

RESULTS FROM THE HIGH ALTITUDE WATER CHERENKOV OBSERVATORY

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Loyola University - Chicago
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OUTLINE

(BRIEF) INTRODUCTION & MOTIVATION

DESCRIPTION OF THE HAWC OBSERVATORY

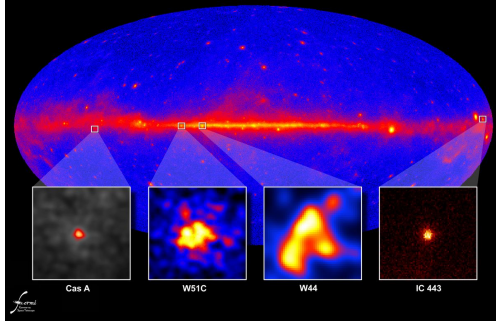
CR-RELATED RESULTS

CONCLUSIONS & OUTLOOK

VHE γ -RAYS



NASA's Fermi telescope resolves supernova remnants at GeV energies

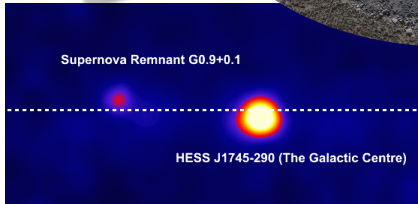
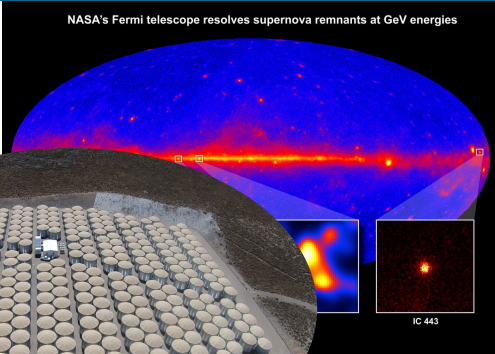


Supernova Remnant G0.9+0.1

HESS J1745-290 (The Galactic Centre)



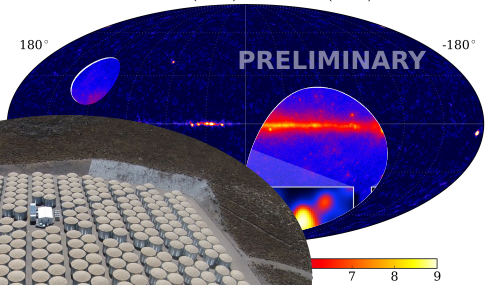
GROUND ARRAYS



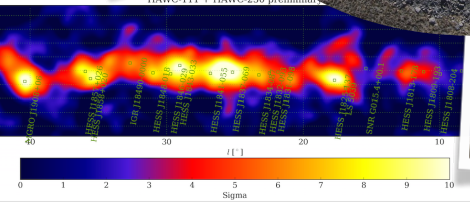
GROUND ARRAYS



HAWC-111 (283 d) + HAWC-250 (105 d)



HAWC-111 + HAWC-250 preliminary

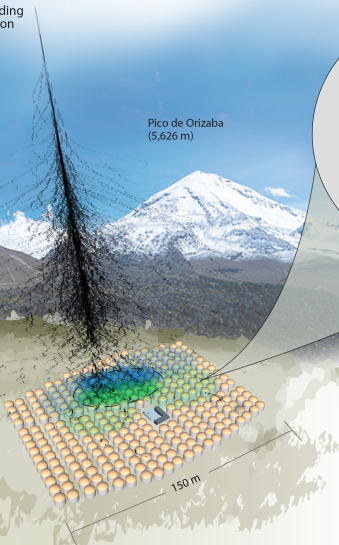


HAWC Observatory

HAWC operates day and night, providing a large field of view for the observation of the highest energy gamma rays.



Pico de Orizaba (5,626 m)

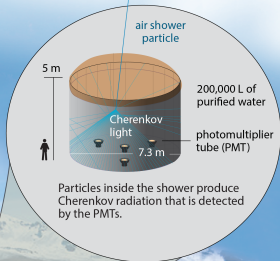


HAWC is located at 4,100 m above sea level, covering an area of 20,000 m².

HAWC

Water Cherenkov tank

HAWC comprises an array of 300 tanks that record the particles created in gamma-ray and cosmic-ray showers.

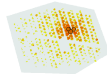


Particles inside the shower produce Cherenkov radiation that is detected by the PMTs.

Gamma rays vs cosmic rays

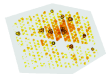
HAWC selects gamma rays from among a much more abundant background of cosmic rays.

gamma-ray shower



"hot" spots concentrate around the core

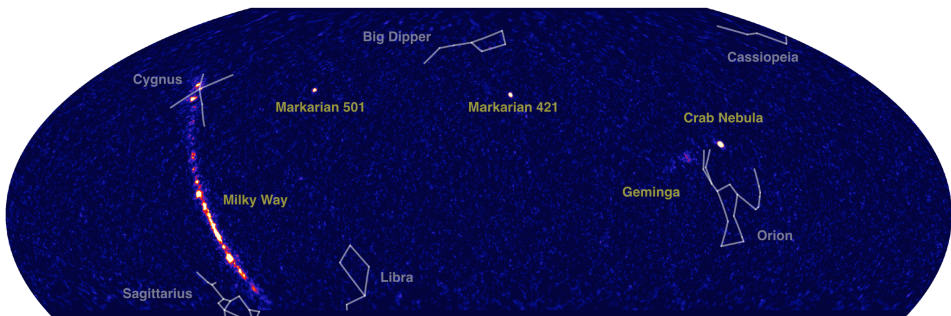
cosmic-ray shower



"hot" spots are more dispersed

HAWC

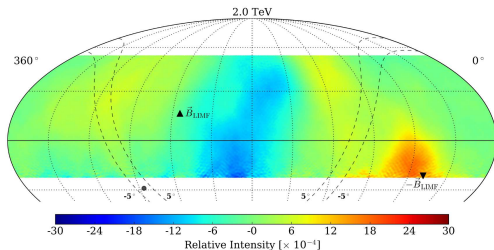
- ▶ large instantaneous sky coverage
- ▶ long, uninterrupted observation periods



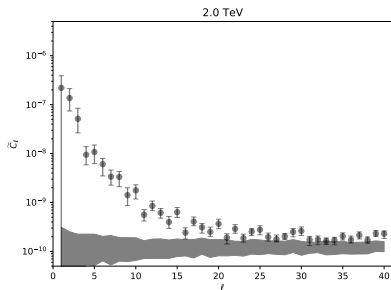
The HAWC Observatory: NIM A1052 (2023) 168253

COSMIC RAY ANISOTROPY (HAWC)

- ▶ Anisotropy in energy bins (from 2.0 to 72.8 TeV)



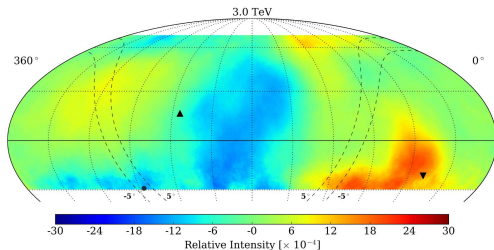
differential relative intensity



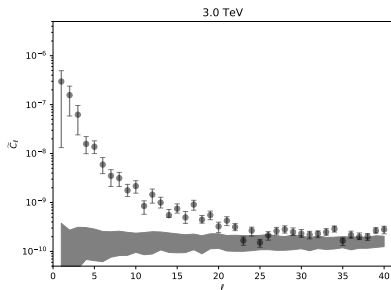
angular power spectrum

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differential relative intensity

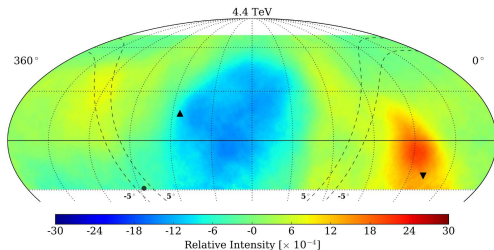


angular power spectrum

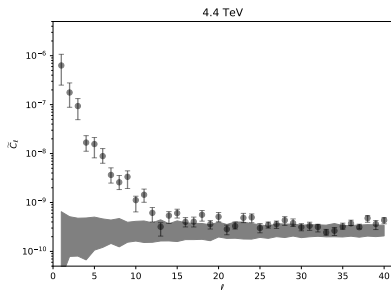
CR anisotropy (2 yrs), HAWC Collaboration: *ApJ* **865** (2018) 57

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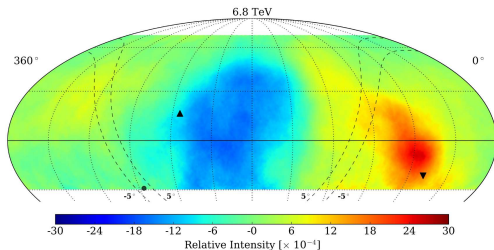
differential relative intensity



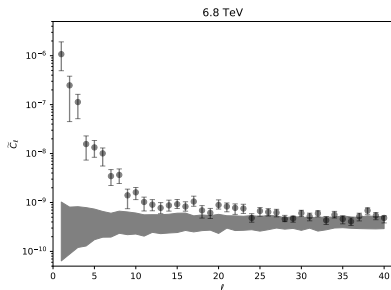
angular power spectrum

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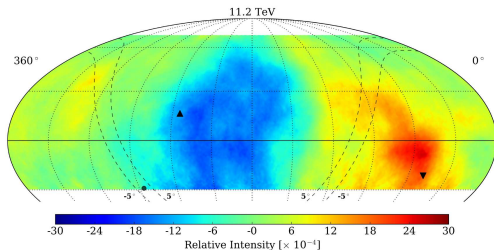
differential relative intensity



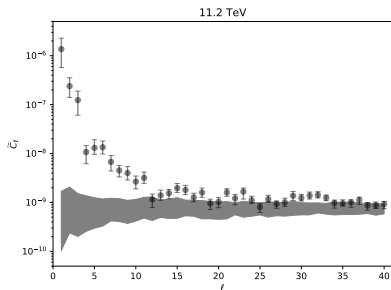
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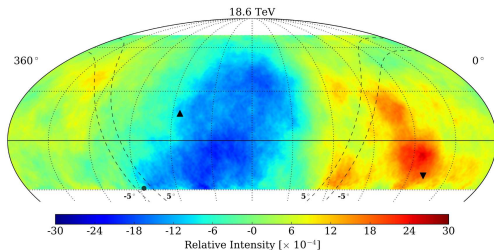
differential relative intensity



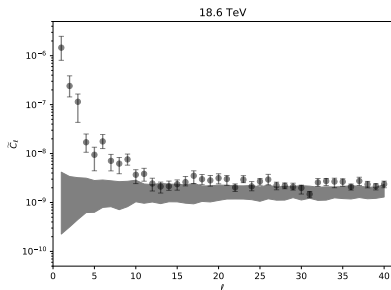
angular power spectrum

COSMIC RAY ANISOTROPY (HAWC)

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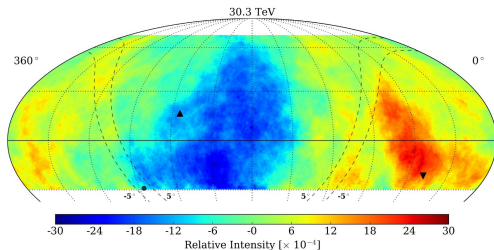
differential relative intensity



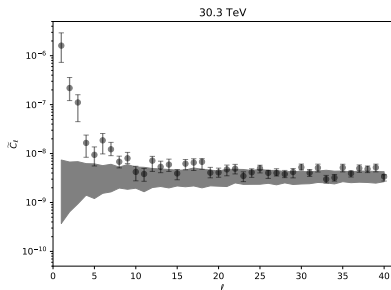
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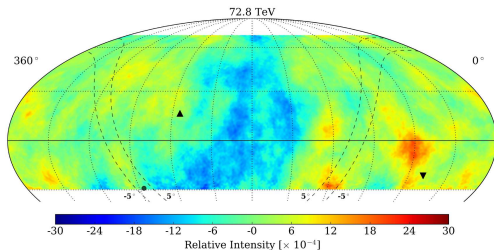
differential relative intensity



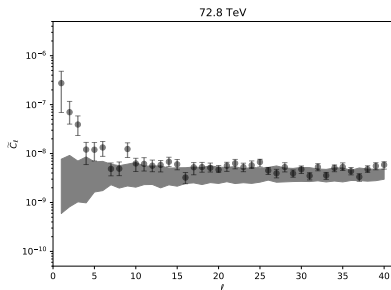
angular power spectrum

COSMIC RAY ANISOTROPY (HAWC)

- ▶ Anisotropy in energy bins (from 2.0 to 72.8 TeV)



differential relative intensity

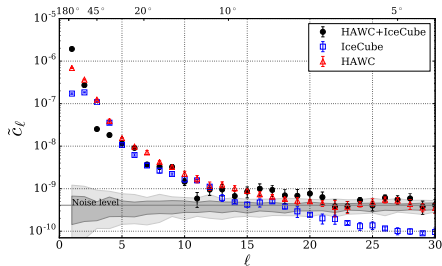


angular power spectrum

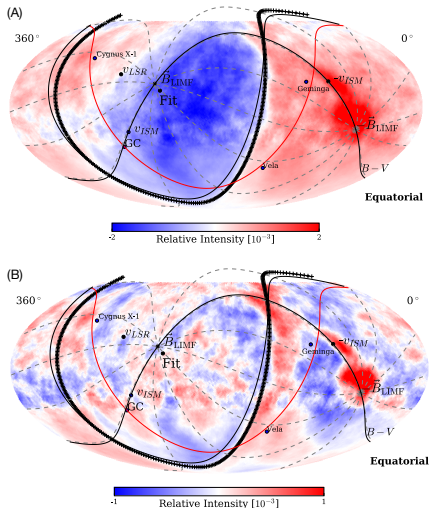
CR anisotropy (2 yrs), HAWC Collaboration: *ApJ* **865** (2018) 57

COSMIC RAY ANISOTROPY (HAWC+ICECUBE)

► Anisotropy at 10 TeV



angular power spectrum

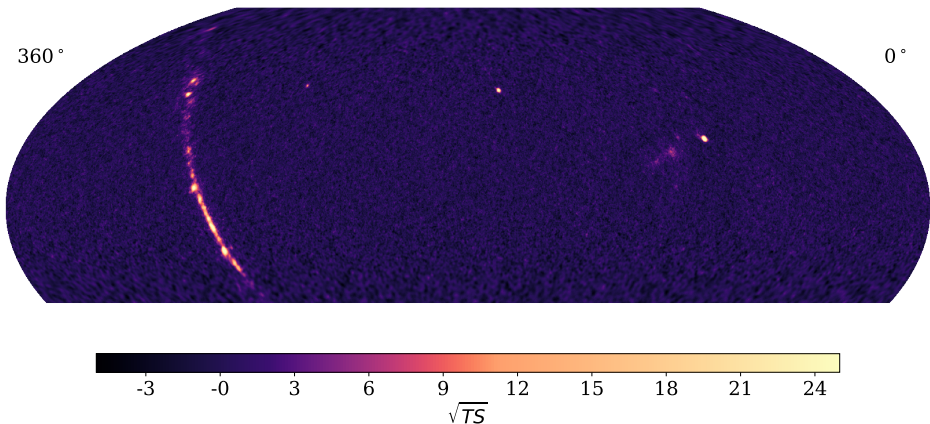


CR anisotropy at 10 TeV, HAWC+IceCube: *ApJ* **871** (2019) 76

THE 3rd HAWC CATALOG OF VHE γ -RAY SOURCES

► Significance map (point-source hypothesis)

All-sky view; 0.0°; 1523 days

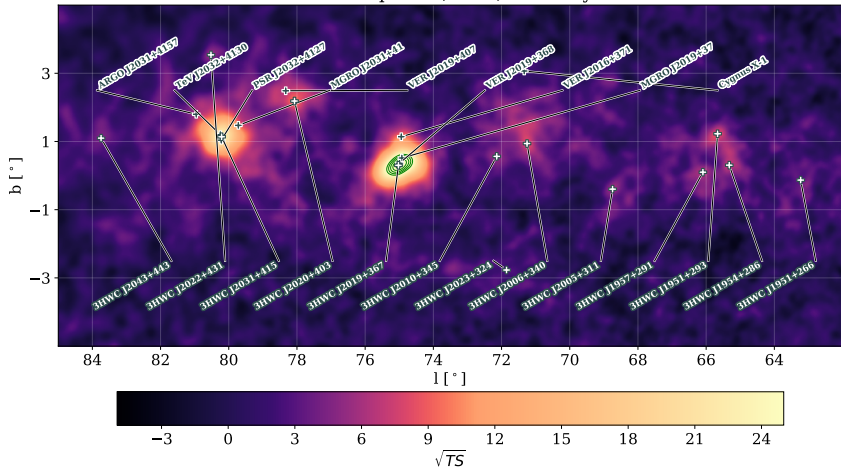


The 3HWC Catalog, HAWC Collaboration: *ApJ* **905** (2020) 76

THE 3rd HAWC CATALOG OF VHE γ -RAY SOURCES

► Significance map (point-source hypothesis)

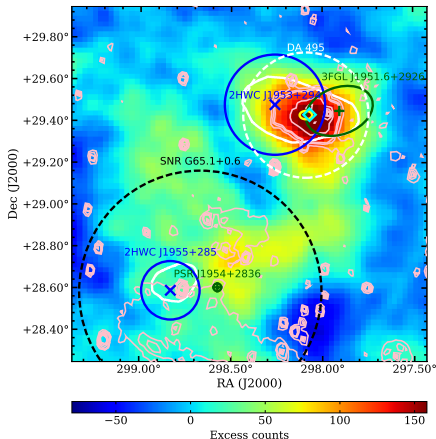
Galactic plane I; 0.0° ; 1523 days



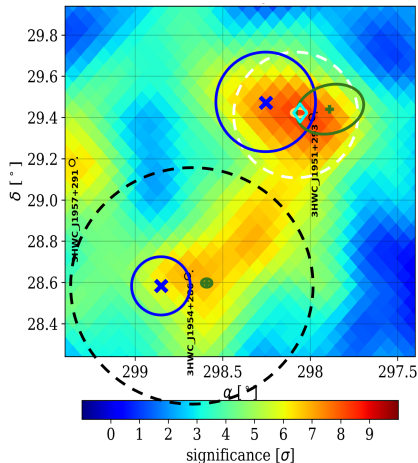
The 3HWC Catalog, HAWC Collaboration: *ApJ* 905 (2020) 76

NEW TEV γ -RAY SOURCES

► PWN DA 495 (2HWC J1953+294)



VERITAS counts map

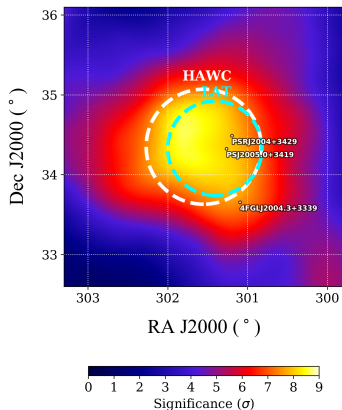


3HWC significance map

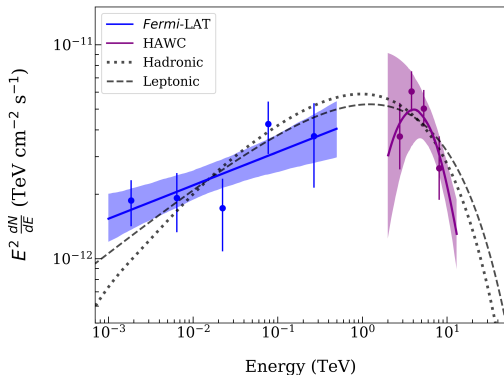
VERITAS+Fermi-LAT+HAWC: *ApJ* **866** (2018) 24

NEW TEV γ -RAY SOURCES

► 2HWC J2006+341



3HWC significance map

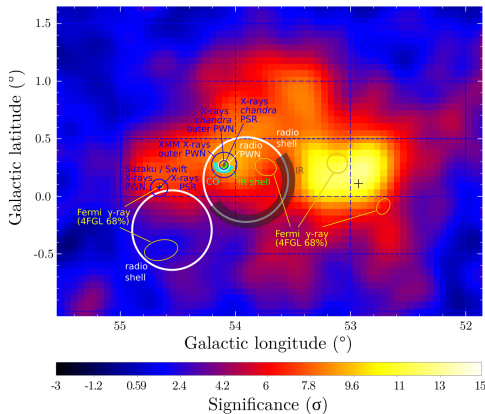


SED from HAWC and LAT data

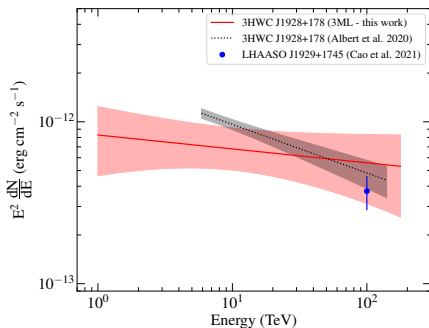
HAWC+Fermi-LAT detection of J2006: *ApJL* **903** (2020) L14

NEW TEV γ -RAY SOURCES

► 3HWC J1928+178 and HAWC J1932+192



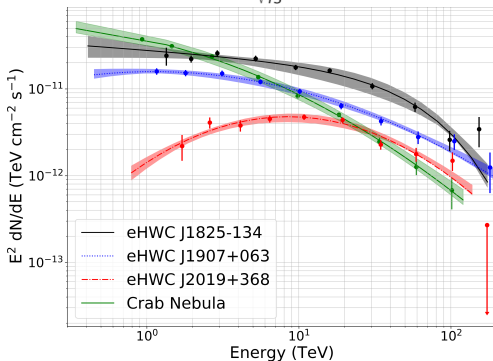
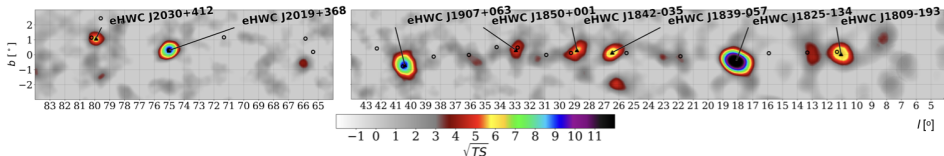
3HWC significance map



SED of 3HWC J1928+178

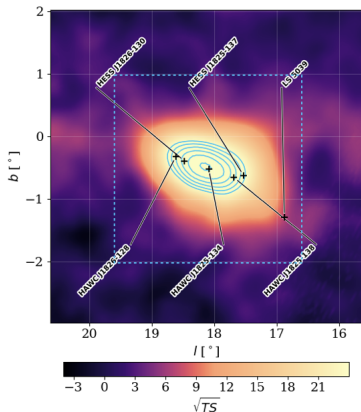
CATALOG OF γ -RAY SOURCES ABOVE 56 TeV

► Significance map ($E > 56$ TeV, 0.5° hypothesis)

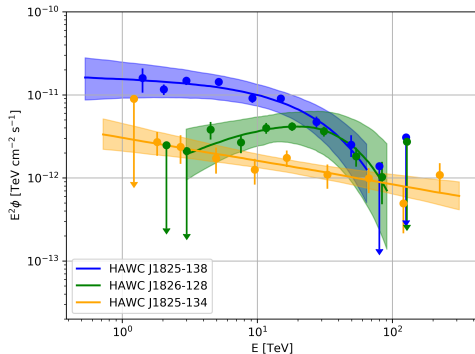


EVIDENCE OF 200 TeV γ RAYS

► eHWC J1825-134



Significance map

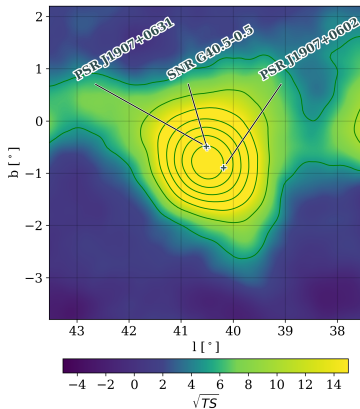


SED from HAWC data

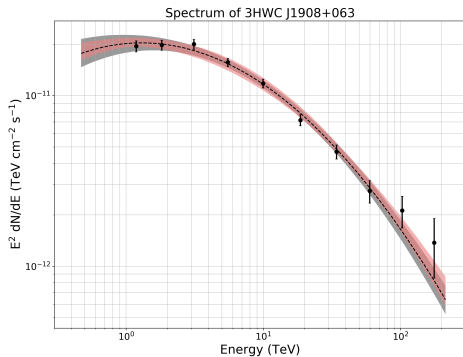
HAWC J1825-134, HAWC Collaboration: *ApJL* 907 (2021) L30

HE γ -RAY SPECTRA

► MGRO J1908+06



Significance map

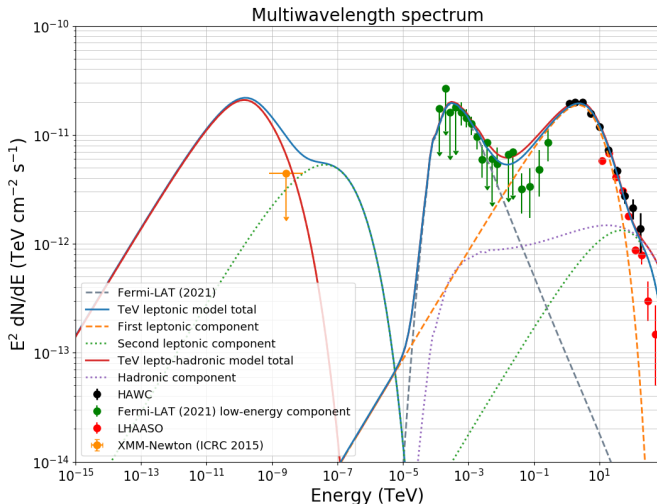


SED from HAWC data

MGRO J1908+06, HAWC Collaboration: *ApJ* **928** (2022) 116

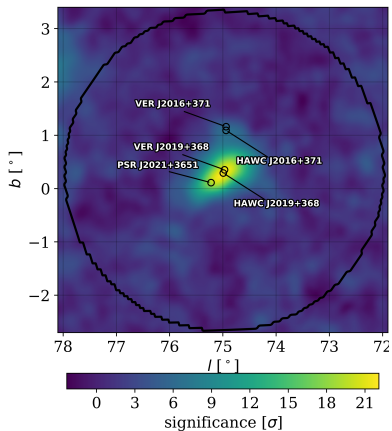
HE γ -RAY SPECTRA

▶ MGRO J1908+06

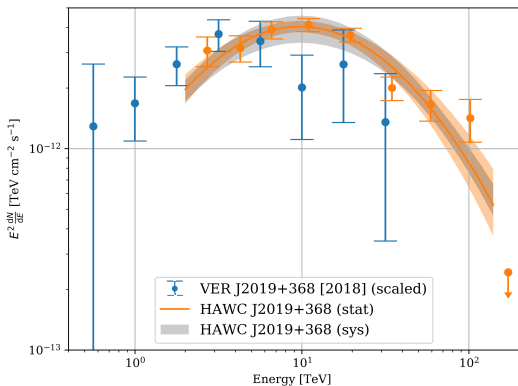
MGRO J1908+06, HAWC Collaboration: *ApJ* **928** (2022) 116

HE γ -RAY SPECTRA & MORPHOLOGY

► HWC J2019+368



Significance map

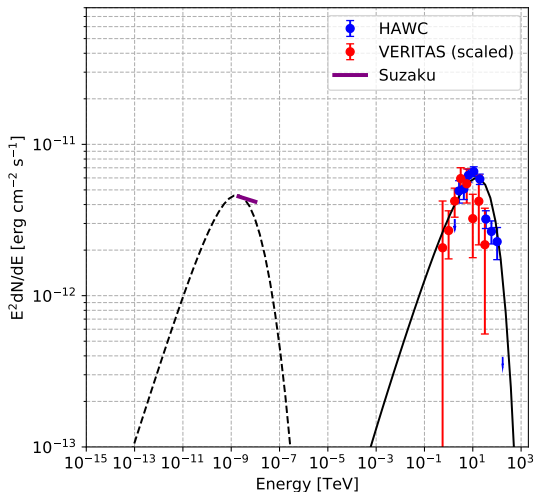


SED from HAWC data

HWC J2019+368, HAWC Collaboration: *ApJ* **911** (2021) 143

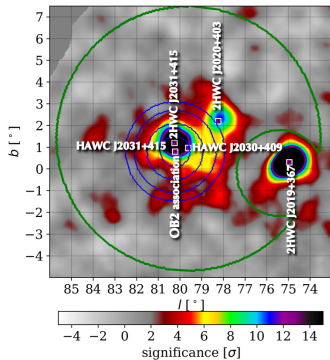
HE γ -RAY SPECTRA & MORPHOLOGY

► HWC J2019+368

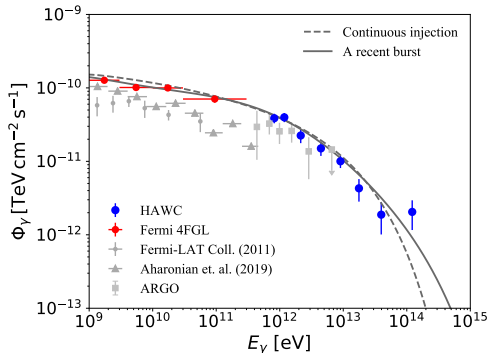


VHE COSMIC-RAY ACCELERATORS

► Cygnus Cocoon



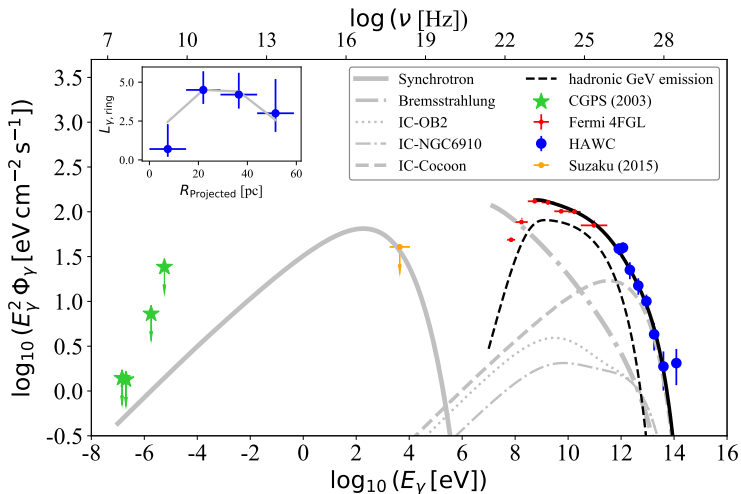
3HWC significance map



SED from HAWC and LAT data

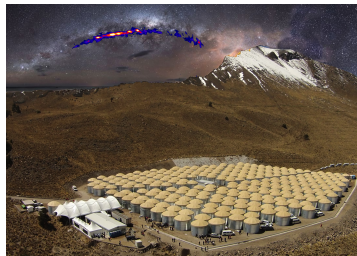
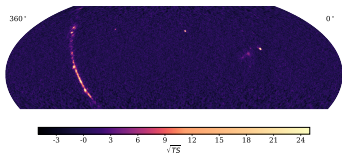
VHE COSMIC-RAY ACCELERATORS

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CONCLUSION & OUTLOOK

- ▶ CR anisotropy
- HAWC catalog **public!**
- New TeV sources
- Pevatron candidates

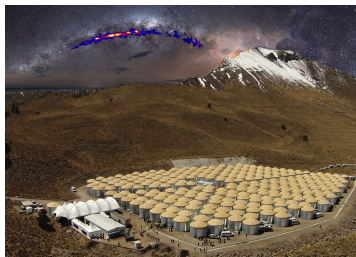


CONCLUSION & OUTLOOK

- ▶ CR anisotropy
HAWC catalog **public!**
New TeV sources
Pevatron candidates

- ▶ Other **science** contributions
Dark matter, CRs, solar physics,
particle physics,
multi-messenger studies,
diffuse emission, extended
regions, EBL, realtime alerts...

- ▶ Outrigger array **completed**
Enhanced sensitivity above
10 TeV



An aerial photograph of a large, conical volcano, likely Mount Fuji, with a snow-capped peak. The mountain's slopes are covered in a dense forest, and a small, rectangular, grid-like structure is visible in the foreground. The sky is a clear, pale blue.

THANK YOU VERY MUCH!

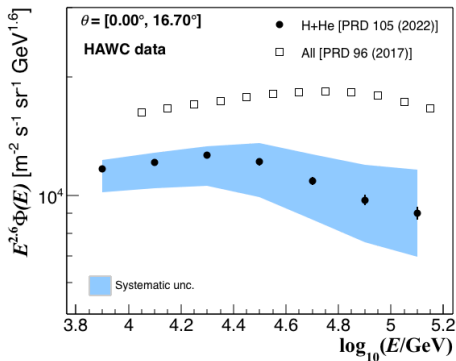
BACK-UP SLIDES

MOST RECENT HAWC PAPERS

- ▶ “The High-Altitude Water Cherenkov Observatory in México: The Primary Detector,” NIM **A1052** (2023) 168253.
- ▶ “Searching for TeV Dark Matter in Irregular Dwarf Galaxies with HAWC Observatory,” ApJ **945** (2023) 25.
- ▶ “Search for Gamma-Ray and Neutrino Coincidences Using HAWC and ANTARES Data,” ApJ **944** (2023) 166.
- ▶ “Validation of standardized data formats and tools for ground-level particle-based gamma-ray observatories,” A&A **667** (2022) A36.
- ▶ “Detailed Analysis of the TeV Gamma-Ray Sources 3HWC J1928+178, 3HWC J1930+188, and the New Source HAWC J1932+192,” ApJ **942** (2023) 96.
- ▶ “Gamma-Ray Emission from Classical Nova V392 Per: Measurements from Fermi and HAWC,” ApJ **940** (2022) 141.

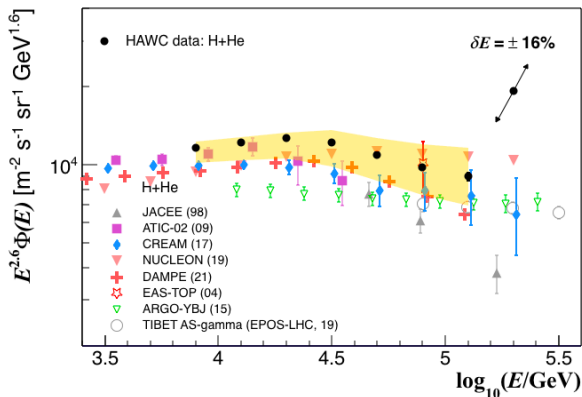
COSMIC RAY SPECTRUM

- ▶ H + He nuclei between 6 and 158 TeV



COSMIC RAY SPECTRUM

- ▶ H + He nuclei compared to other data



COSMIC RAY SPECTRUM

- ▶ all particle spectrum between 10 TeV and 1 PeV

