



ESCAPE: The study of the solar corona from Dome C, Antarctica

G. Capobianco and the ESCAPE team

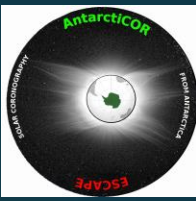
INAF – OATo, Turin (Italy)
UniFi, Florence (Italy)
NASA/JPL, Pasadena (USA)
LATMOS, Paris (France)

gerardo.capobianco@inaf.it

ESCAPE: The study of the solar corona from Dome C, Antarctica

OUTLINE

- The ESCAPE Project
- Observations of the solar corona
- Description of AntarctiCor
- ESCAPE results (Dome C sky brightness)
- ESCAPE results (solar corona data)
- Conclusions



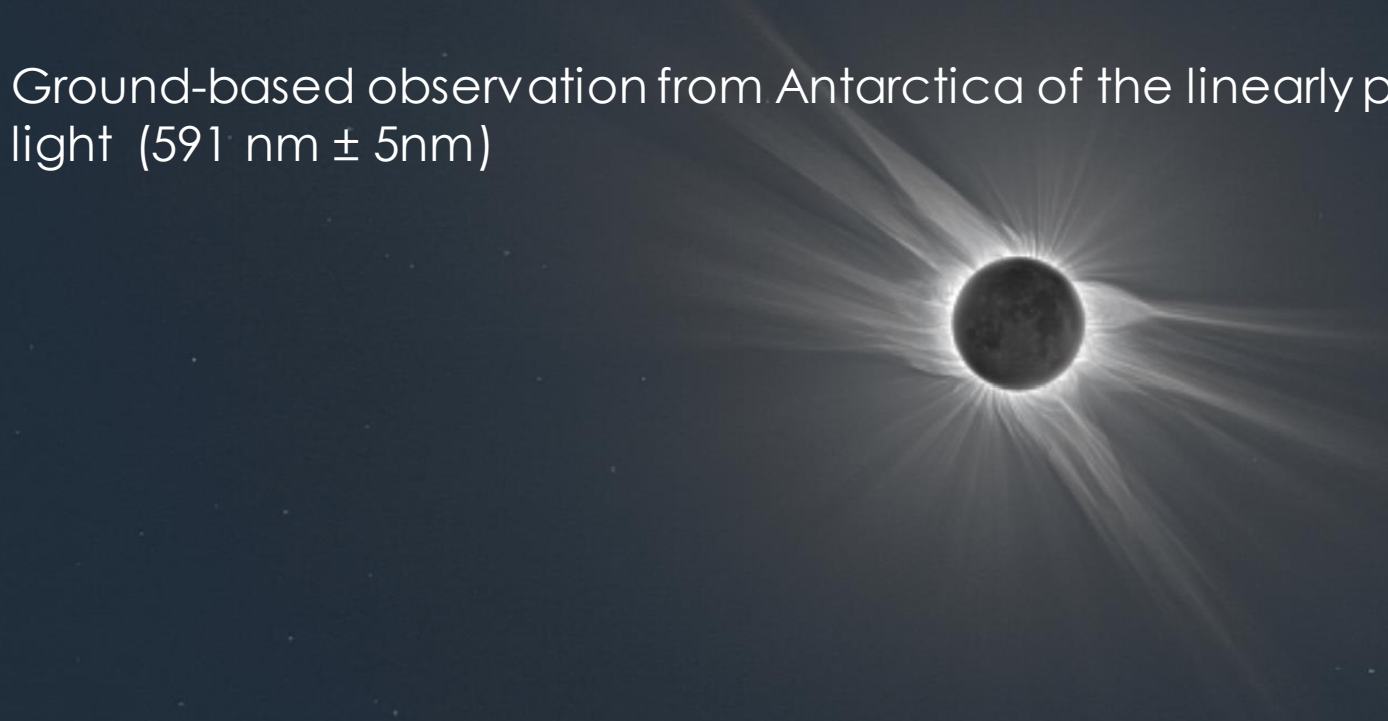
1

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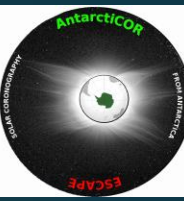
THE ESCAPE PROJECT

ESCAPE: Extreme Solar Coronagraphy Antarctic Program Experiment

Ground-based observation from Antarctica of the linearly polarized solar corona in white-light ($591 \text{ nm} \pm 5 \text{ nm}$)



AntarctiCor is the name of the coronagraph designed and optimized for the purpose

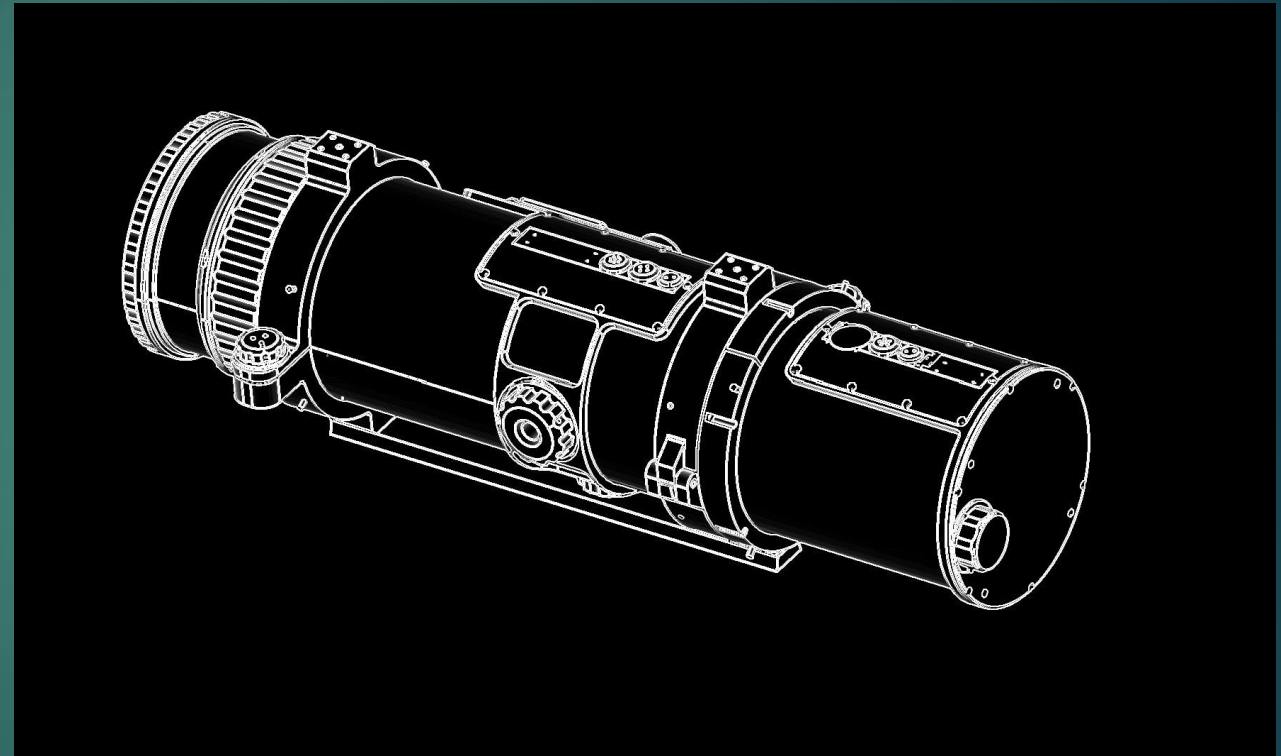
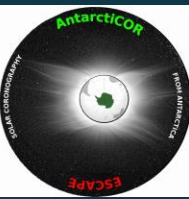


2

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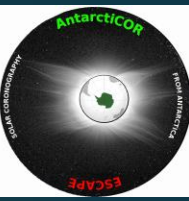
ESCAPE TIMELINE

3

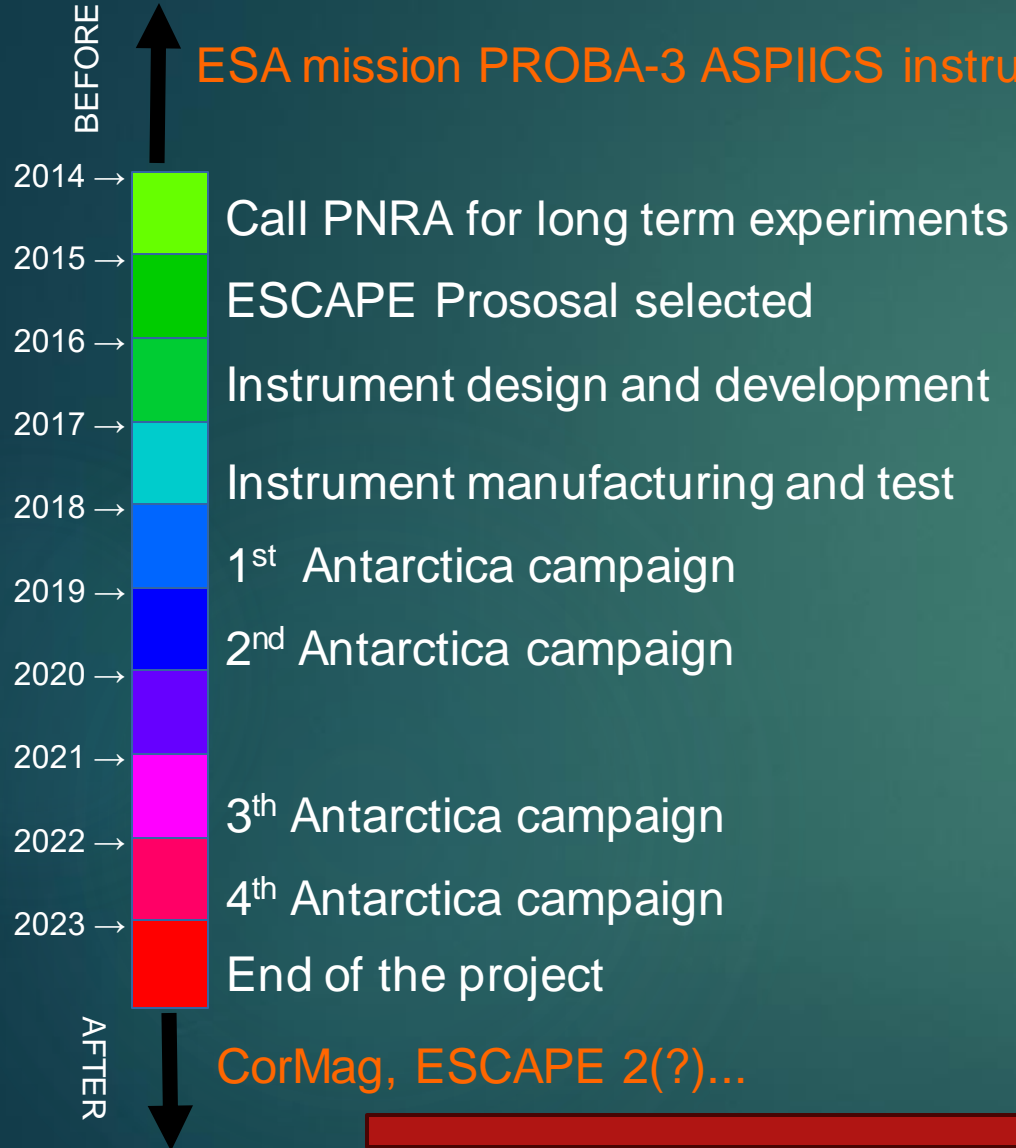


ESCAPE: The study of the solar corona from Dome C, Antarctica

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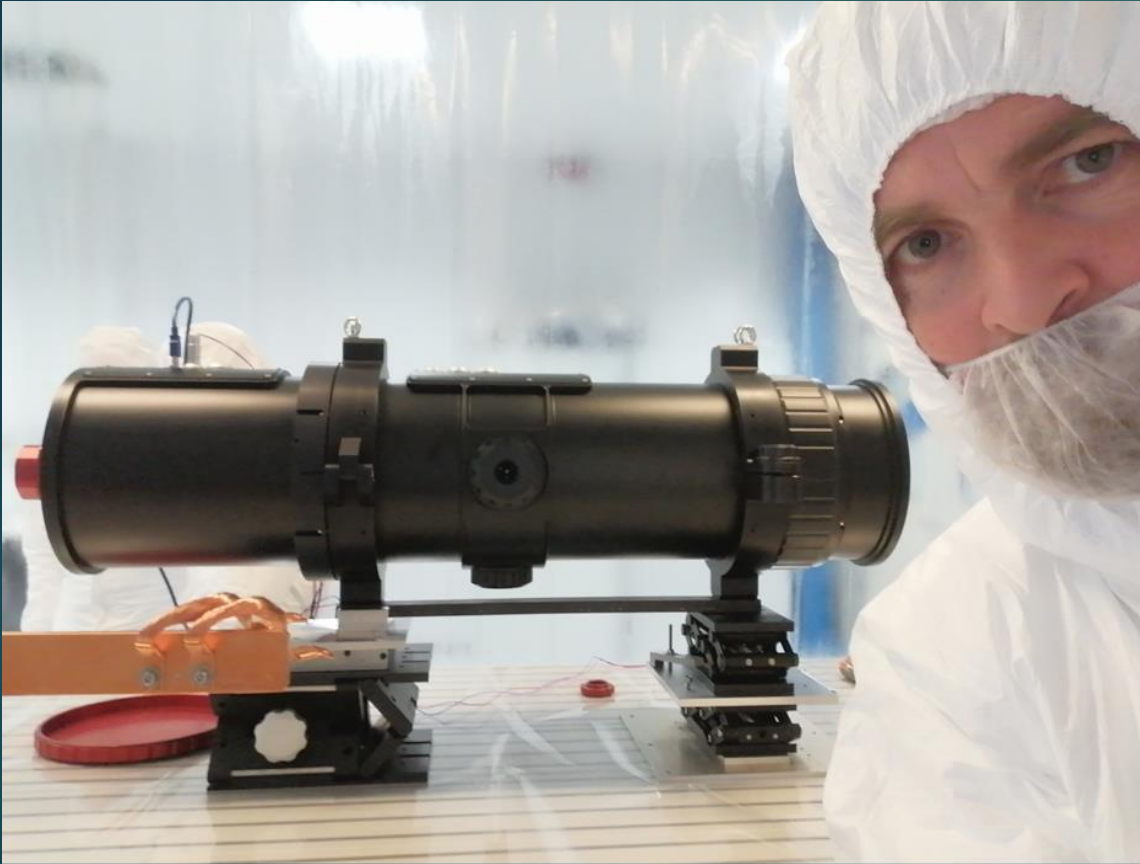
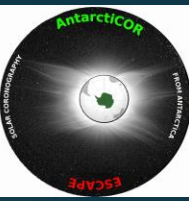
ESCAPE TIMELINE



ESA-PROBA3 Mission Concept – Courtesy of ESA

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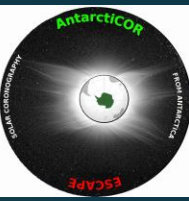


ESCAPE TIMELINE in SNAPSHOT

September 2018 – AntarctiCor tests with the sun simulator in Turin, Italy

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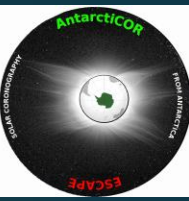
ESCAPE TIMELINE in SNAPSHOT



January 2019 – First installation of ESCAPE in Antarctica on the ASTEP mount

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4



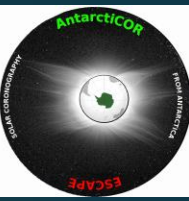
ESCAPE TIMELINE in SNAPSHOT



October 2019 – Test of the mount in the OATo workshop

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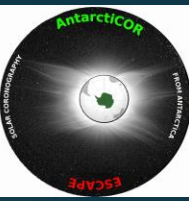
ESCAPE TIMELINE in SNAPSHOT



December 2019 – Installation of ESCAPE in the Baader Planetarium @ Concordia

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ESCAPE TIMELINE in SNAPSHOT



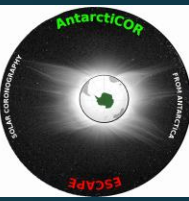
November 2022 – The ESCAPE installation for the 2022/2023 campaign

20/09/2023

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ESCAPE TIMELINE in SNAPSHOT

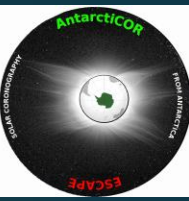
4



January 2023 – Coronagraphic observations in progress...

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SCIENTIFIC OBJECTIVES

1. Polarimetric observations of the solar corona in the spectral broad-band K-corona emission (586 – 596 nm). Retrieve the electron density required by mostly of the coronal diagnostics

2. Demonstrate that Concordia is one of the few existing sites for coronal observations by exploiting Dome C unique low sky brightness

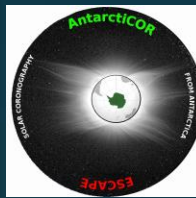
3. Field-test of a space coronagraph (ASPIICS) for a solar mission of the European Space Agency (PROBA-3) to be launched in June 2024 and for the **stratospheric balloon** platform HEMERA launched summer 2022 and summer 2023.



Measurements of interest for Space Weather projects

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SOLAR CORONA

The solar corona is the outermost part of the Sun atmosphere.

$$T_{\text{corona}} (1.E6 \text{ K}) \sim 200 T_{\text{sun}}$$

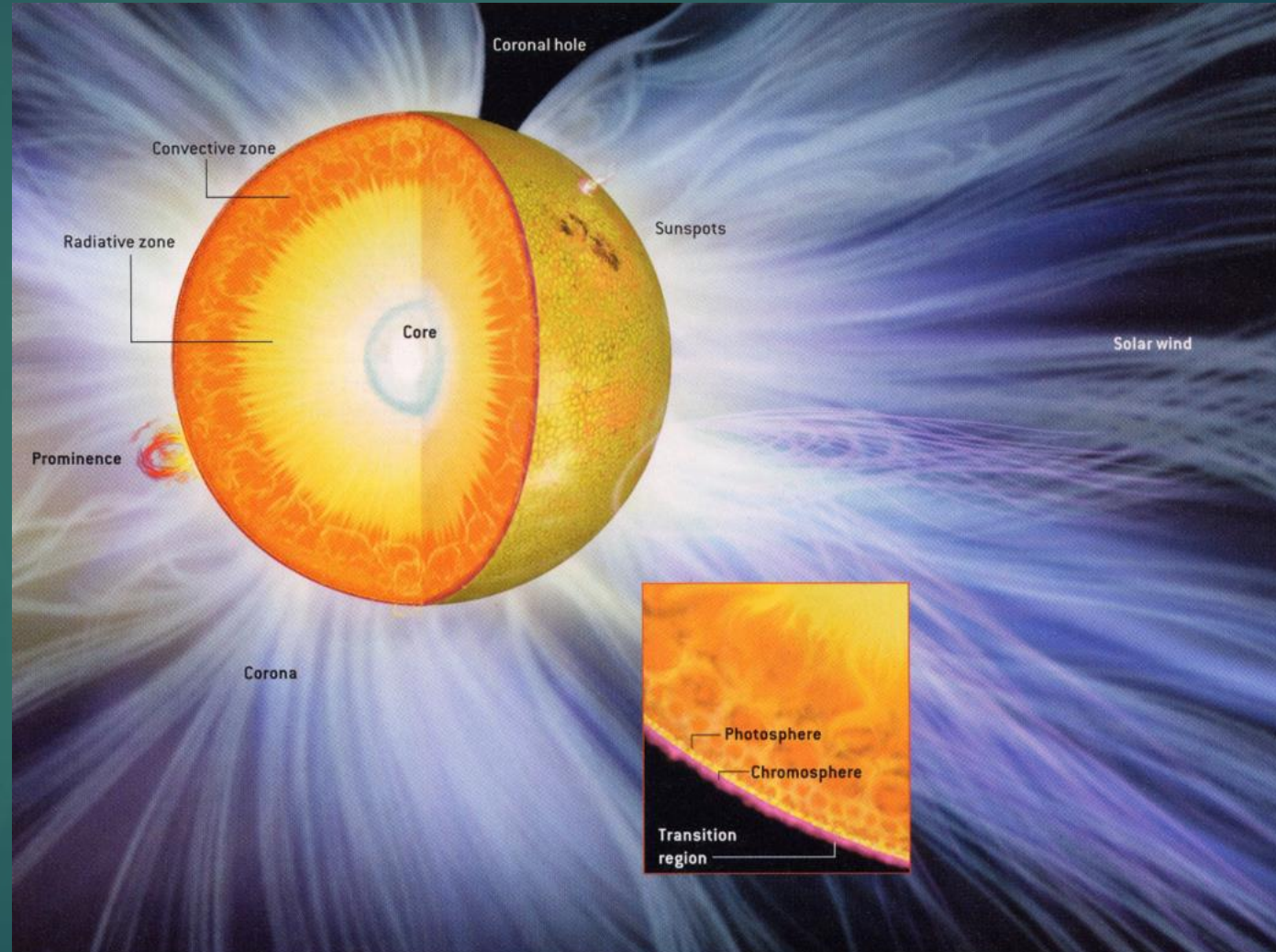
WHY??

$$\rho_{\text{corona}} \sim 1.E-7 \rho_{\text{sun}} \rightarrow$$
$$\rightarrow B_{\text{corona}} \sim 1.E-6 B_{\text{sun}}$$

(difficult to observe)

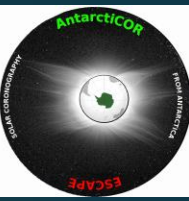
The particles expelled by the Sun (95% p^+ and e^-) are accelerated in the corona (from 300 to 800 km/s) and propagates through the solar system (solar wind).

How this mechanism works?



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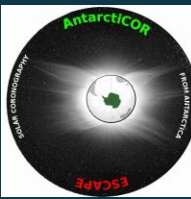
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SOLAR CORONA

The auroras are generated by the interaction of the solar wind with the geomagnetic field



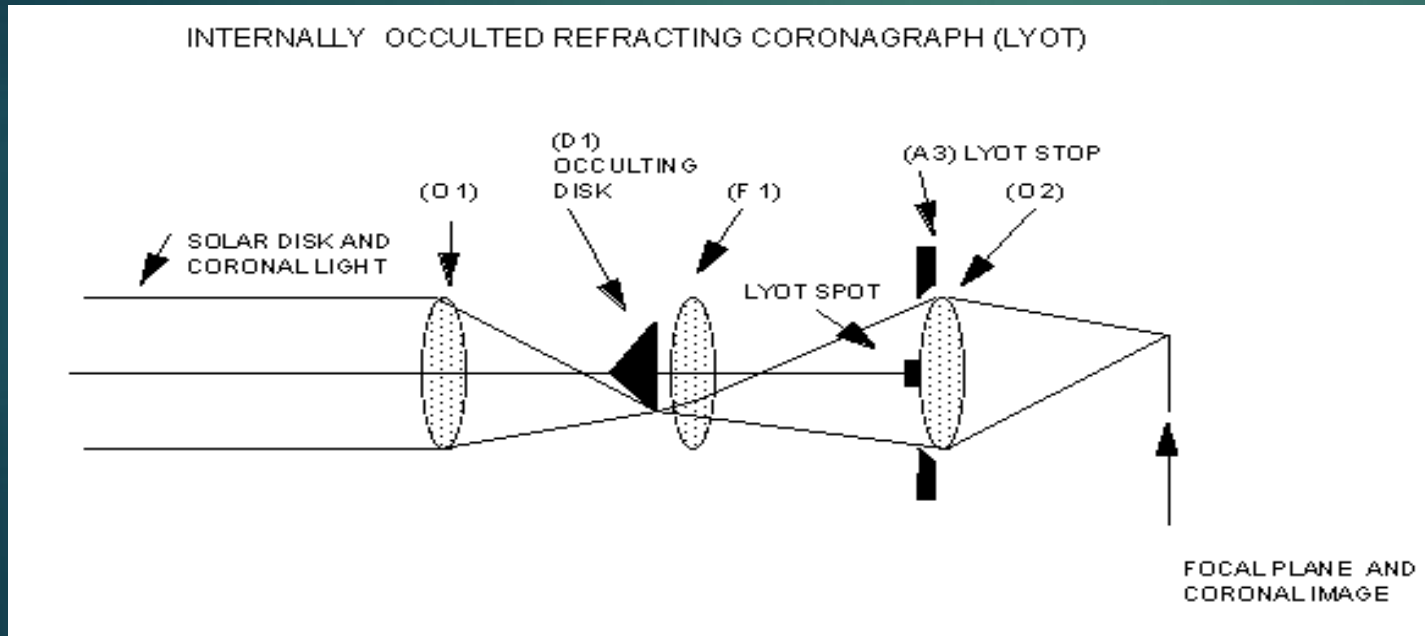


INSTRUMENTATION

In 1930's the french astronomer **Bernard Lyot** invent the coronagraph

Since the coronal light is very faint a coronagraph MUST:

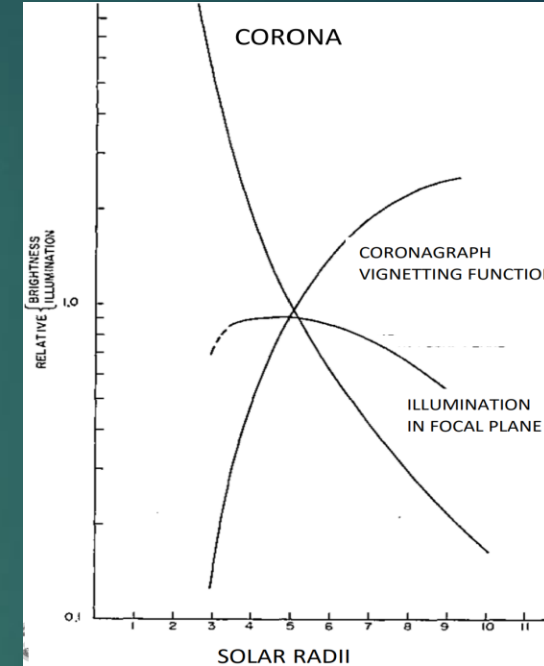
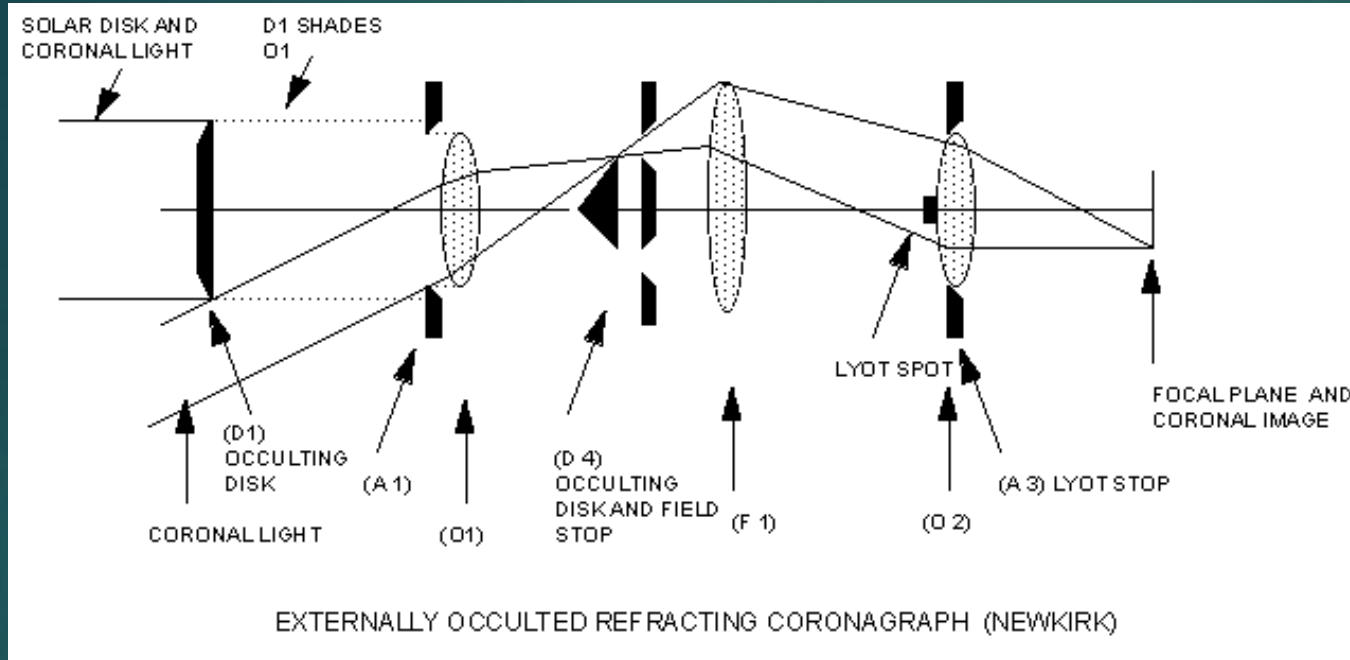
- Block the solar disk light
- Block the light diffracted by the edges of the optical elements



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All the (in operation) space-based solar coronagraphs are externally occulted

INSTRUMENTATION



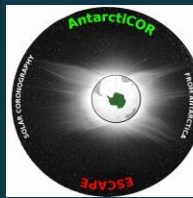
PRO

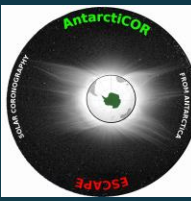
- Very steep vignetting function mitigates the large dynamic range of the coronal brightness
- Only way to achieve low stray light in the outer corona

CONS

- Lower end of FOV not lower than 1.5 Rs
- Low throughput at the lower end of the field of view

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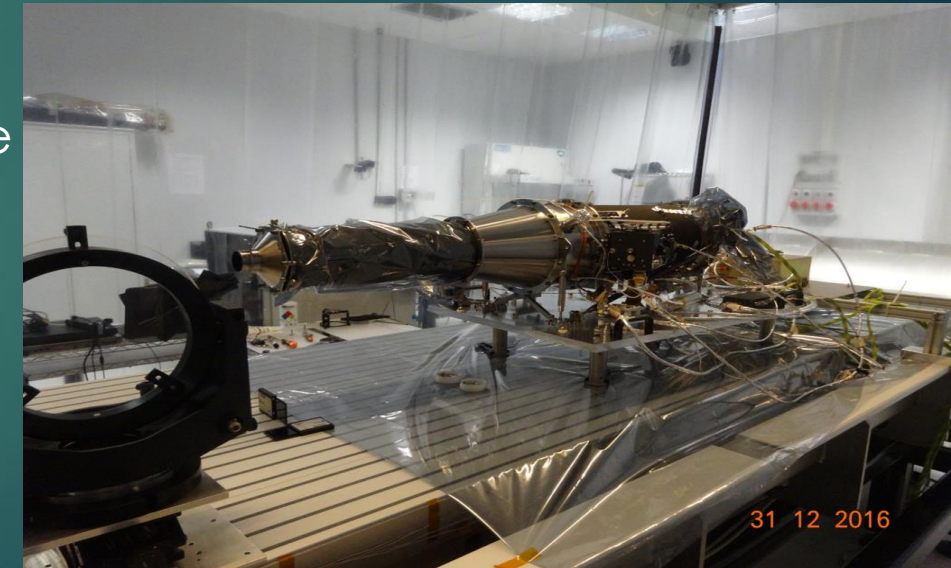


Ground-based solar coronagraphs:

- K-Cor @Mauna Loa (Hawaii) performing a few hours/day observations – closed from November 2022 after the Mauna Loa volcano eruption
- ESCAPE @ Dome C (Antarctica) performing seasonal observation for 12h/day

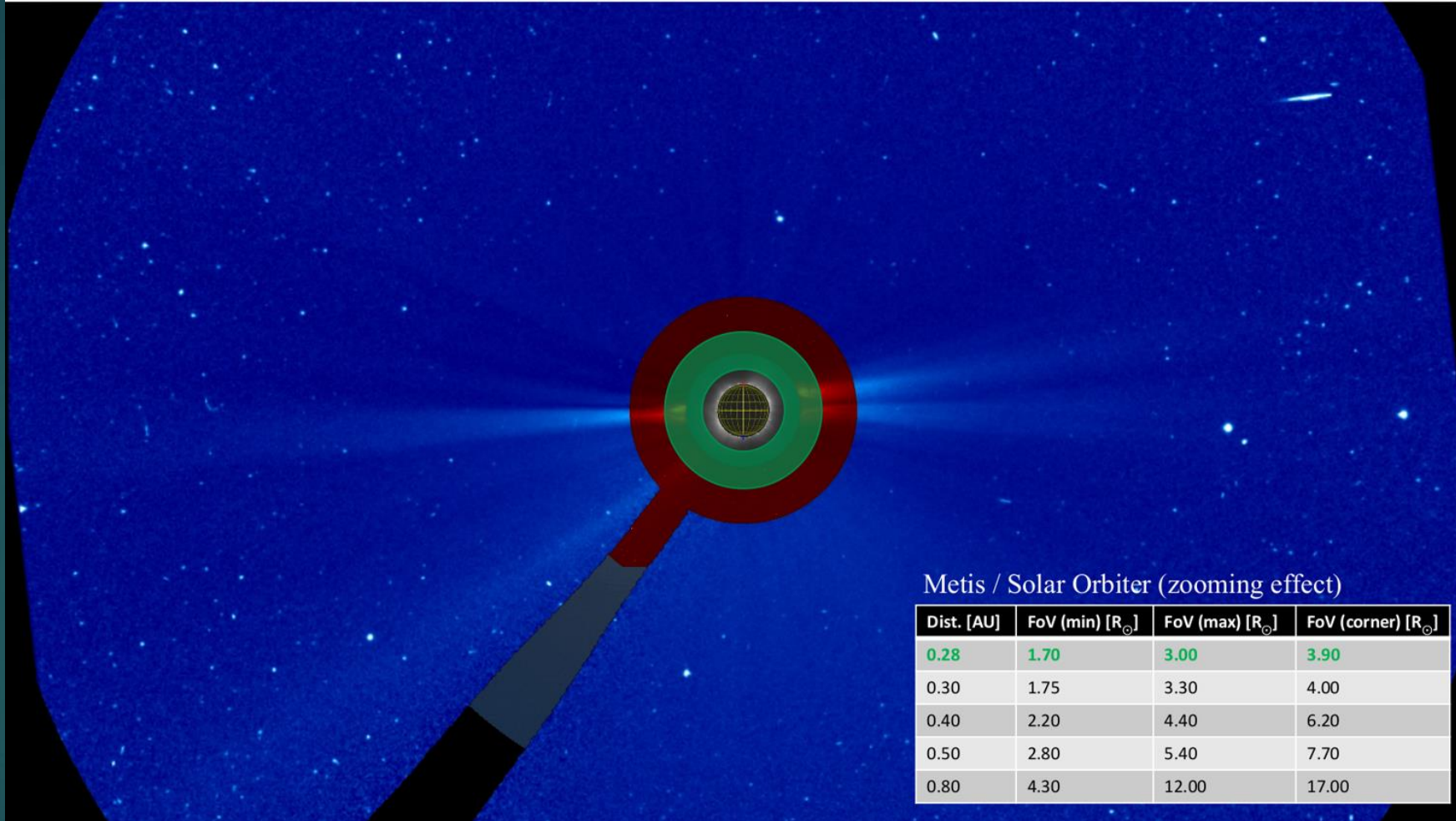
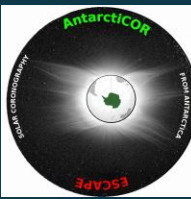
Space-based solar coronagraph:

- LASCO C2/C3 on-board SOHO mission from 1995
- Metis on-board Solar Orbiter launched Feb. 2020
- A few days ago launched Aditya with Visible Emission Line Coronagraph (VELC)

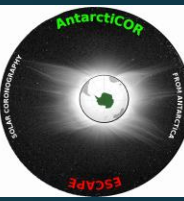
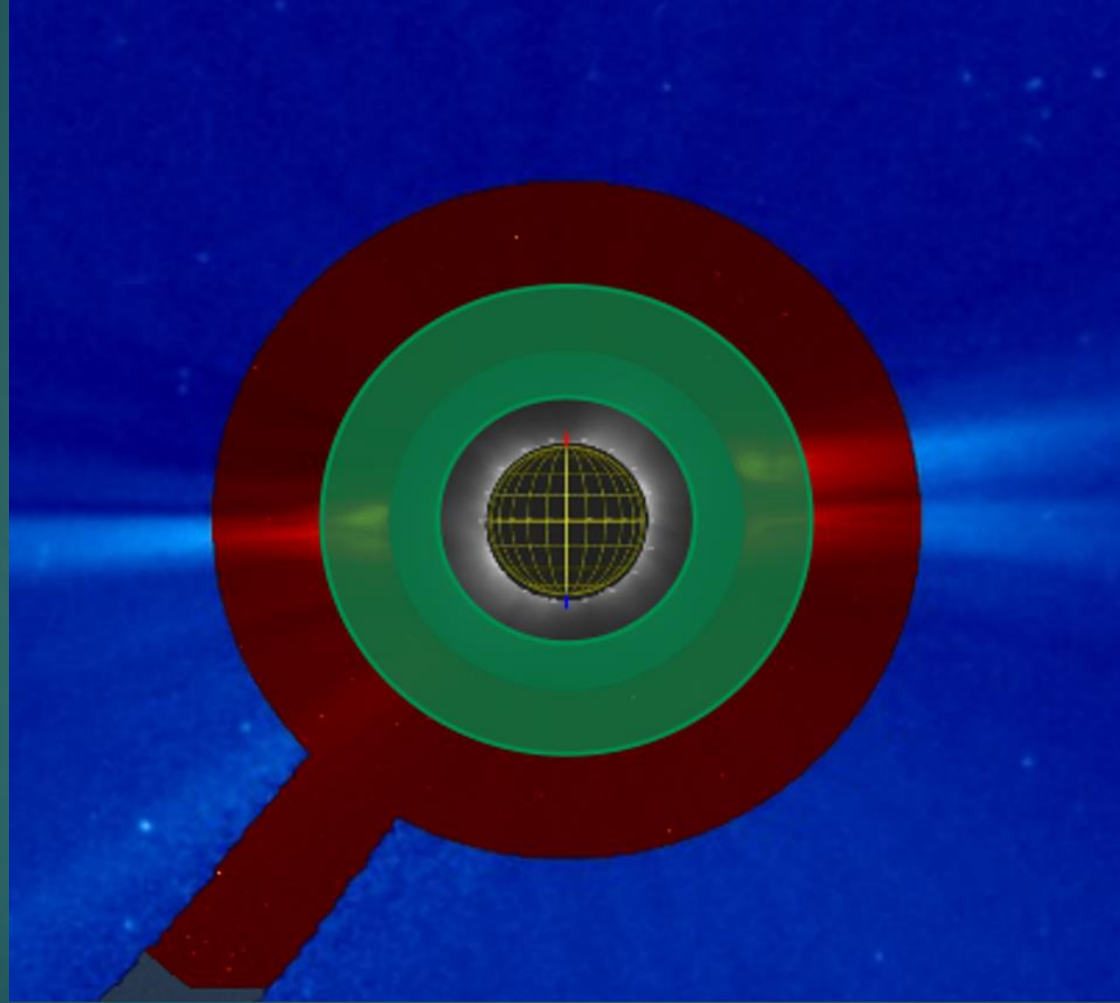


IMPORTANCE of GROUND OBSERVATIONS

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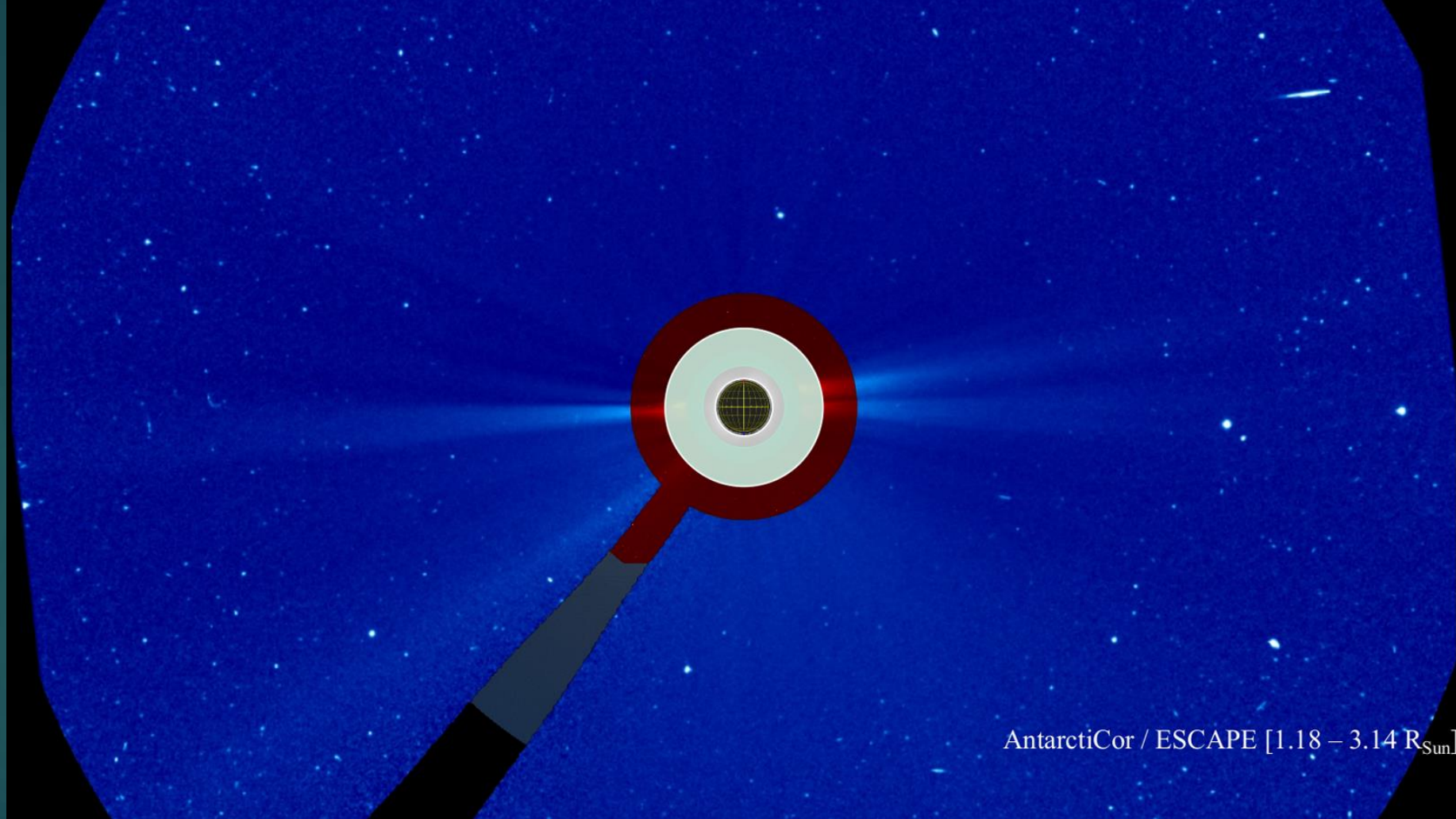


IMPORTANCE of GROUND OBSERVATIONS

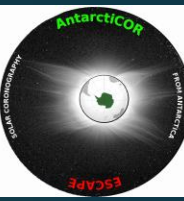


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IMPORTANCE of GROUND OBSERVATIONS



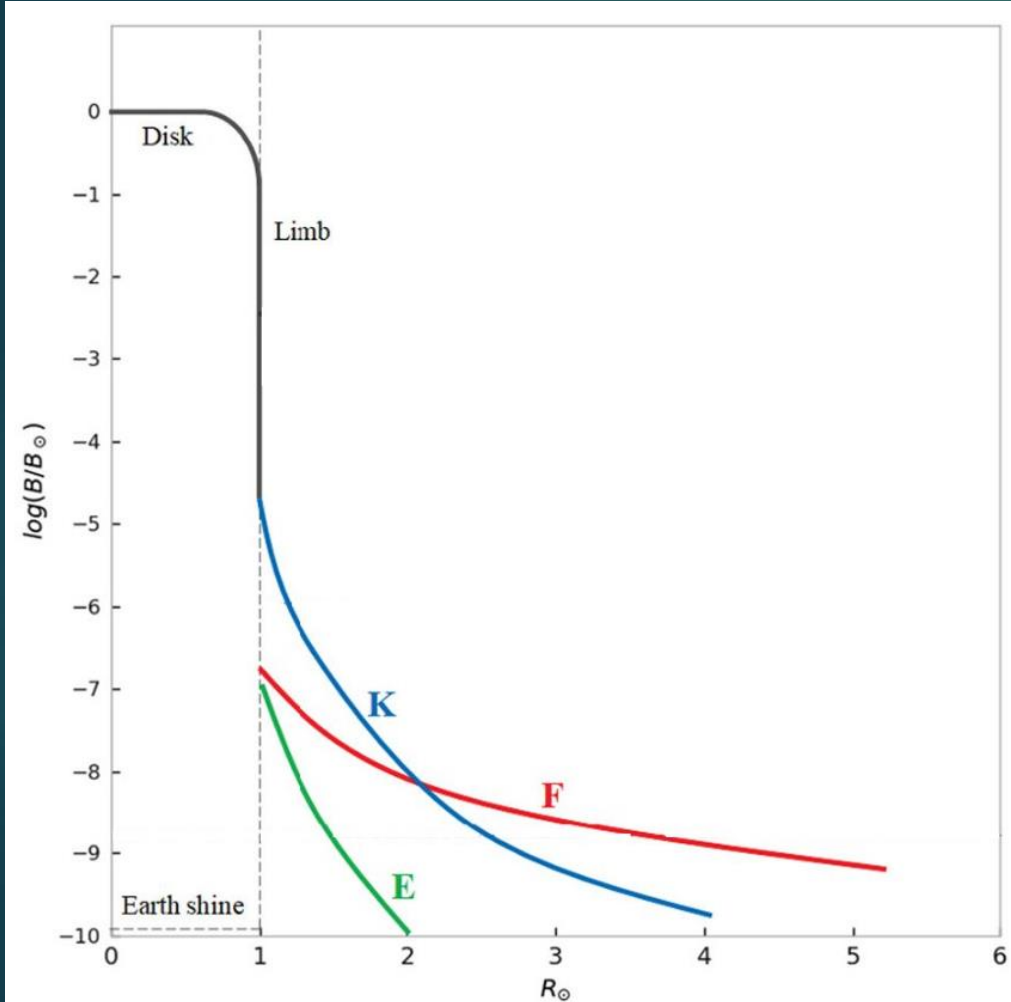
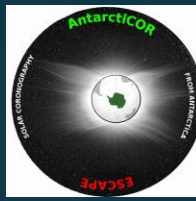
AntarctiCor / ESCAPE [1.18 – 3.14 R_{Sun}]



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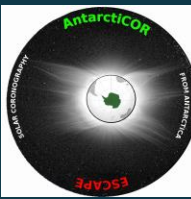
CORONOGRAPHIC SITE SELECTION

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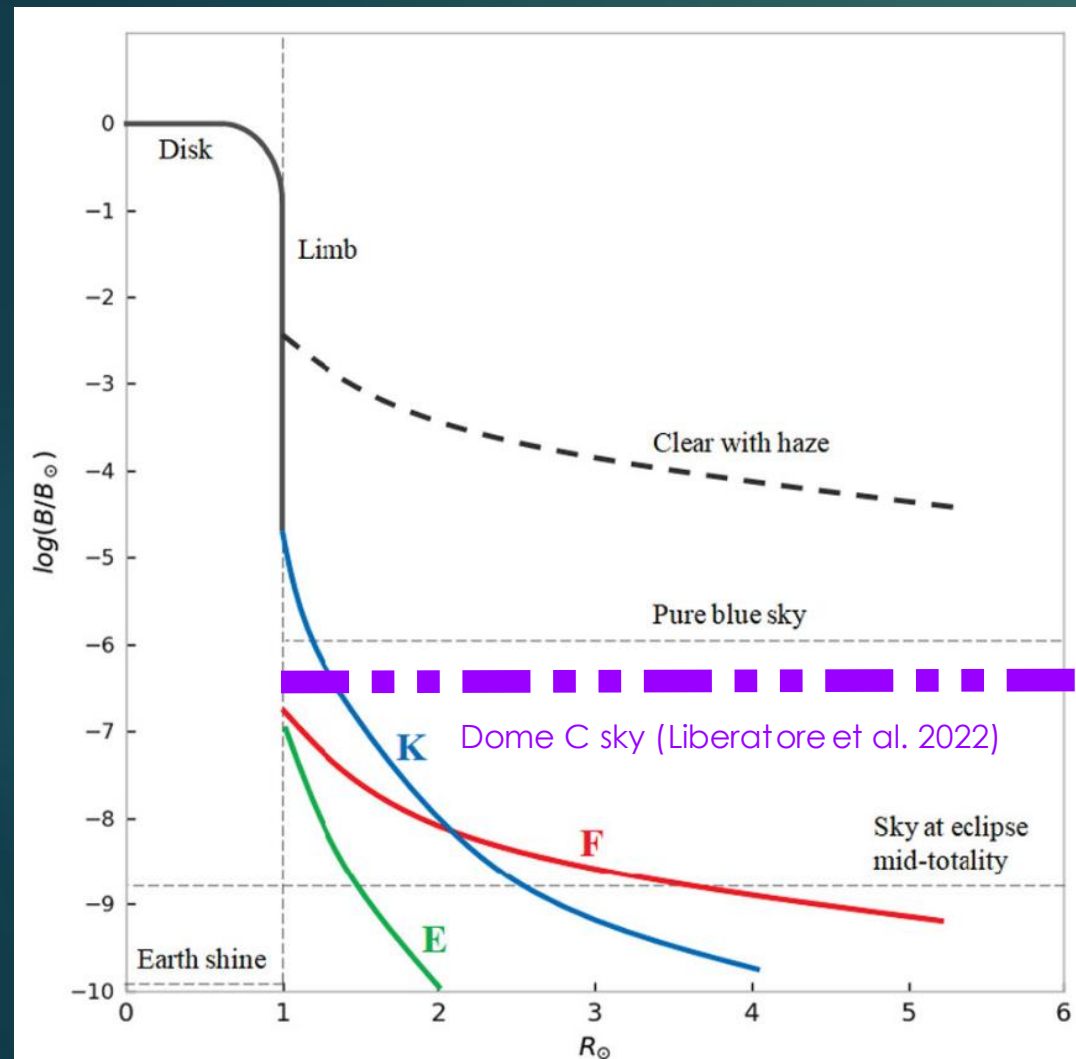
CORONOGRAPHIC SITE SELECTION

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By definition of coronagraphic sky

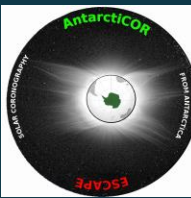
$$B_{\text{sky}} \leq 1.E-6 B_{\text{sun}}$$



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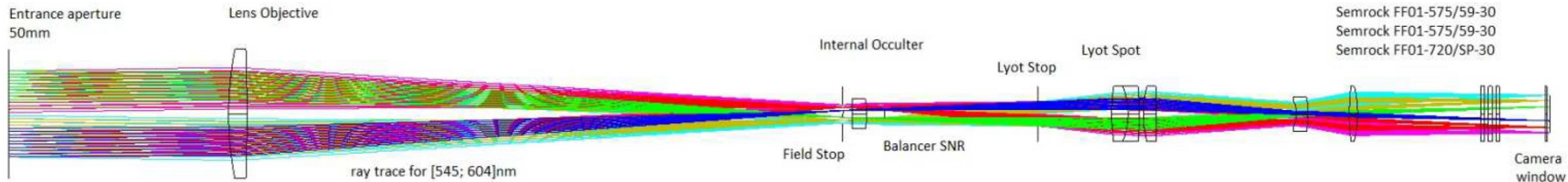
AntarctiCor Optical Design

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| AntarctiCor Instrument Parameters | |
|-----------------------------------|---|
| Telescope design | Classic internally-occulted Lyot coronagraph [8] |
| Aperture | 50 mm |
| Eff. Focal Length | 700 mm |
| Spectral Ranges | 591 nm \pm 5 nm K-corona; 530.3 nm \pm 0.25 nm FeXIV "green line" |
| Camera type | Interline transfer CCD by PolarCam [®] , mod. U4 [10] |
| Camera format | 1950 \times 1950 |
| Pixel size | (7.4 μ m) ² |
| Plate scale | 4.3 arcsec/pixel (8.6 arcsec/polarization super-pixel) |
| Field-of-View | $\pm 0.6^\circ = \pm 2.24 R_\odot$ ($\pm 0.84^\circ = \pm 3.14 R_\odot$ along CCD diagonals) |
| Polarization analysis | spatial modulation by linear micropolarizer array on CCD sensor |

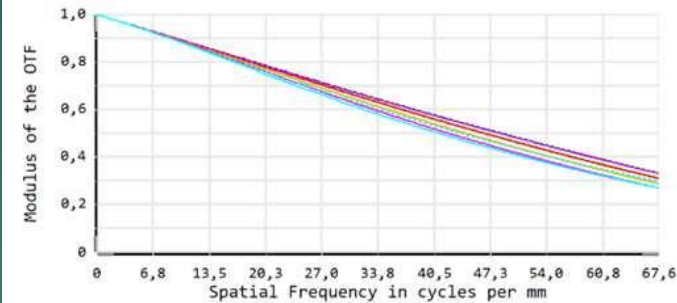
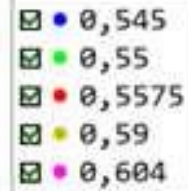
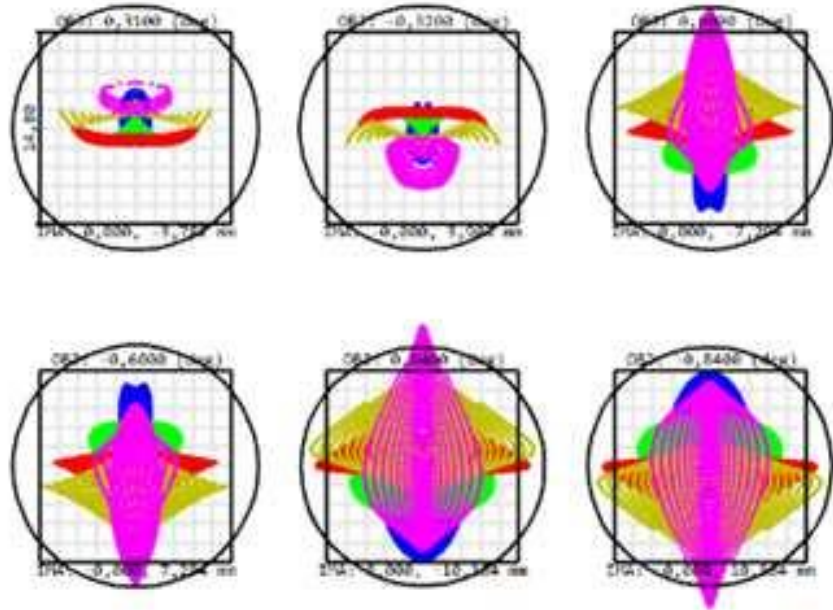
Fineschi et al., 2019



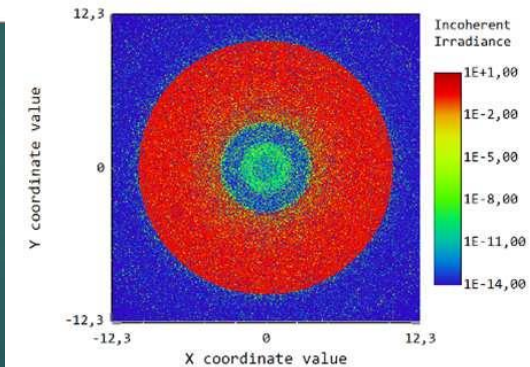
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AntarctiCor Optical Design

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Polychromatic Diffraction MTF
 Refurb. of ASPIICS v15 - OHB Italia
 23/11/2017
 Data for 0,5450 to 0,6040 μm .
 Surface: Image
 Legend items refer to Field positions



Detector Image: Irradiance
 P. Sandri - OHB Italia
 24/11/2017
 Detector 66, HSC Surface 1: FLI
 Size 24,976 W X 24,976 H millimeters, Pixels 1024 X 1024 H, Total Hits = 2852568
 Peak Irradiance : 1,0046980 Watts/cm²
 Total Power : 5,7176-01 Watts

Surface: IMA

Spot Diagram

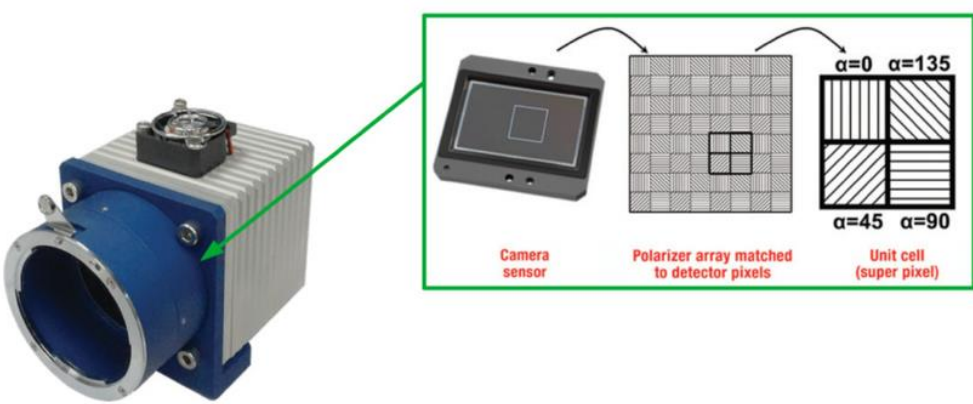
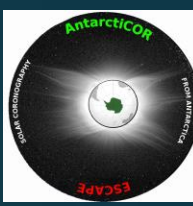
P. Sandri - OHB Italia
 23/11/2017
 Units are μm . Airy Radius: 9,516 μm . Legend items refer to Wavelengths
 Field : 1 2 3 4 5 6
 RMS radius : 1,824 2,072 3,523 3,523 4,300 4,300
 GEO radius : 5,912 5,941 9,186 9,186 10,854 10,854
 Box width : 14,8 Reference : Centroid

Zemax
 Zemax OpticStudio 16 SP2

ANTARTICOR-15112017.imx
 Configuration 2 of 2

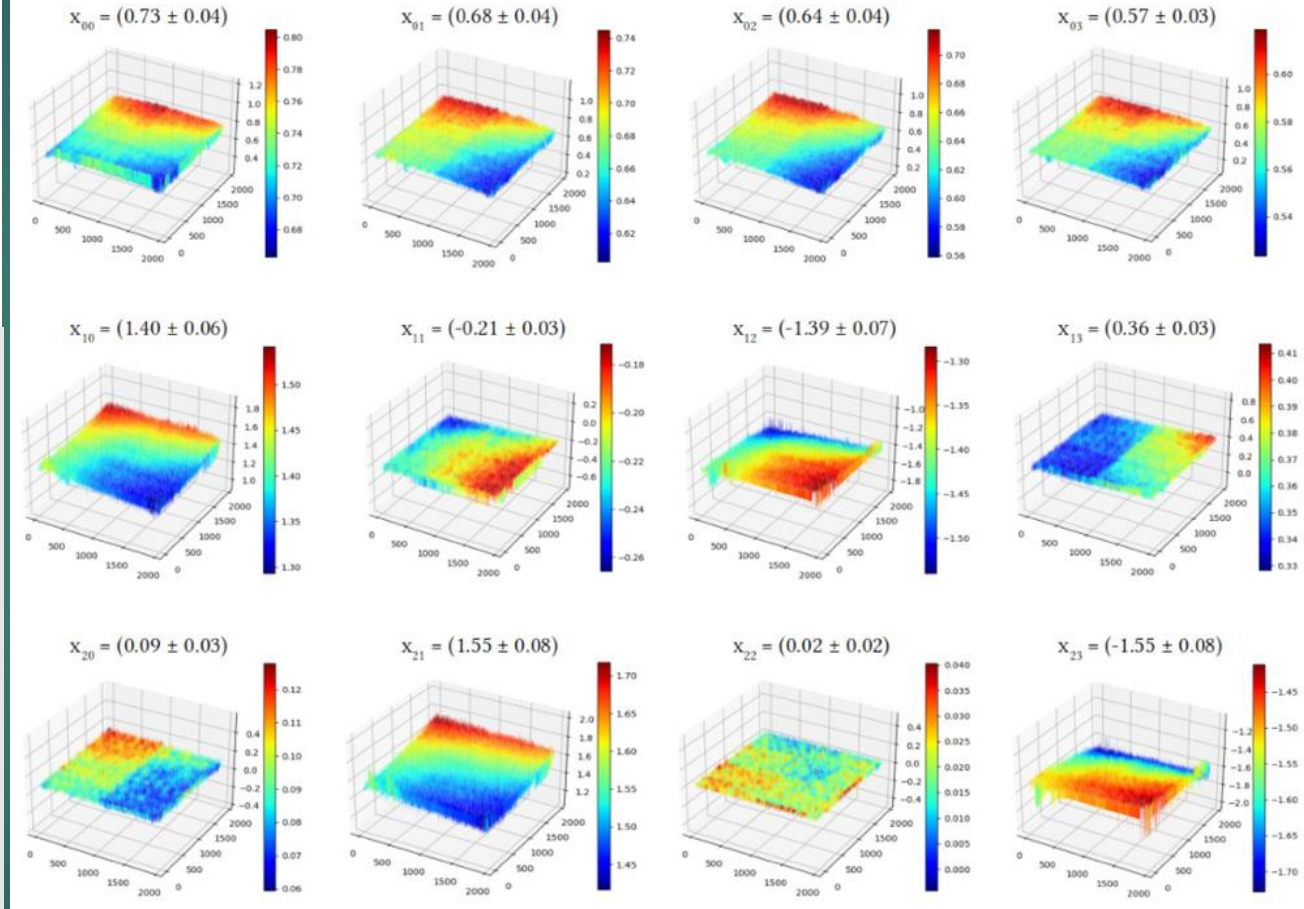
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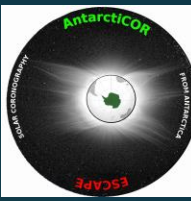


| | |
|-----------------------------|--|
| Sensor Type | Interline transfer CCD ON Semiconductor KAI04070 |
| Micropolarizers orientation | 0°, 45°, 90°, 135° |
| Pixel Size | 7.4 μm \times 7.4 μm |
| Usable Pixels | 1950 \times 1950, 3.8 MP |
| Frame Rate | 14 fps |
| Saturation Capacity | 44 ke- |
| Dark Noise | 3 e-/s |
| Quantum Efficiency | 76% @ 470 nm |
| Physical Envelope | 60 \times 60 \times 95 mm |
| Weight | 210g |
| Power Requirement | 8 W, 12VDC |
| Interface | USB 3.0 |
| Lens Mounting Type | F-Mount |

AntarctiCor Detector



Liberatore et al., 2021



ESCAPE Pointing System

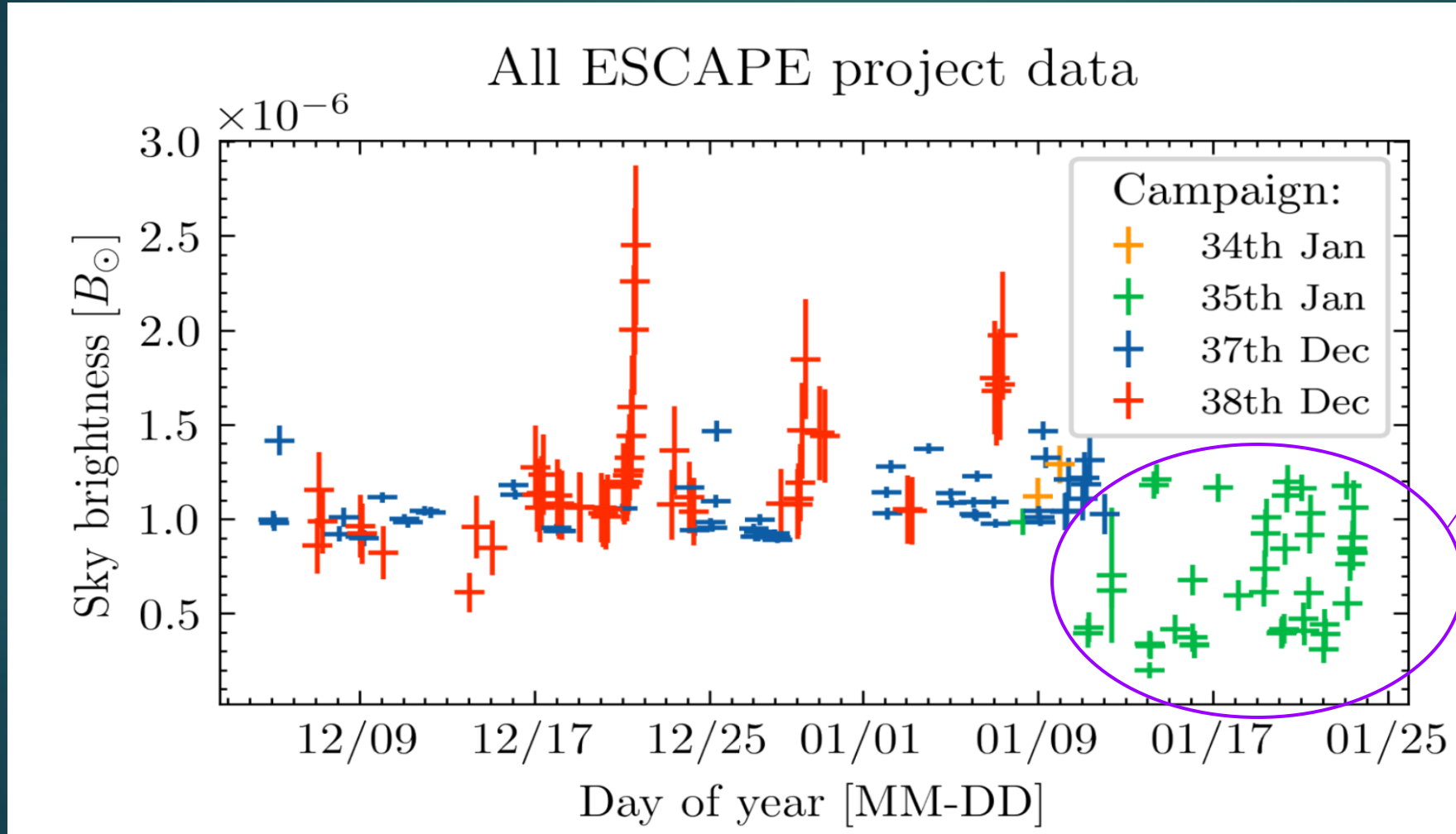
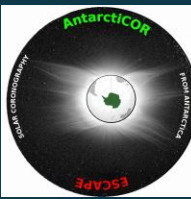
Customized COTS german equatorial mount for extended temperature range (-40°C to +40°C)

Requirements

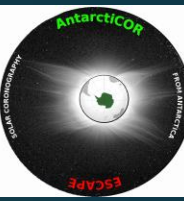
Pointing accuracy: better than 1 arcmin

Tracking accuracy: better than 15 arcsec/min

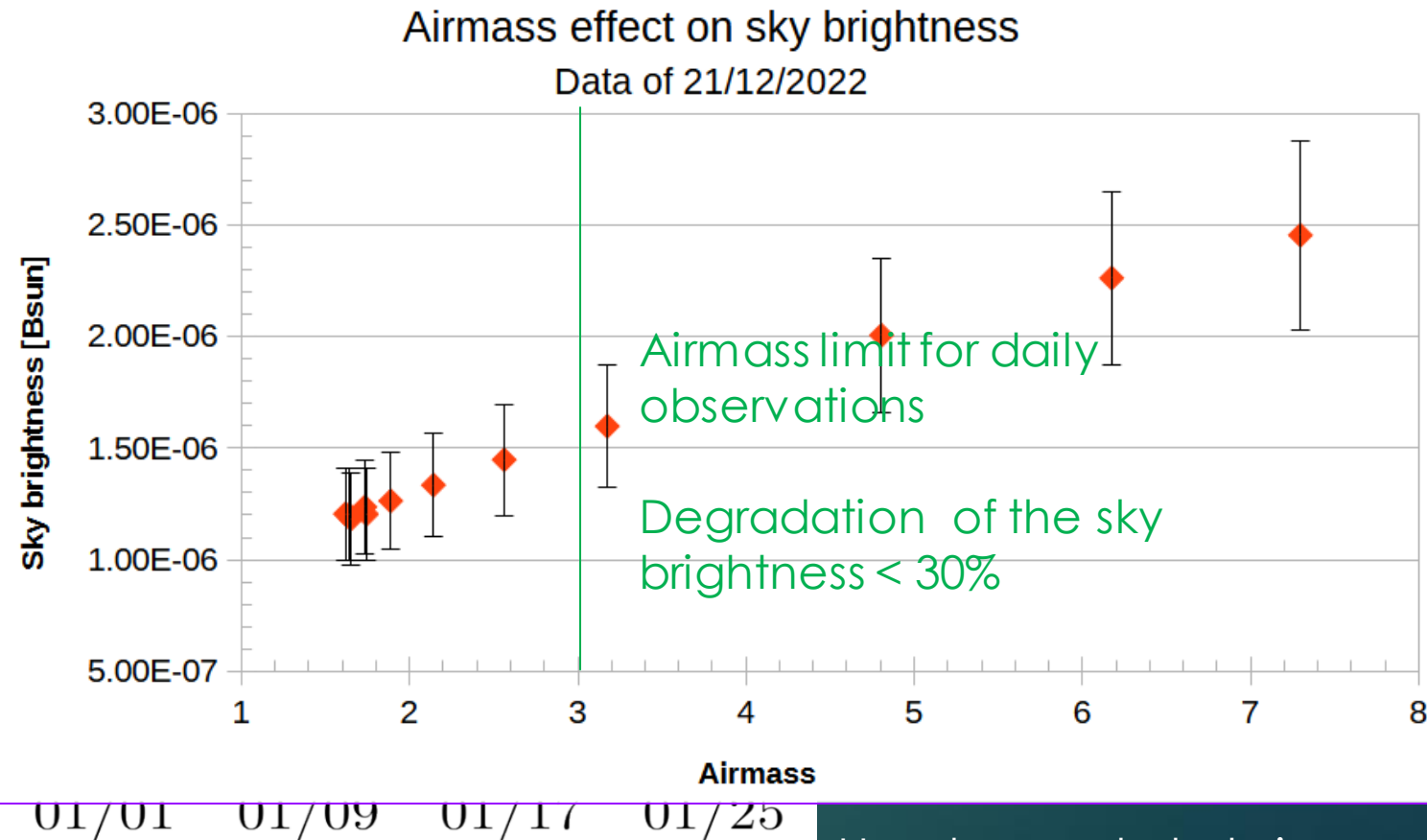
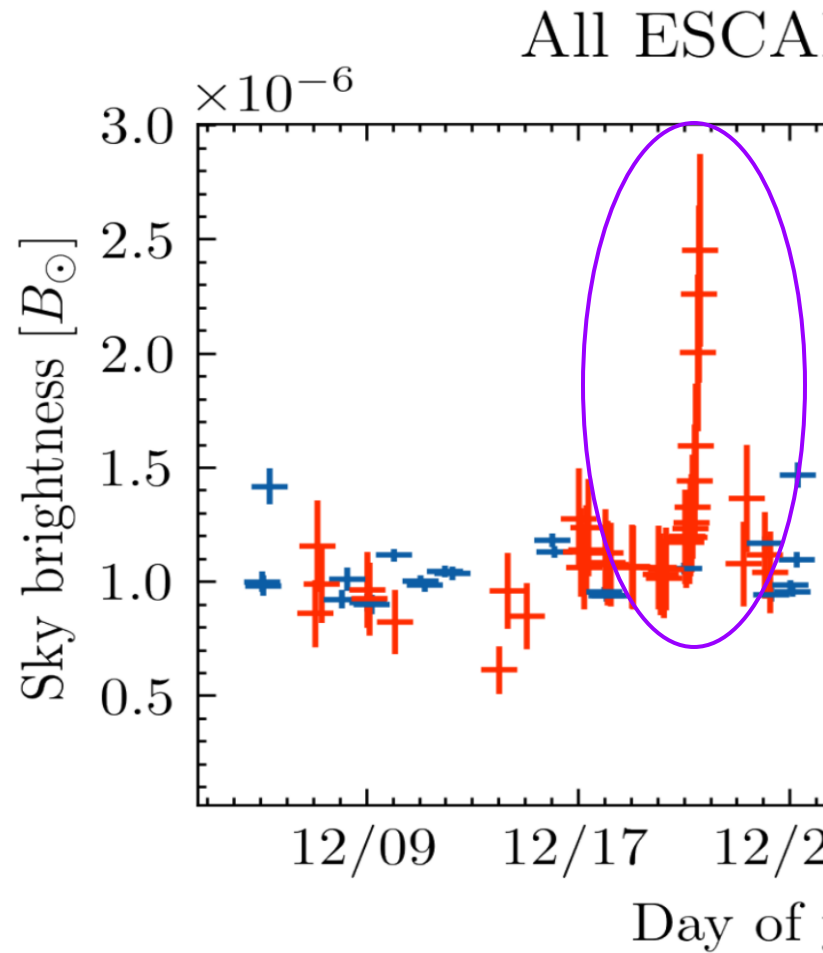




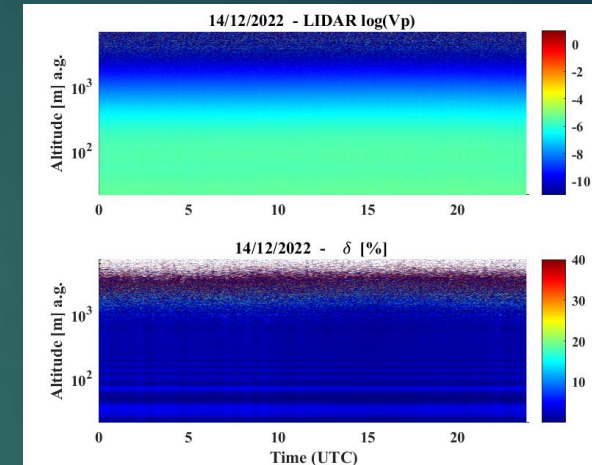
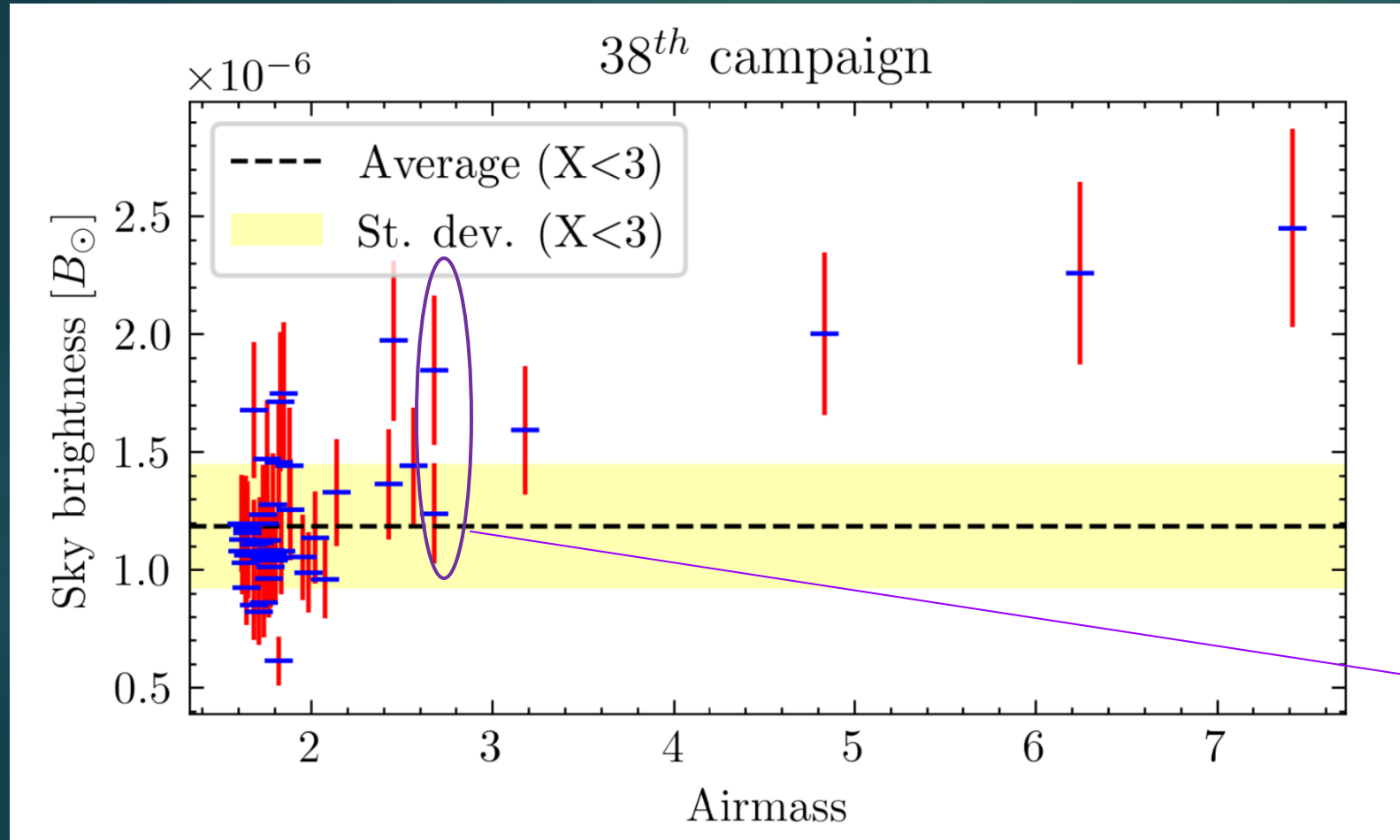
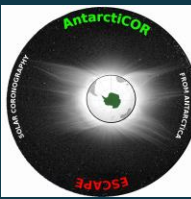
January seems to be the best period for coronal observations



Results – Dome C Sky brightness



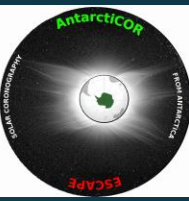
Haudemand et al., in preparation



Check of the atmospheric condition with ICE-OPT data (courtesy of M. Del Guasta)

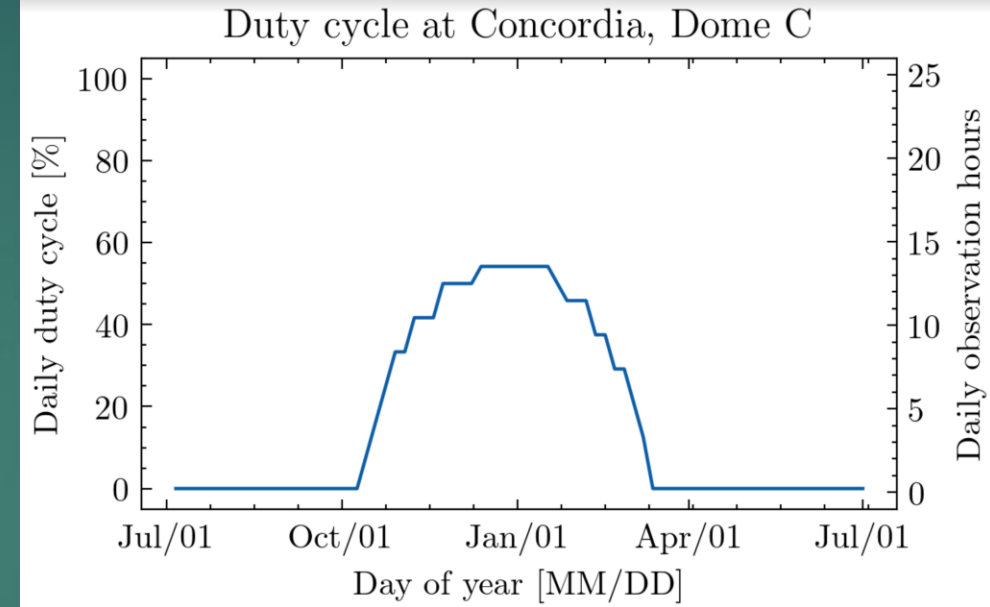
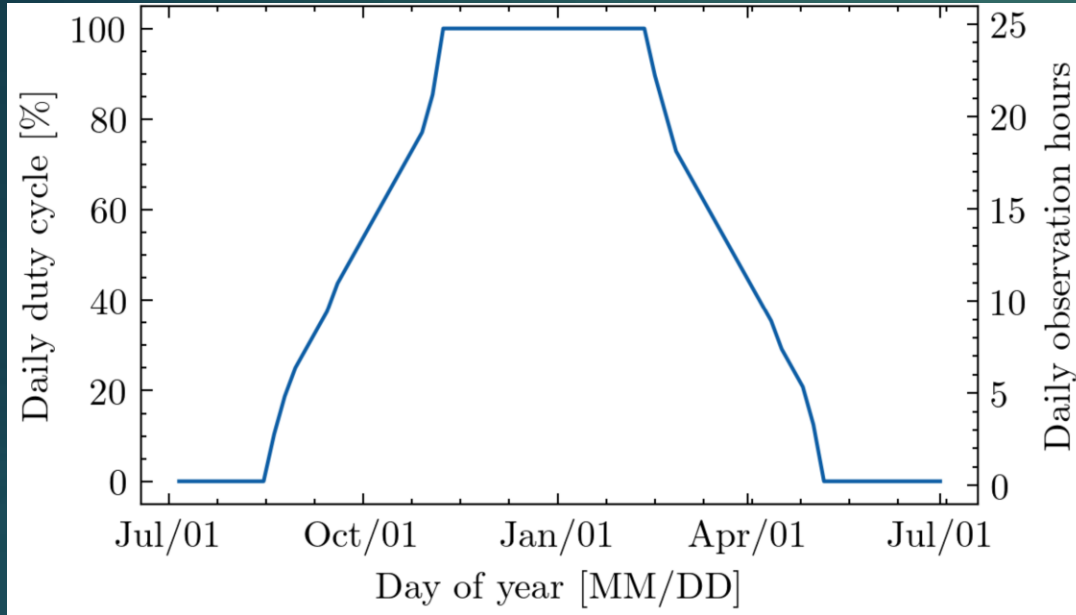
Diamond Dust

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Results – Dome C Sky brightness (duty cycle)

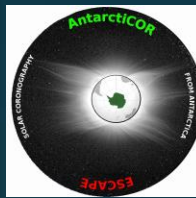


Taking into account the limits in the airmass and the typical conditions we estimate for Dome C a duty cycle of the 10% i.e., approx 850 obs. hours/yr

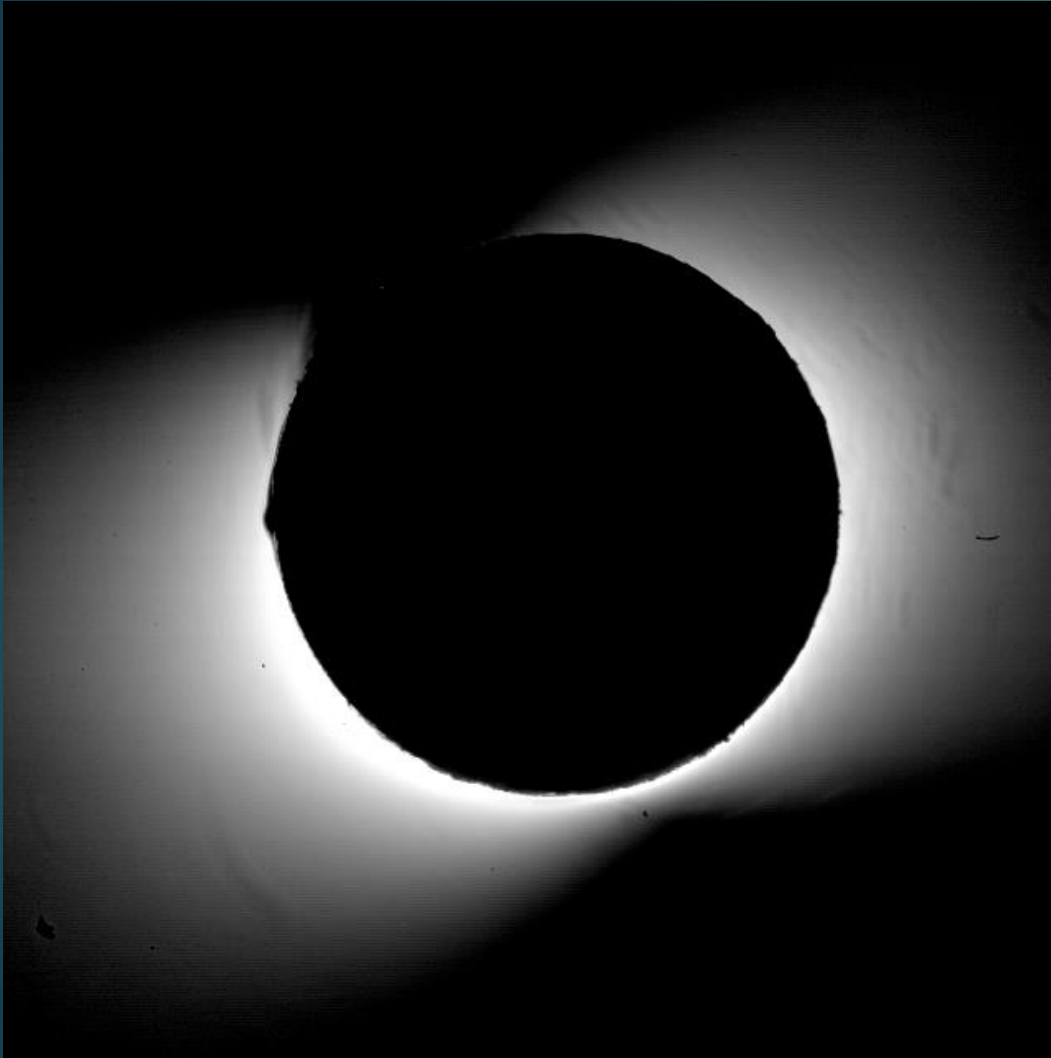


© ENEA-PNRA / IPEV

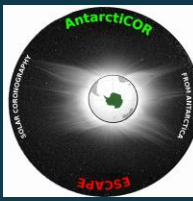




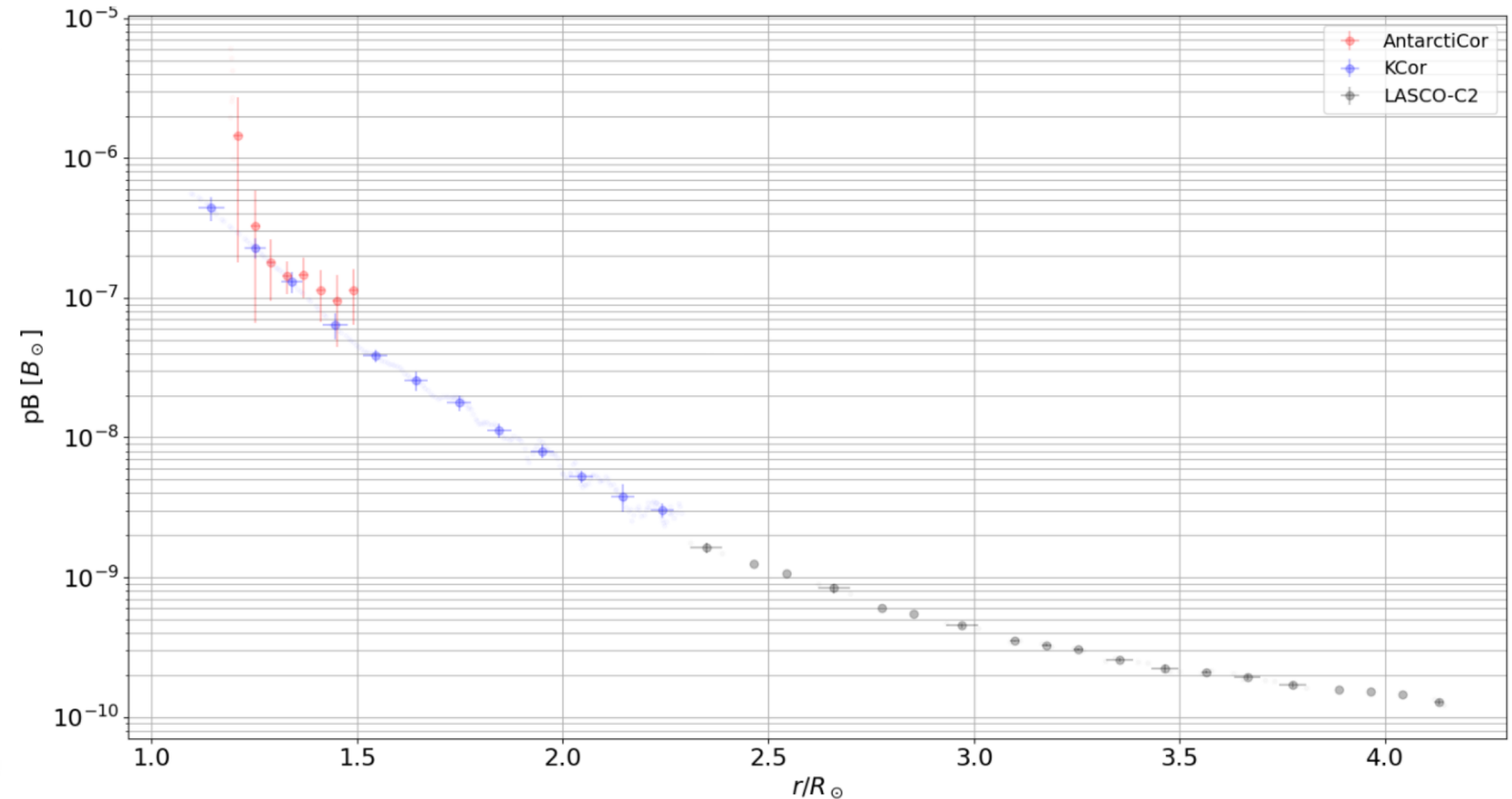
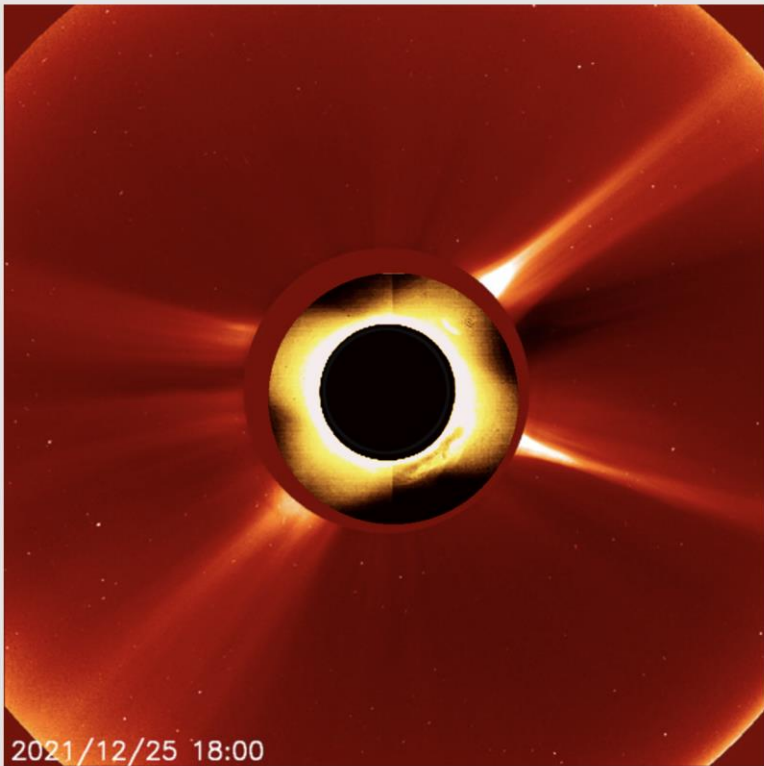
Results – Solar Corona



Solar Corona @ minimum solar activity (23/01/2020)

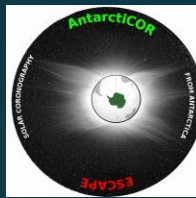


25/12/2021 The new solar cycle (25) is effective!



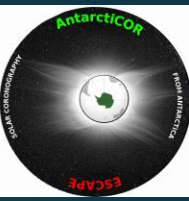
CONCLUSIONS

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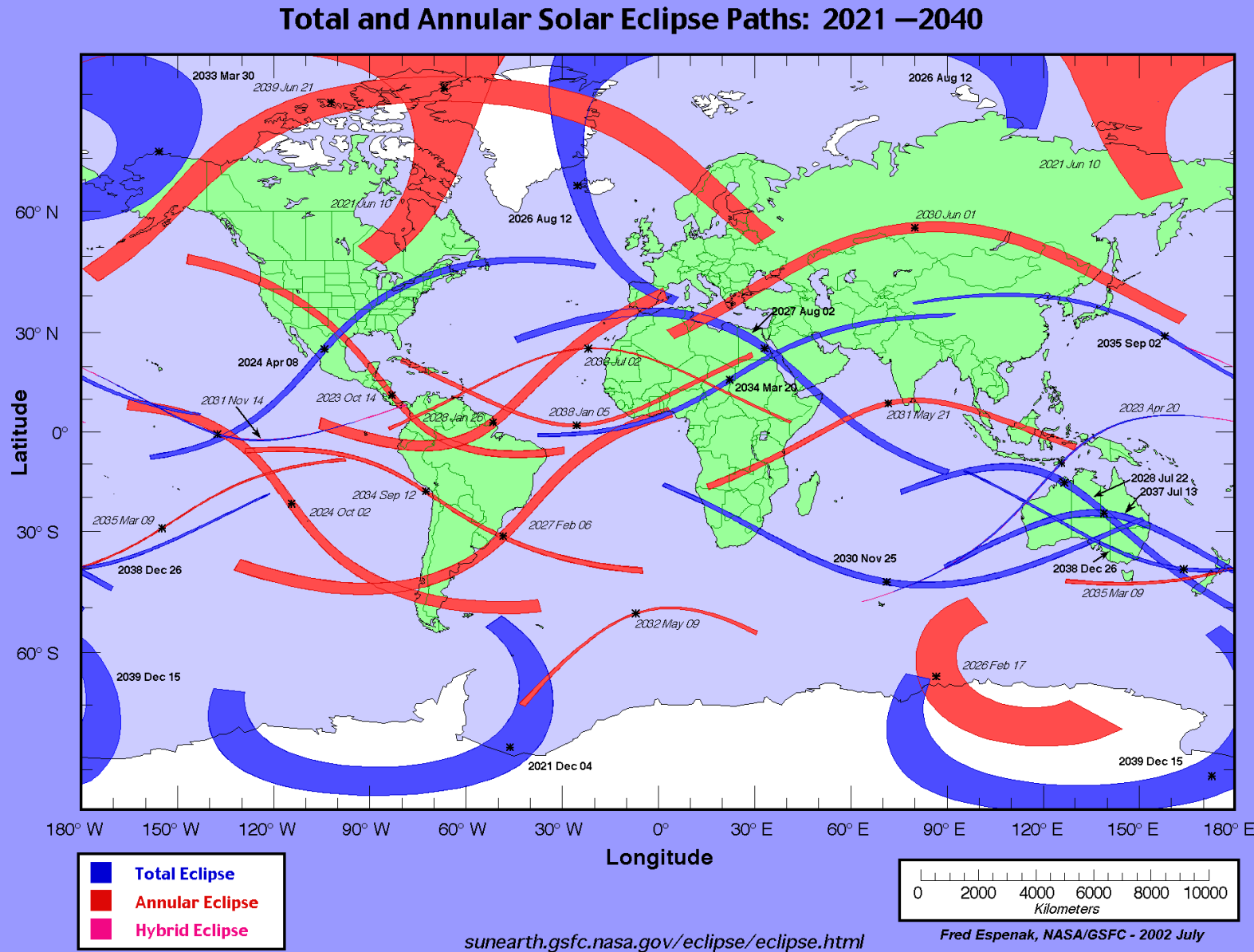
- The solar corona is still a puzzle to be composed
- At the moment the ground-based observations are extremely useful for near limb observations (where the solar magnetic fields recombine and close loops) and are complementary to space-based observations
- The ESCAPE experiment demonstrates that Dome C is one of the few sites in the world with a “coronagraphic” sky. The duty cycle for coronal observation at Dome C is of 10%. A tower should increase the duty cycle up to the 15%
- Ready for the next solar maximum observations (2025-2026) !!! :)

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What's next...



2024 April 08: Total Solar Eclipse

Meeting

20/09/2023

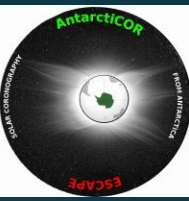
ESCAPE: The study of the solar corona from Dome C, Antarctica



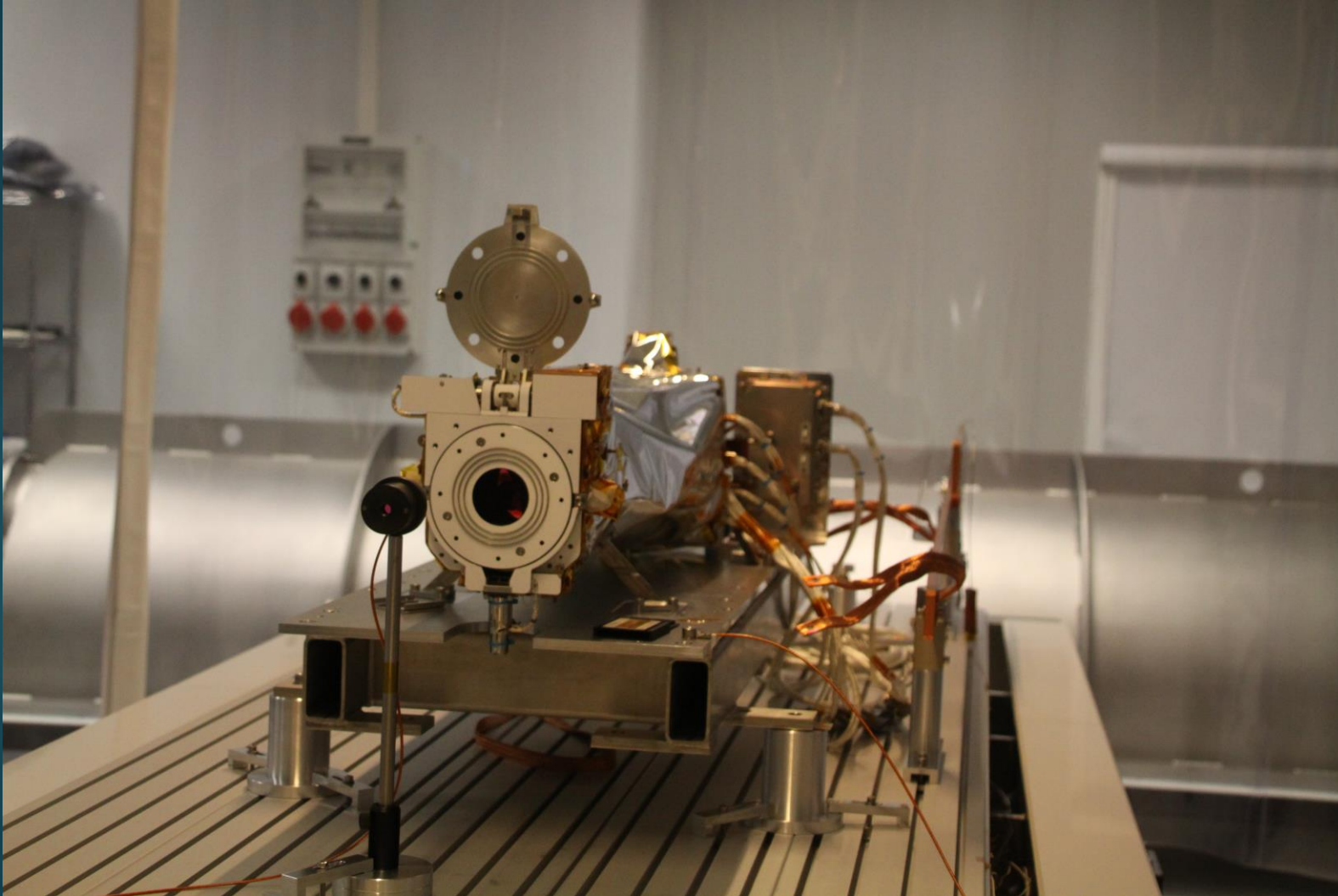
What's next...

June 2024 : NASA CODEX
installed on the ISS

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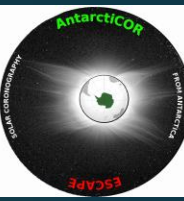


ESCAPE: The study of the solar corona from Dome C, Antarctica



What's next...

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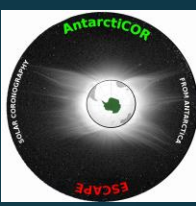
June 2024: PROBA-3 in orbit

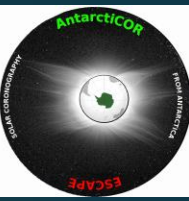
ESCAPE: The study of the solar corona from Dome C, Antarctica

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What's next...

Summer 2025: 3th flight of CorMag





THANKS!!!

