Cross-correlation studies with UHECRs



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We look for UHECR anisotropies

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Harmonic space: global properties

We look for UHECR anisotropies





Auto-correlation: large scales







Auto-correlation: large scales



Cross-correlation with galaxies: small scales







Auto-correlation: large scales



Cross-correlation with galaxies: small scales

A&A 2021, JCAP 2022, +ongoing x3

with D. Alonso, M. Bustamante, S. Camera, K. Tanidis

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we find the autocorrelation $\ensuremath{\mathsf{AC}}$

$$C_{\ell} = \frac{1}{2\ell+1} \sum_{m} |a_{\ell m}^* a_{\ell m}|$$

: observations :



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i. The measured UHECR flux anisotropy expands as $\Delta_{CR}(\hat{n}) = \sum_{\ell m} a_{\ell m}^{CR} Y_{\ell m}(\hat{n})$

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Noise AC $\frac{1}{N_{CR}} \gg \frac{1}{N_{gal}} XC$

The 'observed' UHECR flux anisotropy is given by

$$\Delta_{\rm CR}(\hat{\mathsf{n}}, E_{\rm cut}) = \int \mathrm{d}\chi \, \phi_{\rm CR}(\chi) \, \delta_{\rm s}(z, \chi \hat{\mathsf{n}})$$

where the radial kernel $\phi_{
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The (optional) weights $w(\chi)$ account for

the probability that a galaxy at distance χ is the origin of an UHECR above E_{cut}



 \boldsymbol{z}

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UHECR Sources

 $E_{\rm cut} = 40,63,100$ EeV with 1000, 200, 30 events

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Primaries, three examples: H1, O16, Si28 Injection: $\gamma = 2.60$, $\gamma = 2.10$, $\gamma = 1.50$ • di Matteo & Tinyakov, 2018 \rightarrow for Si28 we cut at $E_{max} = 280$ EeV

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UHECR Kernel

G)

Attenuation $\alpha(z, E_{cut}; \gamma, Z)$ calculated with SimProp

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Magnetic beam

Smearing with $\sigma \sim 1 \text{deg} (40Z/E)/(\sin^2 b + 0.15)$ © Pshirkov, Tinyakov & FU, 2013 \rightarrow we smear with the max angle within a given sky patch

: rotation measures :



© Pshirkov, Tinyakov & FU, 2013





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: hydrogen :



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: hydrogen or iron :



© Tanidis, FU, Camera, 2022

: total power :



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: how much fron :



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...and work in progress...