Open questions in deep learning techniques for the radio detection

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This presentation is given on the Workshop on Machine Learning for Cosmic-Ray Air Showers. I have tried to summarize the progress in radio data analysis and include comments appeared during discussion of the talk.

Quick introduction in radio generation



- Askaryan effect
- Geomagnetic effect



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Open questions in deep learning techniques for the radio detection

Quick introduction in radio detection



Why do we need machine learning for radio?

- Most of the detectors operate in the low signal-to-noise (SNR) environments
- ▶ Background is not completely white \Rightarrow limitation to classical methods
- ▶ Radio-frequency interference (RFI) are diverse \Rightarrow hard to account for all



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First attempts (back in 2015)



arxiv:1701.05158 (Bezyazeekov+ at ECRS2016)

- Simple perceptron (PyBrain): 200 input neurons, 3 layers × 500 each, 1 output
- Amplitudes are not normalized, no augmentation $\xrightarrow{?}$ non-invariant in (t,A)
- Was not tested on real data

Autoencoder approach



Motivation:

- Signal of interest is compact and short comparing to input waveform
- Convolutional filters effectively store signal and background features
- Can be used as denoiser

Implementation decisions:

- Size of input waveforms
- Length of convolutional filters
- "Stack more layers" rule should work with large training set

Autoencoder implementation: Preprocessing of the input waveforms

(based on Shipilov+, but others are very similar)

- Nyquist upsampling
- ► Normalization of amplitudes: $A(t) \rightarrow \overline{A} \in [0, 1]$ ⇒ amplitude-invariant
- Cropping input to $\mathcal{O}(1024)$ + augmentation: signal is randomly shifted \Rightarrow translation-invariant



Autoencoder implementation: design and training

Your input is welcome!

- ► Filter design
 - ► Use physical input: air-shower pulse and RFI duration!
 - Upper and lower boundaries for kernel size?
 - Constant or variable size?
 - Upsampling might help autoencoder: optimal sampling factor
- Input waveforms
 - Median filter for preprocessing?
 - Sensitivity to narrow-band background
 - Optimal size in samples and nanoseconds
- Training strategy
 - Optimal size of dataset
 - Binning in amplitude/SNR?
 - Including pure noise in training?
 - How to properly include samples with strong pulse-like RFI

Response and metrics



- Threshold amplitude \Leftrightarrow 5% tolerance to false positives
- Efficiency: N_{rec.}/N_{tot.}, fraction of events passed the threshold
- ► Purity: N_{hit}/N_{rec.}, fraction of events with reconstructed position of the peak: |t_{rec.} - t_{true}| < 5 ns</p>



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	Shipilov+	Schlüter+	Rehman+
	1812.03347	1901.04079	PoS ICRC2021 (2021) 417
Band (MHz)	30-80	30-80	50-350
Background	Real	Simulated	Simulated*
Normalization	[0;1]	[0;1]	[-1;1]
Input	4096	1000	1024
Bin size (ns)	0.3125	5.6	0.25
Kernel size	$32 { ightarrow} 16 { ightarrow} \dots$	fixed to 5	fixed to 256
Kernel size at 1st layer (ns)	10	28	64
Filters/layer	$16 { ightarrow} 32 { ightarrow} \dots$	$16{ o}32{ o}\dots$	fixed to 8
Layers	3	5	2
Training set	\approx 15k	\approx 70k	\approx 103k+50k
Tested on real data	Yes	No	No

Current state of autoencoder implementations

*improved by Kullgren+ (doi:10.5281/zenodo.6011170)

Hereafter we consider only denoisers

Artificial spectral noise

Artifact when high-rate upsampling is applied (Shipilov+):

- Nyquist upsampling adds zero-amplitude higher frequencies
- Reconstruction contains noise in-between initial bins
 artificial noise at higher frequencies

Less impact for non-upsampled (Schlüter+) or low-rate upsampled (Rehman+) signals:



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Antenna and phase response invariance

Simulated response **Actual response** 1.0 1.0 0.5 0.5 normalized amplitude normalized amplitude ۸ 0.0 0.0 -0.5-0.5-10-1010 20 20 -1010 20 20 time (ns) time (ns)

- Calibration does not describe hardware response precisely
- Autoencoder trained by Shipilov+ works surprisingly well both with and without antenna response
- Indication of phase shift invariance
- Implication of this feature?

Amplitude reconstruction

- At low SNR normalization is done w.r.t. noise peaks
- Due to normalization information about absolute amplitude is erased
- ► Is it fundamental problem for only this architecture or general one?



Hybrid approach: autoencoder + interferometry

Synthesis of signals using reconstructed arrival times







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Next steps

Improvements of the current networks

- ► Alternative architectures, e.g. U-Net, RNN, transformers, LIGO experience
- Adding more channels, wavelets, spectral information
- Integration of modern tools for interpretation, e.g. SHAP
- Spatial information \Rightarrow combination with GNN
- ► Timing information ⇒ combination with CNN encoding sky noise

Compressed neural networks for radio trigger on FPGA

- hls4ml: firmware implementations of machine learning algorithms using high level synthesis language (HLS)
- Intensively used in collider physics
- Successfully tested on old ARIANNA hardware (arXiv:2112.01031)
- ► The only solution for ultra-large scale sparse arrays featuring single antennas

Path to universal network

Search for Holy Grail (and reduce carbon footprint of CoREAS)

Start from GAN?

- High-dimension space: $A(\vec{r}, t)$
- Likely more accurate than analytical models, but still not good enough
- Ideal for design studies and template production
- Can be used for RFI and air-shower pulses generation

Ultimate denoiser

- Universality in background? RFI library?
- ► Train in frequency domain
 - Do they adjustable to every frequency band?
 - CNN will likely not work. Architecture?
 - $F_{\text{true}} = F_{\text{meas}} * F_{\text{IRF}}^{-1}$. How to propagate F_{IRF} to NN?

Conclusion

- Radio community has successfully learned how to use neural networks
- Several architectures are implemented, but only one tested in production
- The recipes for optimal design are not well defined
- ► Few unique features of radio autoencoders have already been discovered
- ► The killer feature of technology is FPGA trigger for stand-alone antennas
- Training in the frequency domain is almost unexplored:
 - Representation of signals: complex numbers? phase-amplitudes?
 - Treatment of noise? CNN for dynamic spectra? LIGO approach?
 - Efficient phase unwrapping?