

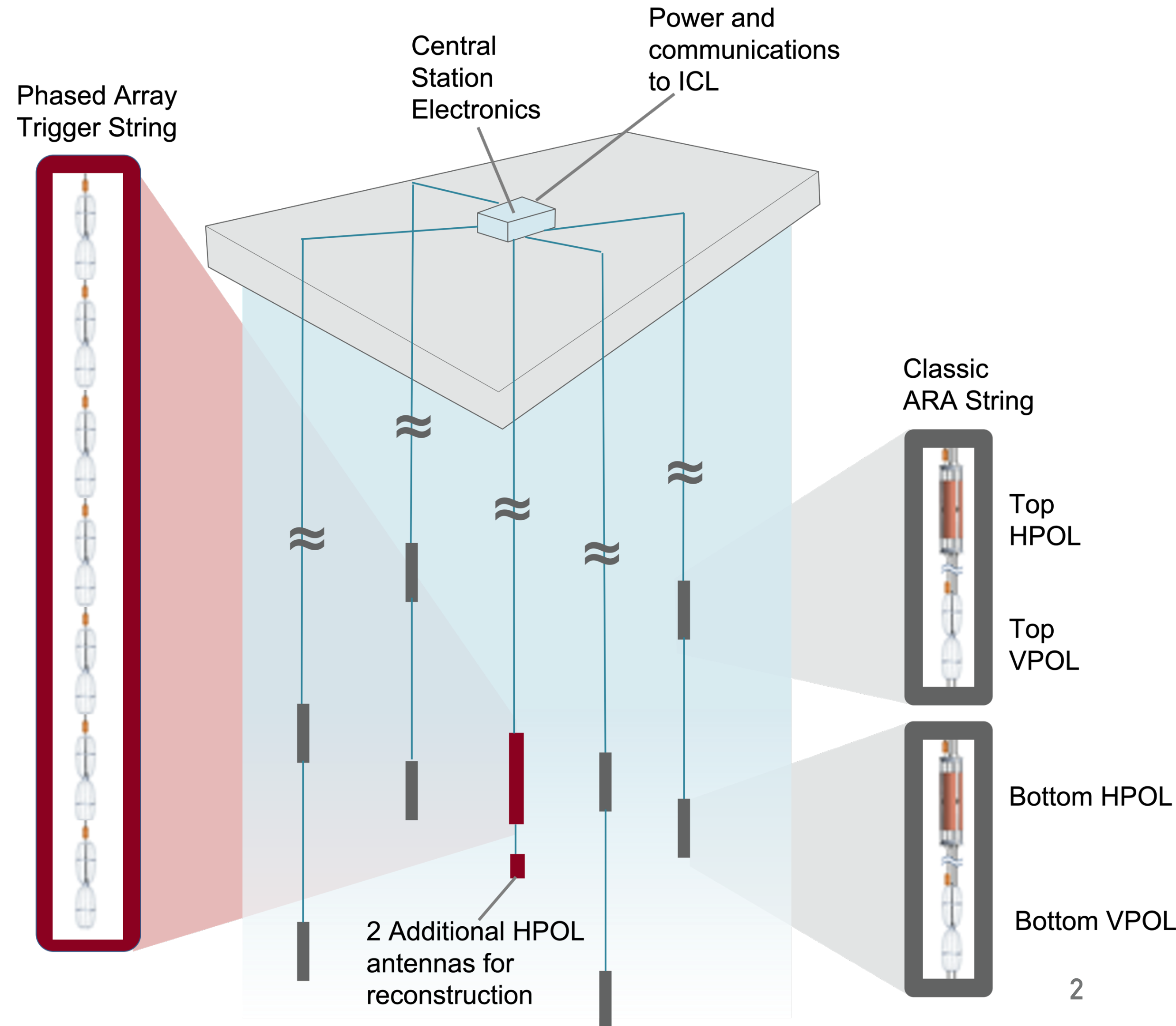
# DETERMINING DEEP ANTENNA POSITIONS AND UNCERTAINTIES WITH ARA STATION 5

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IceCube Calibration Workshop  
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# ARA STATION 5

- Two separate (but connected) experiments located in the same spot
- Phased Array: one string with 7 VPol antennas and 2 HPol antennas. It uses a phased array trigger on VPol antennas only.
- A5 Station: traditional ARA station with 4 strings. Each string has two VPol and two HPol antennas ~30 m apart. It uses a power threshold trigger
- One local calibration pulser
- Goal: determine antenna locations to within 10 cm



# WHAT WE NEED FOR GEOMETRY CALIBRATION

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	Description
<b>Ice model</b>	Index of refraction changes as a function of depth; big impact on distant calibration sources
<b>Cable Delays</b>	Relative delays due to different cable lengths + differences in bulk cable
<b>Calculated time delays using calibration waveforms</b>	Dependent on the quality of the digitizers, similarity of antenna impedances, and available in-situ measurements
<b>Calibration Sources</b>	Ideally, want sources at a variety of angles, as well as both local and distant. For ARA5 we have two sources: 1 local cal pulser + 1 SPIceCore drop

# FITTING PROCESS FOR CALIBRATING ARA 5

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1. Using only phased array antennas, solve simultaneously for best ice model and relative phased array cable delays
2. Using best ice model, calculate expected and measured time delays from:
  - Local calibration pulser
  - SPiceCore pulser
3. Input those time delays in a Minuit optimizer to find best antenna locations and ARA channel cable delays

# ASSUMPTIONS MADE FOR ARA STATION 5

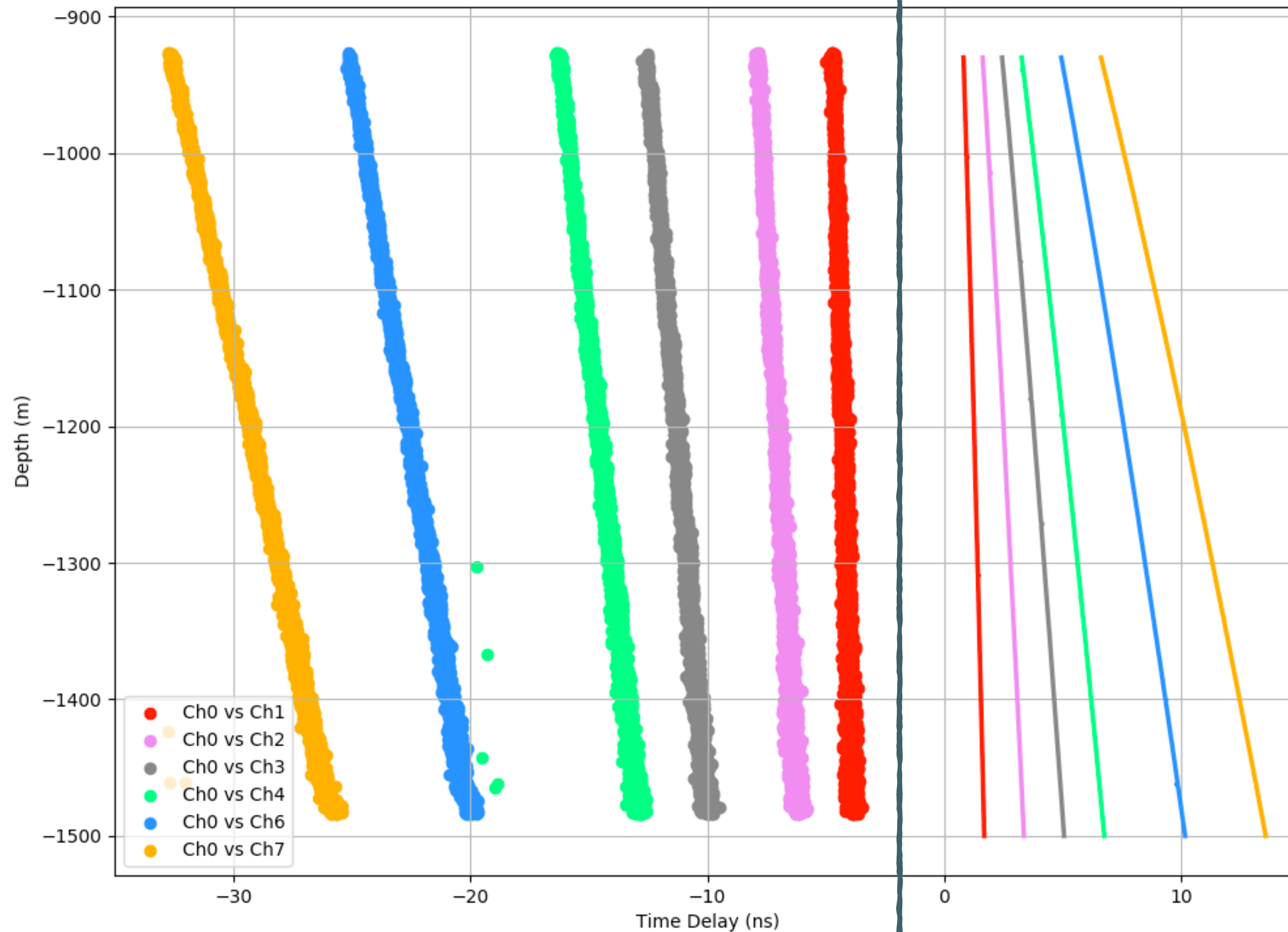
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- We know the depth of the phased array exactly.
  - This is a reasonable assumption because the phased array antennas are deployed compactly, within  $\sim 10$  m of each other, so likely the relative depths are very accurate.
- We know the depth of the SPIceCore pulser as a function of time.
  - This was measured carefully during the SPIceCore run
  - Any uncertainties here are absorbed by the fit for ice model
- The Phased Array and the SPIceCore holes are approximately parallel.

# USING SPICECORE TO FIND BEST ICE MODEL AND CABLE DELAYS

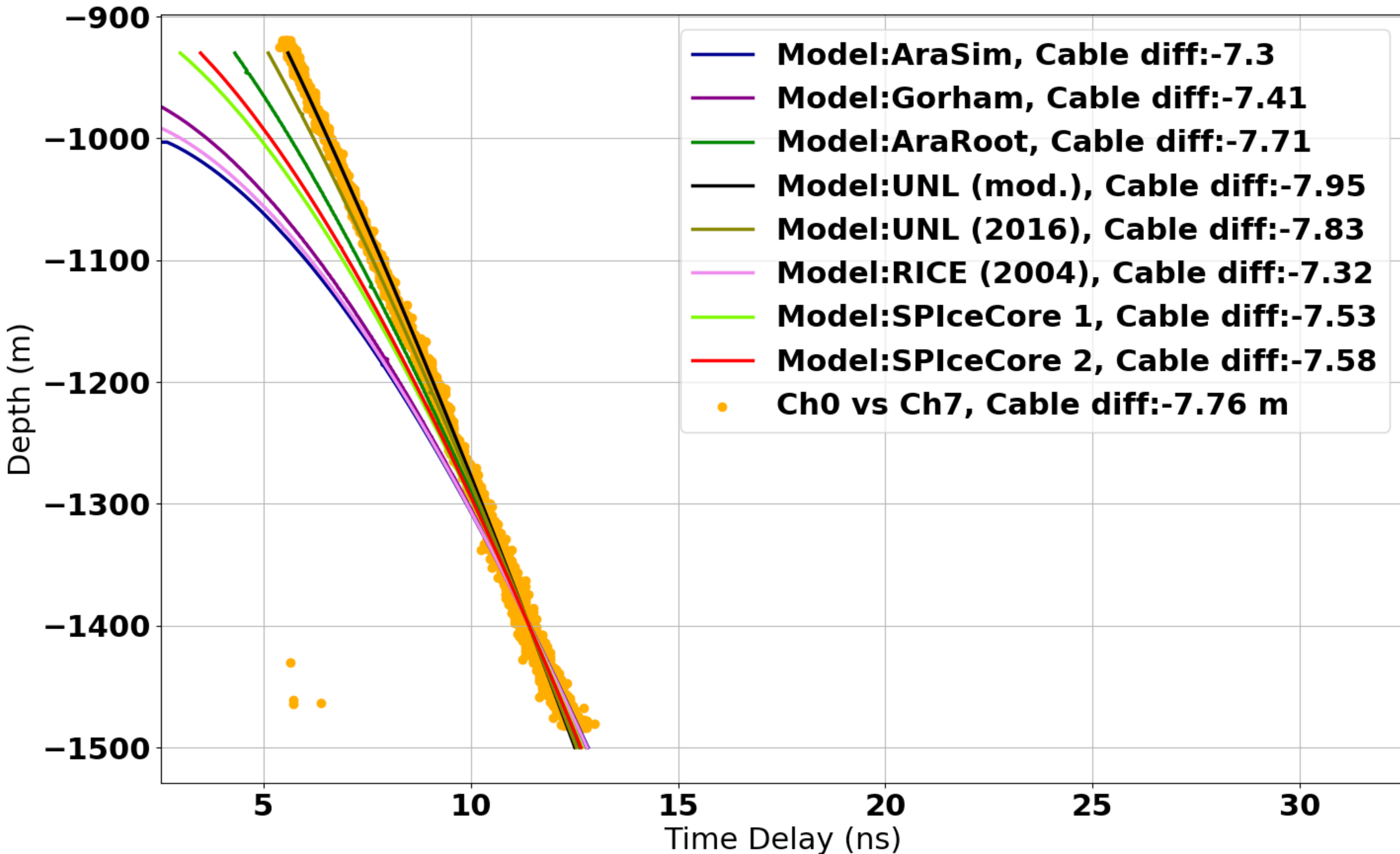
Phased Array Data (Without cable delays)

UNL (2016) Model Predictions



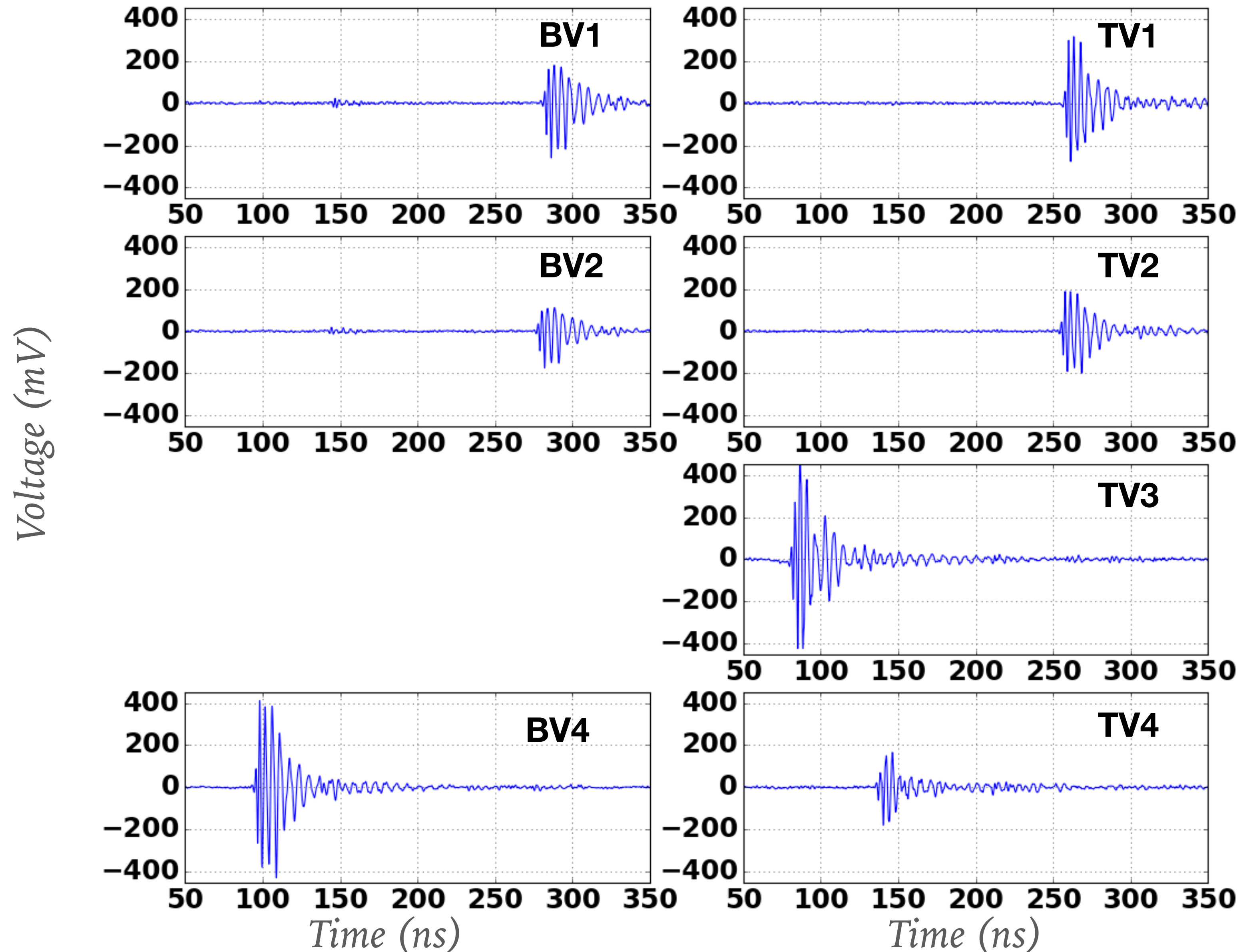
- Because we didn't have individual cable delay measurements, we had to use SPiceCore data to solve for ice model and cable delays simultaneously
- As the SPiceCore pulser dropped, the time delays between channels changes
- Plotted here are time delays as a function of depth- no cable delays included. Data is on left, model prediction is on the right
- Difference between model (right) and data (left) is one way to find the cable delay
  - For example, difference between red data and model is 5.1 ns. Using nominal speed = 0.202 m/ns we can get cable length of 1.03 m (compared to 1.0 m predicted)

# COMPARING MULTIPLE MODELS



- ▶ While a shift in the x axis can account for the cable delay, the shape of the curve determines the ice model
- ▶ Here, I've forced all models to overlap data at -1400 m to easily compare
- ▶ Only UNL (2016) gets close to describing data accurately
- ▶ Solution: modify UNL (2016) to get new best ice model

# AVERAGE LOCAL CALIBRATION PULSER- A5 ANTENNAS

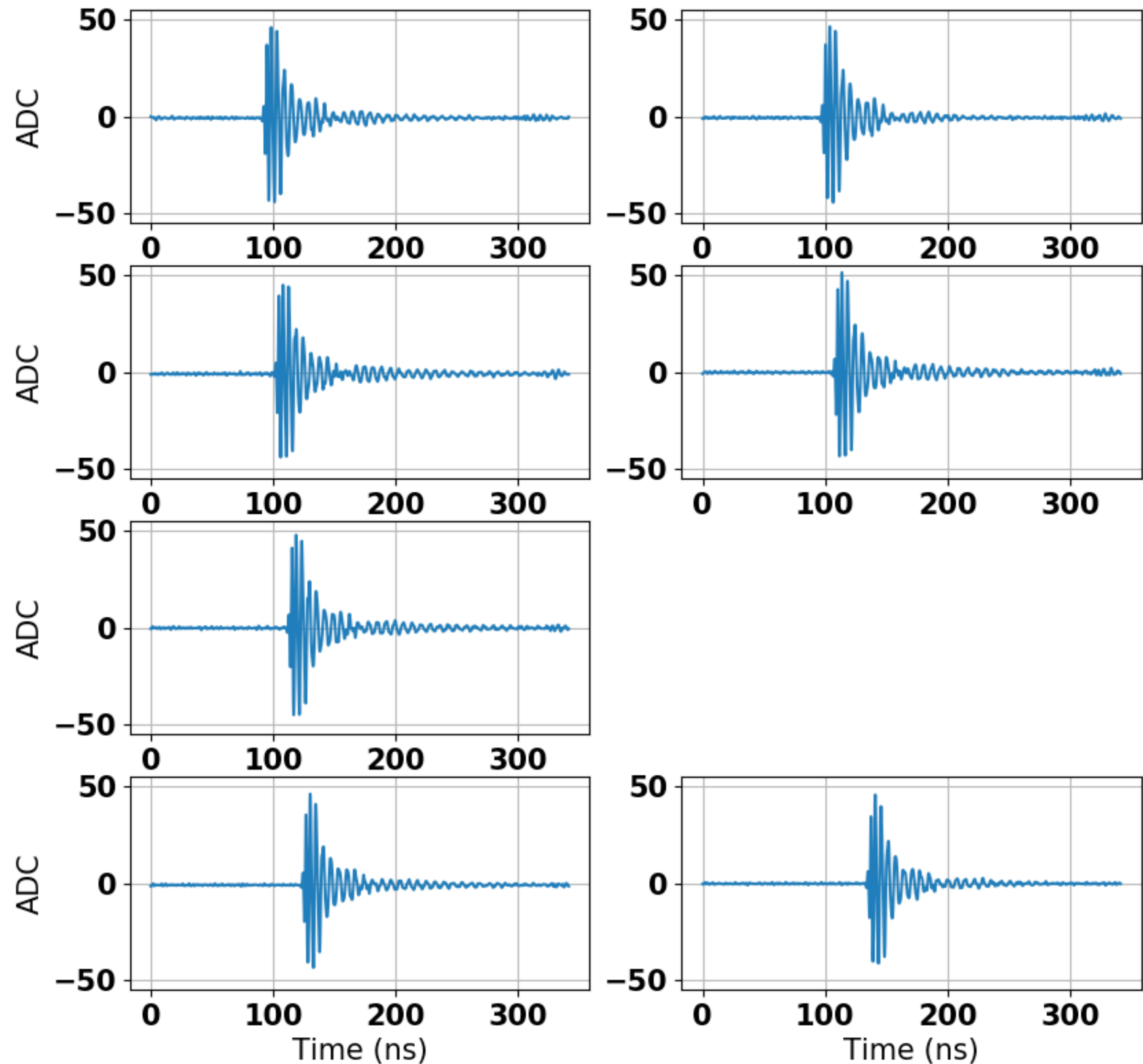


- Averages for 7/8 VPol antennas for a local calibration pulser
- Differences between channels caused by two main things:
  - Non-identical impedances and non-identical impulse responses
  - Angular dependence of both transmitting + receiving antenna
- Average calculated time delay error: 82 ps



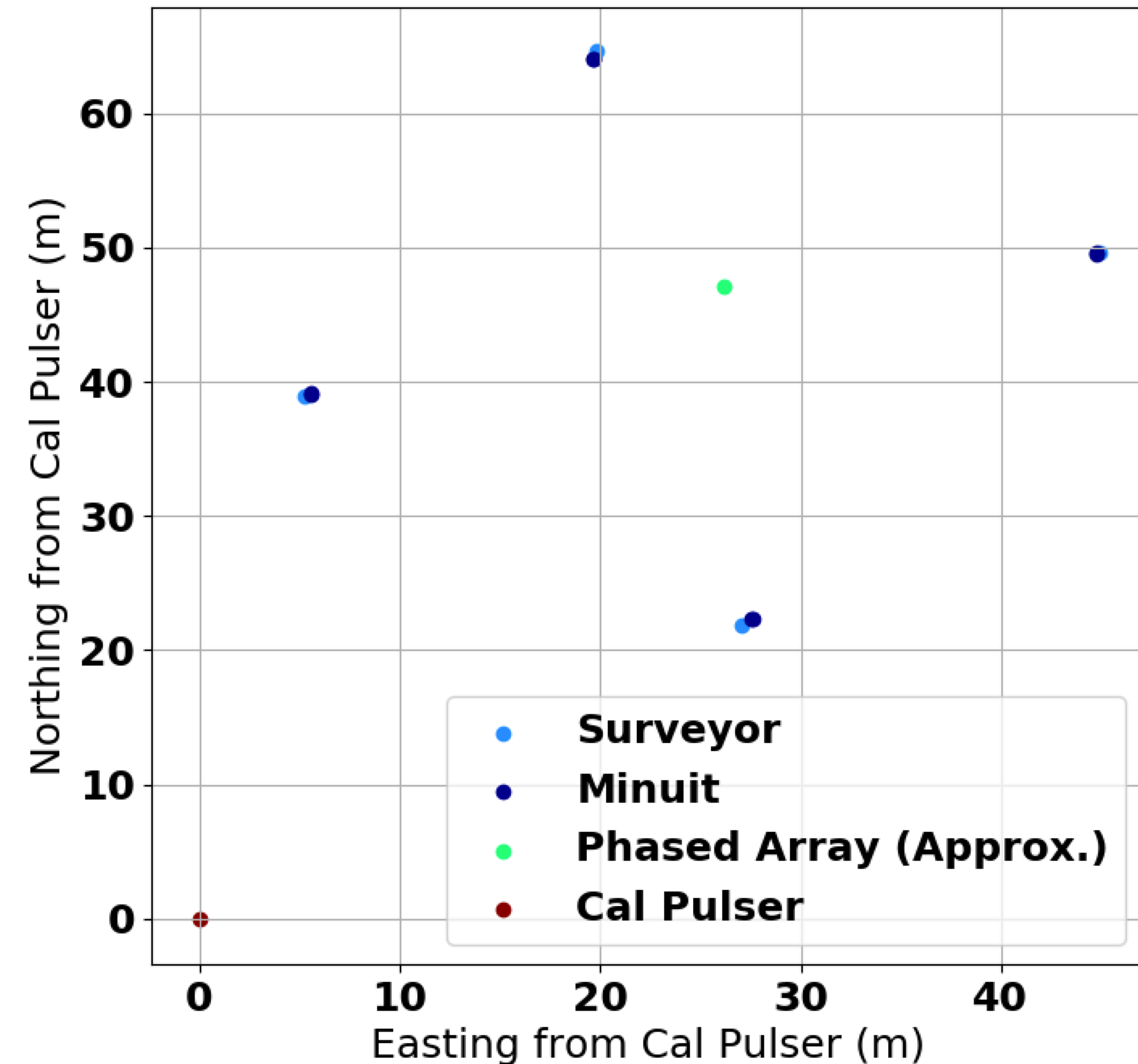
# PHASED ARRAY WAVEFORMS

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- What changed?
- Better digitizers
- Antenna impedances and impulse responses extremely similar
- Compact array means antenna angular response less pronounced
- Average time delay error: 30 ps

# USING MINUIT TO FIND POSITIONS



Measured time delay from cross correlation (including cable delays)

Expected time delay from RayTracer

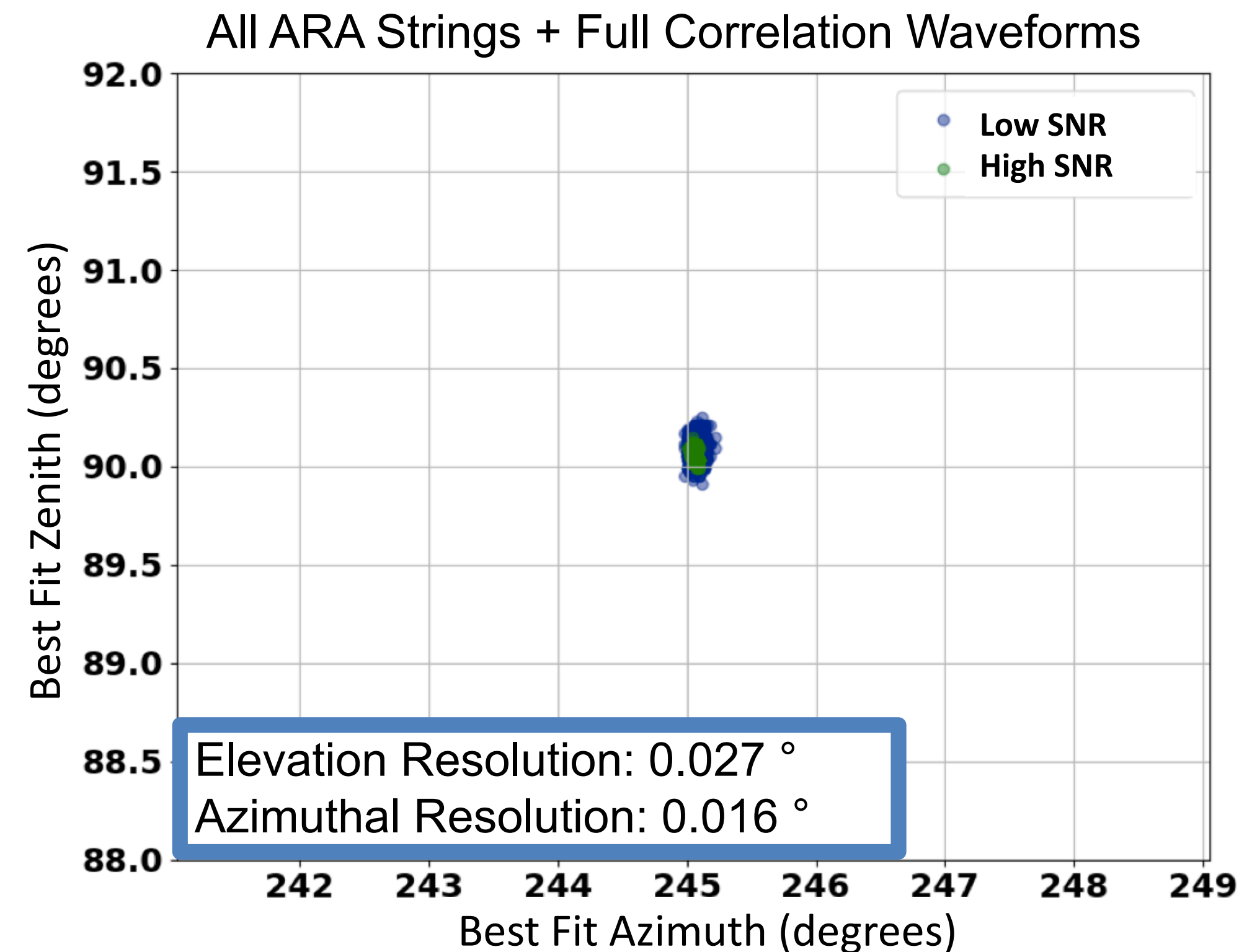
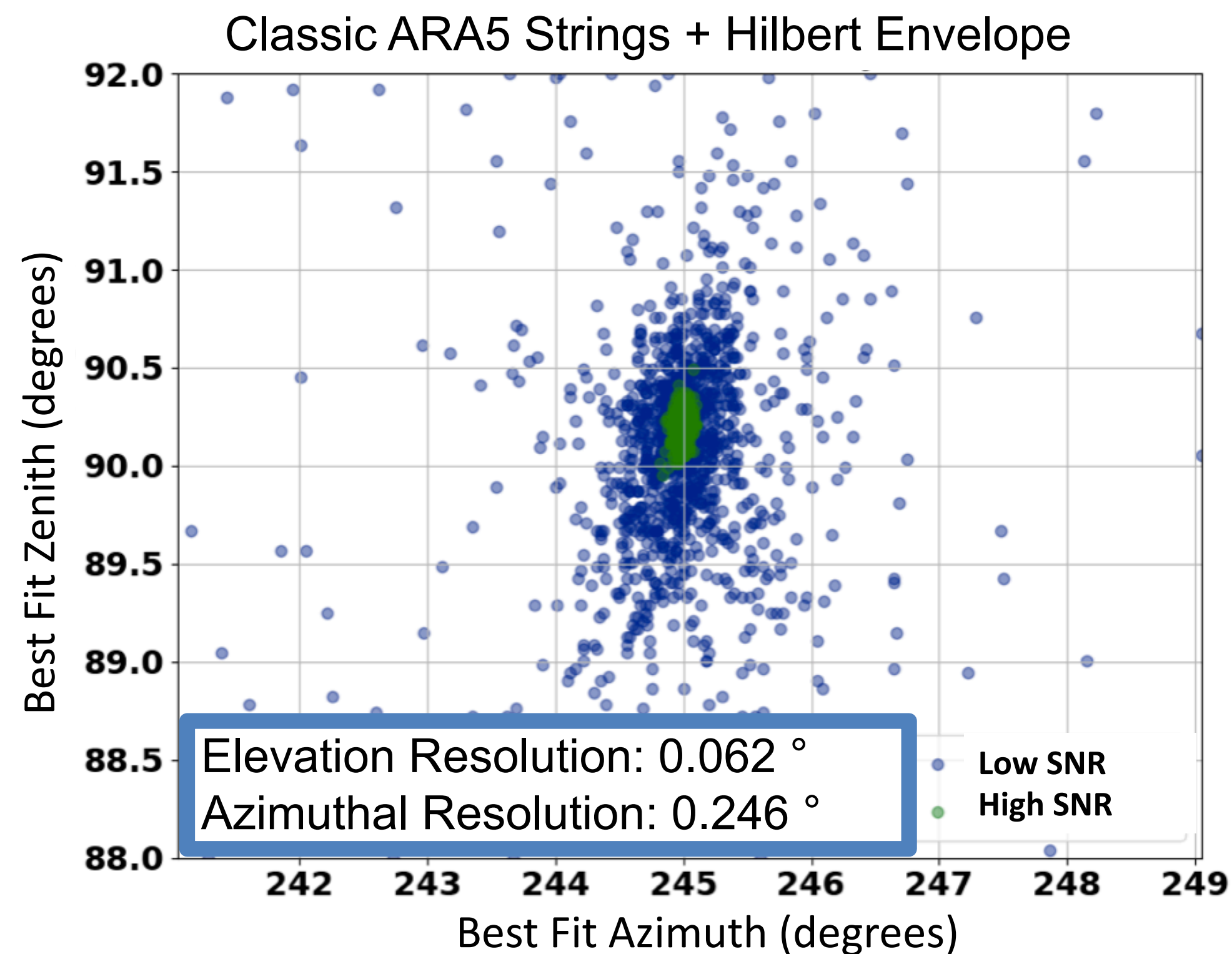
$$\chi^2 = \sum_{pairs} \frac{(\Delta t - RayTracer(r_1, r_2, z_1, z_2))^2}{\sigma^2}$$

Variance from time delay calculation

- Two main steps:
  1. Build a library of time delays, using all available calibration sources
  2. Minimize the above equation in two steps: once assuming holes are straight/parallel, then allowing holes to tilt
- Errors on fit calculated using built-in Hessian error functionality- all below 5 cm

# RESULTING RESOLUTION: POINTING BACK TO THE LOCAL CALIBRATION PULSER

- Very successful at pointing events back to local calibration pulser
- But, this is expected, since we used this data for calibrating!
- Ideally, would have a third, separate calibration source to test the geometry



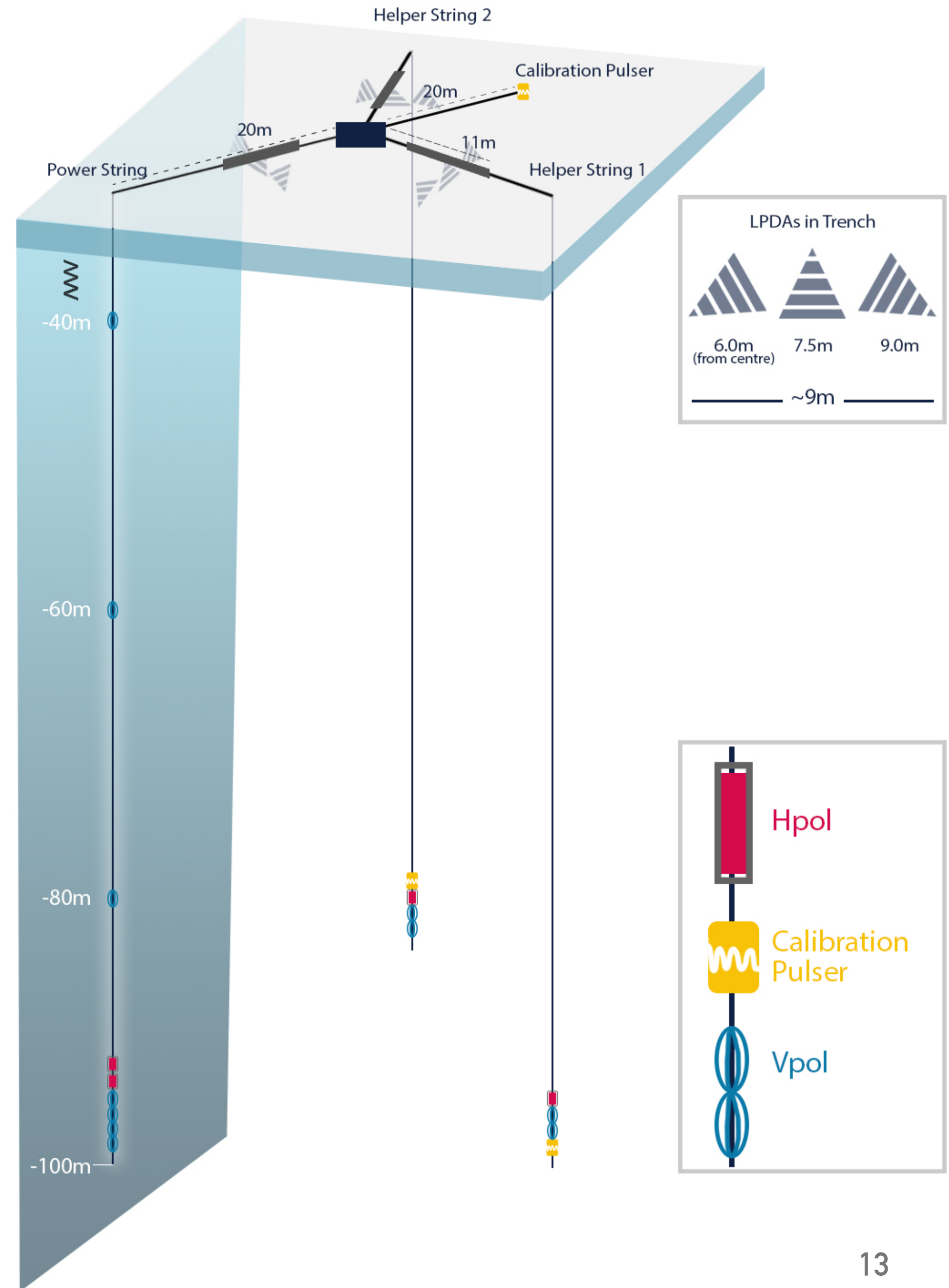
# SUMMARY OF PRECISION

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	Estimated error	How to improve
Ice model	+/- 0.05 ns; additional systematic offset	More precise measurements of cable lengths and cable speeds at cold temperatures prior to deployment
A5 calibration pulse time delays	+/- 0.082 ns	More identical antenna impedances; better in situ measurements; improved digitizers
Phased array calibration pulse time delays	+/- 0.03 ns	
Minuit fit	<5 cm	Orthogonal calibration pulsers would likely make it easier to minimize

# FUTURE PLANS

- RNO-G: first stations to be deployed this summer
- Upgrades:
  - Multiple local calibration pulsers
  - All fibers and cables measured
  - New VPol and HPol antenna designs
  - Better digitizers
- Other calibration sources also planned (DISC borehole, snowmobile surface pulsing, maybe others)



**BACKUP**

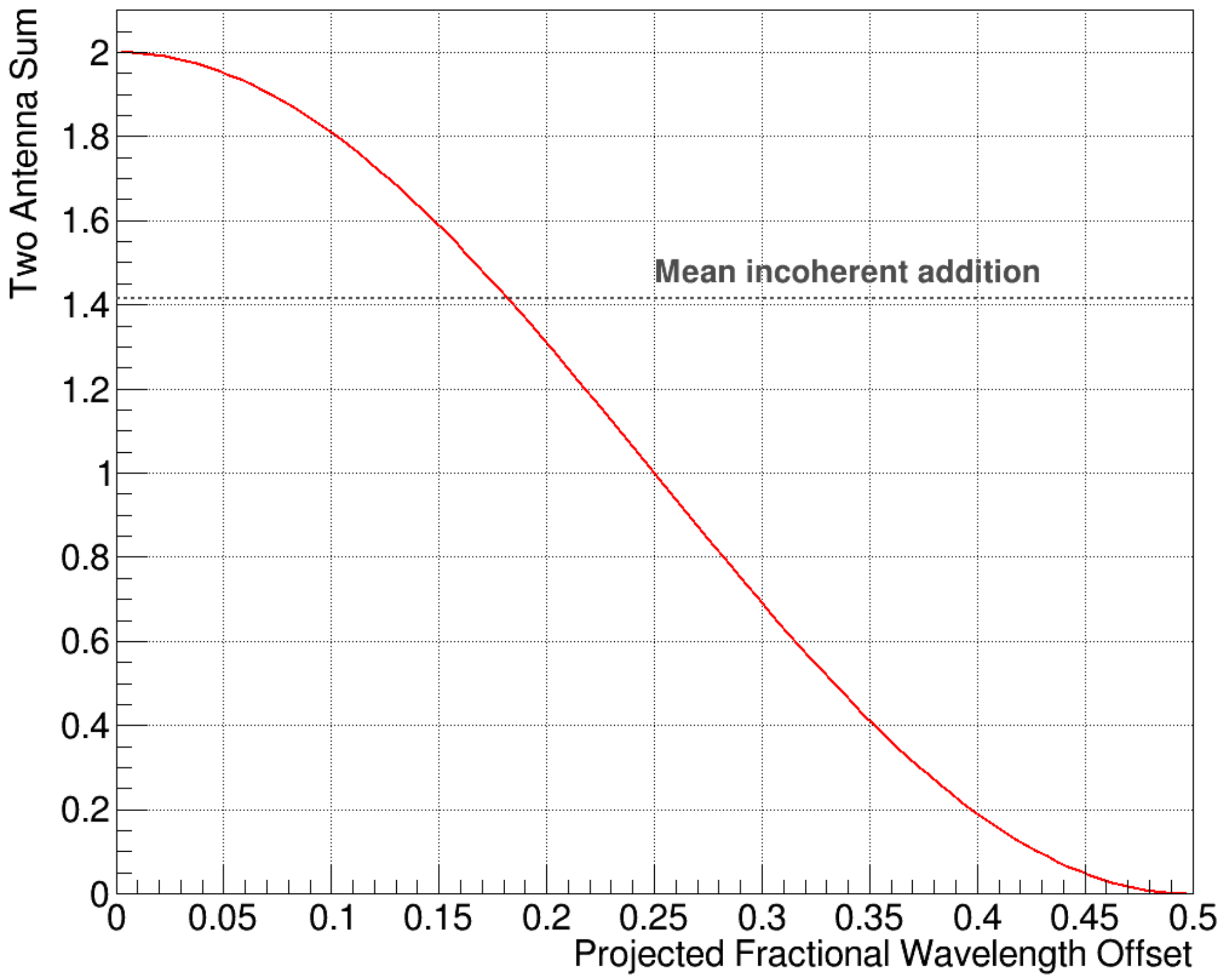
# UNKNOWN AND KNOWN PARAMETERS

Quantity	Unknown Parameters	Known Parameters
(x, y, z) of each A5 VPol antenna	24	
(x, y, z) of local Calibration Pulser	3	
(x, y) of SPIceCore hole	2	
(x, y) of Phased Array	2	
Cable delays for each A5 VPol antenna	8	
Cable delays for each Phased Array VPol antenna	7	
Ice Model	3	
Time Delays from A5 baselines + local calibration pulser		28
Time Delays from A5 baselines + SPIceCore		28
(z) of Phased Array		7
z(t) of SPIceCore pulser		1
Total	49	64

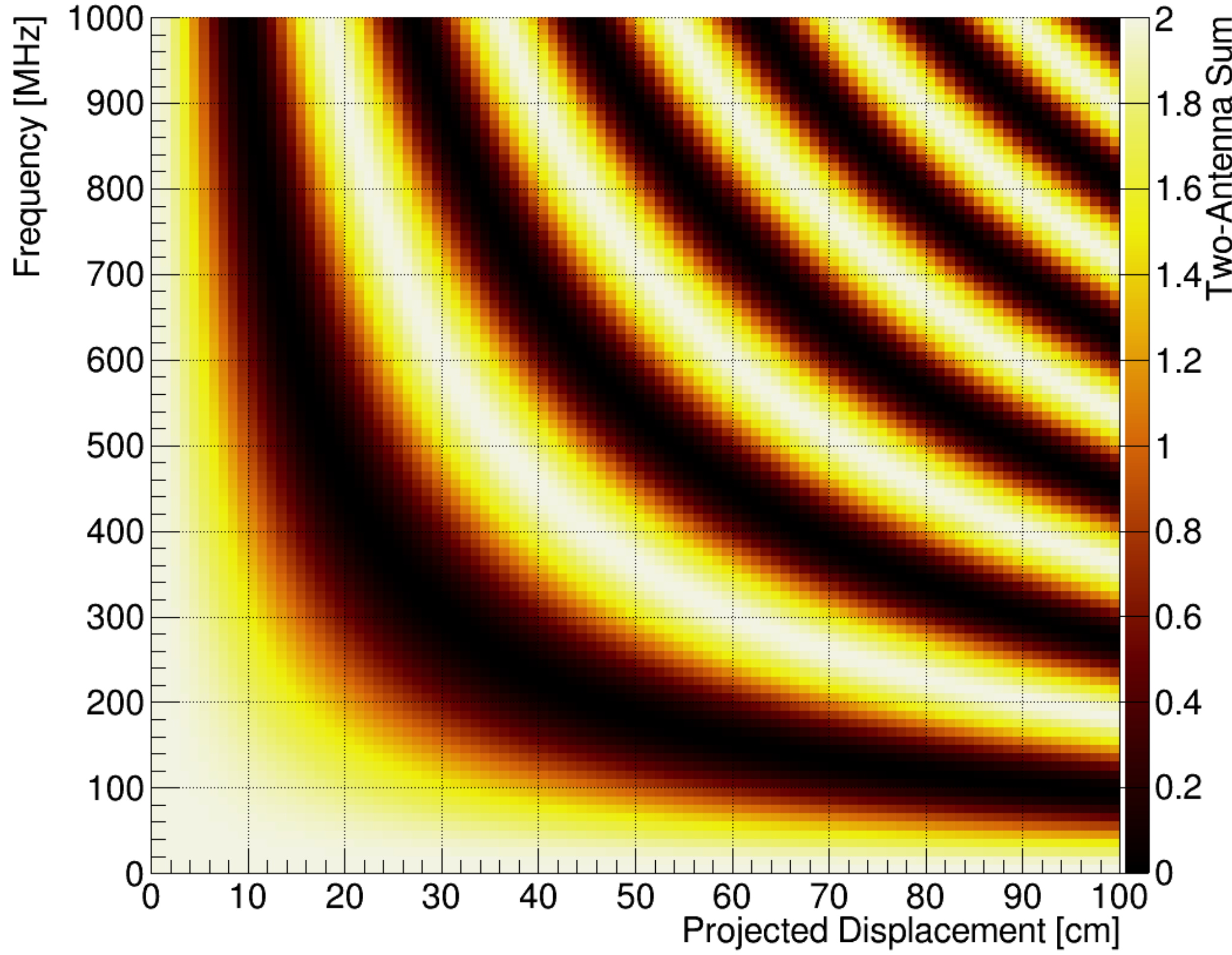
The relationship between time delays and position is not linear- so very hard to know when you have “enough” known parameters

# HOW GOOD DOES THE POSITION FIT NEED TO BE?

Two-Antenna Sum vs. Wavelength Mismatch



n=1.7 Two-Antenna Sum vs. Displacement



To use full waveforms, need to have errors below 10 cm



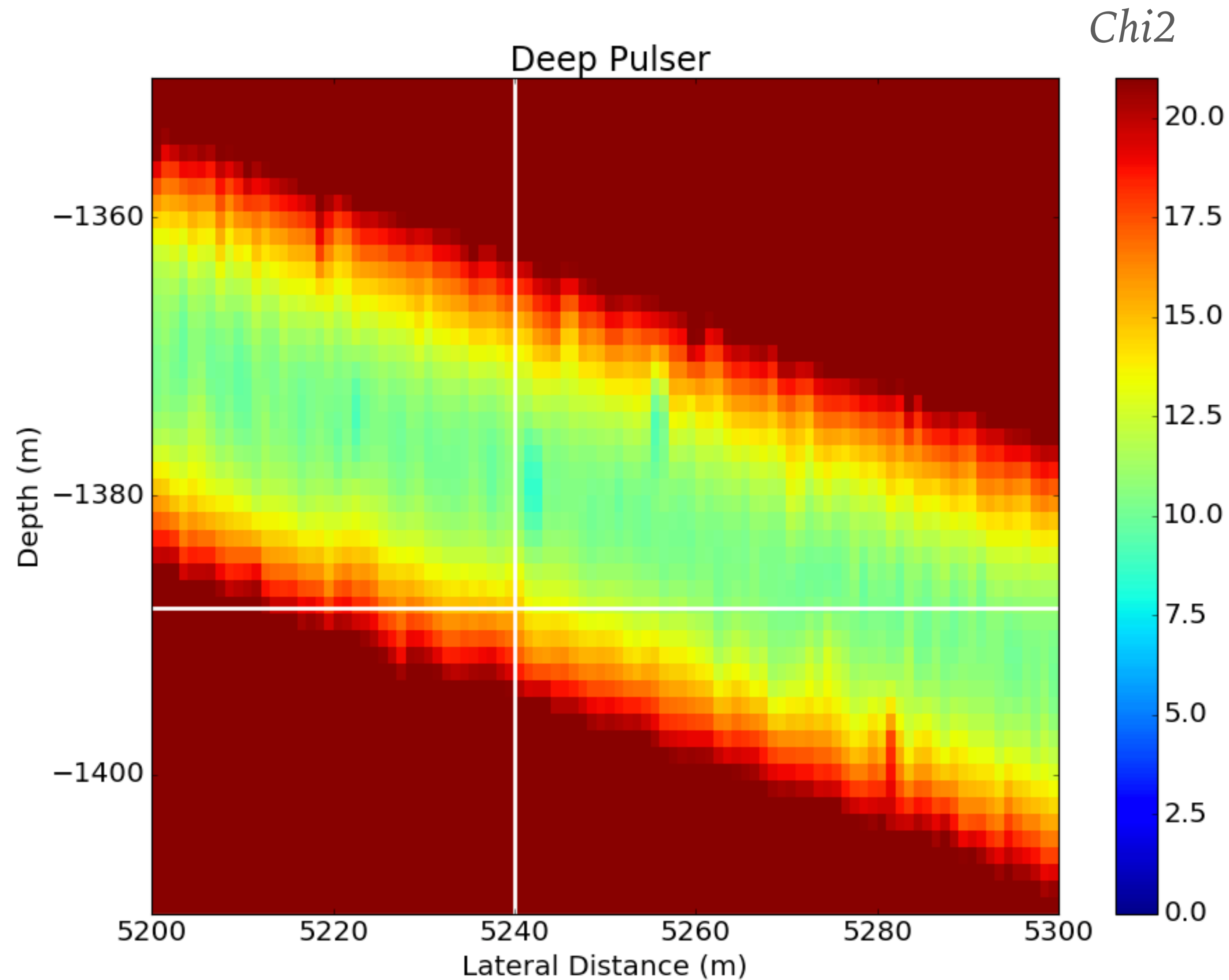
# POTENTIAL ICE MODELS

$$n(z) = A - (A - B)e^{Cz}$$

Model Name	A	B	C
AraSim	1.78	1.35	0.0132
Gorham	1.788	1.325	0.0140
AraRoot	1.78	1.353	0.0160
UNL 2016	1.78	1.326	0.0182
UNL 2016, modified	1.78	1.326	0.0202
RICE (2004)	1.78	1.36	0.0132
SPIceCore 1	1.774	1.293	0.0154
SPIceCore 2	1.774	1.249	0.0163
Uzair	1.78	1.17054	0.0171774

I use this model

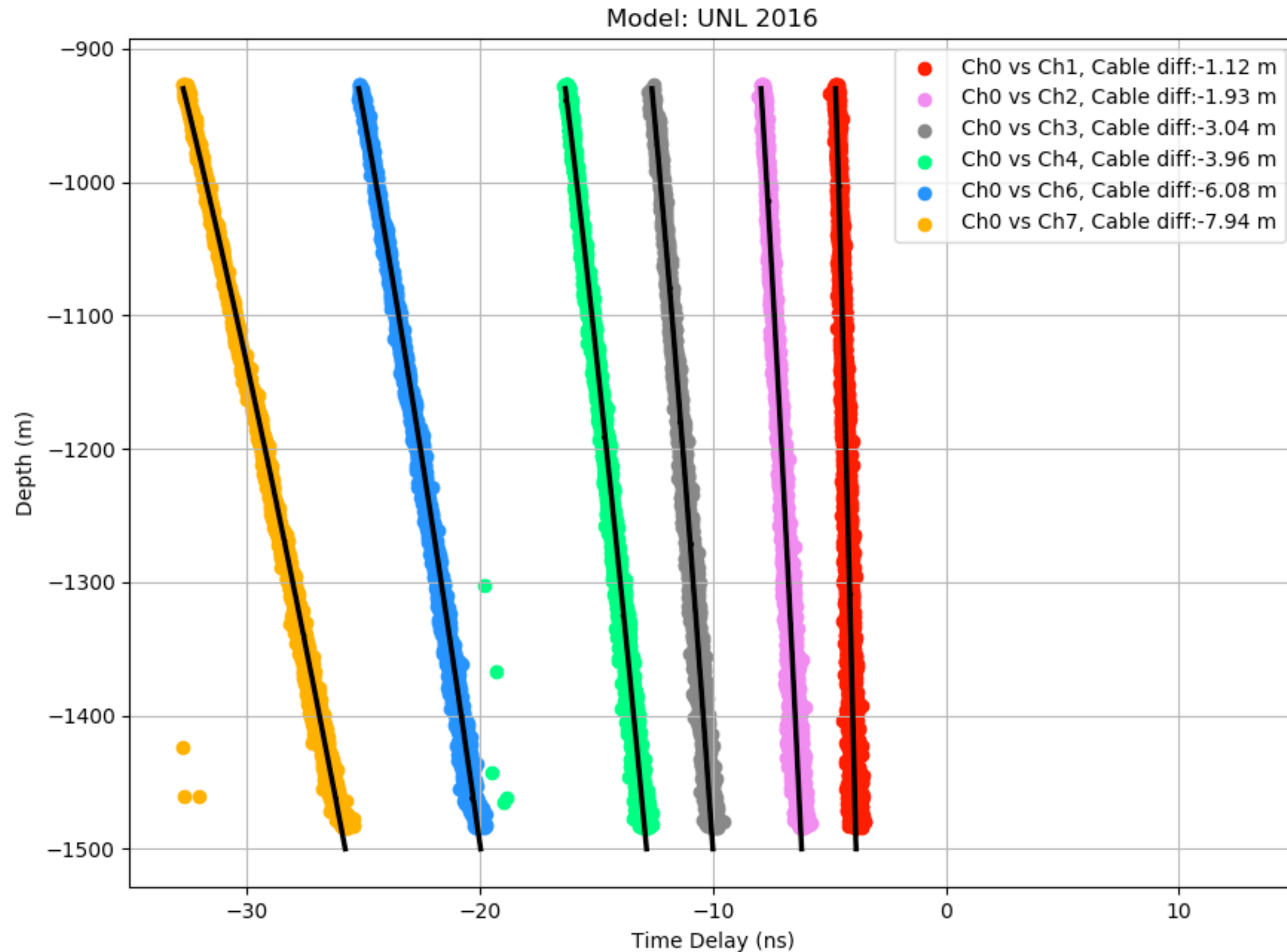
# DOES DEEP PULSER RECONSTRUCT WITH THIS ICE MODEL? YES



- Ice model: UNL (2016) modified
- Using only phased array, deep pulser reconstructs to approximately correct location

DoF = 21

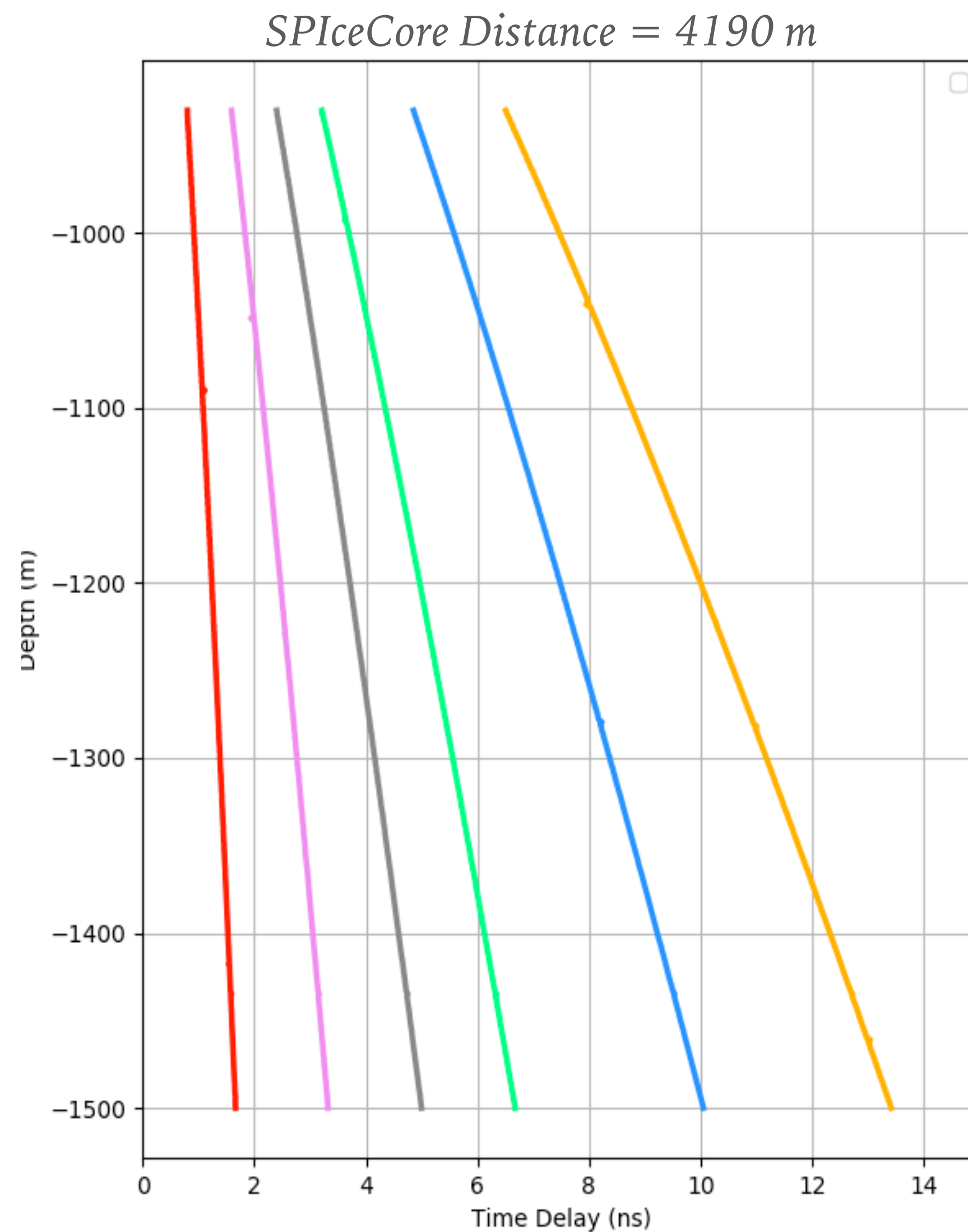
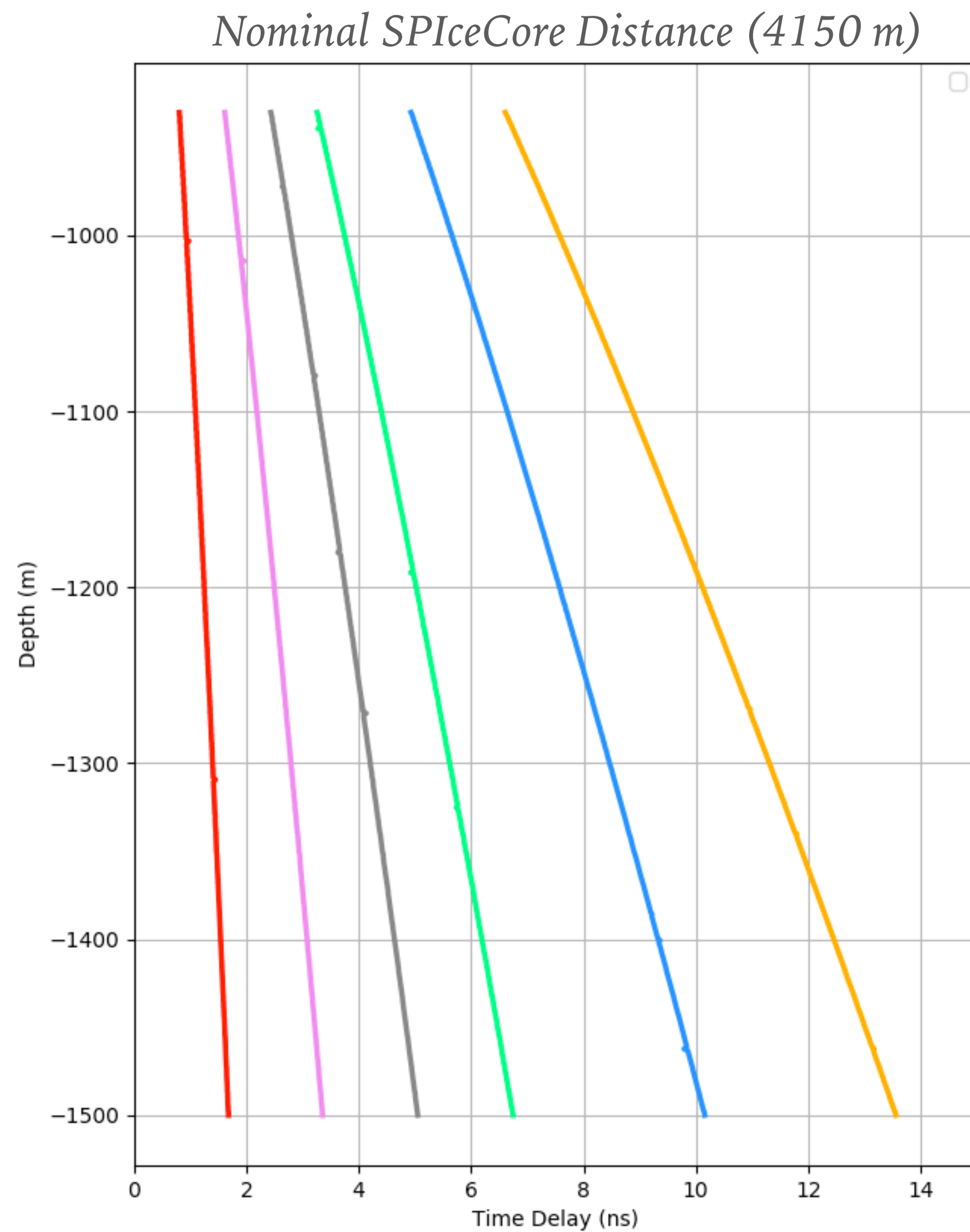
# UNL (2016) MODIFIED: HOW THIS LOOKS ACROSS CHANNELS



- Modified UNL fits all pairs of channels
- Six pairs of channels shown here

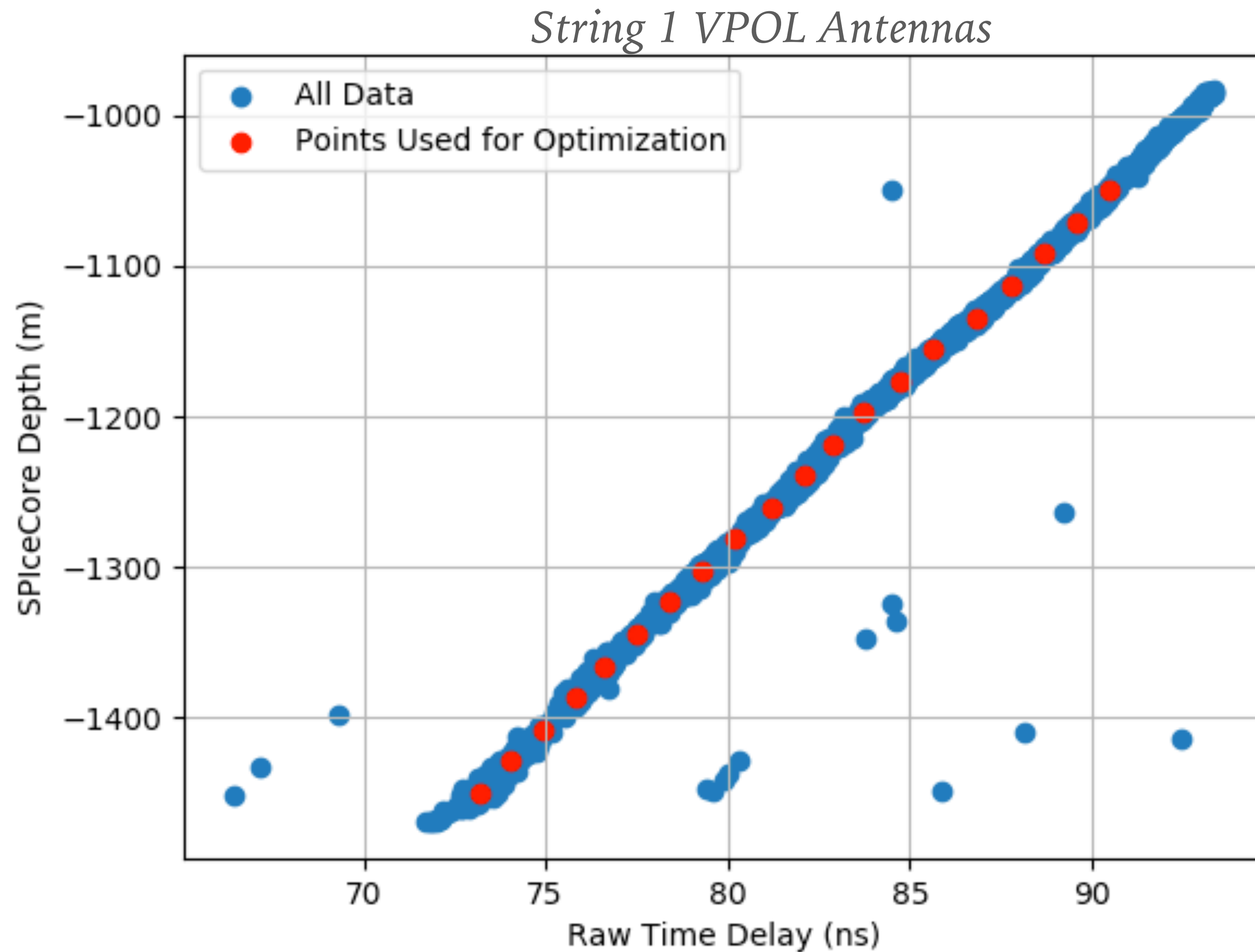
# COULD THIS BE A DISTANCE EFFECT? NOT PRIMARILY

Plotting the model for two different SPiceCore distances does not change answer much



# USING SPICECORE FOR CALIBRATION- A5 CHANNELS

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- For a single antenna pair, plot all time delays as a function of depth of SPiceCore
- Interpolate to find 20 points at specific depths
- Repeat for all baselines
- Note: cable delays here have not been added

# MAP OF ARA STATION 5

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