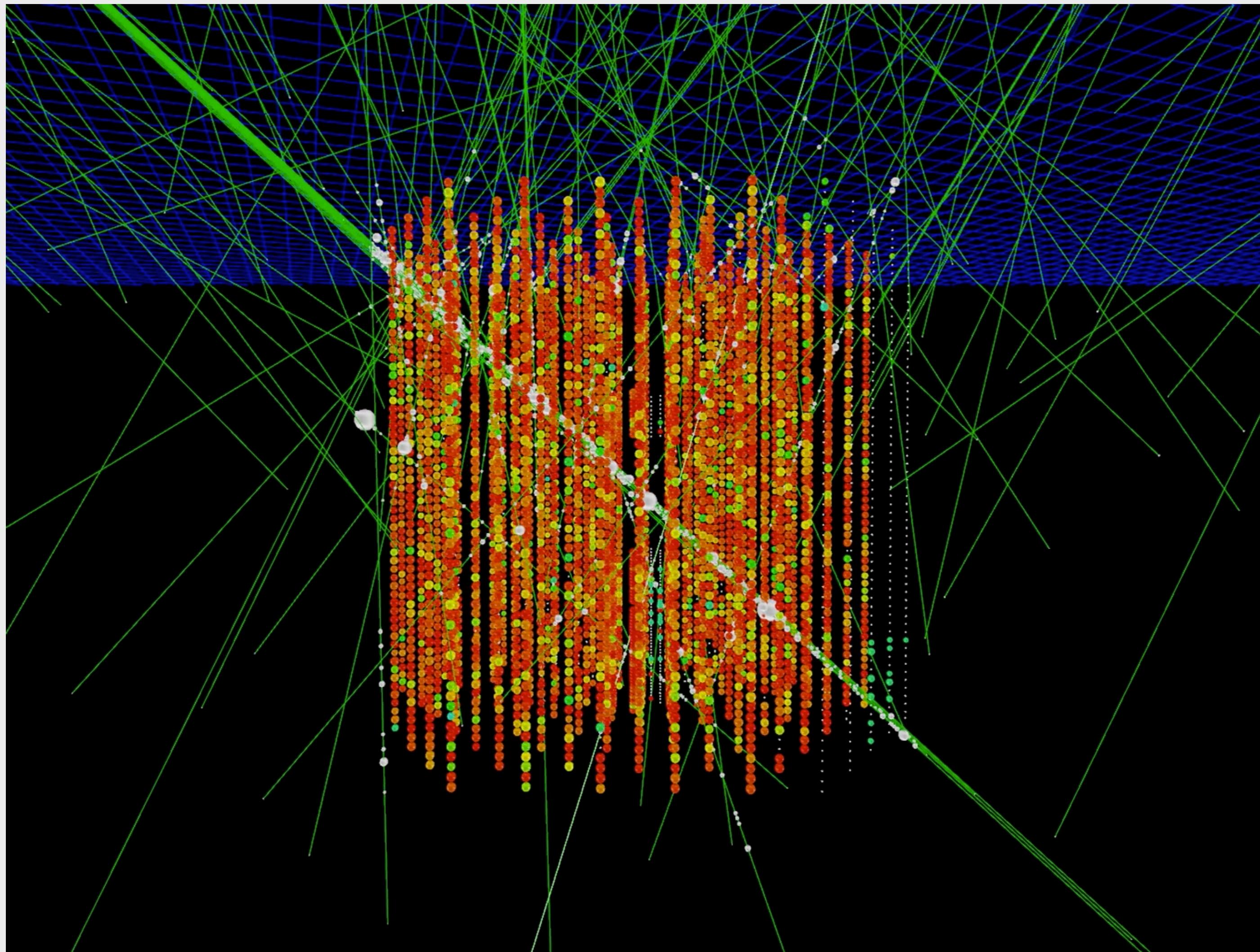
- Removes events that don't produce light in the detector and removes branches of I3MCTrees whose particles don't produce enough PEs in the detector,
- Q-frames split into P-frames based on timing, topology



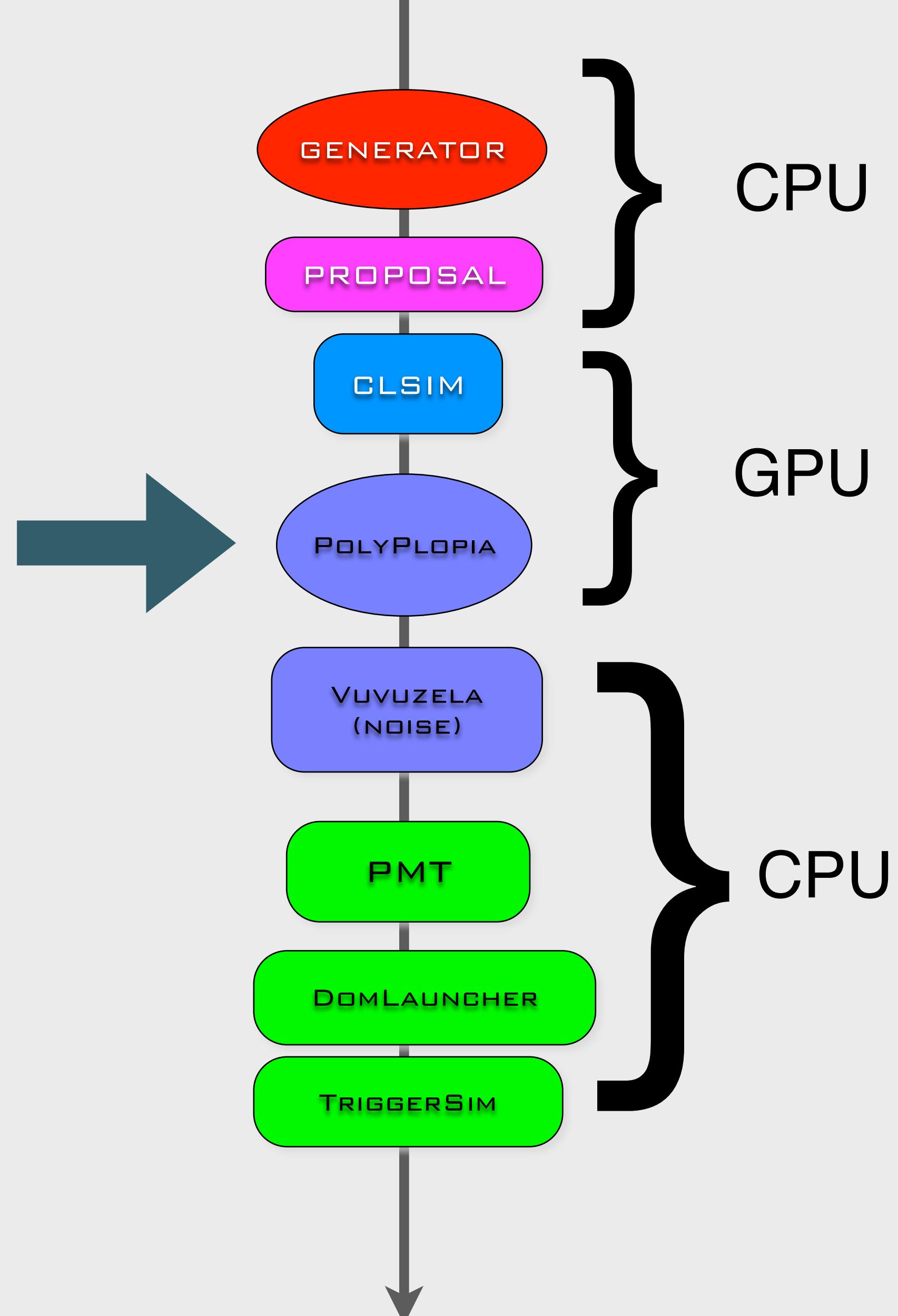
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Coincident atmospheric shower events in IceCube

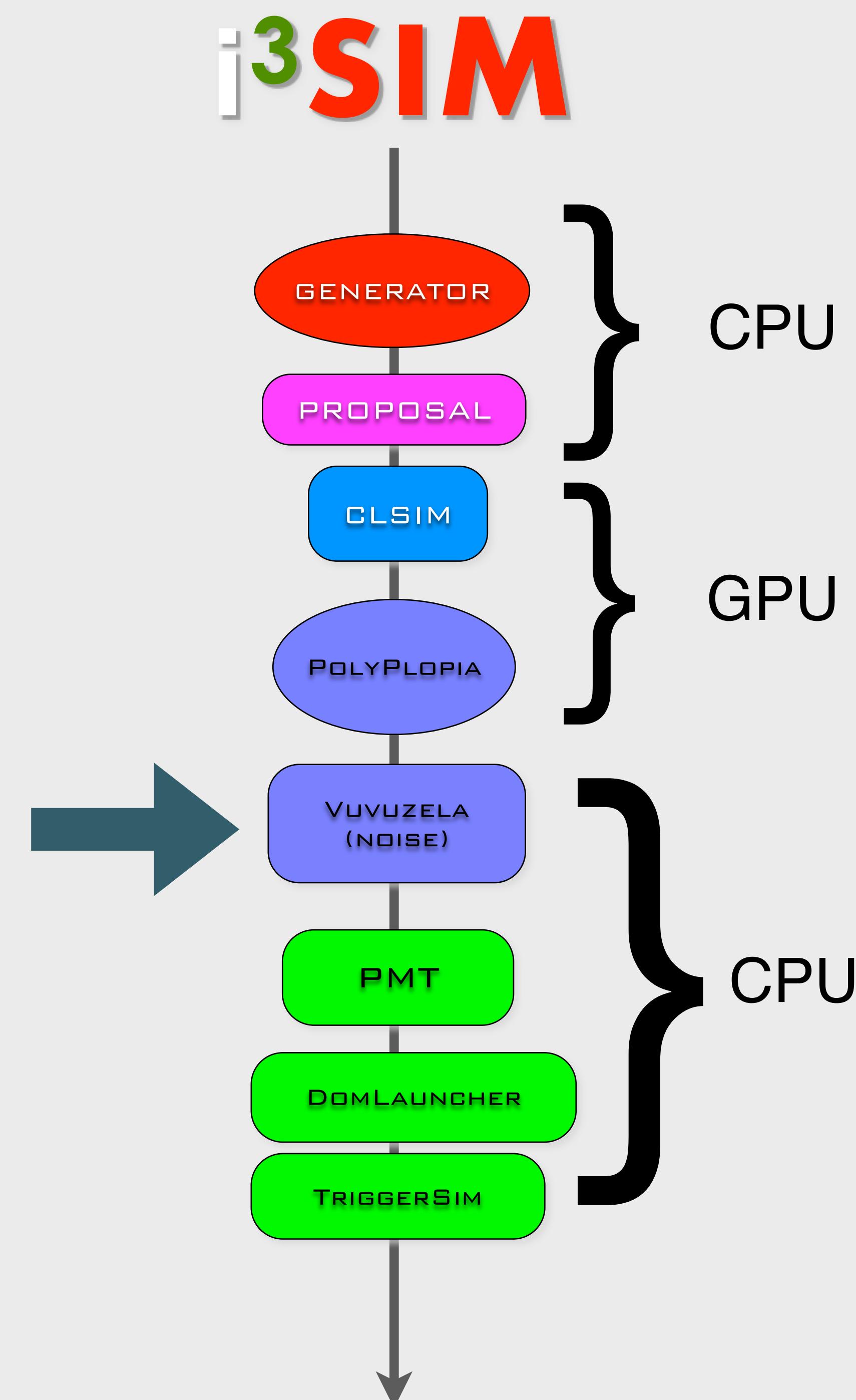
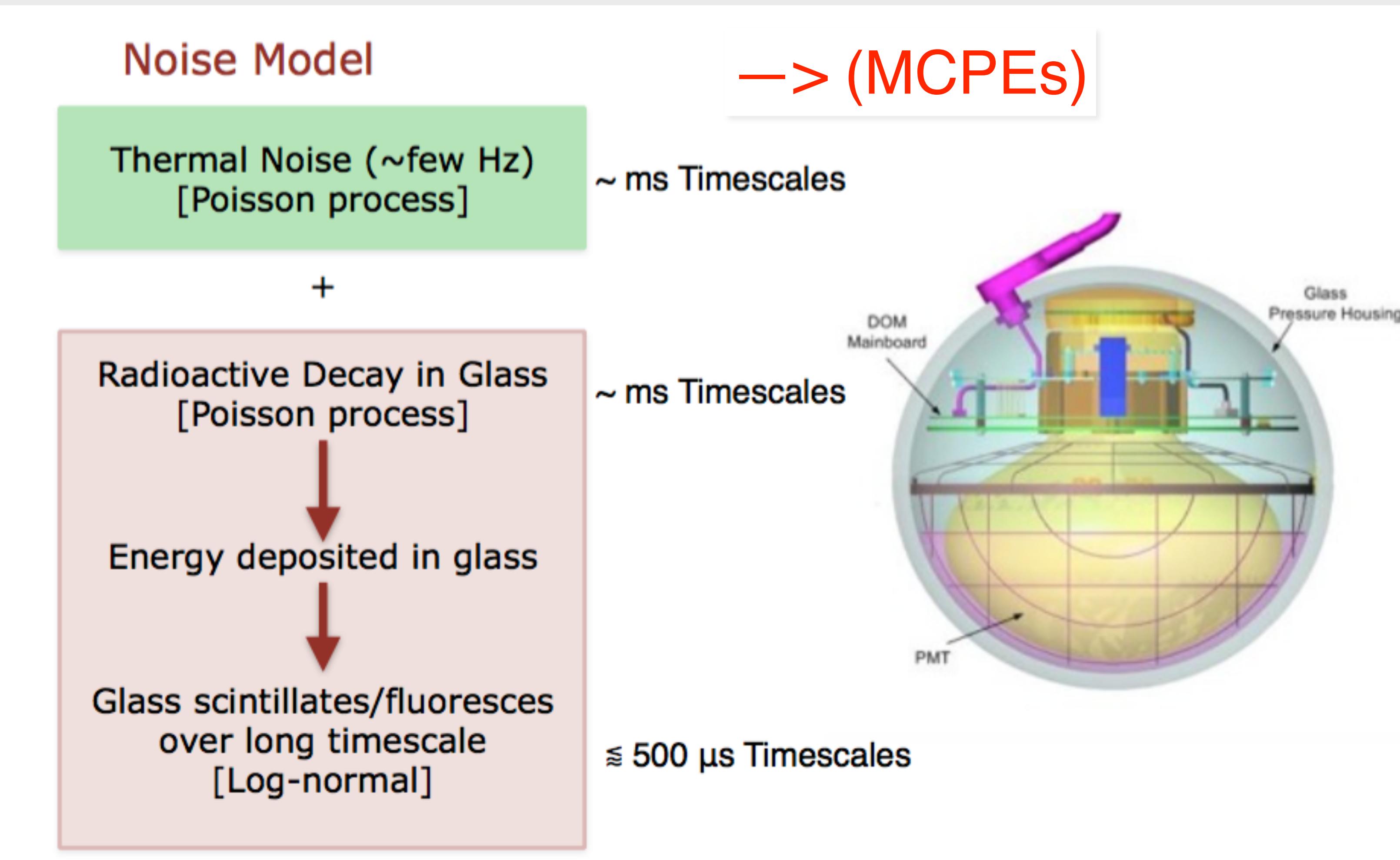


i3SIM



Vuvuzela (Noise)

i3SIM



PMTResponseSimulator

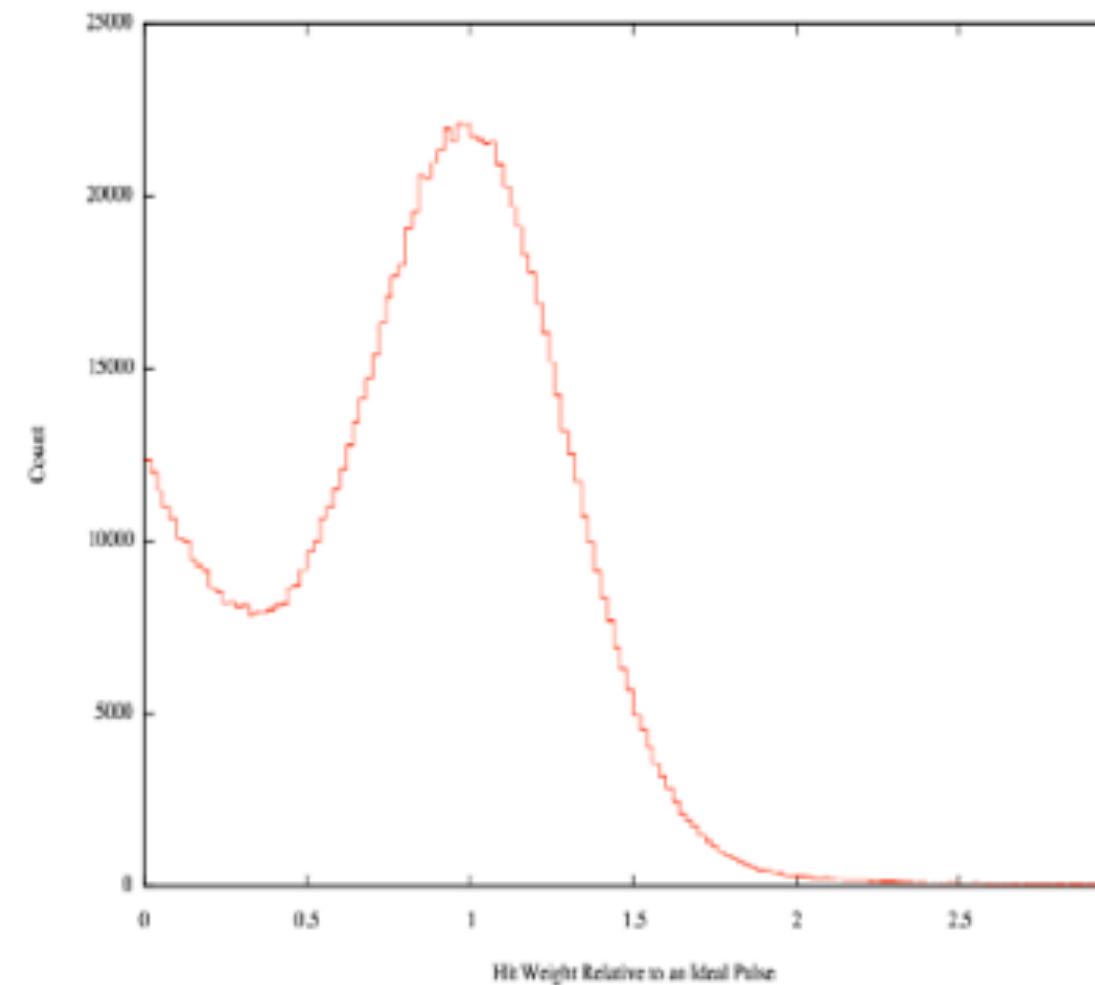


In- and output:

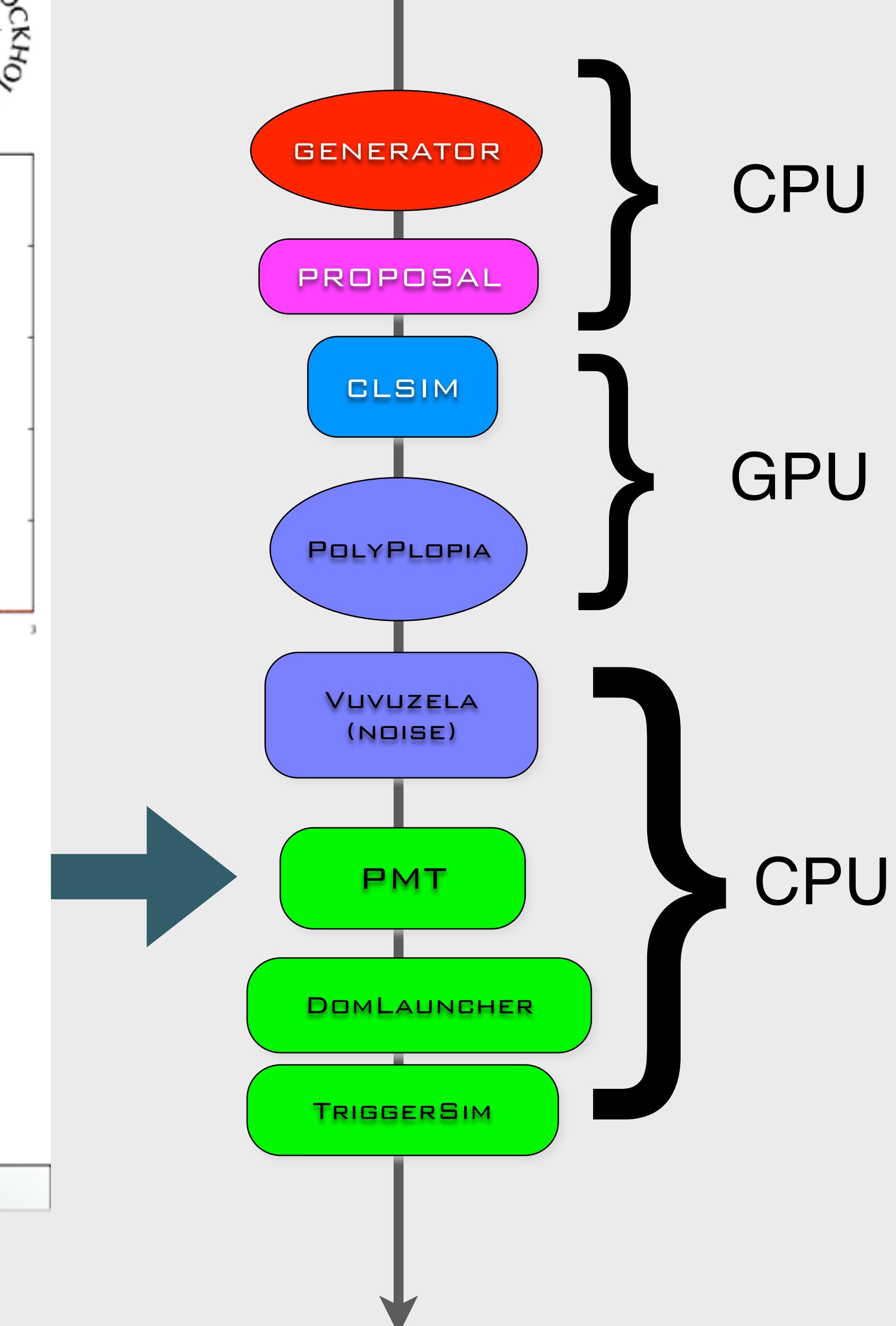
- Takes MCHits as input.
- Produces weighted MCHits.

Processing MCHits:

- Give each MCHit a weight → corresponding to the pulse charge that photon would yield.
- Generate prepulses, late pulses and after pulses.
- Apply time jitter.
- Simulate the effect of saturation.



<http://software.icecube.wisc.edu/documentation/projects/DOMLauncher/index.html>



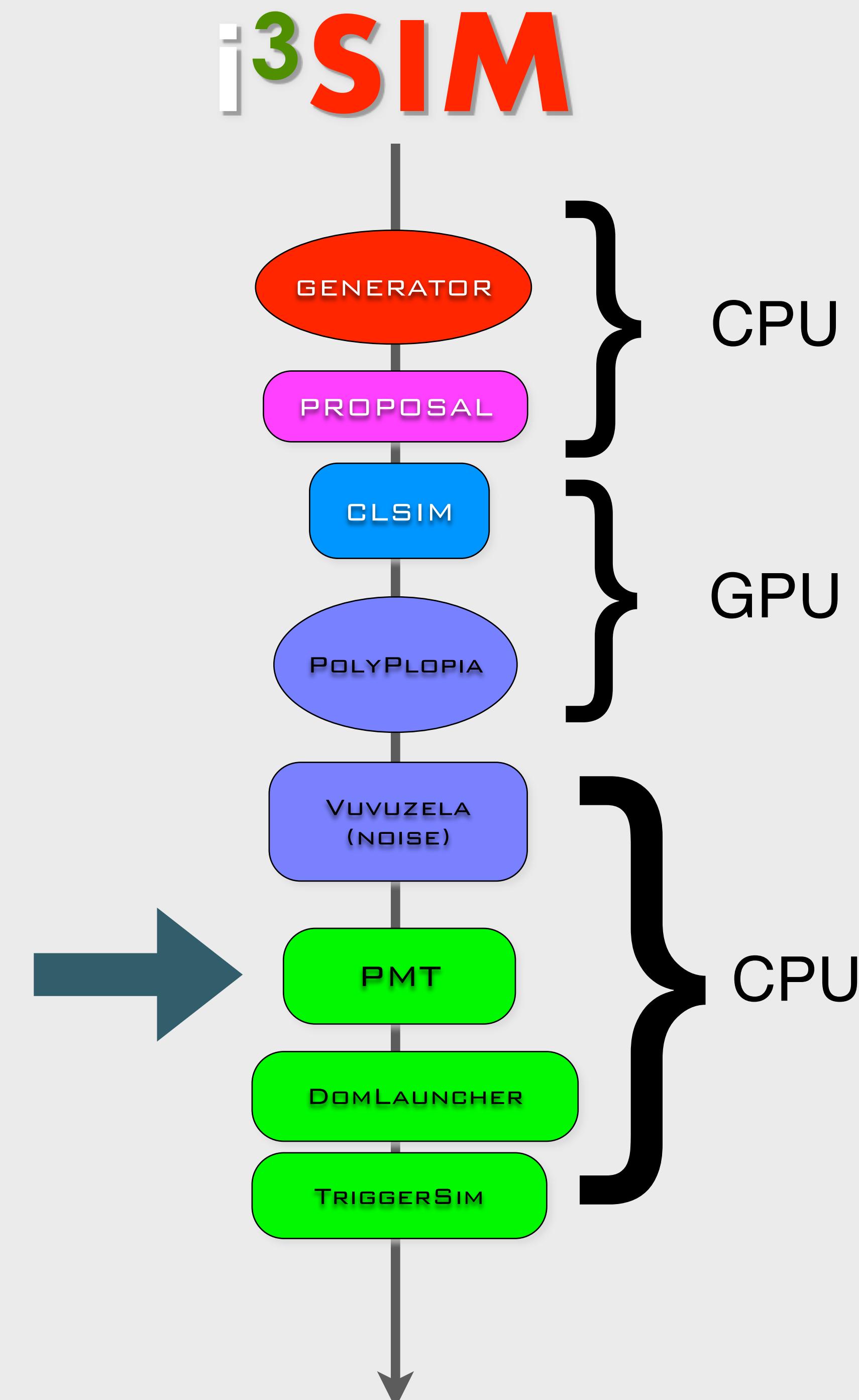
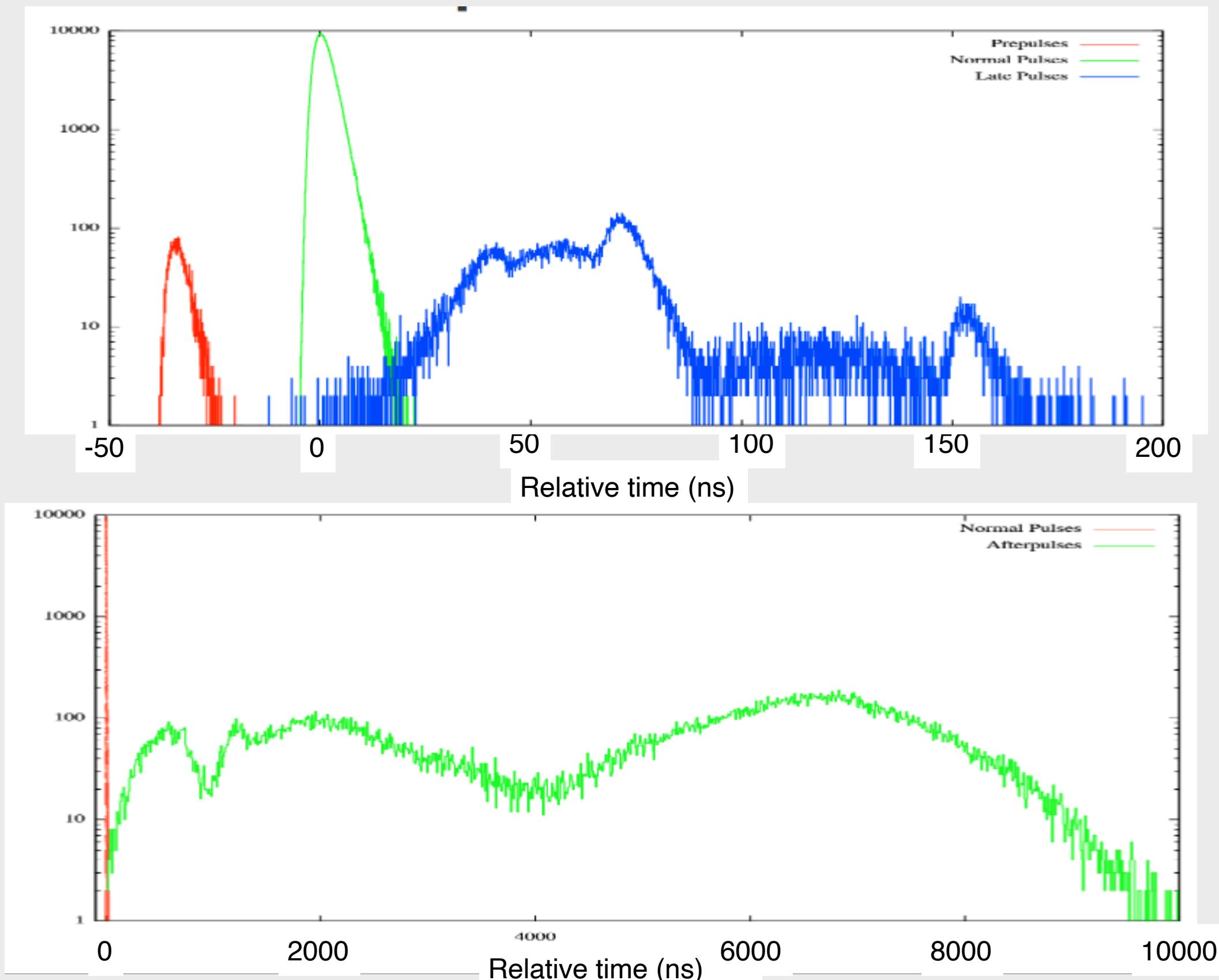
PMT Simulation

i3SIM

Pre-pulses: photoelectrons ejected from the first dynode,

Late pulses: electrons backscatter from dynode to cathode.

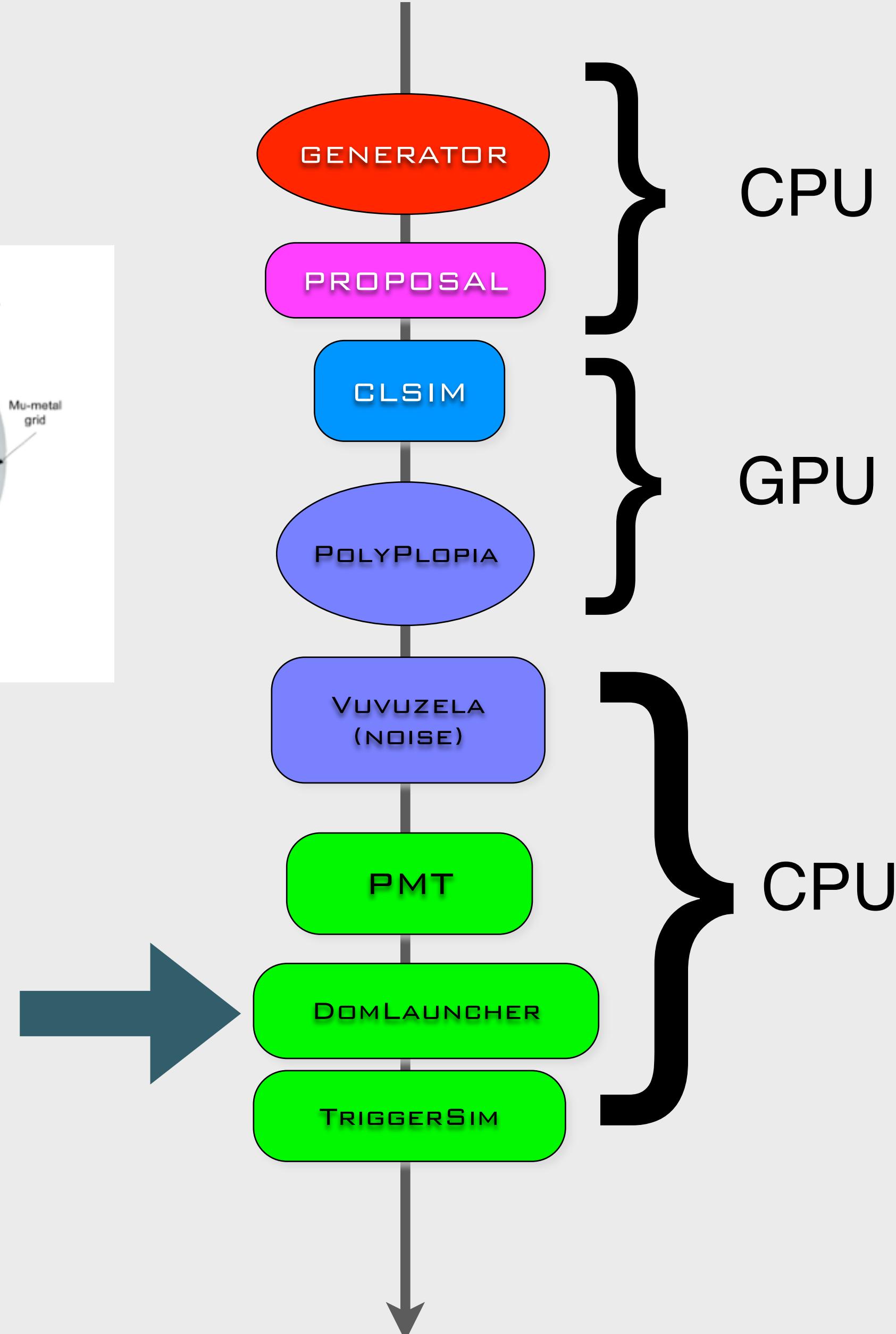
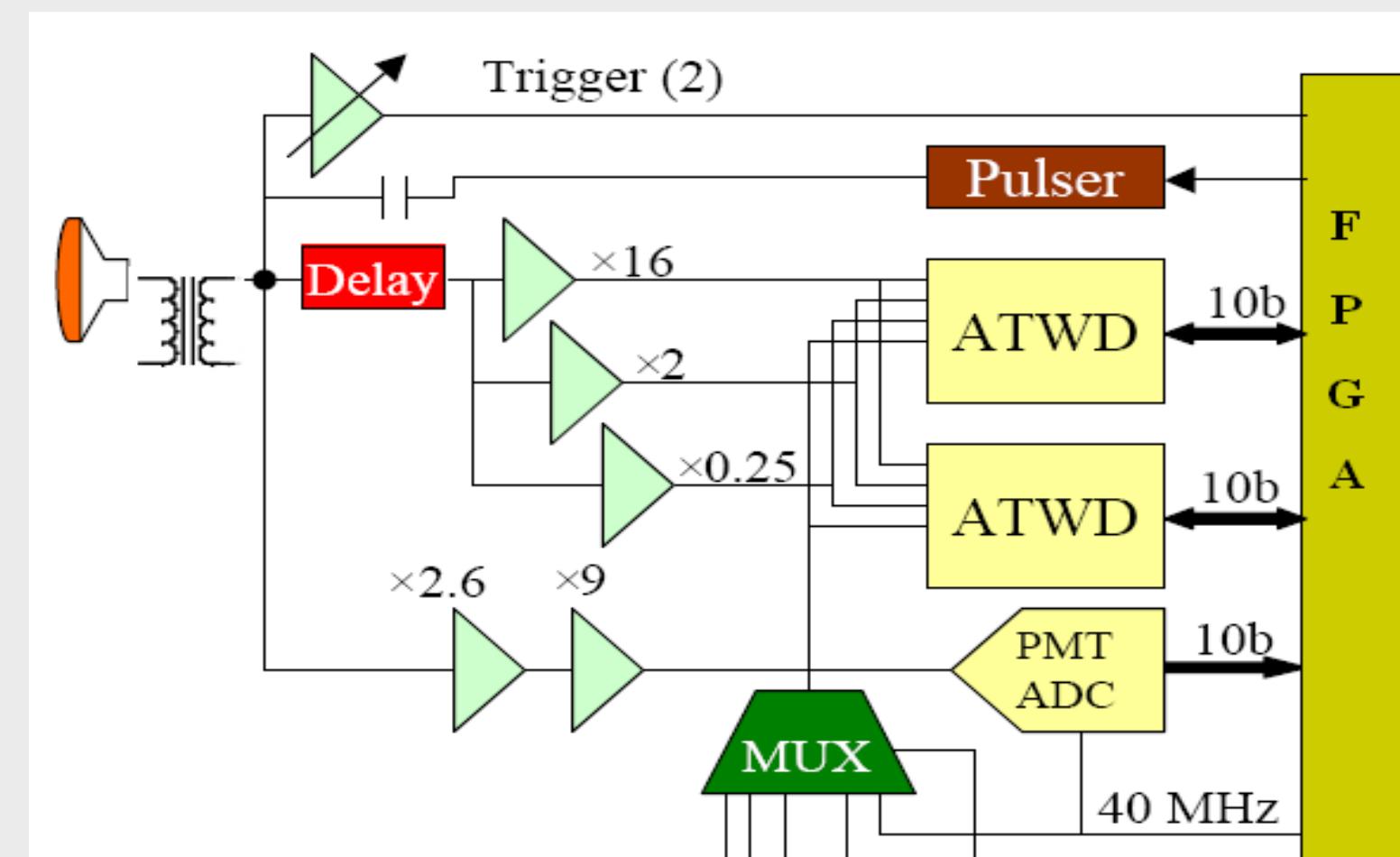
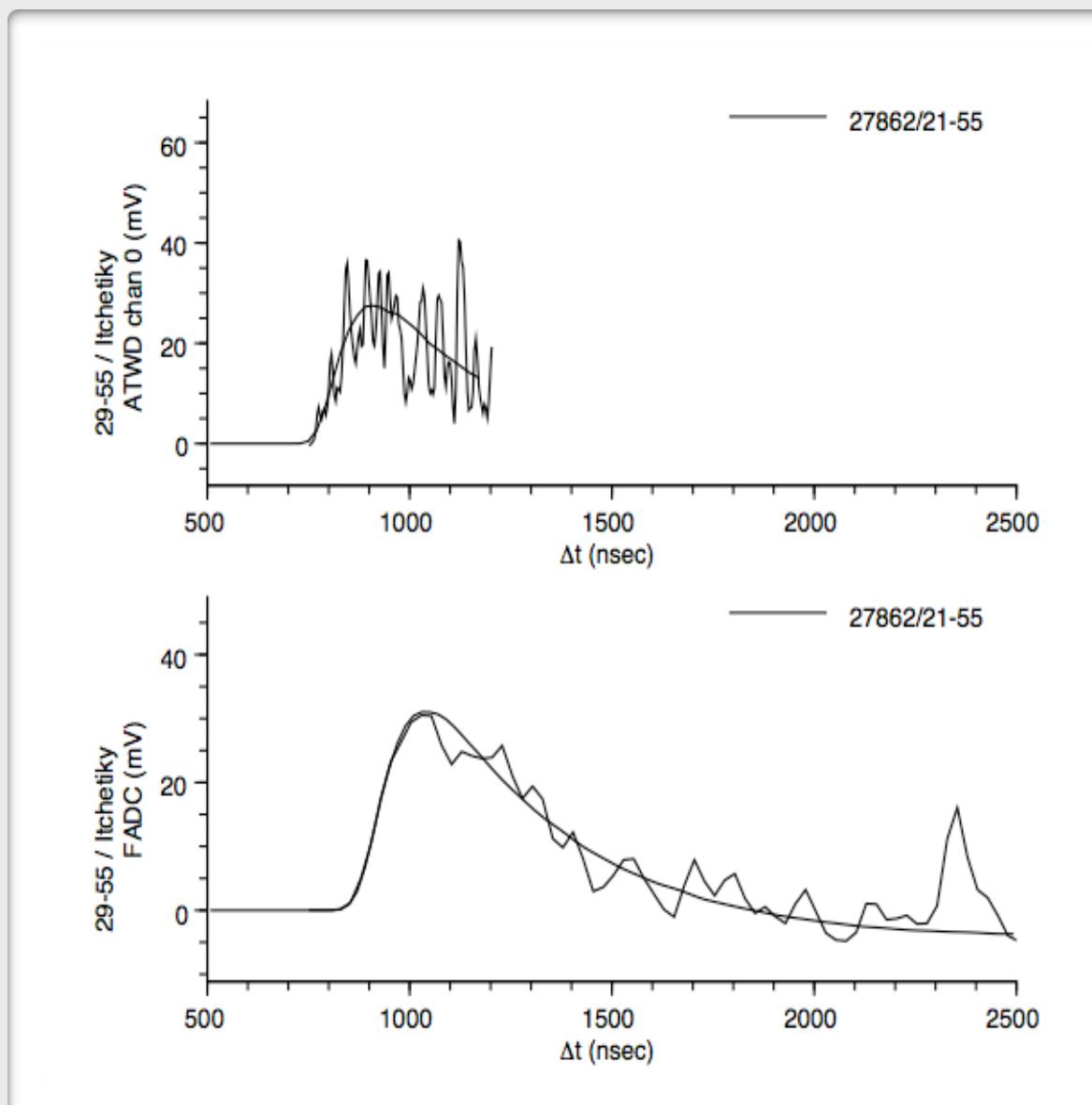
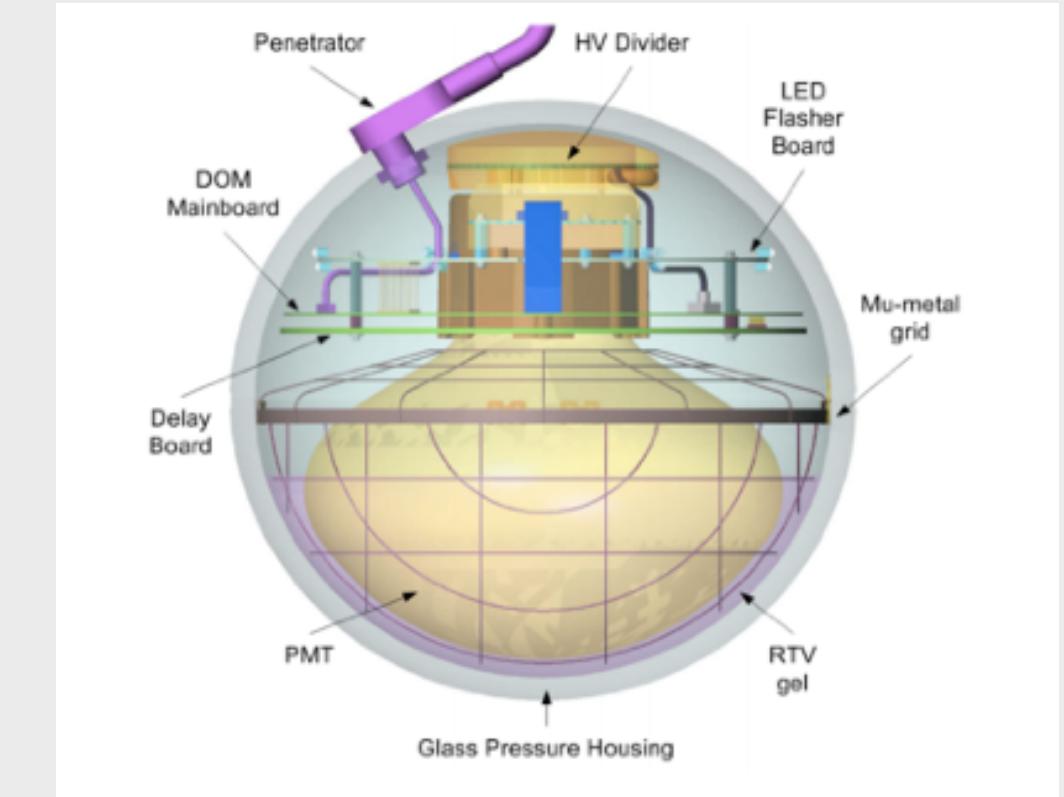
After-pulses: ionization of residual gases by electrons accelerated in the space between dynode.



DOMLauncher: DOM electronics simulation

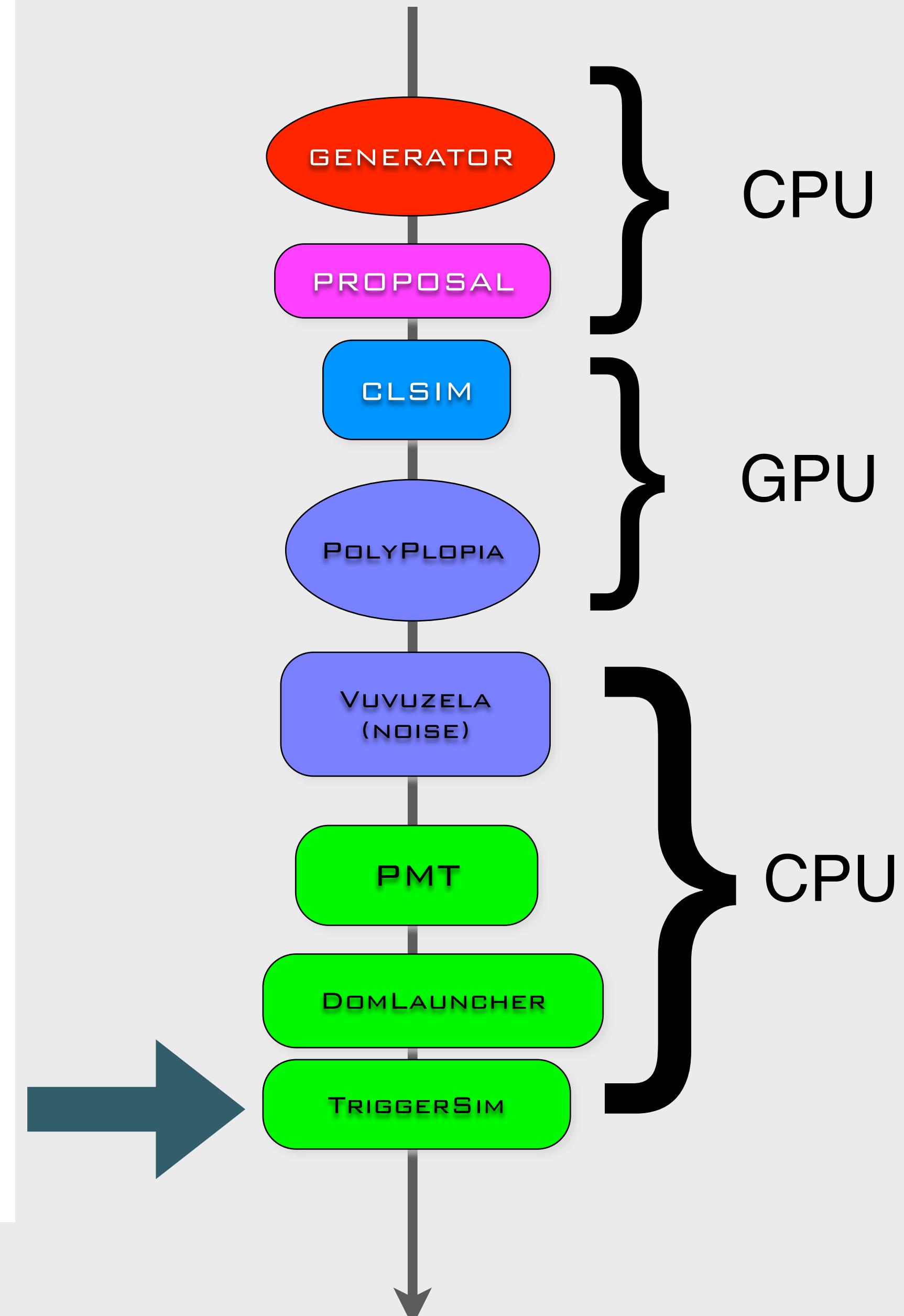
i3SIM

- Discriminator
- LC-logic
- Digitization
- Simulated effects
 - Electronic noise in the digitizers
 - Beacon launches (CPU triggered launches)
 - The FPGA Clock phase
 - RAPcal time uncertainty



Trigger Types

- **Simple Multiplicity Trigger (SMT)**
 - N HLC hits or more in a time window
 - Example: InIce SMT8 with $N_{\text{hits}} \geq 8$ in $5 \mu\text{s}$
 - readout window around this captures early and late hits ($-4 \mu\text{s}, +6 \mu\text{s}$)
- **String trigger** (a.k.a. Cluster trigger in DAQ-land)
 - N HLC hits out of M DOMs on a string in a time window
 - Example: 5 hits from a run of 7 adjacent DOMs in a time window of 1500 ns
- **Volume** trigger (a.k.a Cylinder trigger in DAQ-land)
 - simple majority of HLC hits (SMT4) with volume element including one layer of strings around a center string
 - cylinder height is 5 DOM-layers (2 up and down from the selected DOM).
- **Slow Particle** trigger (SLOP)
 - slow-moving hits along a track
 - lengths of the order of $500 \mu\text{s}$ and extending up to milliseconds
- **Fixed Rate** trigger, **Minimum Bias** trigger, **Calibration** trigger



simprod-scripts

<https://github.com/icecube/icetray/tree/main/simprod-scripts>

Collection of scripts, tray segments used in simulation production.
Central place with standard segments for running simulation in both
official production and private simulations.

simprod-scripts

Scripts:

\$I3_SRC/simprod-scripts/resources/scripts

(run the individual pieces as broken down by production tasks)

```
$ python nugen.py -h
```

```
Usage: nugen.py [options]
```

```
Options:
```

- h, --help show this help message and exit
- no-execute boolean condition to execute
- outputfile=OUTPUTFILE
Output filename
- summaryfile=SUMMARYFILE
XMLSummary filename
- mjd=MJD MJD for the GCD file
- seed=RNGSEED RNG seed
- UseGSLRNG

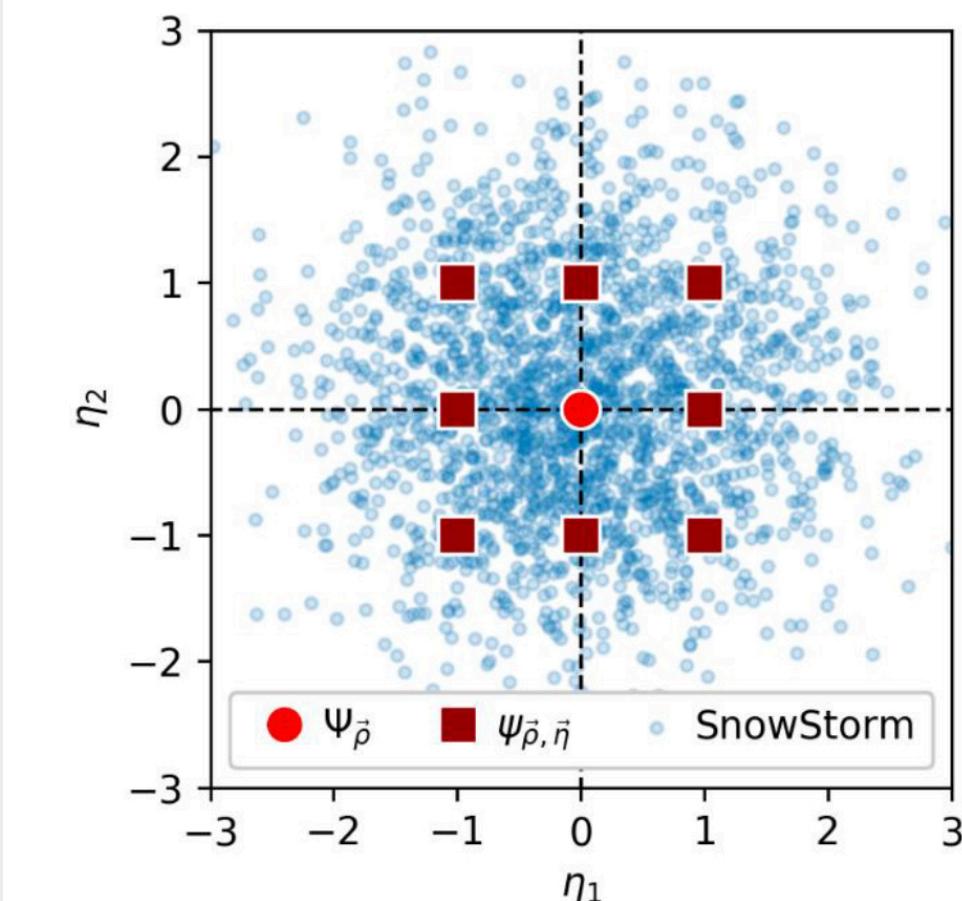
```
...
```

simprod-scripts (SnowSuite)

Scripts: `$I3_SRC/simprod-scripts/resources/scripts`

(run the individual pieces as broken down by production tasks)

```
simprod-scripts/resources/scripts/SnowSuite/
├── 1-process-Gen.py
├── 2-Polyplopia.py
├── 2-Propagate.py
├── 3-Snowstorm.py
├── 4-process-Weight.py
├── demo
└── SimpleInjector.py
iceprod2
├── iceprod_config_example.json
├── load_env
└── source_env
jobs
└── condor_manager.py
Snowstorm_Cfg.yml
Snowstorm_FullSystematics.yml
Snowstorm_Nominallce_Cfg.yml
utils.py
```



Journal of Cosmology and Astroparticle Physics

Efficient propagation of systematic uncertainties from calibration to analysis with the SnowStorm method in IceCube

M.G. Aartsen¹⁶, M. Ackermann⁵⁴, J. Adams¹⁶, J.A. Aguilar¹², M. Ahlers²⁰, C. Alispach²⁶, B. Al Atoum⁴, K. Andeen³⁷, T. Anderson⁵¹, I. Ansseau¹² [+Show full author list](#)

Published 21 October 2019 • © 2019 IOP Publishing Ltd and Sissa Medialab

[Journal of Cosmology and Astroparticle Physics, Volume 2019, October 2019](#)

simprod-scripts

Example: Running scripts:

```
icecube@M16:~$ ssh submitter
[submitter]$ condor_submit -interactive 'request_gpus=1'
Submitting job(s).
1 job(s) submitted to cluster 120263704.
Waiting for job to start...
Welcome to slot1@gtx-00.icecube.wisc.edu!

[gtx-00]$ cd $CONDOR_SCRATCH_DIR
[gtx-00]$ /cvmfs/icecube.opensciencegrid.org/py3-v4.3.0/icetray-env icetray/v1.10.0
*****
*                                         *
*          W E L C O M E   t o   I C E T R A Y   *
*                                         *
*          Version icetray.stable      git:f5d21802  *
*                                         *
*          You are welcome to visit our Web site    *
*                  http://icecube.umd.edu             *
*                                         *
*****
[gtx-00]$ python ${I3_BUILD}/simprod-scripts/resources/scripts/nugen.py \
--outputfile nutau.i3 --nevents 100 \
--seed=123 --procnum 0 --nproc=1 \
--FromEnergy 1e5 --ToEnergy 1e6 --NuFlavor NuTau --UseGSLRNG

[gtx-00]$ dataio-shovel nutau.i3
```

simprod-scripts

Example: Running scripts:

```
[gtx-00]$ dataio-shovel nutau.i3
```

simprod-scripts

Example: Running scripts:

```
[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/clsim.py \
    --gcdfile /cvmfs/icecube.opensciencegrid.org/data/GCD/
GeoCalibDetectorStatus_2020.Run134142.Pass2_v0.i3.gz \
    --inputfilelist nutau.i3 --outputfile mcpes.i3 \
    --seed 123 --procnum 0 --nproc 1 --no-RunMPHitFilter \
    --UseGPUs --UseGSLRNG
```

```
[gtx-00]$ dataio-shovel mcpes.i3
```

```
[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/detector.py \
    --gcdfile /cvmfs/icecube.opensciencegrid.org/data/GCD/
GeoCalibDetectorStatus_2020.Run134142.Pass2_v0.i3.gz \
    --infile mcpes.i3 --outfile det.i3 \
    --seed 123 --procnum 0 --nproc 1 --RunID 123 --UseGSLRNG
```

simprod-scripts

Example: Running scripts:

```
[gtx-00]$ dataio-shovel mcpes.i3
```

I3 Data Shovel	Press '?' for help	
Name	Type	Bytes
I3MCPESeriesMap	I3Map<OMKey, vector<I3MCPE> >	41
I3MCPESeriesMapParticleIDMap	I3Map<OMKey, map<I3ParticleID, vector<unsigned int...>>	41
I3MCTree	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3...>>	6878
I3MCTree_preMuonProp	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3...>>	1538
I3MCTree_preMuonProp_RNGState	I3 GSL Random Service State	85
I3MCWeightDict	I3Map<__cxx11::string, double>	1400
MMCTrackList	I3Vector<I3MMCTrack>	40
NuGPrimary	I3Particle	150

Key: 1/8
Frame: 3/102 (2%)
Stop: DAQ
Run/Event: (n/a)
SubEvent: (n/a)

StartTime: (n/a)
Duration: (n/a)

simprod-scripts

Exercise: Running scripts:

```
[gtx-00] $ dataio-shovel det.i3
```

I3 Data Shovel

Press '?' for help

Name	Type	Bytes
BeaconLaunches	I3Map<OMKey, vector<I3DOMLaunch> >	46
I3EventHeader	I3EventHeader	99
I3MCPESeriesMap	I3Map<OMKey, vector<I3MCPE> >	113286
I3MCPESeriesMapParticleIDMap	I3Map<OMKey, map<I3ParticleID, vector<unsigned int...>>	36649
I3MCPESeriesMapWithoutNoise	I3Map<OMKey, vector<I3MCPE> >	109543
I3MCPulseSeriesMap	I3Map<OMKey, vector<I3MCPulse> >	82000
I3MCPulseSeriesMapParticleI...	I3Map<OMKey, map<I3ParticleID, vector<unsigned int...>>	40743
I3MCPulseSeriesMapPrimaryIDMap	I3Map<OMKey, map<I3ParticleID, vector<unsigned int...>>	27299
I3MCTree	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3...>>	10730
I3MCTree_preMuonProp	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3...>>	422
I3MCTree_preMuonProp_RNGState	I3GSLRandomServiceState	85
I3MCWeightDict	I3Map<__cxx11::string, double>	1400
I3TriggerHierarchy	I3Tree<I3Trigger>	792
I3Triggers	I3Tree<I3Trigger>	414
IceTopRawData	I3Map<OMKey, vector<I3DOMLaunch> >	46
InIceRawData	I3Map<OMKey, vector<I3DOMLaunch> >	44640
MMCTrackList	I3Vector<I3MMCTrack>	2864
NuGPrimary	I3Particle	150
TimeShift	I3PODHolder<double>	36

simprod-scripts

Exercise: Running scripts:

```
[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/corsika.py \
    --nshowers 10000 --outputfile corsika_bg.i3 --seed 1234 \
    --CORSIKAseed=123 --ranpri 2 \
    --corsikaVersion v6960-5comp \
    --corsikaName dcorsika --UseGSLRNG \
    --skipoptions compress

[gtx-00]$ dataio-shovel corsika_bg.i3

[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/polyplopia.py \
    --gcdfile /cvmfs/icecube.opensciencegrid.org/data/GCD/
GeoCalibDetectorStatus_2020.Run134142.Pass2_V0.i3.gz \
    --inputfile mcpes.i3 --outputfile merged_pes.i3 \
    --seed 1234 \
    --backgroundfile corsika_bg.i3 --mctype NuTau \
    --UseGSLRNG

[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/detector.py \
    --gcdfile /cvmfs/icecube.opensciencegrid.org/data/GCD/
GeoCalibDetectorStatus_2020.Run134142.Pass2_V0.i3.gz \
    --inputfile merged_pes.i3 --outputfile det_wcoinc.i3 \
    --seed 123 --RunID 123 --UseGSLRNG

[gtx-00]$ dataio-shovel det_wcoinc.i3
```

simprod-scripts

Exercise: Running scripts:

```
[gtx-00] $ dataio-shovel corsika_bg.i3
```

I3 Data Shovel — Press '?' for help

Name	Type	Bytes
CorsikaInteractionHeight	I3P0DHolder<double>	36
CorsikaWeightMap	I3Map<__cxx11::string, double>	484
I3CorsikaInfo	I3CorsikaInfo	109
I3MCTree	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3...>>	546
I3MCTree_preSampling	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3...>>	546

Key: 1/5
Frame: 3/111+
Stop: DAQ
Run/Event: (n/a)
SubEvent: (n/a)

simprod-scripts

Exercise: Running scripts:

```
[gtx-00] $ dataio-shovel merged_pes.i3
```

I3 Data Shovel

Press '?' for help

Name	Type	Bytes
BackgroundI3MCPESeriesMap	I3Map<OMKey, vector<I3MCPE> >	41
BackgroundI3MCPESeriesMapPa...	I3Map<OMKey, map<I3ParticleID, vector<unsigned int> > >	41
BackgroundI3MCTree	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3ParticleID>>	32
BackgroundI3MCTreePEcounts	I3Map<unsigned int, unsigned int>	47
BackgroundI3MCTree_preMuonProp	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3ParticleID>>	32
BackgroundI3MCTree_preMuonP...	I3GSLRandomServiceState	85
BackgroundMMCTrackList	I3Vector<I3MMCTrack>	40
I3MCPESeriesMap	I3Map<OMKey, vector<I3MCPE> >	41
I3MCPESeriesMapParticleIDMap	I3Map<OMKey, map<I3ParticleID, vector<unsigned int> > >	41
I3MCTree	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3ParticleID>>	2902
I3MCTree_preMuonProp	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3ParticleID>>	422
I3MCTree_preMuonProp_RNGState	I3GSLRandomServiceState	85
I3MCWeightDict	I3Map<__cxx11::string, double>	1424
MMCTrackList	I3Vector<I3MMCTrack>	40
NuGPrimary	I3Particle	150
PhotonSeriesMap	I3Map<ModuleKey, I3Vector<I3CompressedPhoton> >	53
PolyplopiaInfo	I3Map<__cxx11::string, int>	135
PolyplopiaPrimary	I3Particle	150
SignalI3MCPEs	I3Map<OMKey, vector<I3MCPE> >	41
SignalI3MCTree	TreeBase::Tree<I3Particle, I3ParticleID, i3hash<I3ParticleID>>	2902

Simulating the online filter and L2 processing

```
[gtx-00]$ python filterscripts/resources/scripts/SimulationFiltering.py -h
```

```
usage: SimulationFiltering.py [-h] [-i INFIL] [-g GCDFILE] [-o OUTFILE]
                               [-n NUM] [--qify]
                               [--MinBiasPrescale MINBIASPRESCALE]
                               [--photronicsdir PHOTONICSDIR] [--enable-gfu]
                               [--log-level LOG_LEVEL] [--log-filename LOGFN]
                               [--needs_wavedeform_spe_corr]
```

optional arguments:

-h, --help show this help message and exit

-i INFIL, --input INFIL

Input i3 file(s) (use comma separated list for
multiple files)

-g GCDFILE, --gcd GCDFILE

GCD file for input i3 file

-o OUTFILE, --output OUTFILE

Output i3 file

-n NUM, --num NUM Number of frames to process

--qify Apply QConverter, use if file is P frame only

--MinBiasPrescale MINBIASPRESCALE

Set the Min Bias prescale to something other than
default

--photronicsdir PHOTONICSDIR

Directory with photonics tables

--enable-gfu Do not run GFU filter

--log-level LOG_LEVEL

Sets the logging level (ERROR, WARN, INFO, DEBUG,
TRACE)

--log-filename LOGFN If set logging is redirected to the specified file.

--needs_wavedeform_spe_corr

apply_spe_corection in wavedeform.

Simulating the online filter and L2 processing

```
[gtx-00]$ python filterscripts/resources/scripts/offlineL2/process.py -h

usage: process.py [-h] [-s] [-i INFILe] [-g GCDFILE] [-o OUTFILE] [-n NUM]
                  [--dstfile DSTFILE] [--gapsfile GAPSFILE]
                  [--icetopoutput ICETOPOOUTPUT] [--eheoutput EHEOUTPUT]
                  [--slopoutput SLOPOOUTPUT] [--rootoutput ROOTOUTPUT]
                  [--photonicssdir PHOTONICSDIR] [--log-level LOG_LEVEL]
                  [--log-filename LOGFN]

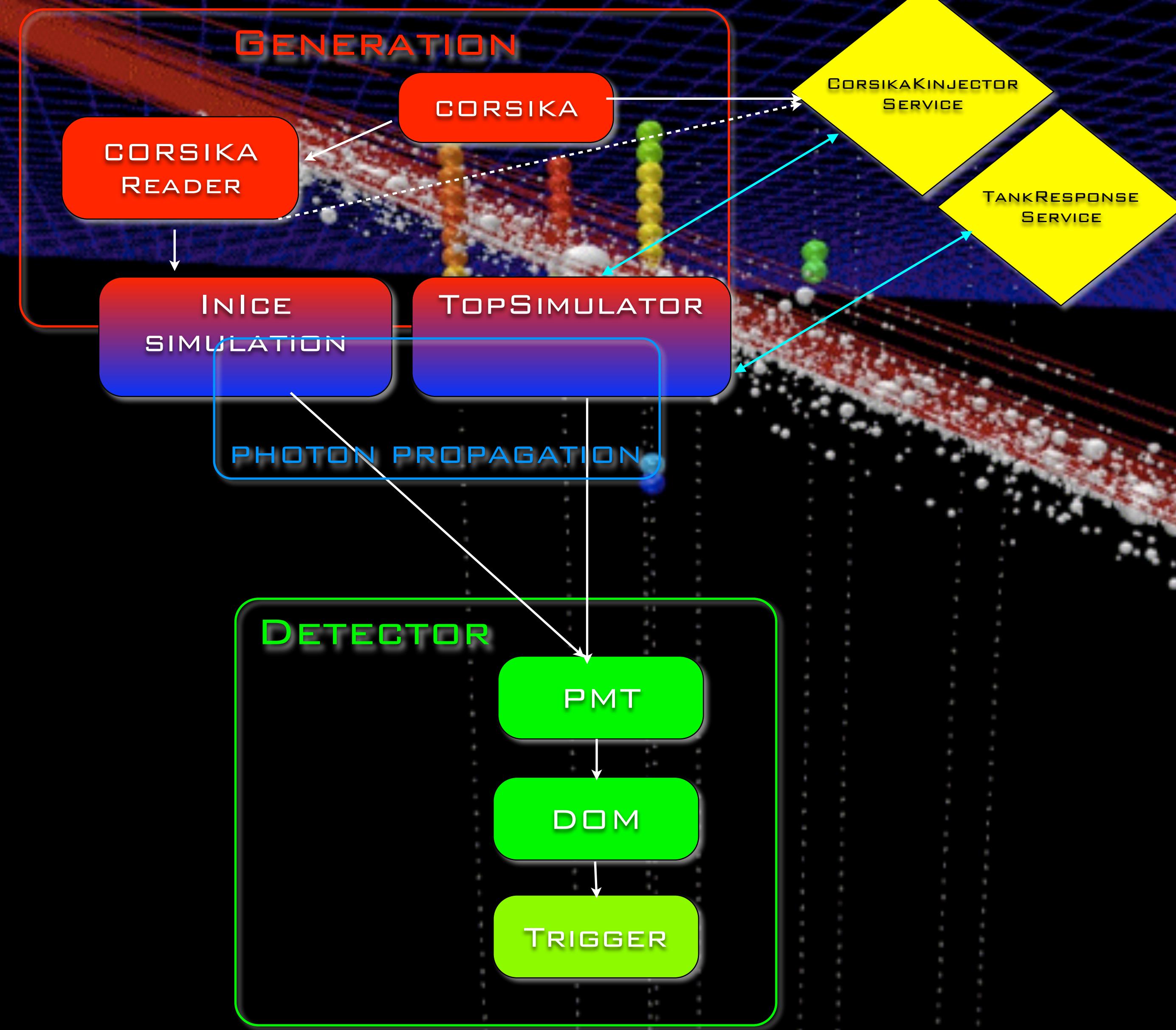
optional arguments:
  -h, --help            show this help message and exit
  -s, --simulation      Mark as simulation (MC)
  -i INFILe, --input INFILe
                        Input i3 file(s) (use comma separated list for
                        multiple files)
  -g GCDFILE, --gcd GCDFILE
                        GCD file for input i3 file
  -o OUTFILE, --output OUTFILE
                        Output i3 file
  -n NUM, --num NUM    Number of frames to process
  --dstfile DSTFILE    DST root file (should be .root)
  --gapsfile GAPSFILE  gaps text file (should be .txt)
  --icetopoutput ICETOPOOUTPUT
                        Output IceTop file
  --eheoutput EHEOUTPUT
                        Output EHE i3 file
  --slopoutput SLOPOOUTPUT
                        Output SLOP file
  --rootoutput ROOTOUTPUT
                        Output root file
  --photonicssdir PHOTONICSDIR
                        Directory with photonics tables
  --log-level LOG_LEVEL
                        Sets the logging level (ERROR, WARN, INFO, DEBUG,
                        TRACE)
  --log-filename LOGFN  If set logging is redirected to the specified file.
```

More on Simulation

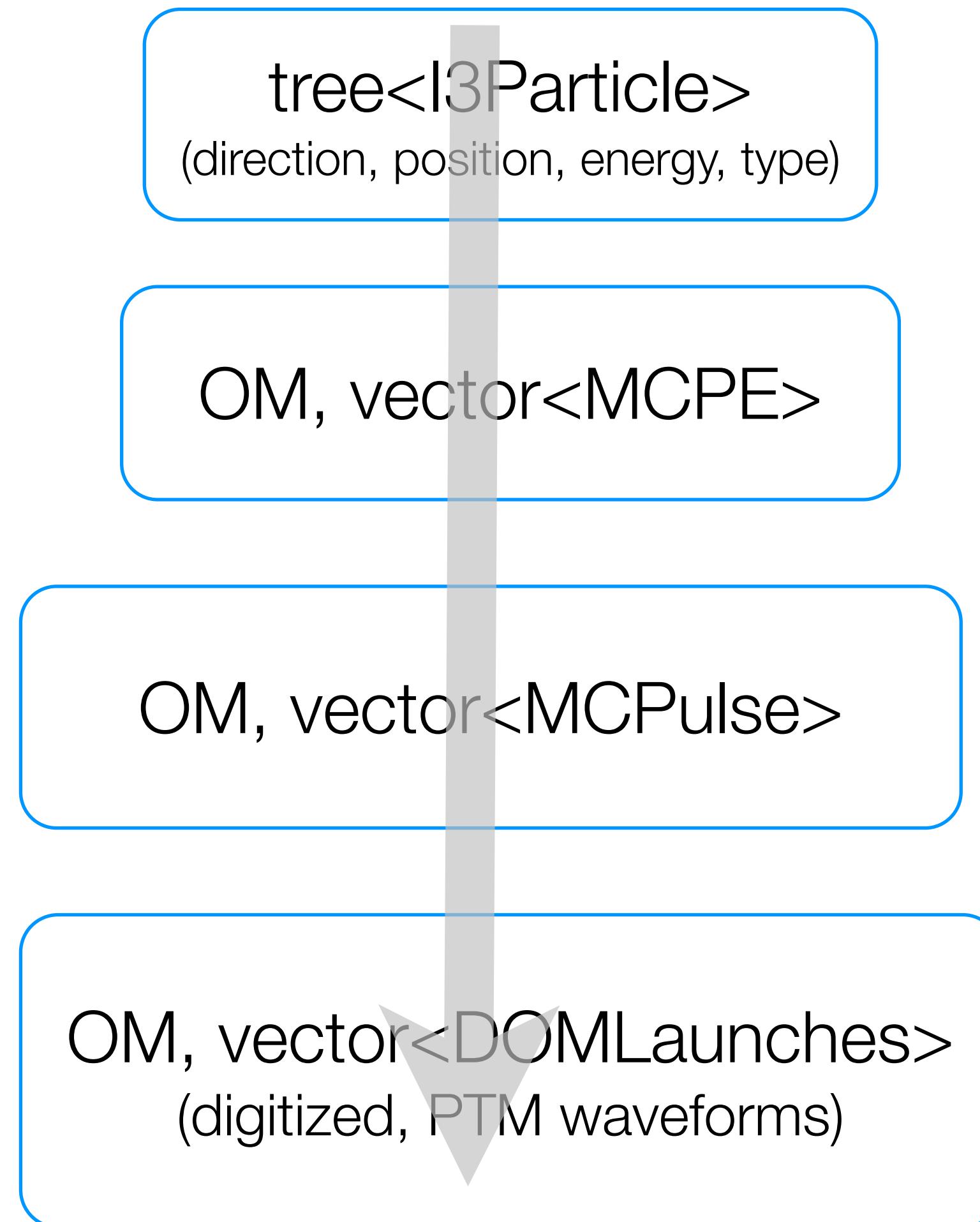
- GitHub repository: <https://docs.icecube.aq/icetray/main>
- SimProd Portal: <http://simprod.icecube.wisc.edu>
- IceProd Docs: <https://wipacrepo.github.io/iceprod/>
- JSON Templates: <https://github.com/icecube/simprod-templates>
- Weighting: <https://docs.icecube.aq/simweights/main/>
- SLACK: [#simulation](#)

Backup

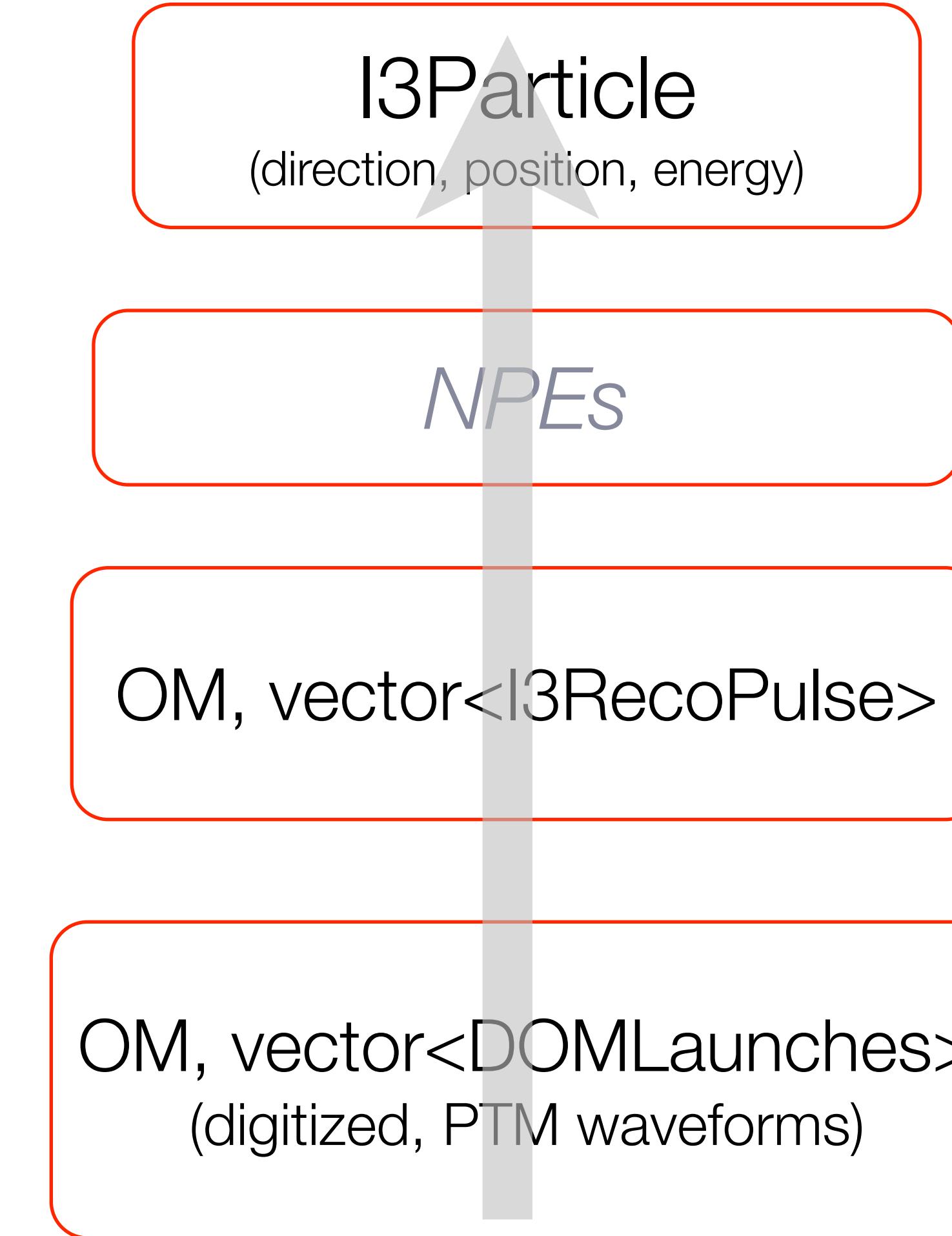
simulaton chain (IT)



Simulation



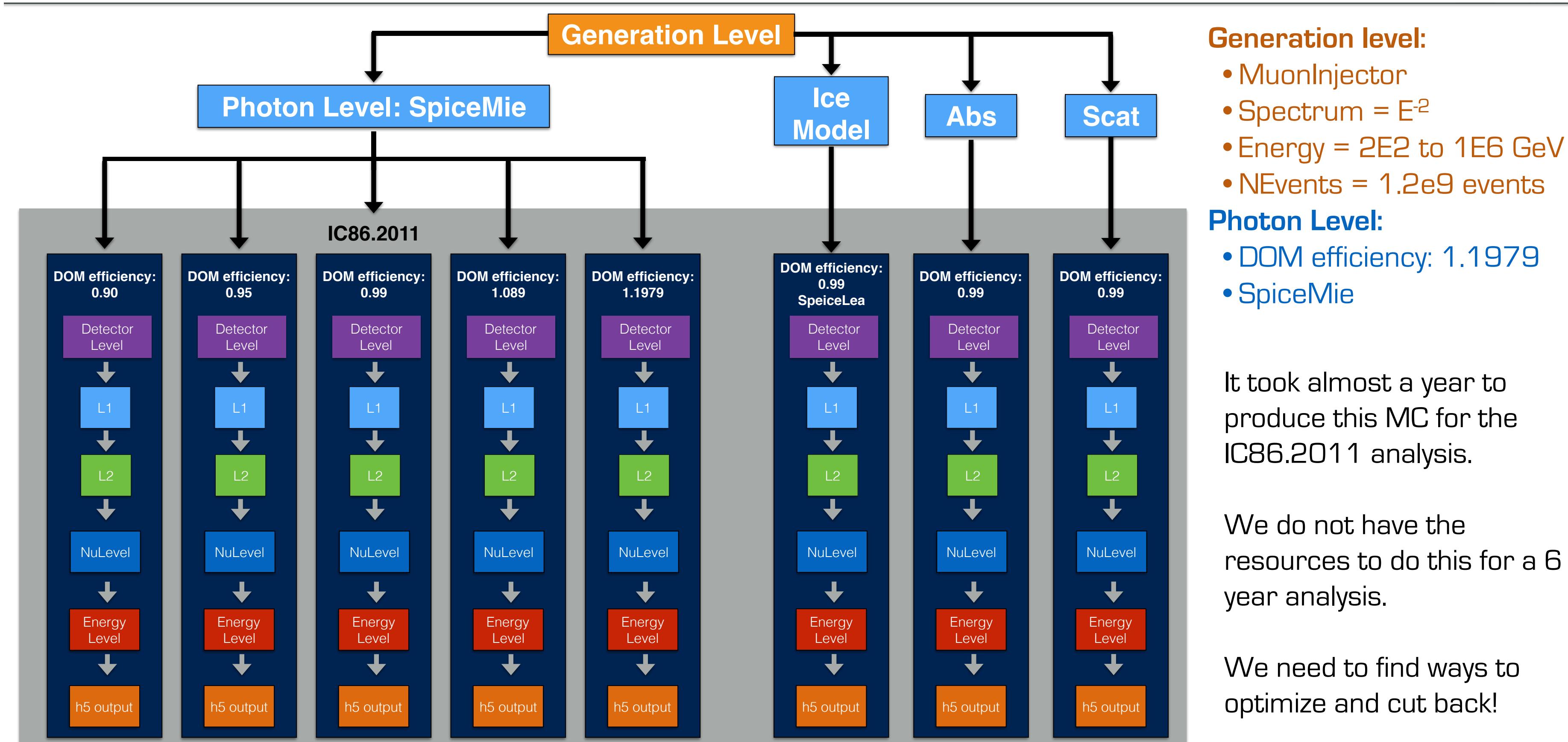
Reconstruction



Simulating Systematic Uncertainties

Example: High-Energy Sterile Neutrino MC Generation

Spencer N. Axani



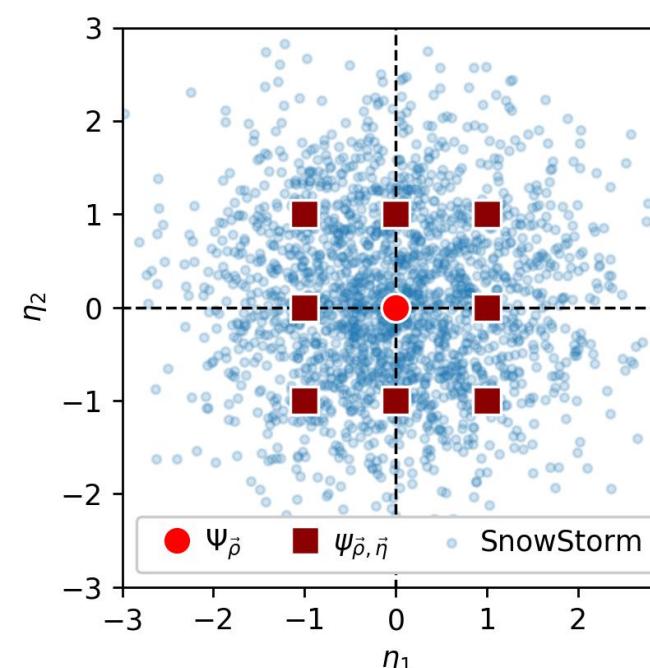
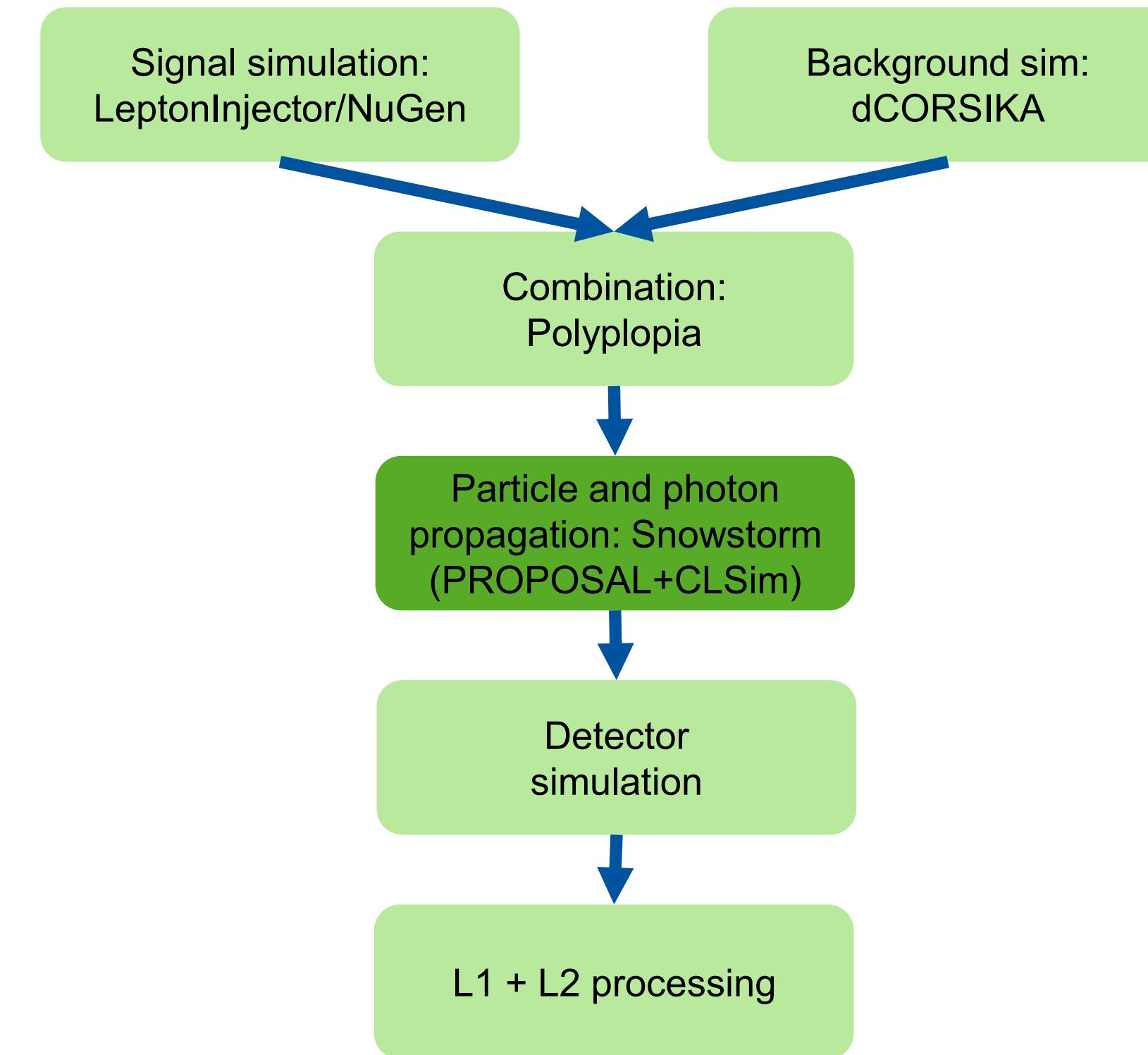
SnowStorm

https://events.icecube.wisc.edu/event/118/contributions/6499/attachments/5362/6082/DiffuseParallel_Brussels_SnowStormMCGlobalfit.pdf

Erik Ganster

SnowStorm Simulation Chain – SnowStorm

- Based on “standard” simulation chain
- Merge of signal+background I3MCTrees before any particle or photon propagation
→ Ensures that all particles get treated/propagated with the exact same parameters/settings further on
 - Main SnowStorm simulation step:
 - Particle (muon) propagation with PROPOSAL
 - Photon propagation using CLSim
 - Perturbing the ice model properties for chunks of frames using the *SnowStorm perturber*



SnowStorm short: Continuous variation of nuisance parameters (detector systematics) (blue) instead of discrete sets for specific values (red)

flow of experimental and simulation data

