

# Neutrino Physics

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*(4<sup>th</sup> year grad student)*

# Outline

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- Neutrinos in the Standard Model
- Neutrino Oscillations
- Open Questions

# Neutrinos in the Standard Model

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# Why study neutrinos?

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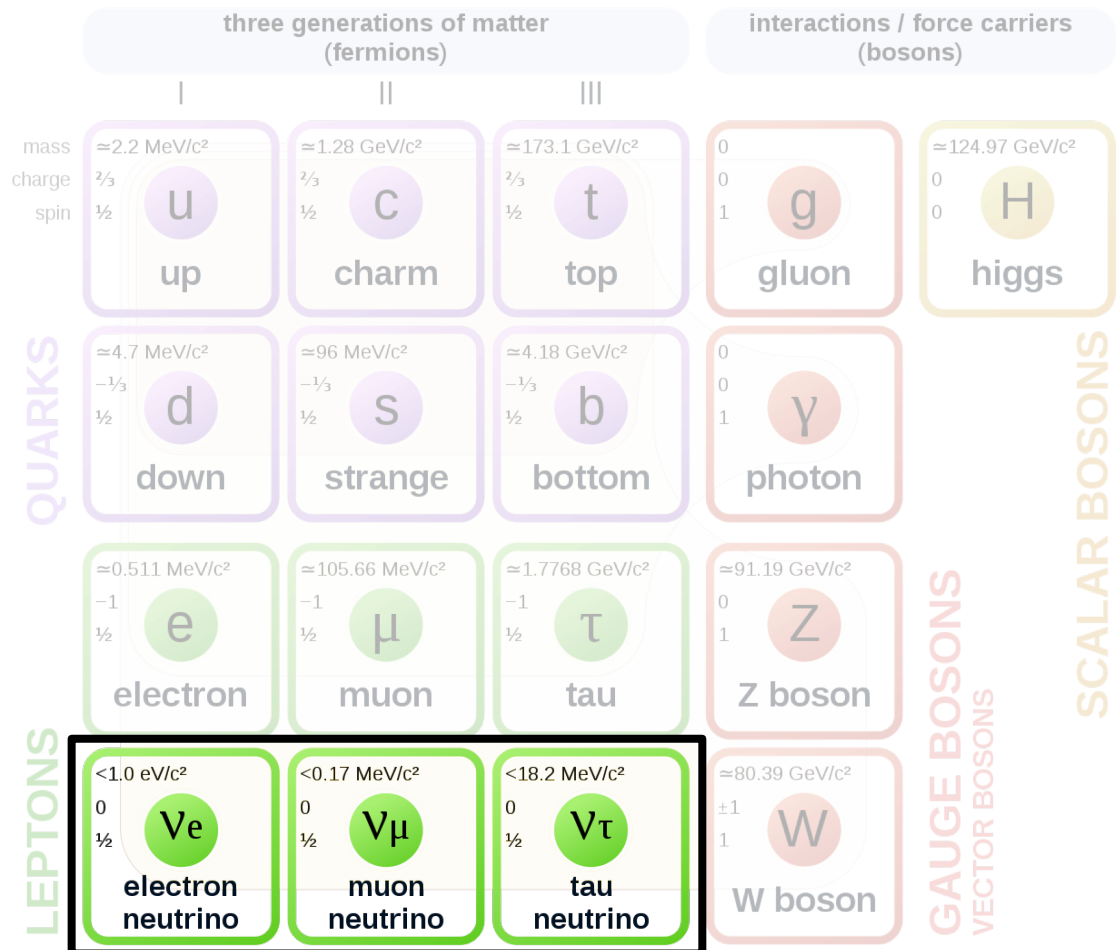
- One of the most abundant particles in the Universe
- Yet also one of the least well understood
- Properties that confirm the existence of physics Beyond the Standard Model (neutrino oscillations)



# Neutrinos in the Standard Model

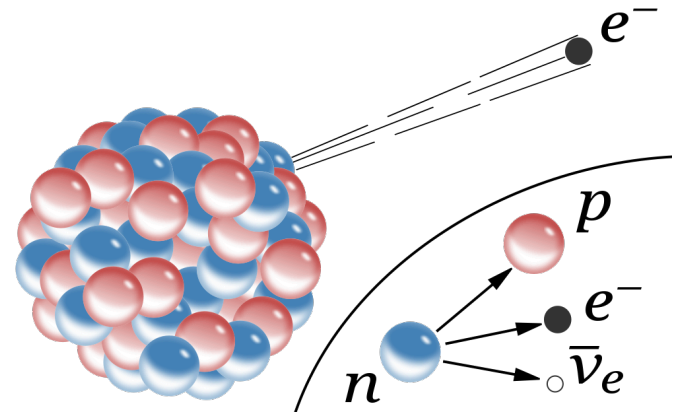
- Leptons
- no charge
- “no mass”
- 3 flavors
- Each flavor has a charged lepton as a sister particle

## Standard Model of Elementary Particles

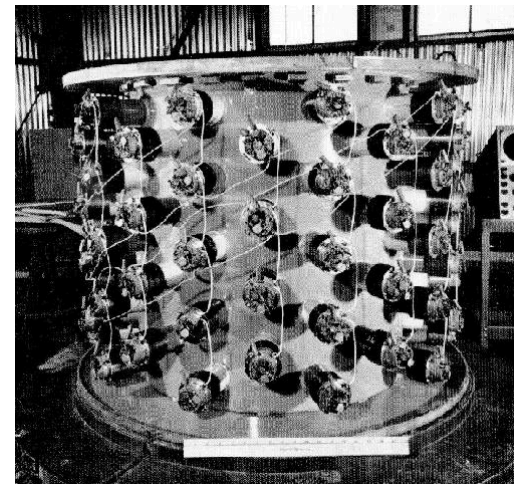


# History

- 1931 - Wolfgang Pauli proposes electrically neutral, invisible particle called a “neutron” to explain missing energy in beta decays
- 1934 - Enrico Fermi proposes beta decay theory, including “neutrino”
- 1956 - Frederick Reines & Clyde L. Cowan discover neutrino at Savannah River Site with underground reactor experiment

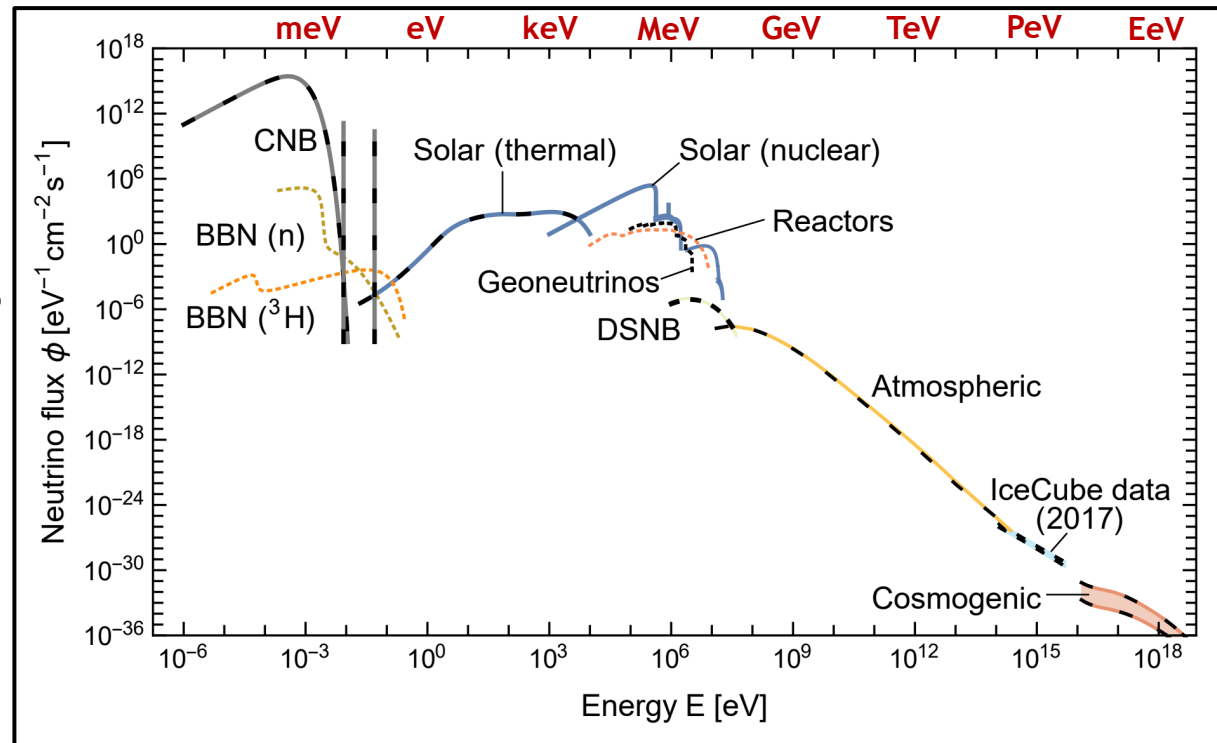


1995



# Where do neutrinos come from?

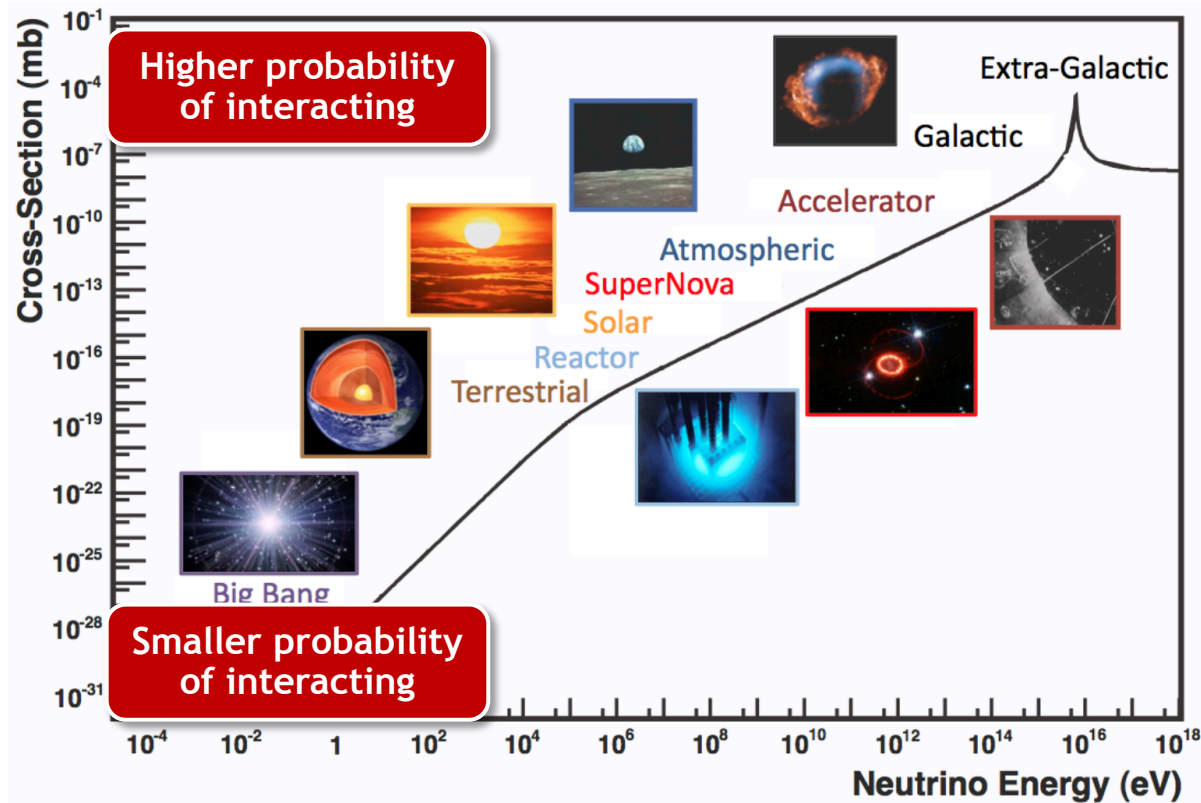
- Natural
  - Big Bang
  - Sun
  - Earth (Geo.)
  - Earth Atmosphere
  - Galactic
  - Extragalactic
- Man-made
  - Nuclear reactors
  - Accelerators



[arXiv:1910.11878](https://arxiv.org/abs/1910.11878)

# Cross Section

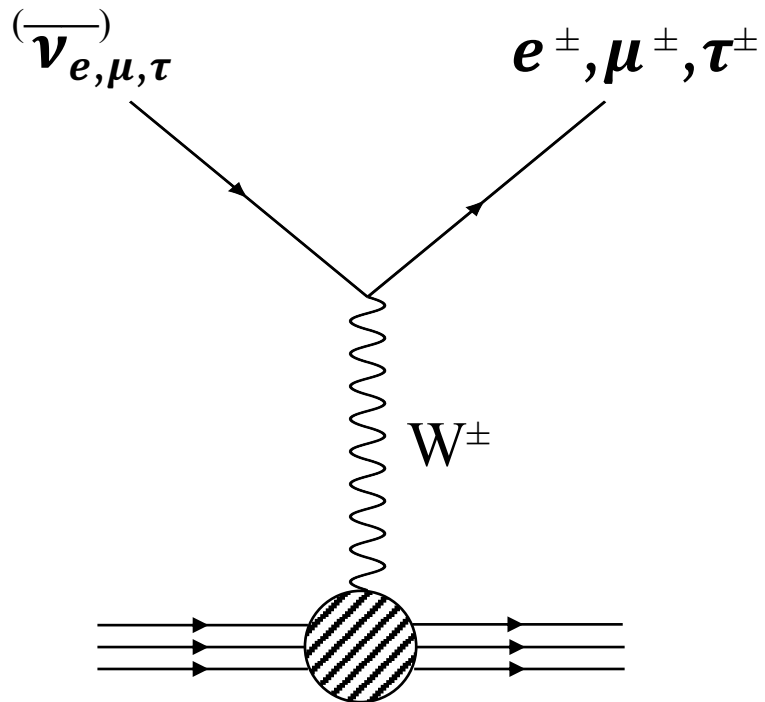
- Neutrinos rarely interact with matter because of their small cross section
- Cross section scales with energy





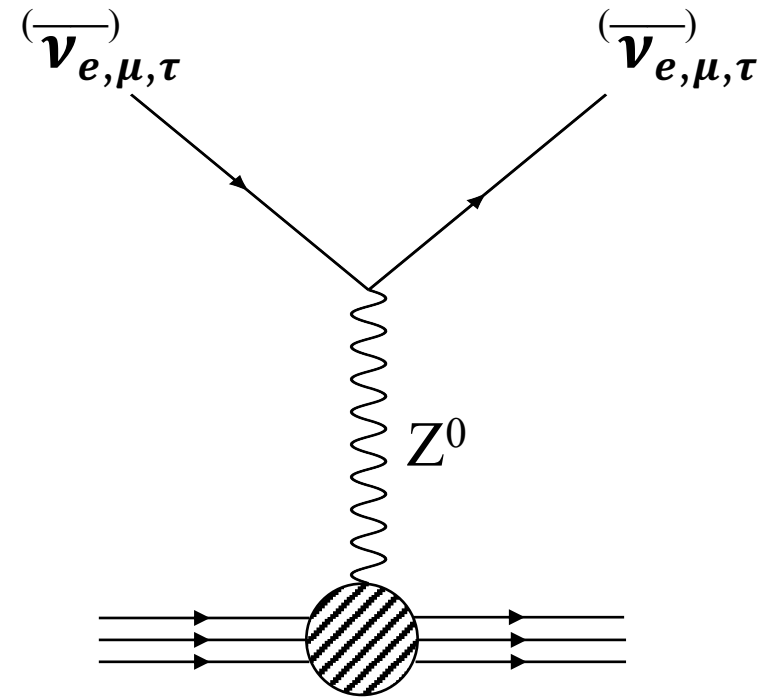
# Neutrino Interactions

## Charged Current



Neutrino in,  
Charged particle out

## Neutral Current



Neutrino in,  
Neutrino out

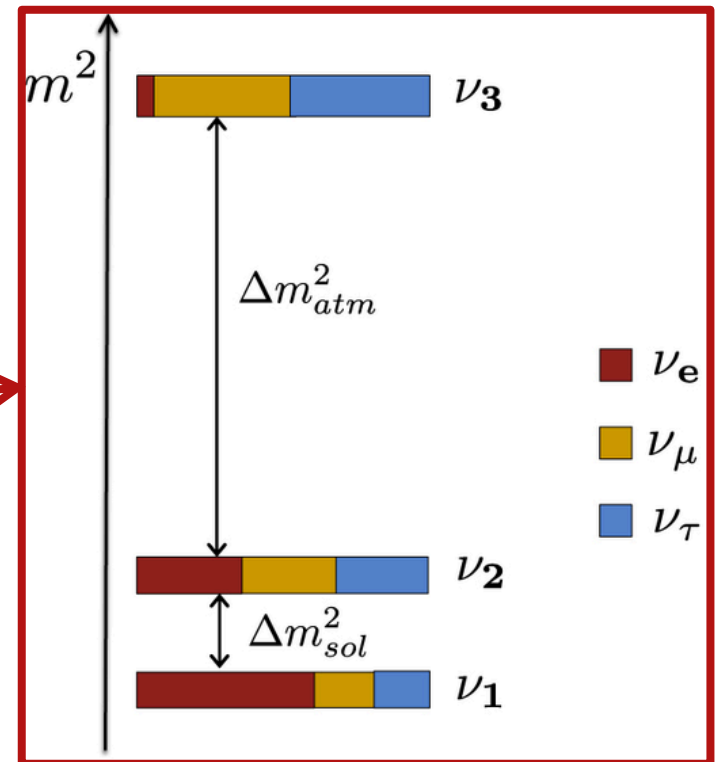
# Neutrino Oscillations

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# Neutrino Masses

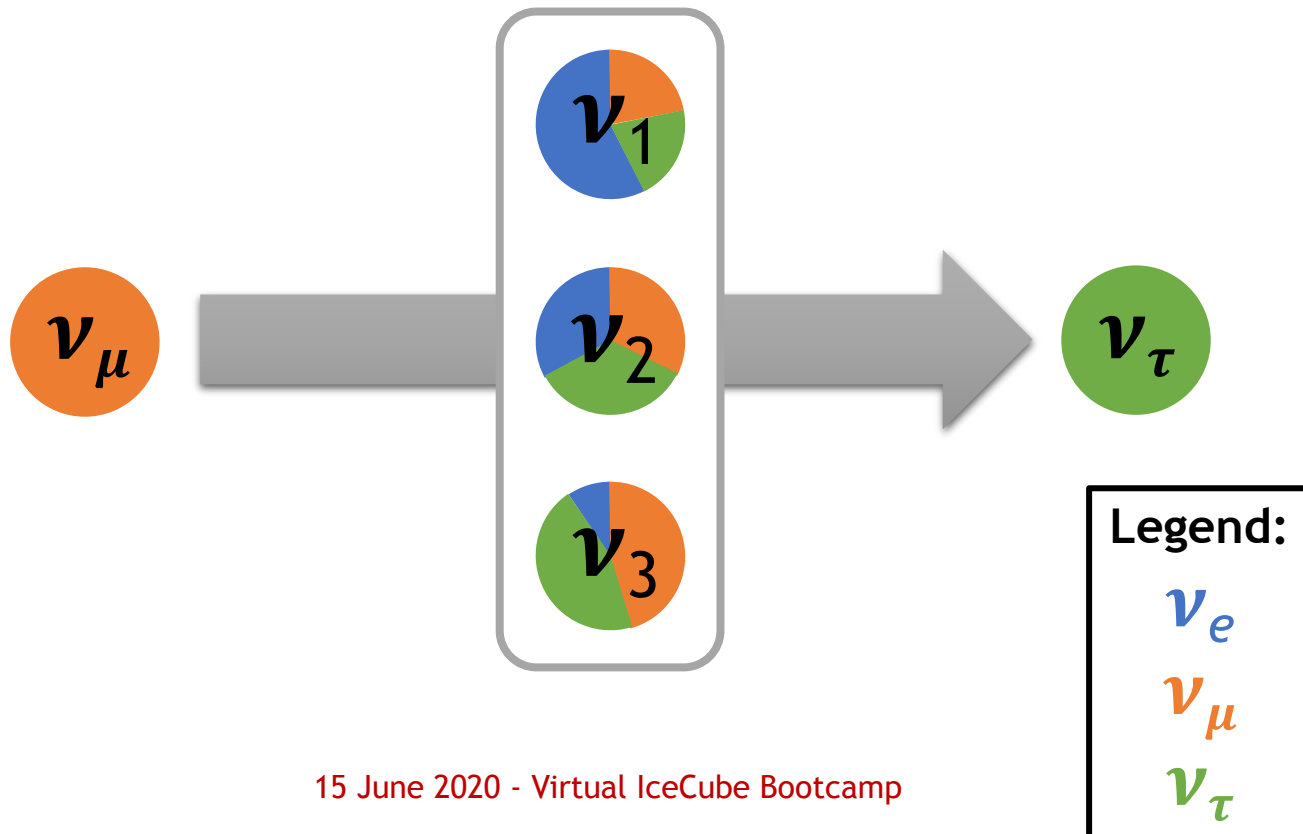
- There are 3 “orthogonal” neutrinos.
- But there are different bases to describe them.
- The bases do not have the same eigenstates.
  - Mass Basis:  $\nu_1$ ,  $\nu_2$ ,  $\nu_3$
  - Flavor Basis:  $\nu_e$ ,  $\nu_\mu$ ,  $\nu_\tau$

**Mass states can be expressed as linear combination of flavor states.**



# Neutrino Oscillations

- Neutrinos are created in their flavor state
- Neutrinos travel in their mass states
- Neutrinos are detected in their flavor state



# Describing Neutrino Oscillations

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- We want to perform a change of basis.
  - Mass Basis:  $\nu_1$  ,  $\nu_2$  ,  $\nu_3$
  - Flavor Basis:  $\nu_e$  ,  $\nu_\mu$  ,  $\nu_\tau$
- The matrix that performs this transformation is known as the PMNS Matrix (Pontecorvo-Maki-Nakagawa-Sakata).

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{bmatrix} = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{bmatrix}$$

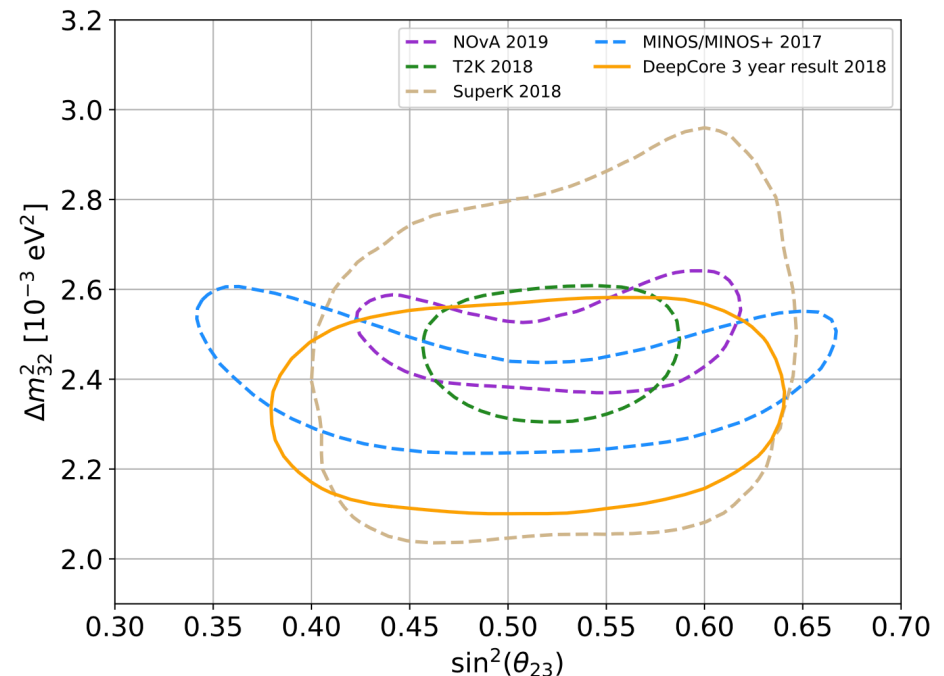
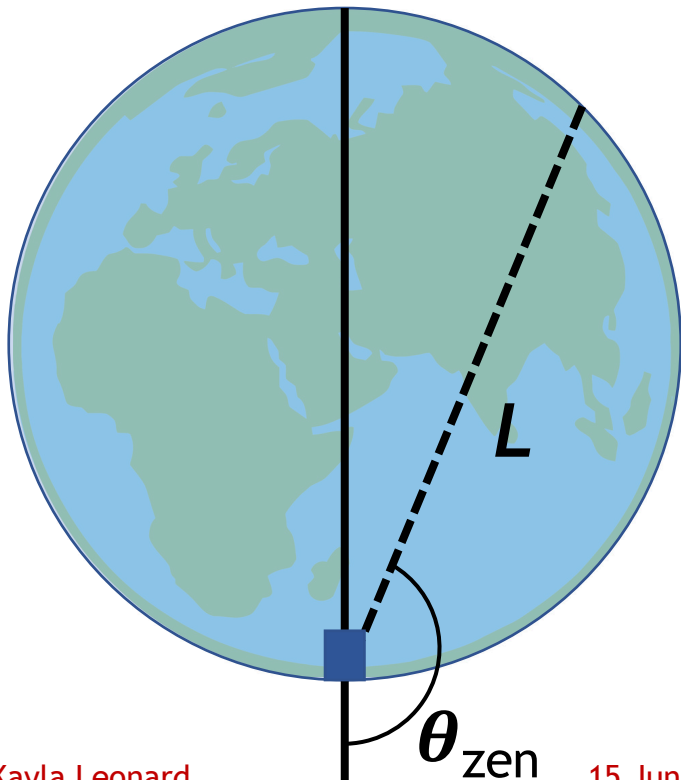
- We can use PMNS matrix to calculate the probability of starting in one flavor state, and ending in a different flavor state:

$$P_{\alpha \rightarrow \beta, \alpha \neq \beta} = \sin^2(2\theta) \sin^2 \left( 1.27 \frac{\Delta m^2 L \text{ [eV}^2\text{] [km]}}{E \text{ [GeV]}} \right)$$

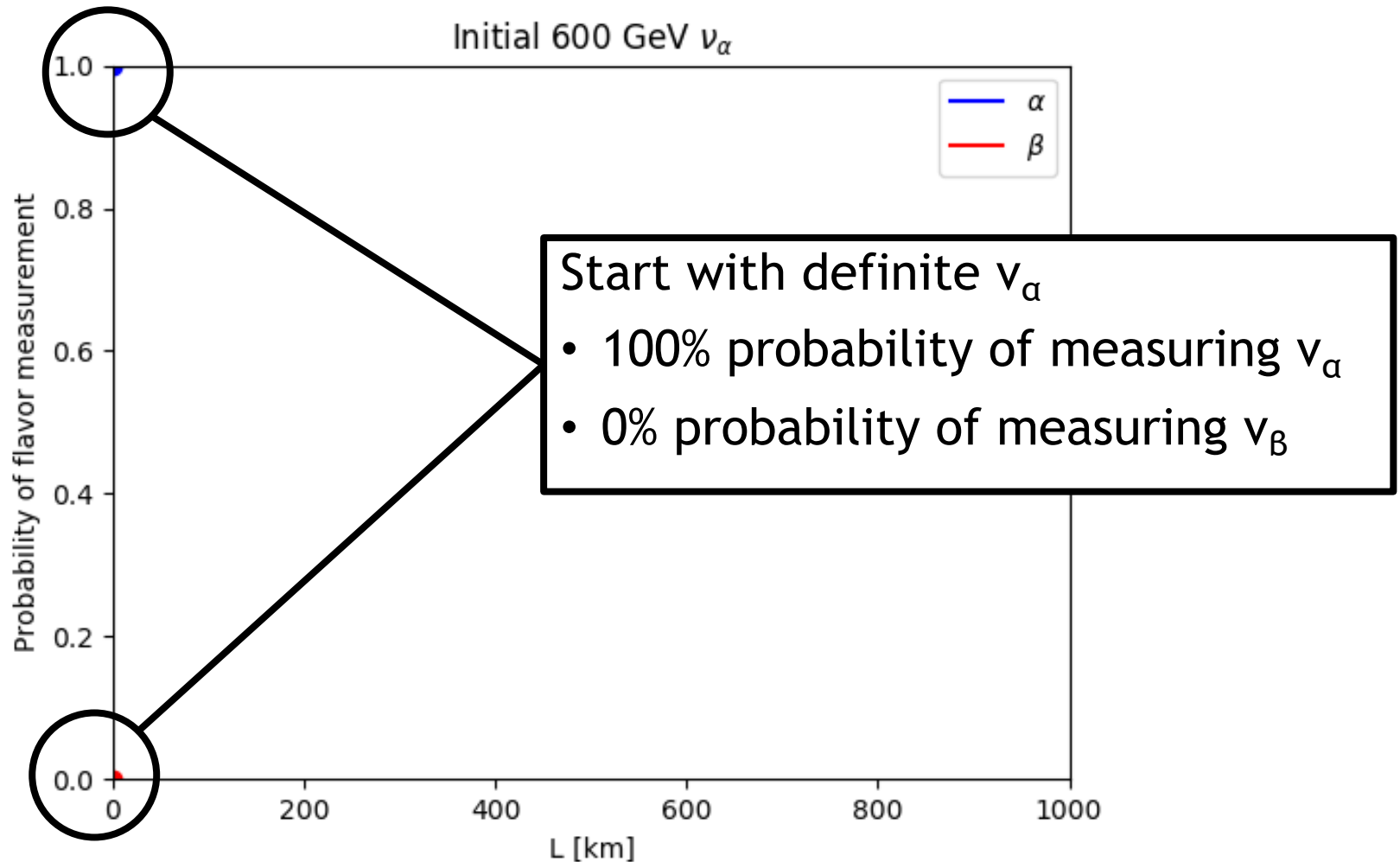
# Describing Neutrino Oscillations

$$P_{\alpha \rightarrow \beta, \alpha \neq \beta} = \sin^2(2\theta) \sin^2 \left( 1.27 \frac{\Delta m^2 L [\text{eV}^2] [\text{km}]}{E [\text{GeV}]} \right)$$

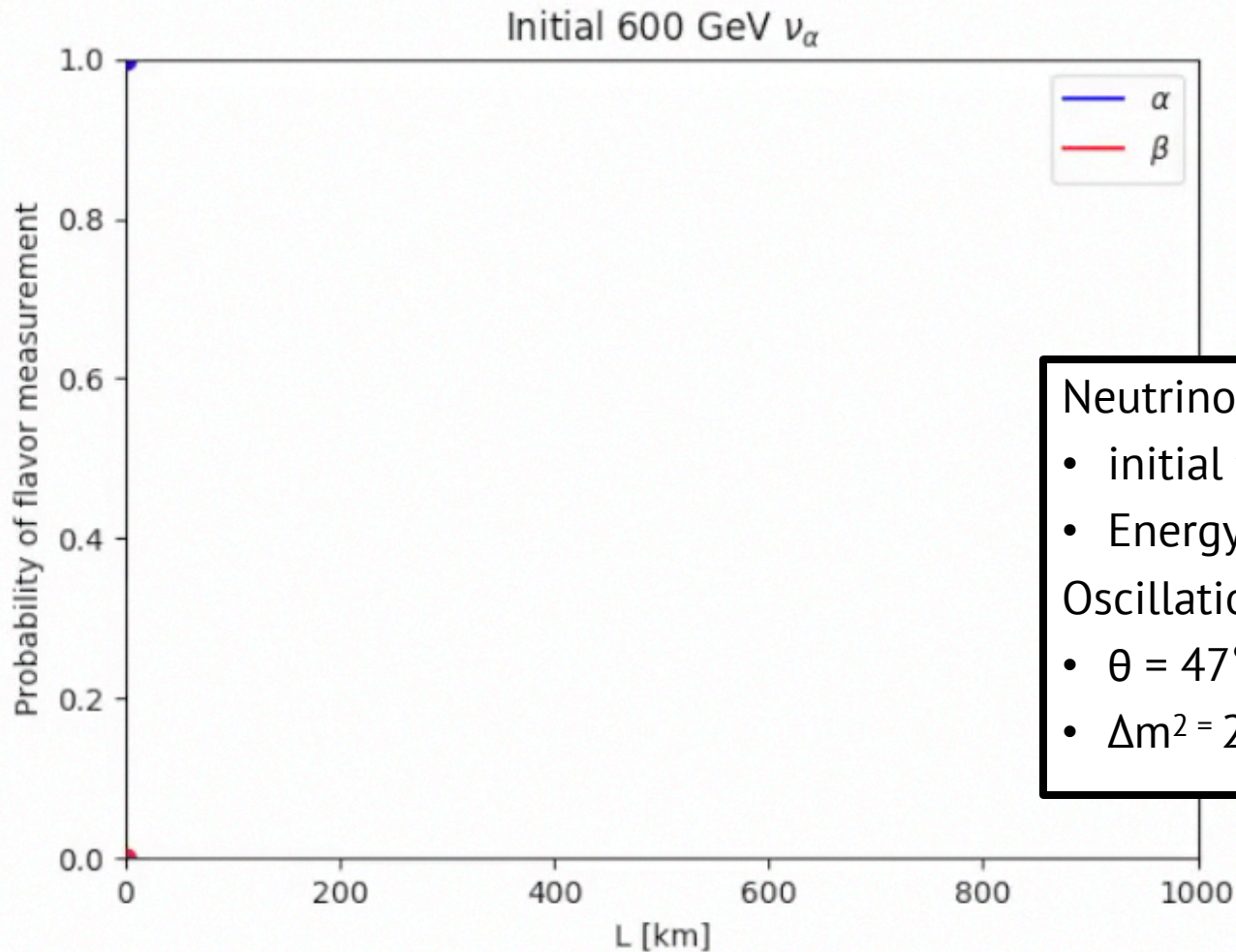
- We can estimate the **Energy** and the **Length** travelled
- We want to measure  $\sin^2(\theta)$  and  $\Delta m^2$  (physics parameters)



# Toy Model



# Toy Model

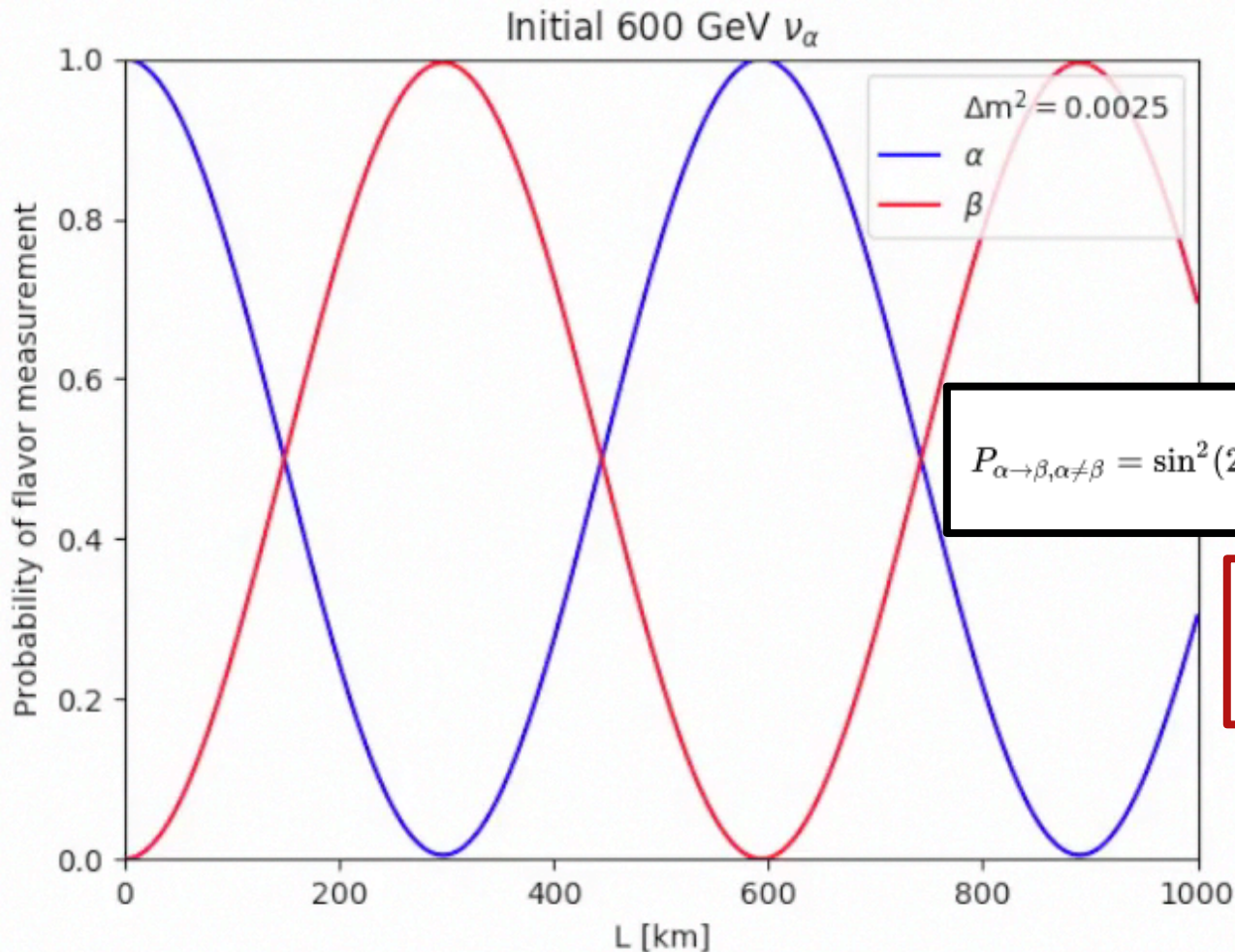


## Neutrino Properties:

- initial  $\nu_\alpha$
  - Energy: 600 GeV
- ## Oscillation Parameters:
- $\theta = 47^\circ$
  - $\Delta m^2 = 2.5 \cdot 10^{-3} \text{ eV}^2$



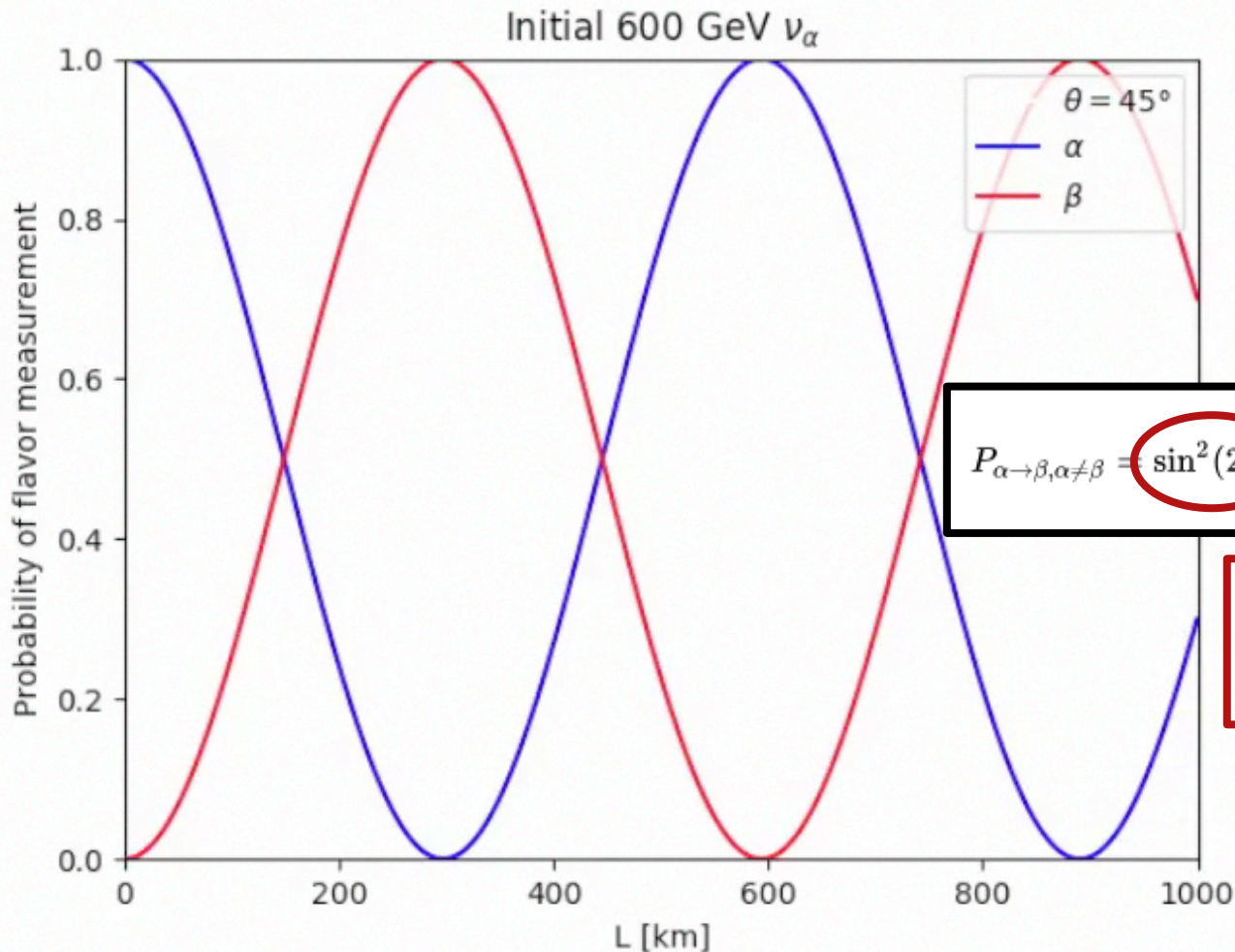
# Toy Model



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$\Delta m^2$  controls period  
 $\theta$  controls amplitude

# Toy Model



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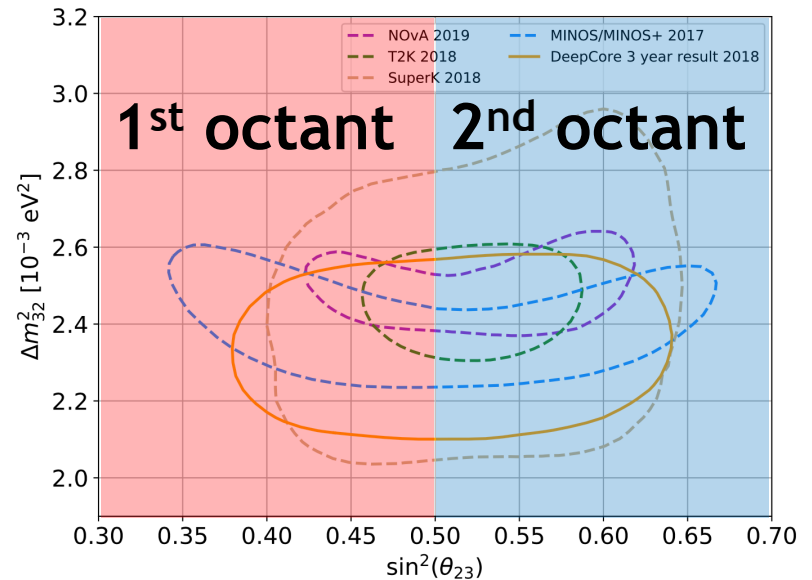
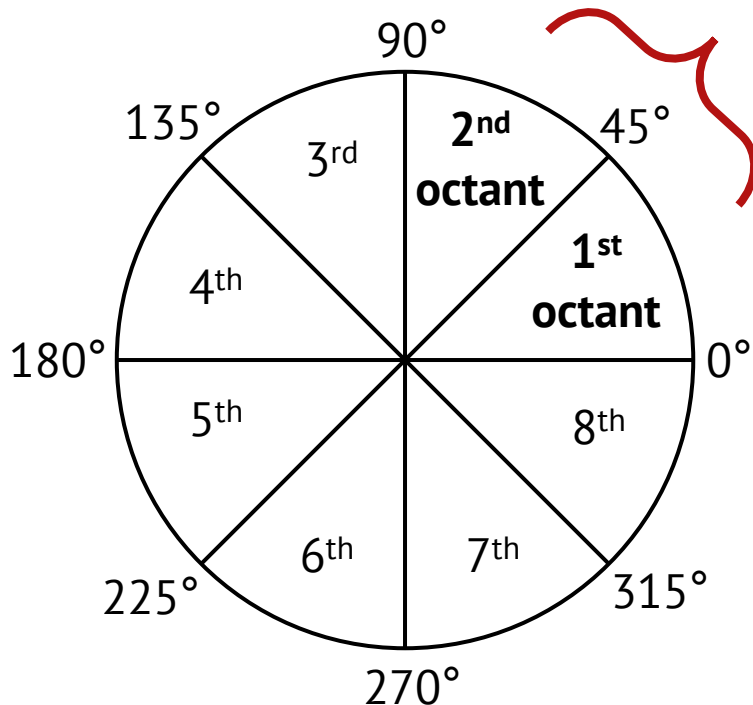
$\Delta m^2$  controls period  
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# Open Questions

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# Octant?

- Oscillation probabilities have terms with  $\sin^2(2\theta_{23})$  so there is a degeneracy near  $45^\circ$  if non-maximal.
  - ex.  $\sin^2(2 \cdot 44^\circ) = \sin^2(2 \cdot 46^\circ)$
- So if we measure a certain amount of oscillation, how do we know if  $\theta_{23}$  is a little above  $45^\circ$  or a little below  $45^\circ$ ?

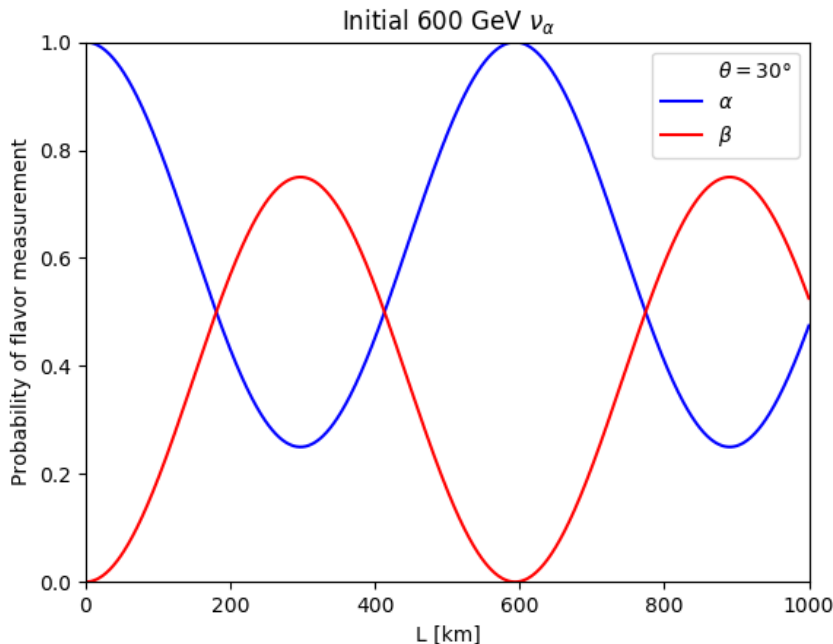


# Maximal Mixing

- Experiments tell us the true value of  $\theta_{23}$  is near  $45^\circ$ .
- There is still tension about whether or not it is *exactly* maximal.

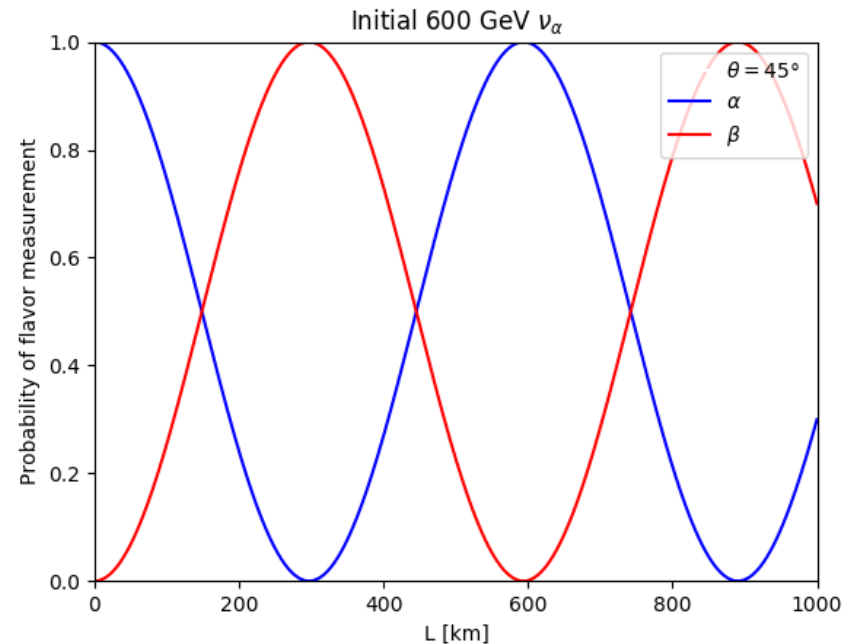
$$\theta_{23} \neq 45^\circ$$

“non-maximal mixing”



$$\theta_{23} = 45^\circ$$

“maximal mixing”

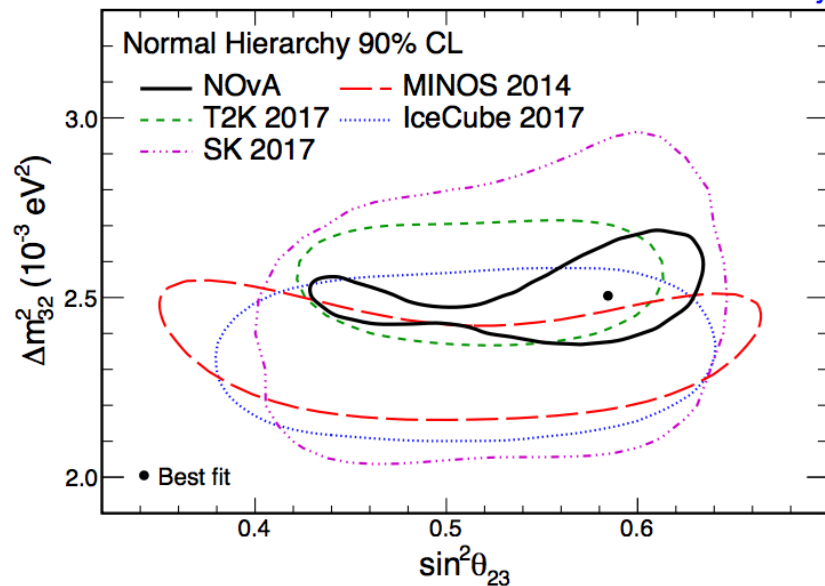


# Maximal Mixing

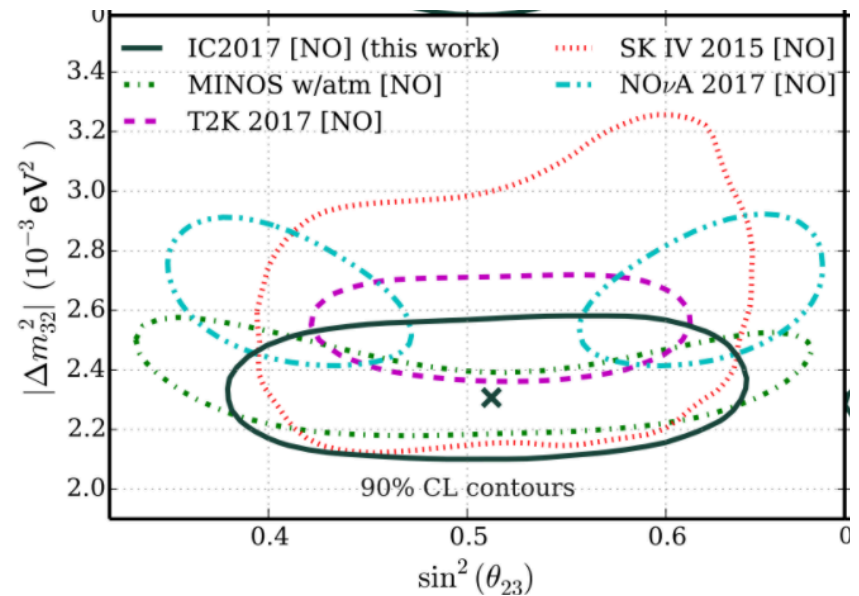
- Experiments tell us the true value of  $\theta_{23}$  is near  $45^\circ$ .
- There is still tension about whether or not it is *exactly* maximal.

NO $\nu$ A prefers non-maximal [3]

NO $\nu$ A Preliminary



Others centered closer to maximal



# Beyond the Standard Model

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- Are there additional neutrino flavors called “Sterile Neutrinos”? These proposed particles don’t interact with regular matter, but the other flavors could oscillate to and from the sterile state.
- Could Sterile Neutrinos be Dark Matter?
- Much more!

# Thanks for listening!

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## Questions?