

IceCube-Gen2 Radio Calibration Workshop



**In-situ calibration device for the
measurement of the snow accumulation
and the index-of-refraction profile**

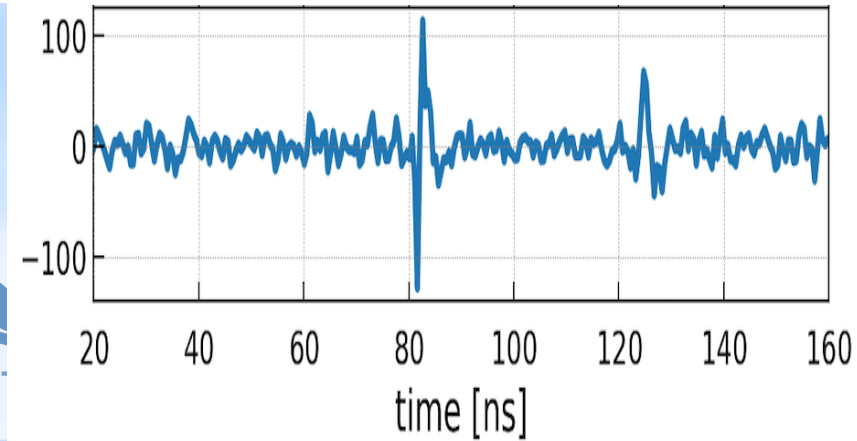
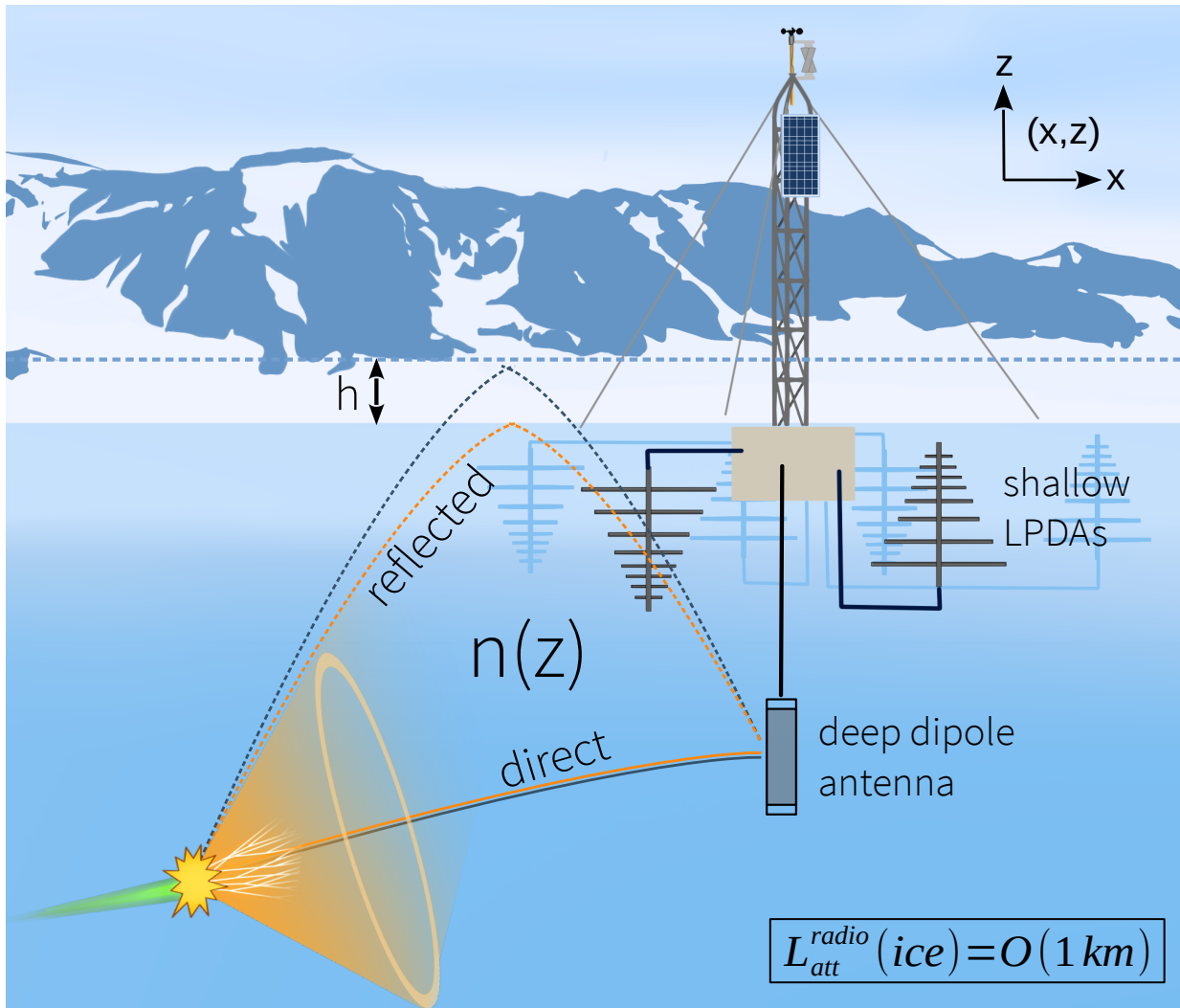
Jakob Beise and Christian Glaser

Uppsala University, Sweden & Humboldt University, Germany

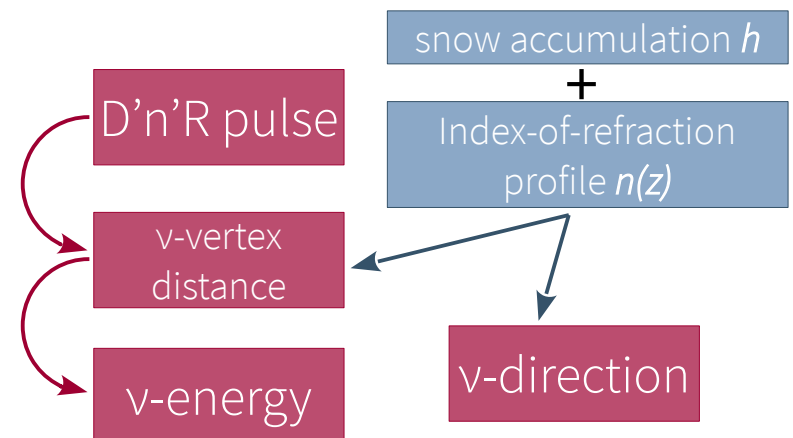
Thursday, April 8th 2021

Firn properties and D'n'R technique

Credit: ARIANNA Collaboration,
Journal of Cosmology and Astroparticle Physics 11(2019)030

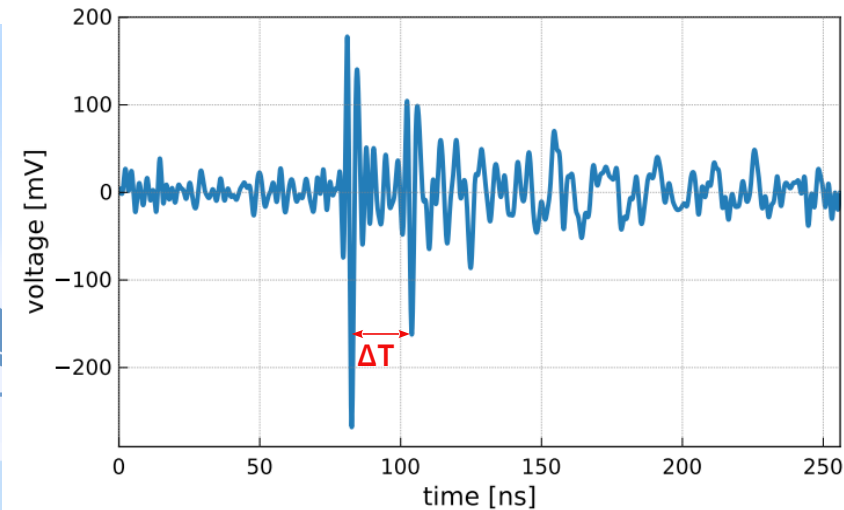
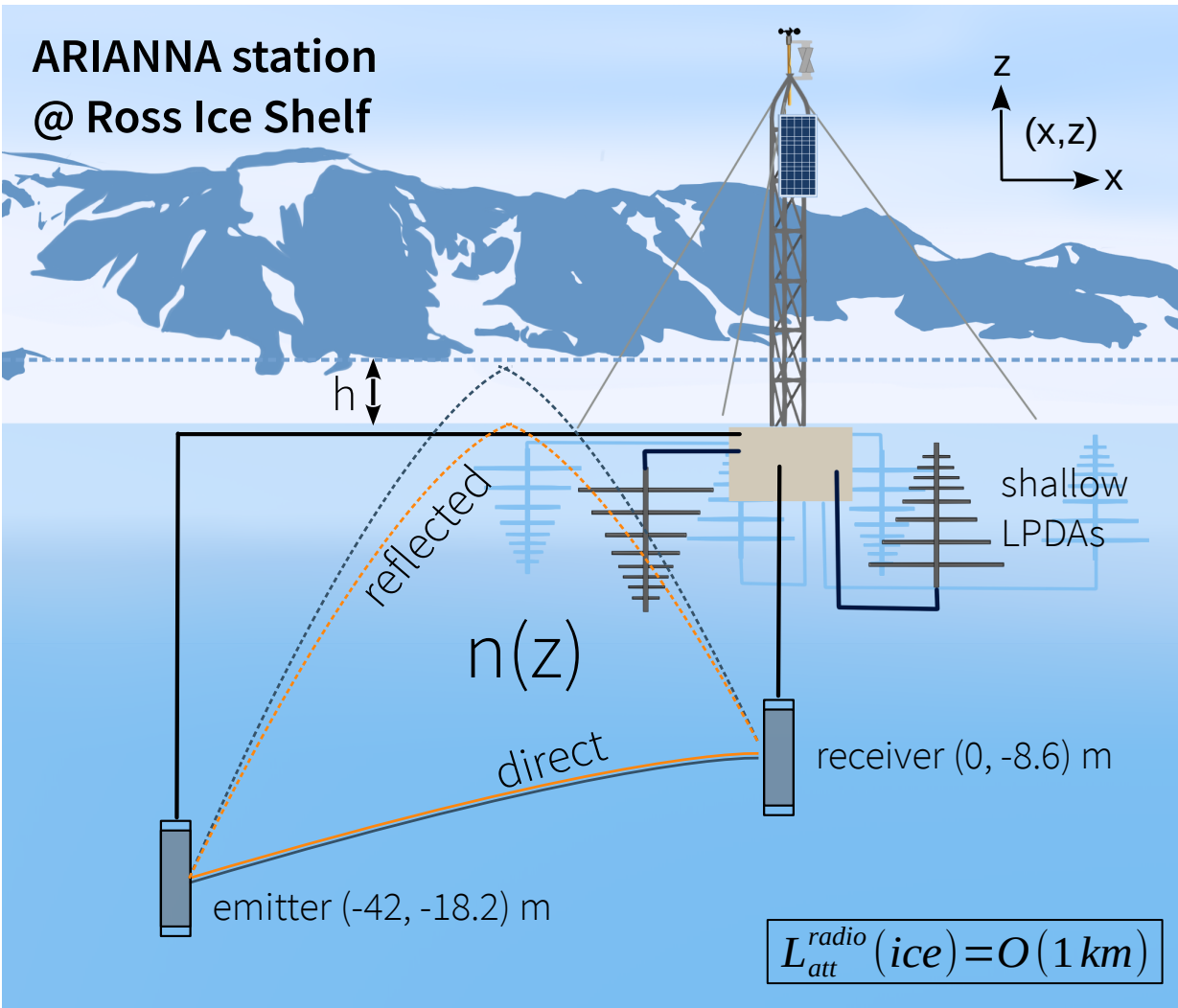


D'n'R technique:



Set-up for the snow accumulation measurement

Credit: ARIANNA Collaboration,
Journal of Cosmology and Astroparticle Physics 11(2019)030

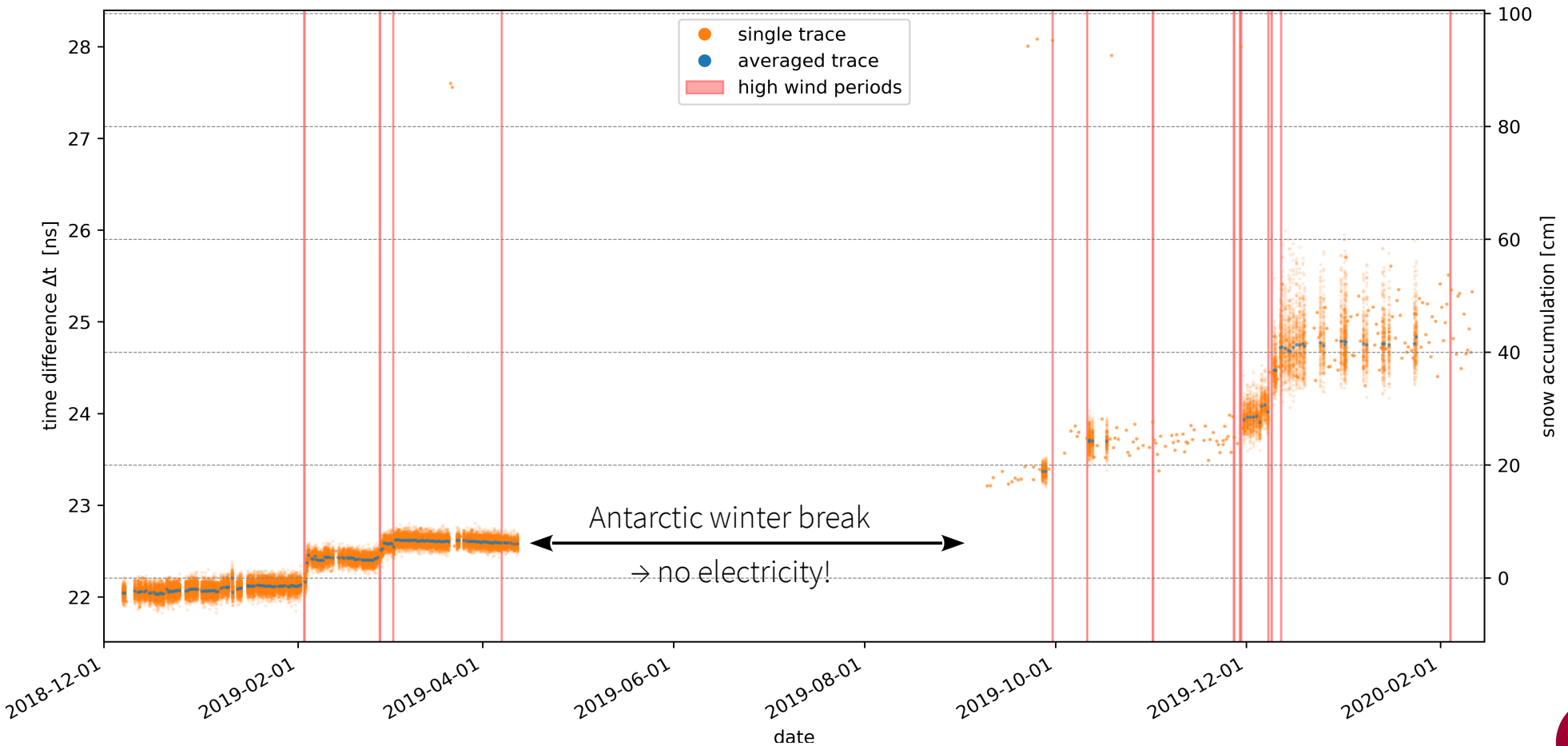


- Pulsar triggers emitter generating a burst of 150 events every 12 h
- $h \sim \Delta T$ using known geometry and $n(z)$

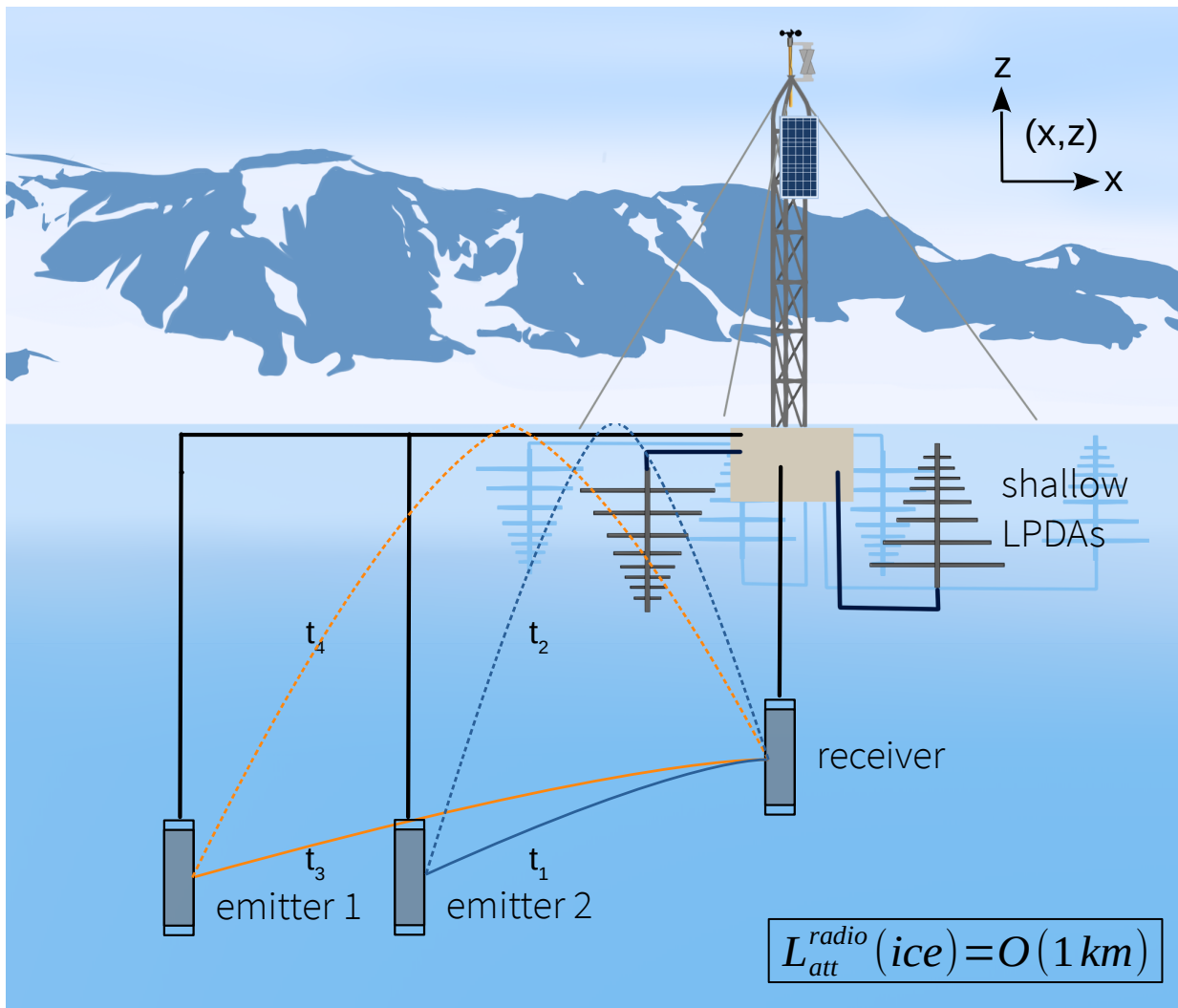
Snow accumulation measurement

Credit: ARIANNA Collaboration,
Journal of Cosmology and Astroparticle Physics 11(2019)030

- In-situ measurement for one ARIANNA station at Ross Ice-Shelf
- Scatter for averaged traces ~ 1 mm (5 ps) to ~ 5 mm (25 ps)
- Jumps in snow accumulation correlate to periods of high winds



Extension to the measurement of the $n(z)$ profile



Refractive index with depth:

$$n(z) = n_0 - \Delta n \cdot \exp(-z/z_0)$$

1.78

Δn : change of refractive index

z_0 : characteristic length

➤ Reconstruct $h, \Delta n, z_0$

➤ 3 independent observables

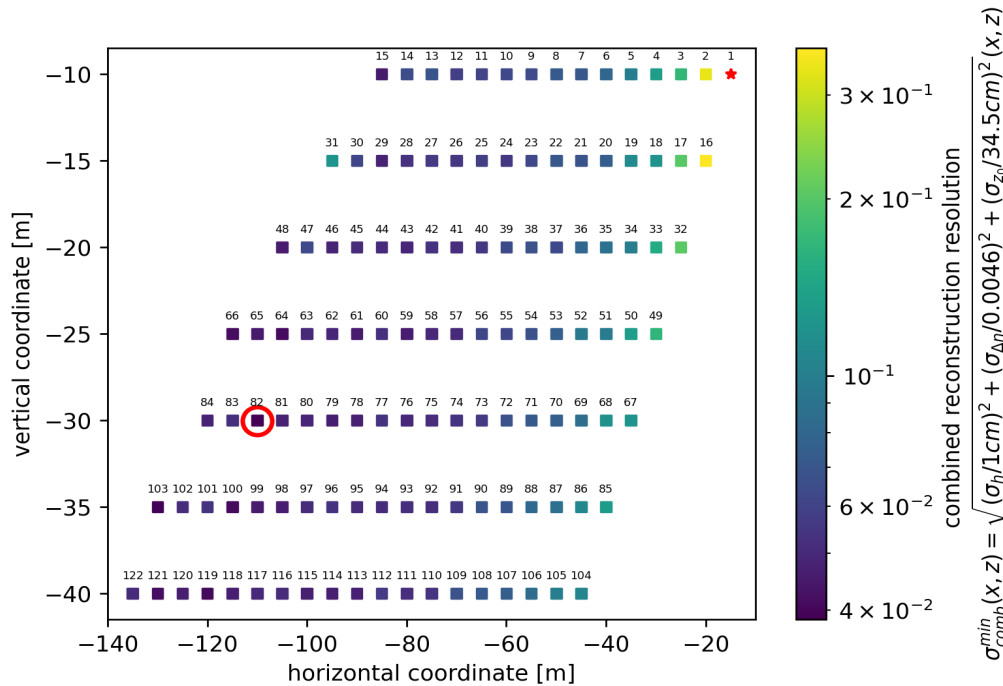
$$\Delta t_{21} = t_2 - t_1, \Delta t_{31} = t_3 - t_1, \Delta t_{41} = t_4 - t_1$$

➤ Optimize emitter configuration

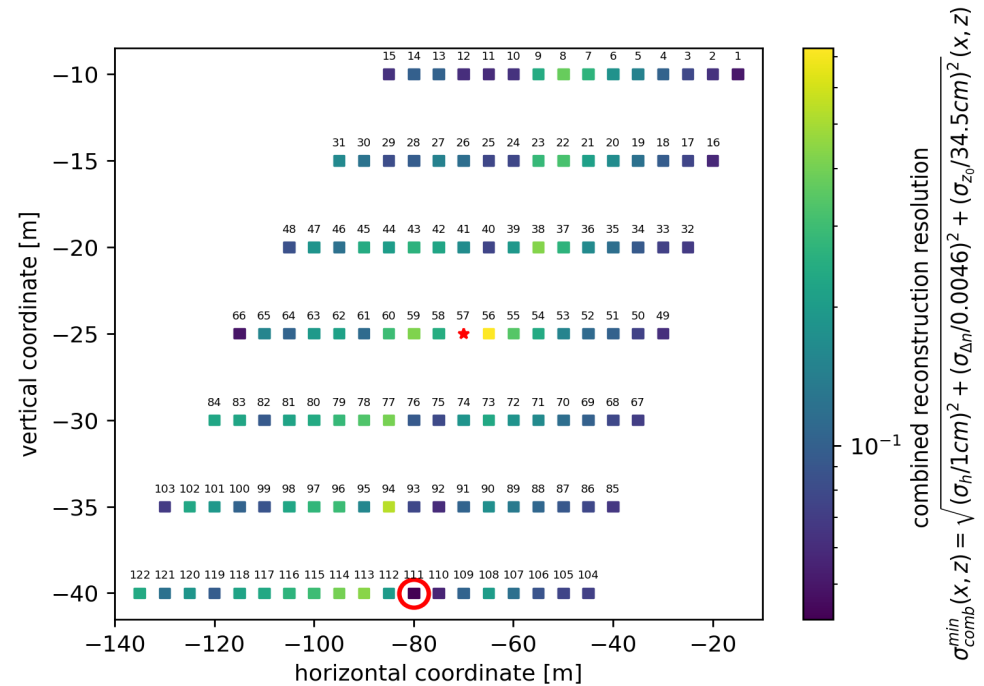
Optimizing emitter position

- Simulate 10,000 measurements with event uncertainty of 0.2 ns
- Grid of possible conditions is constrained

for E1 @ (-15,-10) m, optimal E2 @ (-110,-30) m



for E1 @ (-70,-25) m, optimal E2 @ (-80,-40) m



$$\sigma_h \sim O(cm), \sigma_{z_0} \sim O(m) \Rightarrow \sigma_{comb} = \sqrt{(\sigma_h/1cm)^2 + (\sigma_{\Delta n}/0.0046)^2 + (\sigma_{z_0}/34.5cm)^2}$$

Achievable precision

- Find configuration of emitters where **reconstruction resolution** and **correlation** between parameters is minimal:
- Reconstruction resolution independent of E1 position, E2 can always be placed so that combined resolution is good
↳ Correlation provides a clear minimum

optimal configuration: E1: [-115, -20] m, E2: [-30, -25] m
(per 100 events uncertainty)

(relative) Reconstruction Resolution	$\sigma_h = 4 \text{ mm}$	$\sigma_{\Delta n} = 0.08\%$	$\sigma_{z_0} = 0.12\%$
Correlation	$\rho_{h,\Delta n} = -4\%$	$\rho_{h,z_0} = -3\%$	$\rho_{\Delta n,z_0} = 15\%$

Systematic uncertainty

- A) Emitter deployment uncertainty in \mathbf{x} of ± 1 cm

$$\sigma(\Delta n) \sim 0.022\%, \sigma(\mathbf{z}_0) \sim 0.004\%$$

- B) Emitter deployment uncertainty in \mathbf{z} of ± 1 cm

$$\sigma(\Delta n) \sim 0.08\%, \sigma(\mathbf{z}_0) \sim 0.09\%$$

- C) Do we have a sensitivity to deviation from the exponential $n(z)$ profile?

Summary

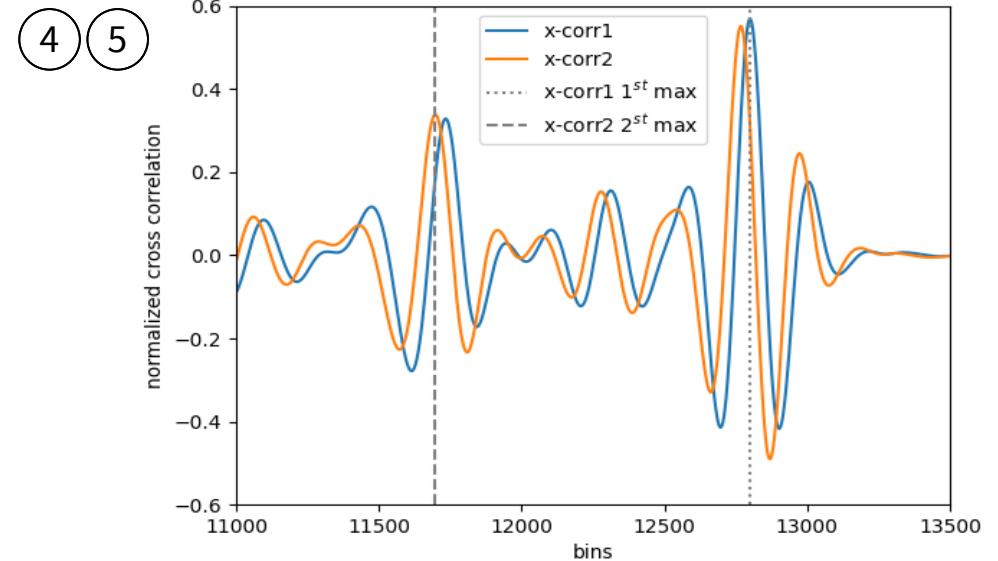
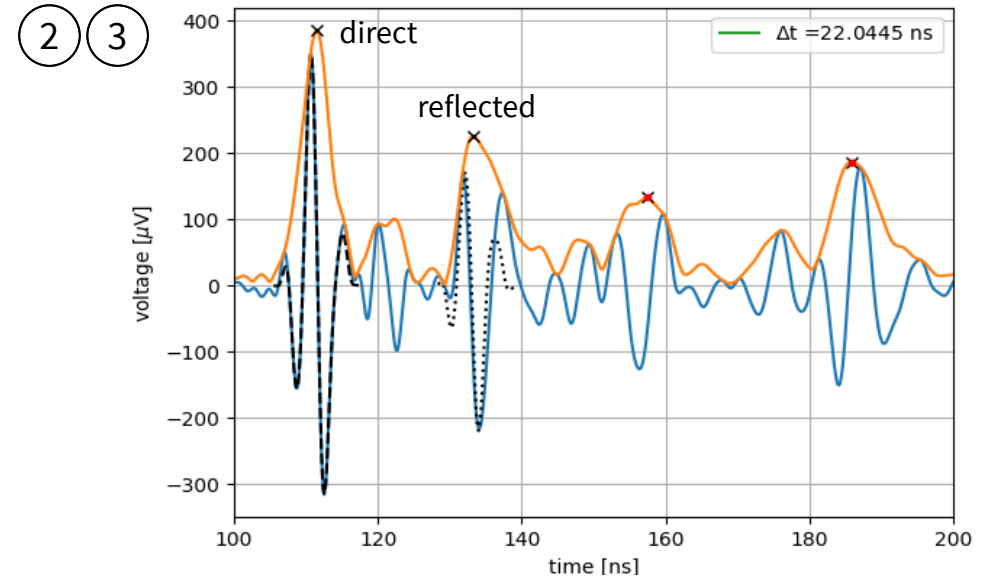
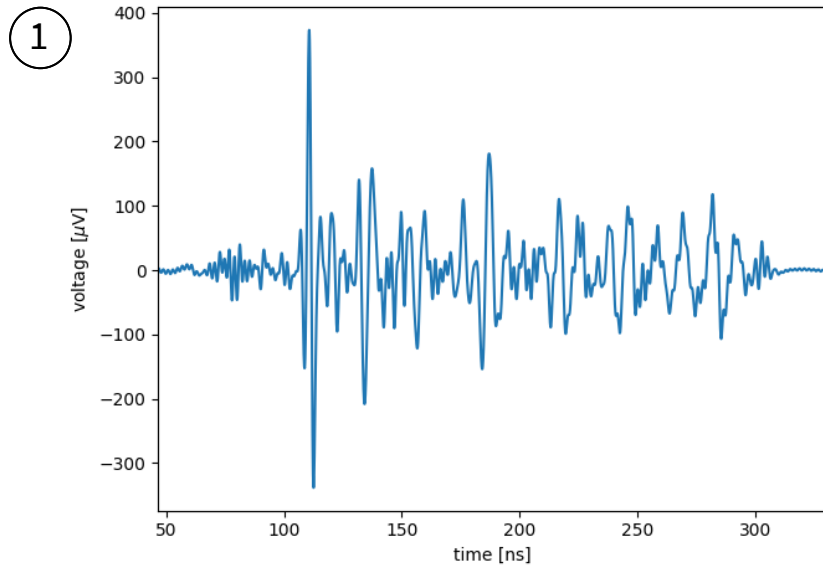
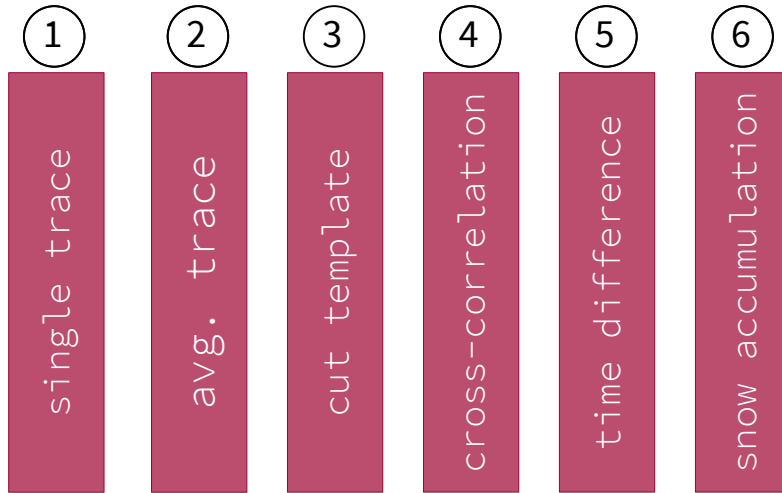
- In-situ calibration device for snow accumulation measurement
- Extension to a two emitter calibration device
 - $\sigma(h) = 4 \text{ mm}$
 - $\sigma(\Delta n) = 0.08\%$
 - $\sigma(z_0) = 0.12\%$
- Study of systematic uncertainties likely the dominating factor
- Impact of $n(z)$ uncertainty on the v-direction reconstruction

Proposal: Equip every station with at least one emitter and check variation in $n(z)$ profile for a few stations equipped with the two emitter calibration system. Adapt layout if necessary.

BACKUP SLIDES

Details of snow accumulation analysis

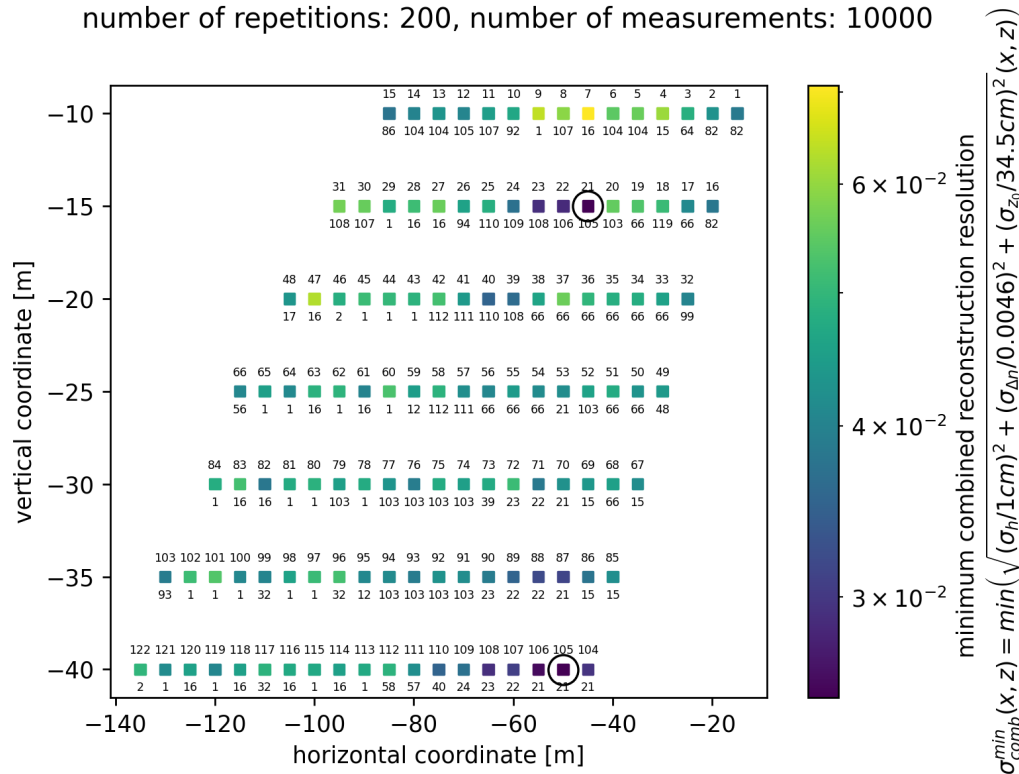
Credit: ARIANNA Collaboration,
Journal of Cosmology and Astroparticle Physics 11(2019)030



Reconstruction Resolution and Correlation

Reconstruction Resolution

number of repetitions: 200, number of measurements: 10000



Correlation

number of repetitions: 200, number of measurements: 10000

