

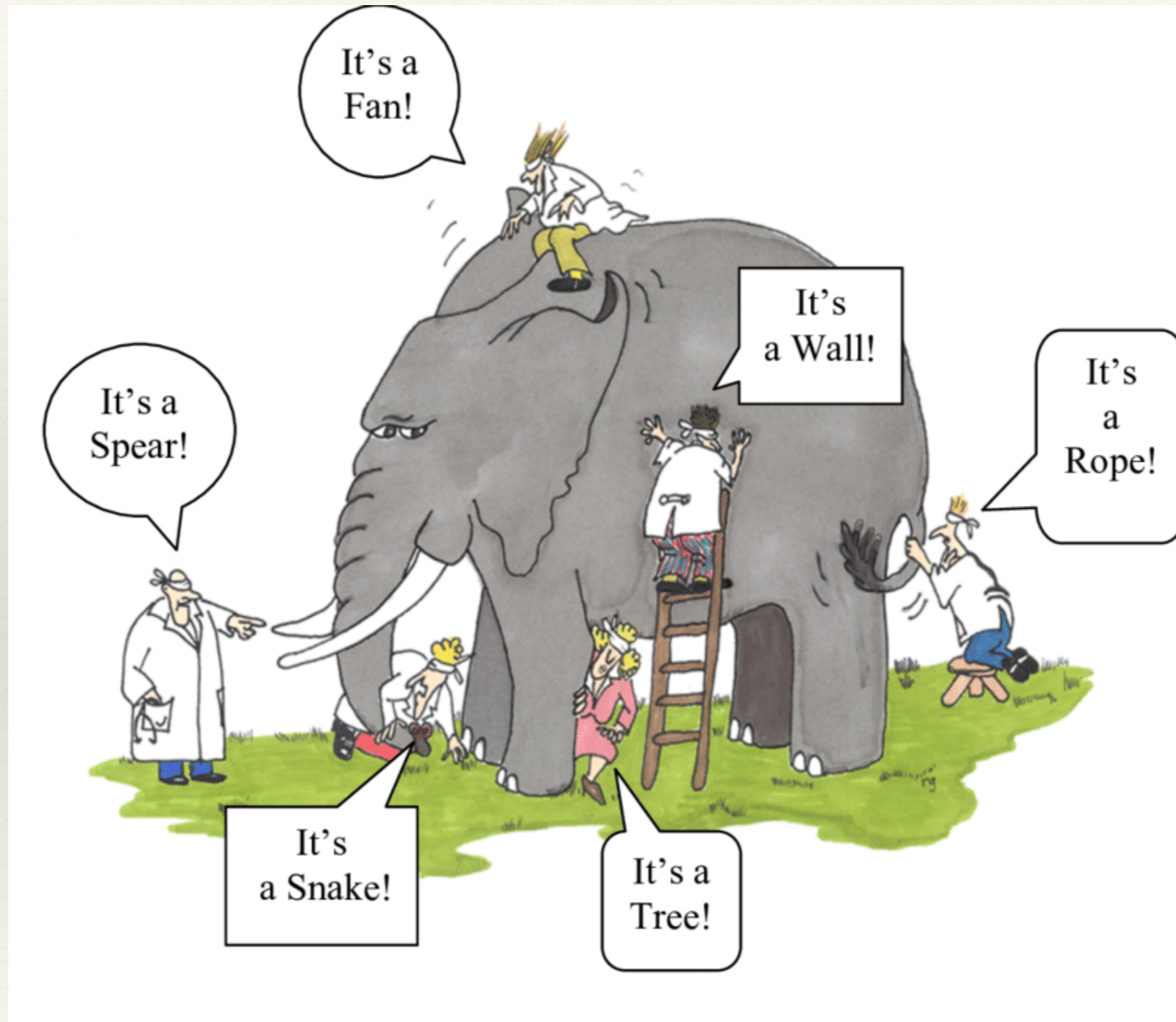


Multi-messenger Astrophysics with IceCube

Abhishek Desai
IceCube Bootcamp 2020

The Multi-messenger Picture

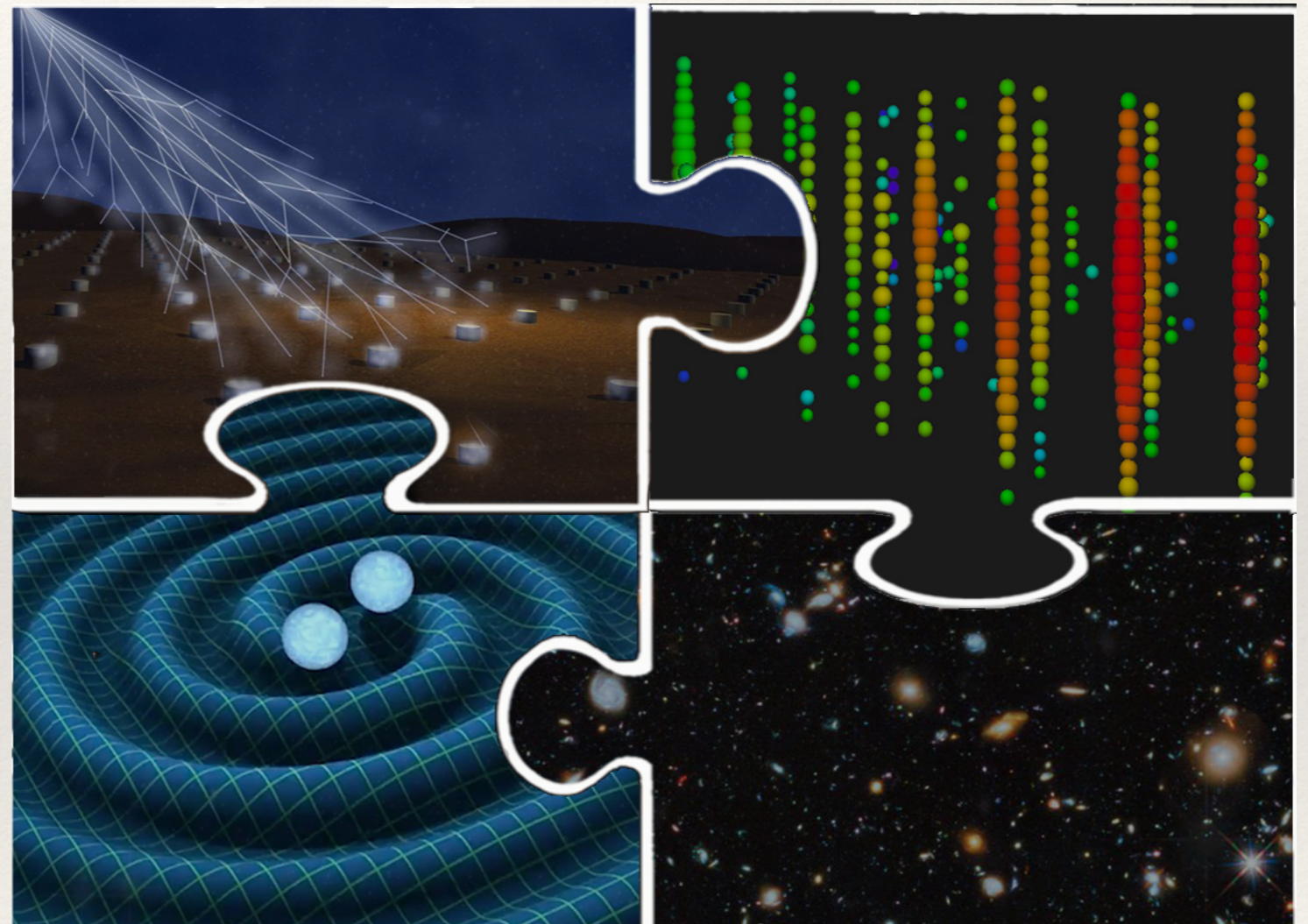
The Blind Men And The Elephant



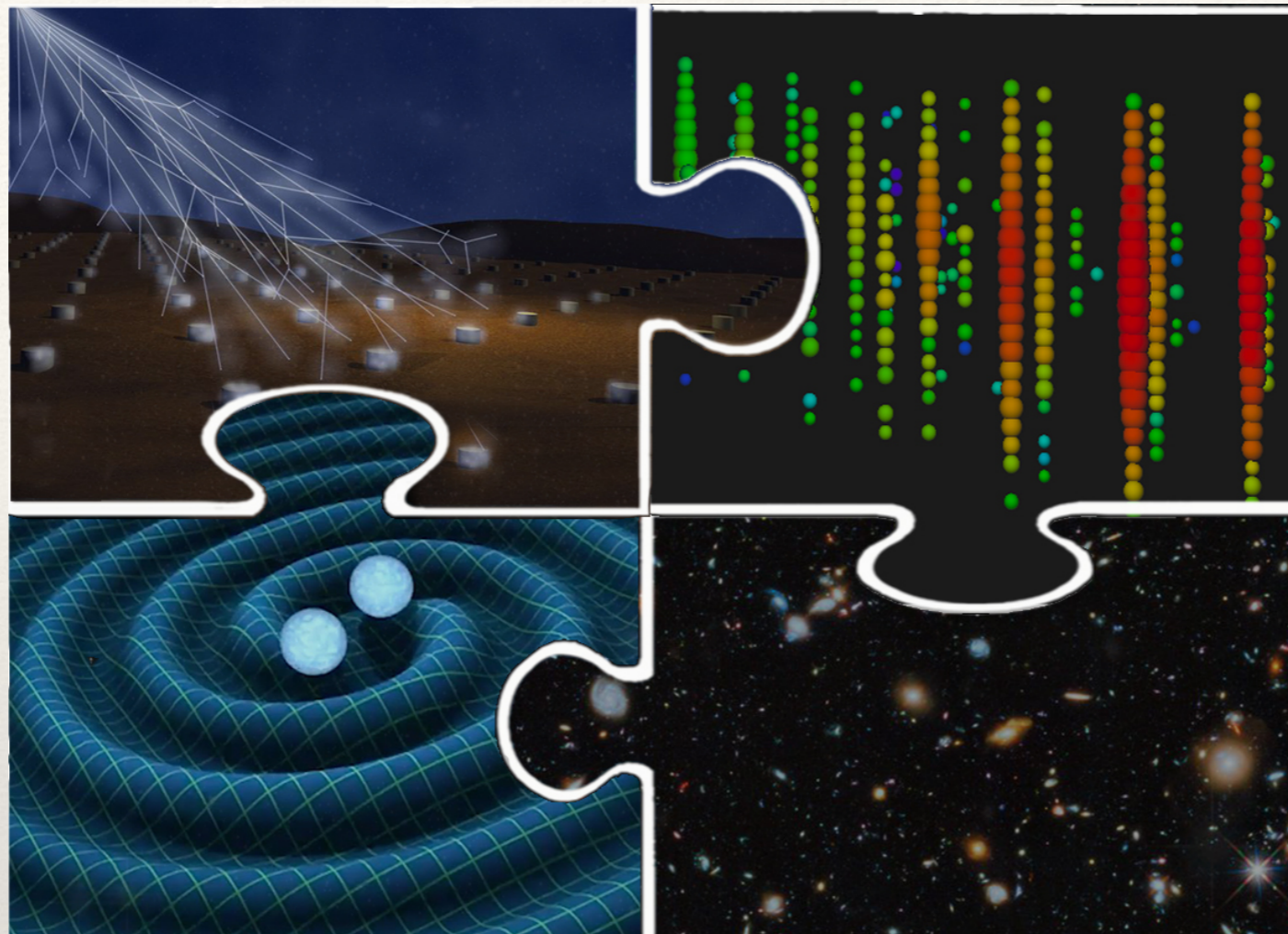
The Multi-messenger Picture

- ❖ Definition: Observations of a single source producing distinct signals associated with two or more of the 4 fundamental forces.

Force	Messenger
EM	Photons
Gravity	Gravitational waves
Strong	Protons, Nuclei
Weak	Neutrino



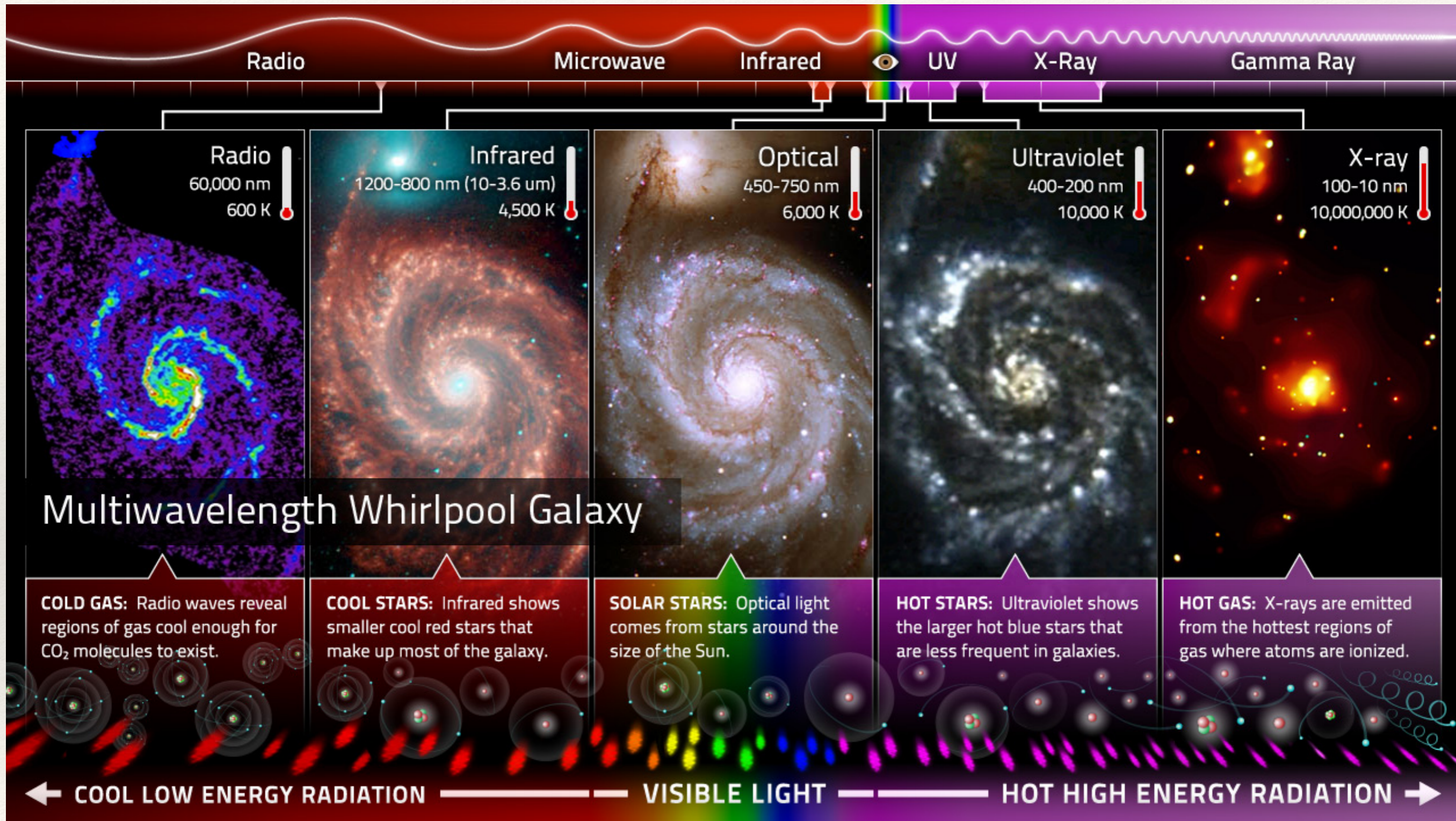
The Multi-messenger Picture



Photons

- ❖ The electromagnetic force is responsible for generating visible light as well as radiation in other wavebands not detectable by the human eye.
- ❖ The electromagnetic radiation can be thermal or non thermal

The Multi-wavelength Picture



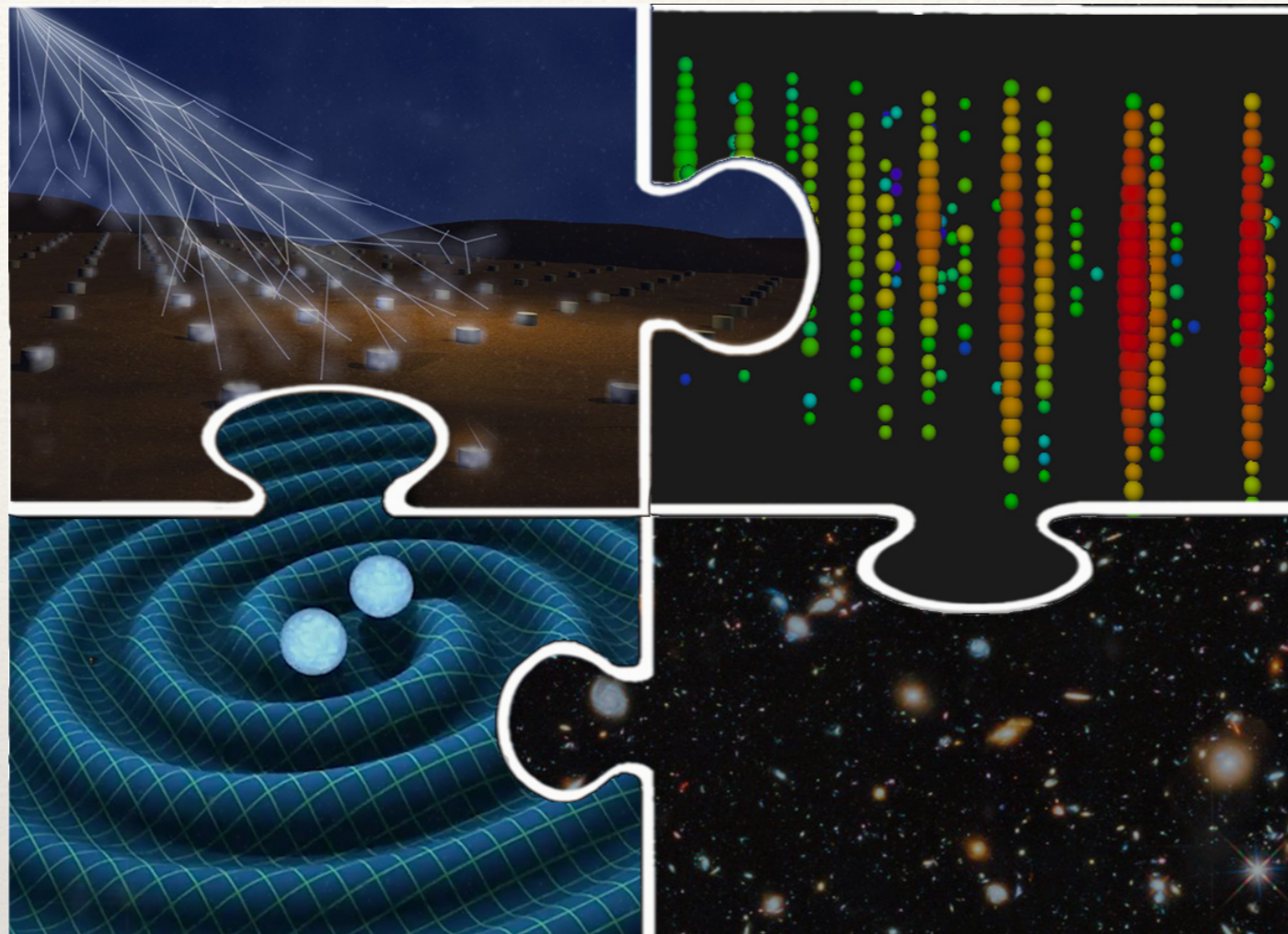
<http://ecuip.lib.uchicago.edu/multiwavelength-astronomy/astrophysics/05.html>

Topics Of Interest:

- ❖ Multiwavelength astronomy: <http://ecuip.lib.uchicago.edu/multiwavelength-astronomy/index.html>

The Multi-messenger Picture

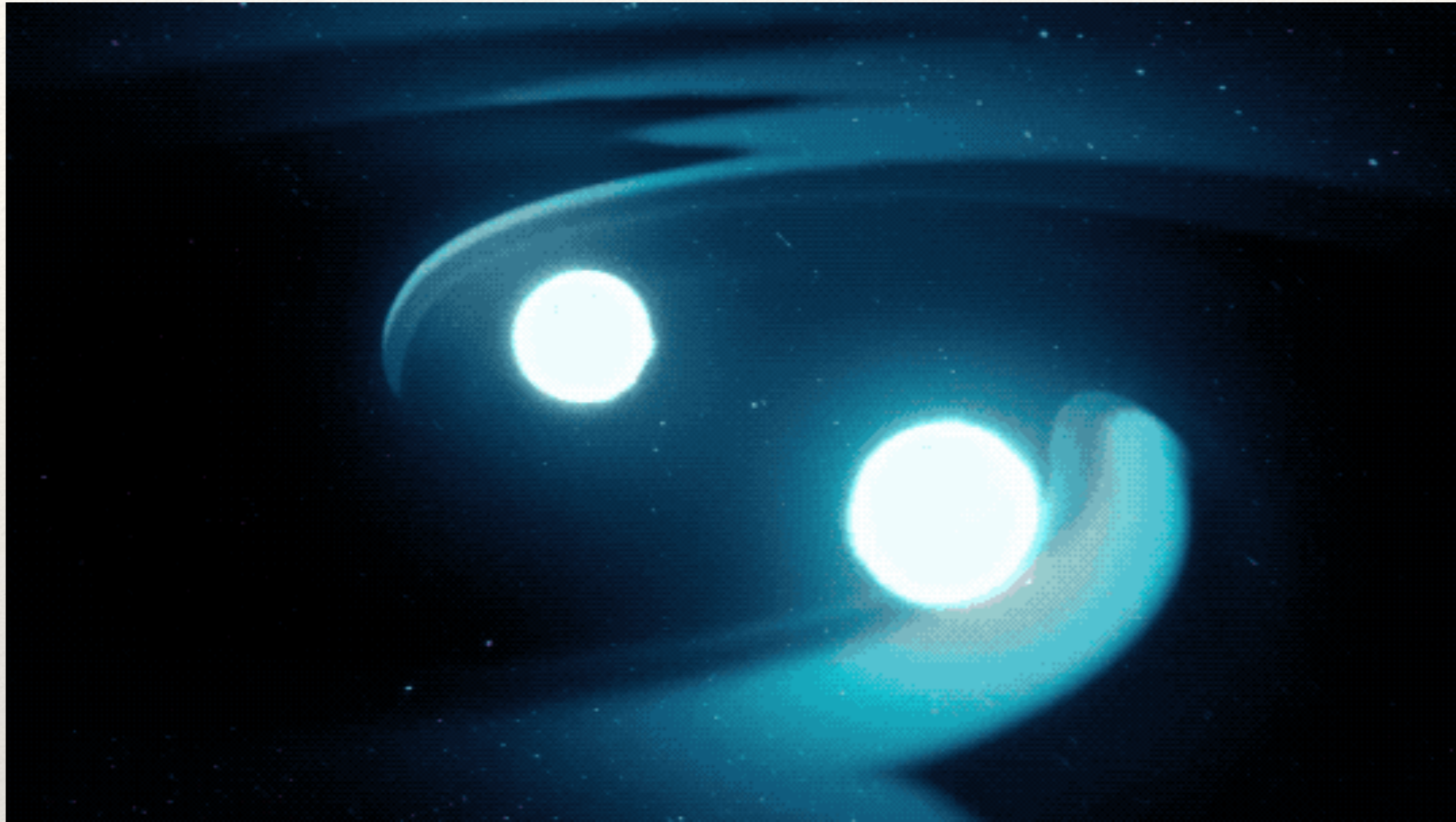
Gravitational
Waves



Photons

- ❖ According to general relativity, mass causes space-time to curve in a describable manner, manifesting as gravity.
- ❖ When masses move, corresponding changes in the gravitational field move through the cosmos as gravitational waves at the speed of light, like ripples across a pond.

Gravitational Waves



- ❖ However gravity is an extremely weak force so even the sources of the biggest gravitational waves, like the cataclysmic collisions of black holes, would only produce the tiniest of wiggles by the time they reached Earth.
- ❖ Using detectors like LIGO these tiny gravitational waves wiggles as detected as messengers

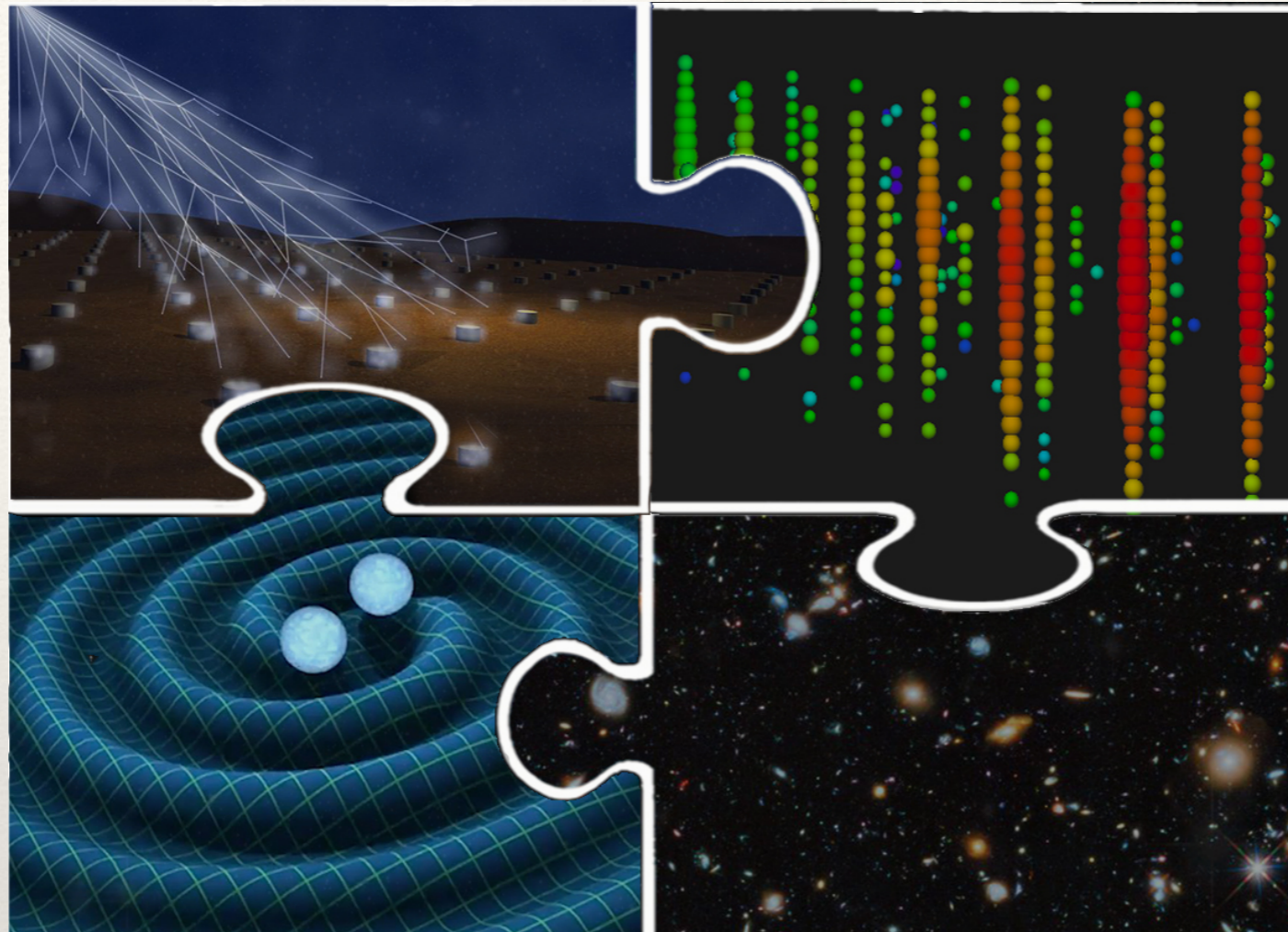
Topics Of Interest:

- ❖ Multiwavelength astronomy: <http://ecuip.lib.uchicago.edu/multiwavelength-astronomy/index.html>
- ❖ Gravitational Waves: <https://youtu.be/hhbMpe17fzA>

The Multi-messenger Picture

Cosmic Rays

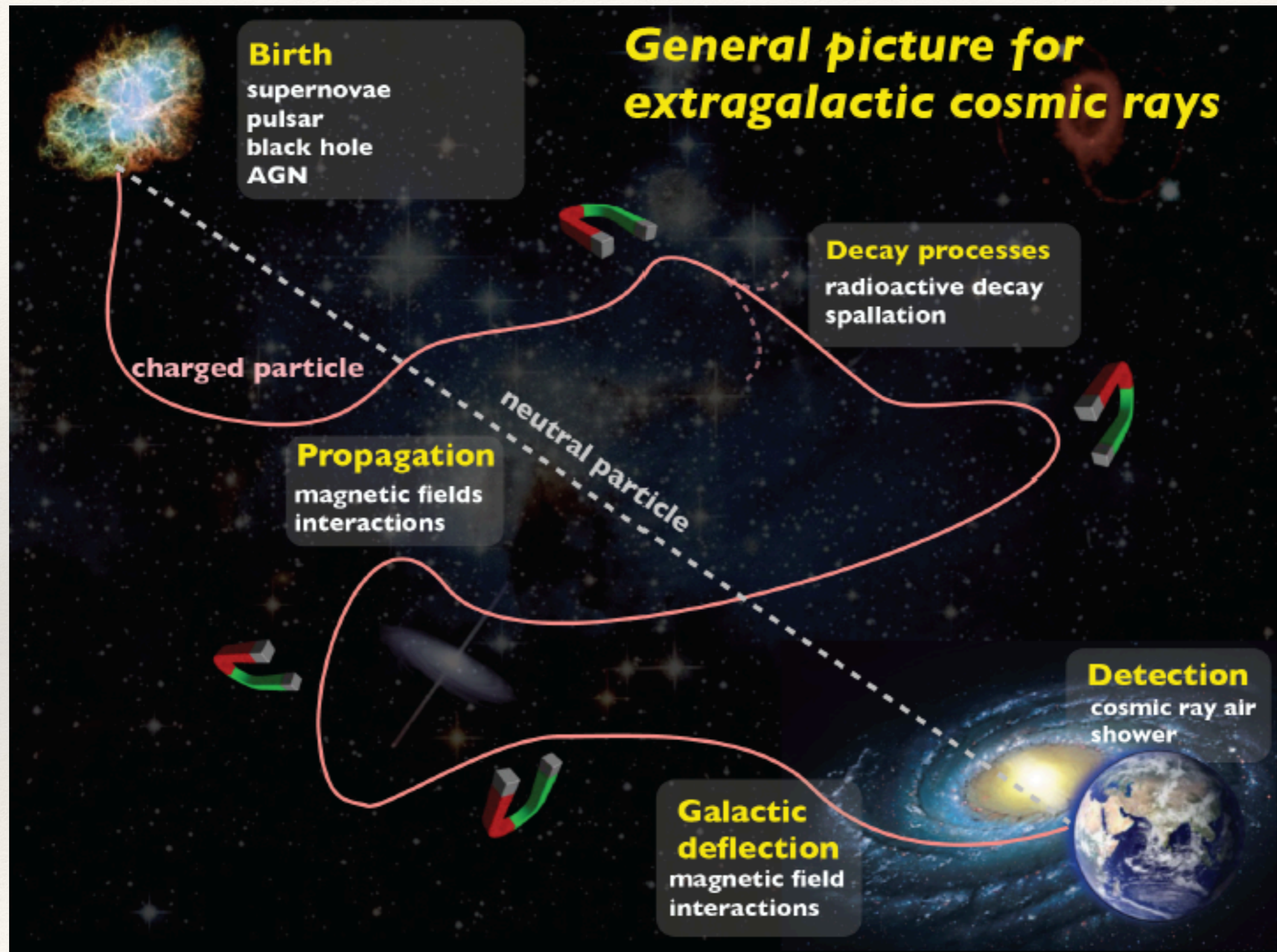
Gravitational
Waves



Photons

- ❖ Cosmic Rays are high energy particles that move through space at nearly the speed of light.
- ❖ These particles come from outer space and from our own solar system. Scientists first called these particles “rays” because they thought they were a form of electromagnetic radiation.

Cosmic Rays

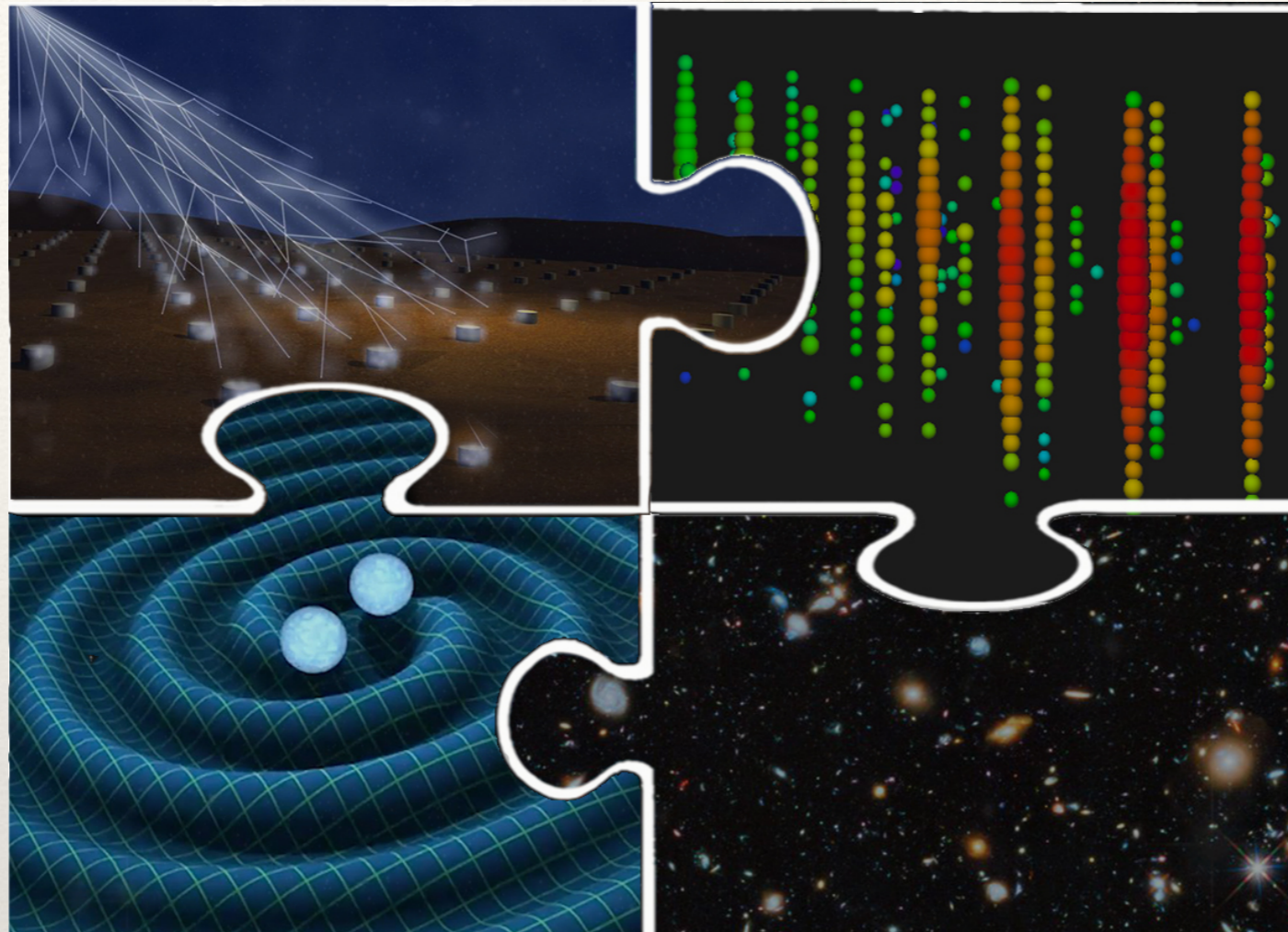


Topics Of Interest:

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- ❖ Gravitational Waves: <https://youtu.be/hhbMpe17fzA>
- ❖ Cosmic Rays: <https://svs.gsfc.nasa.gov/cgi-bin/details.cgi?aid=11209>

The Multi-messenger Picture

Cosmic Rays



Neutrinos

Gravitational
Waves

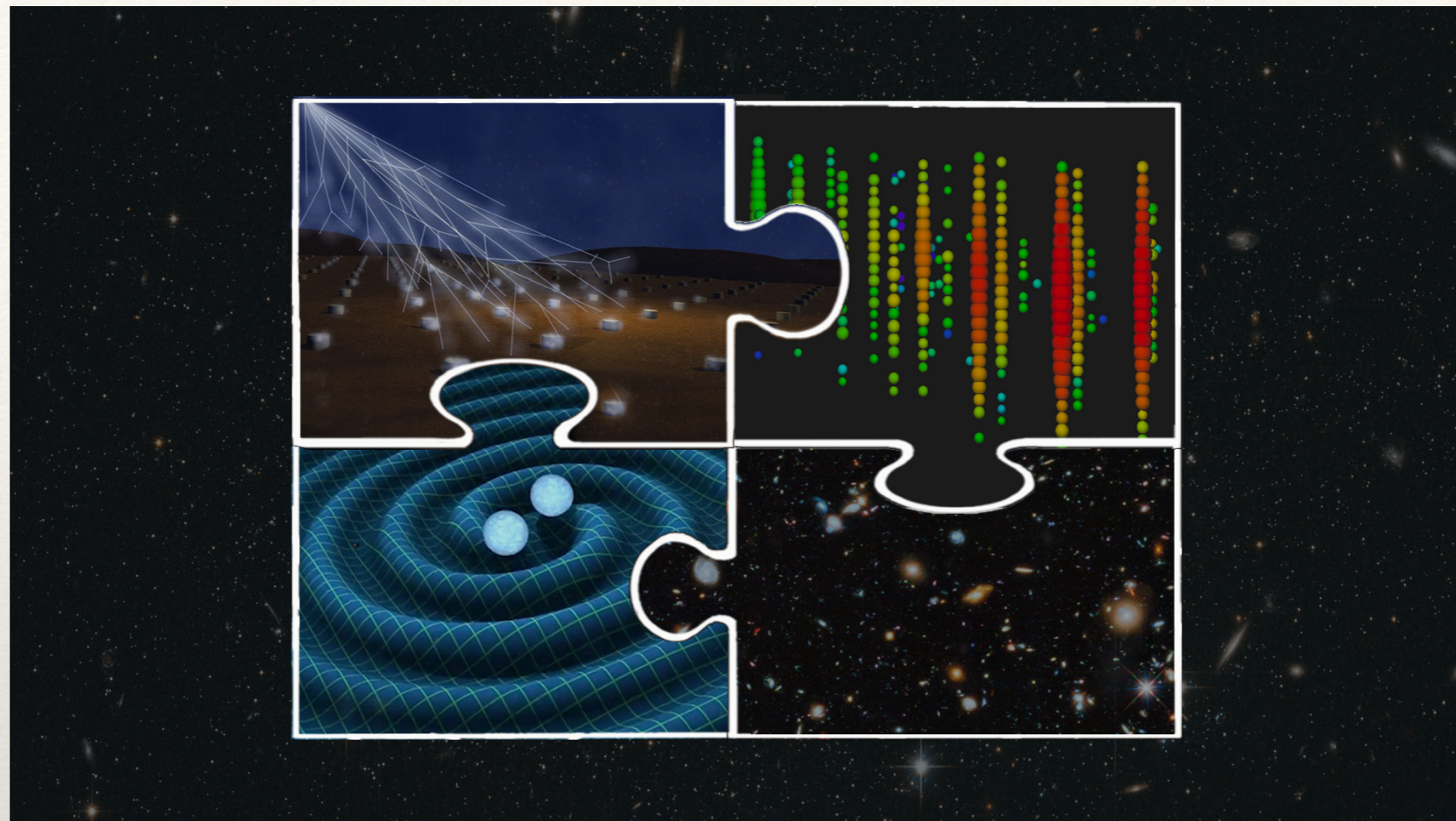
Photons

- ❖ Neutrinos are nearly massless and have no electric charge. Therefore, unlike the other particles, they only interact via the weak nuclear force.
- ❖ Since the weak nuclear force only acts at short ranges, neutrinos can pass through massive objects without interacting with them.

The Multi-messenger Picture

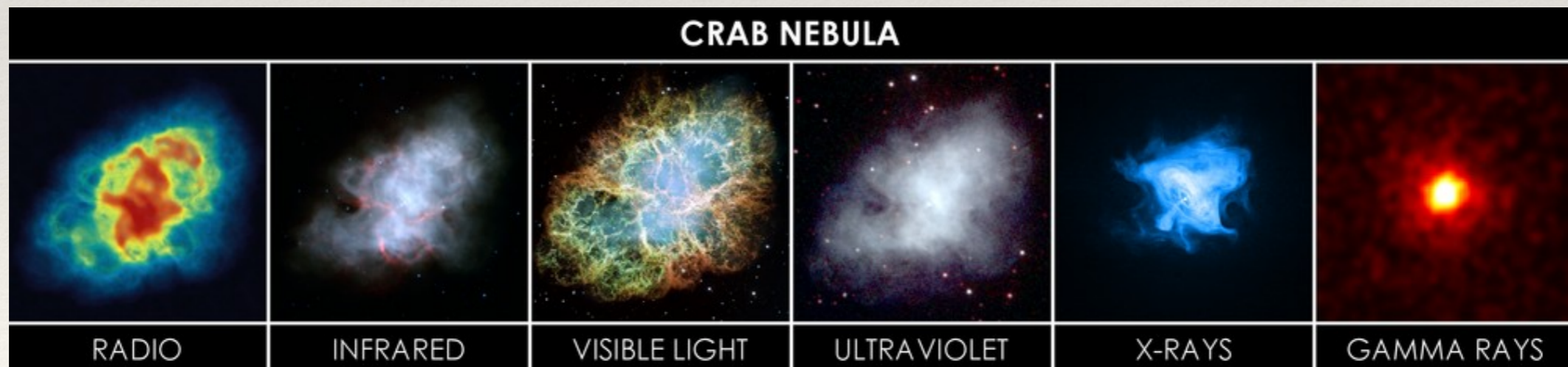
Cosmic Rays

Gravitational
Waves



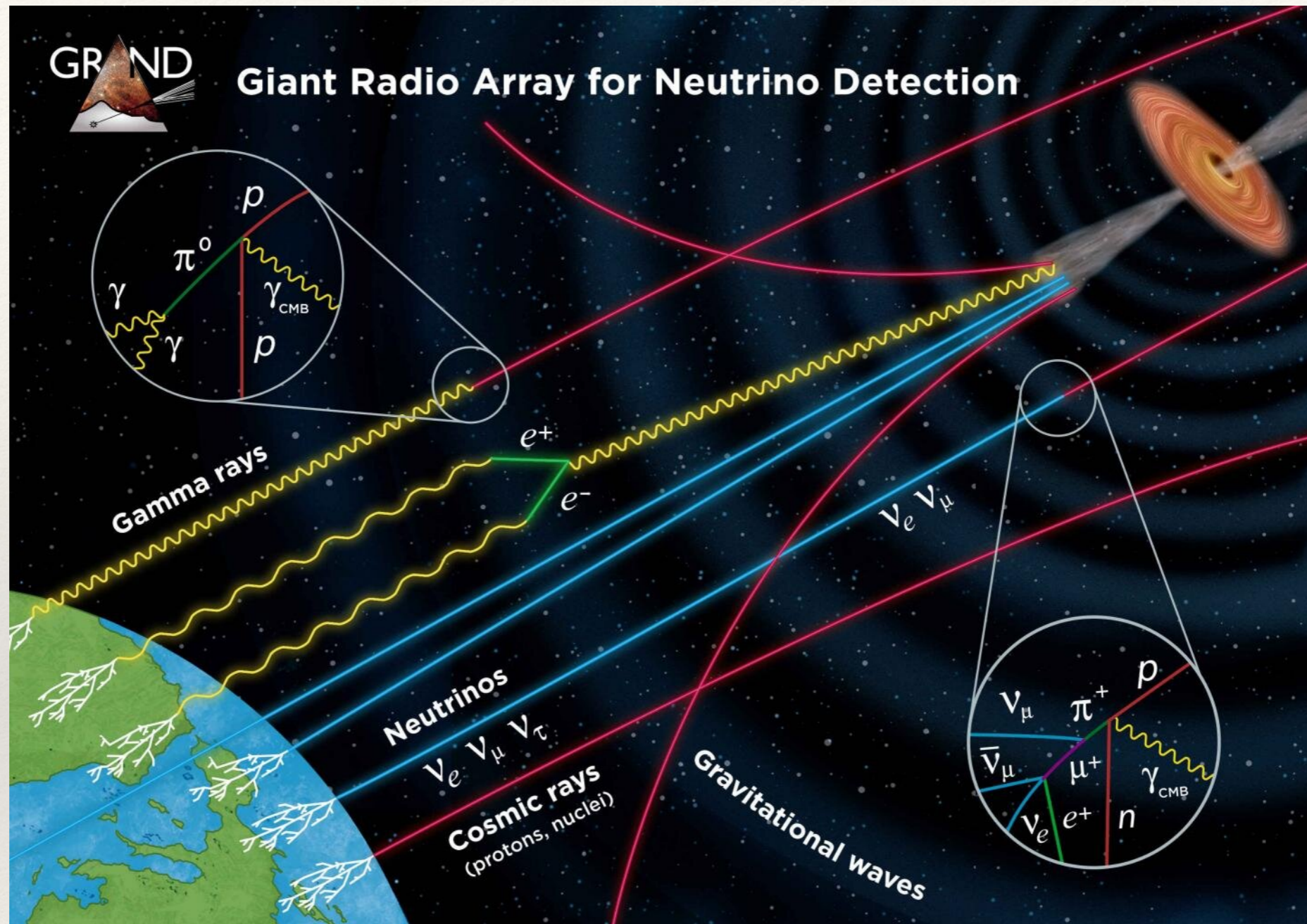
Neutrinos

Photons



Multi-wavelength Observations

The Multi-messenger Picture



Creation and propagation of ultra-high energy particles in the Universe.

Credit: Science China Press

Diffuse Background

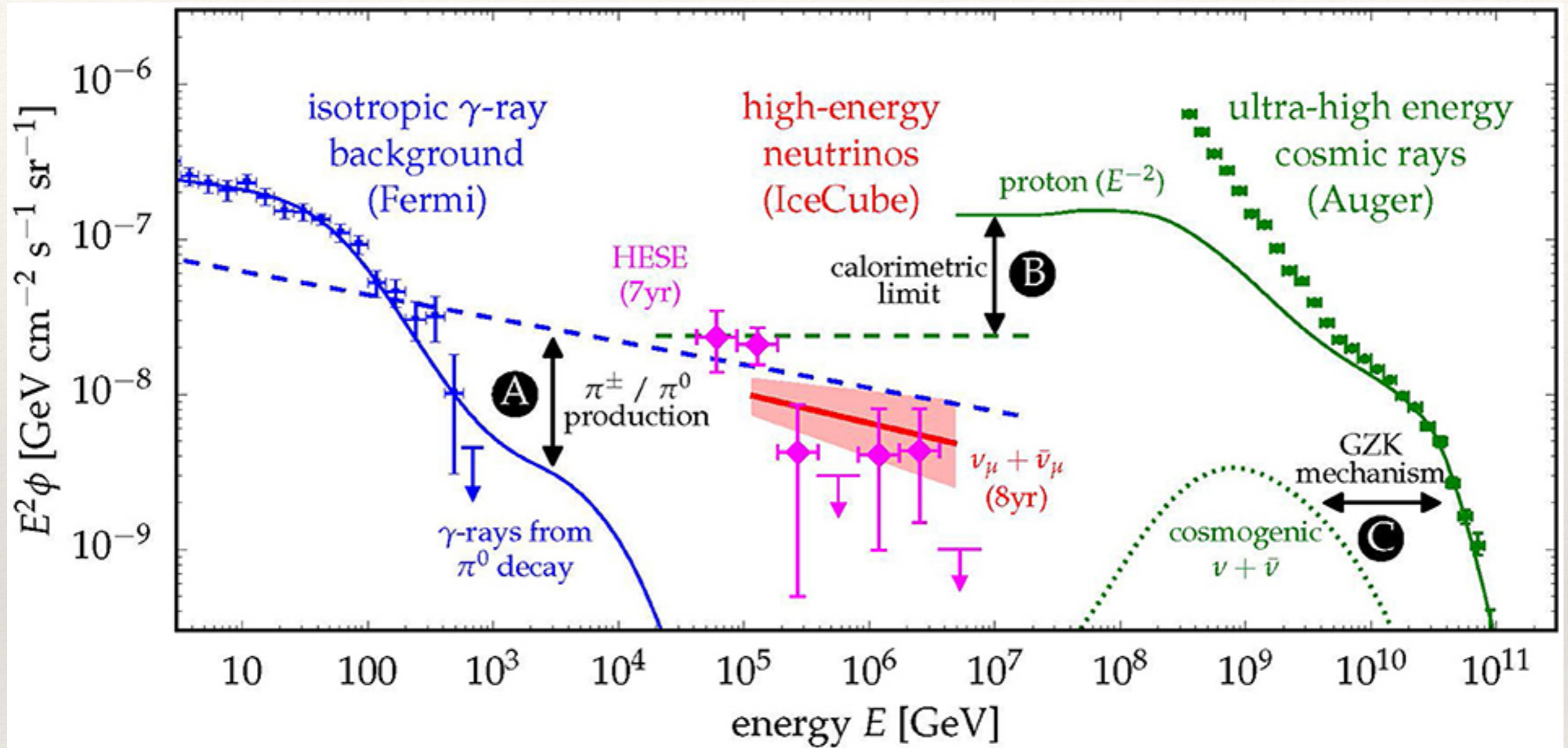


Figure from Ahlers and Halzen (2018)

Diffuse Background

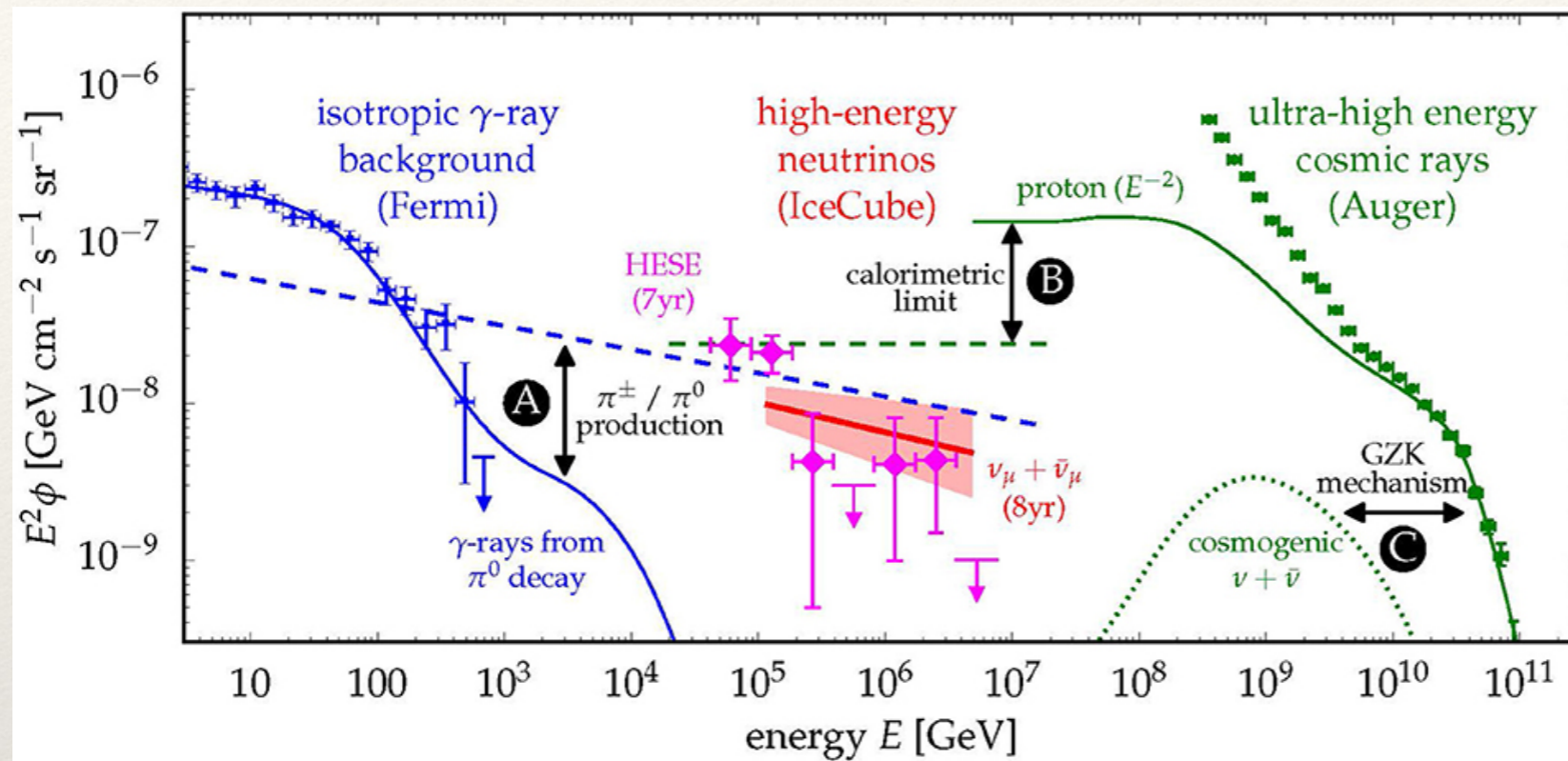


Figure from Ahlers and Halzen (2018)

- ❖ Photons are inevitably produced in association with neutrinos when accelerated cosmic rays produce neutral and charged pions.
- ❖ The pionic gamma rays should then accompany neutrinos at the site of production. However, since they undergo EBL absorption, they will appear at lower energies.
- ❖ Multimessenger interface (A) The joined production of charged pions and neutral pions in cosmic-ray interactions leads to the emission of neutrinos (dashed blue) and gamma rays (solid blue), respectively.

Diffuse Background

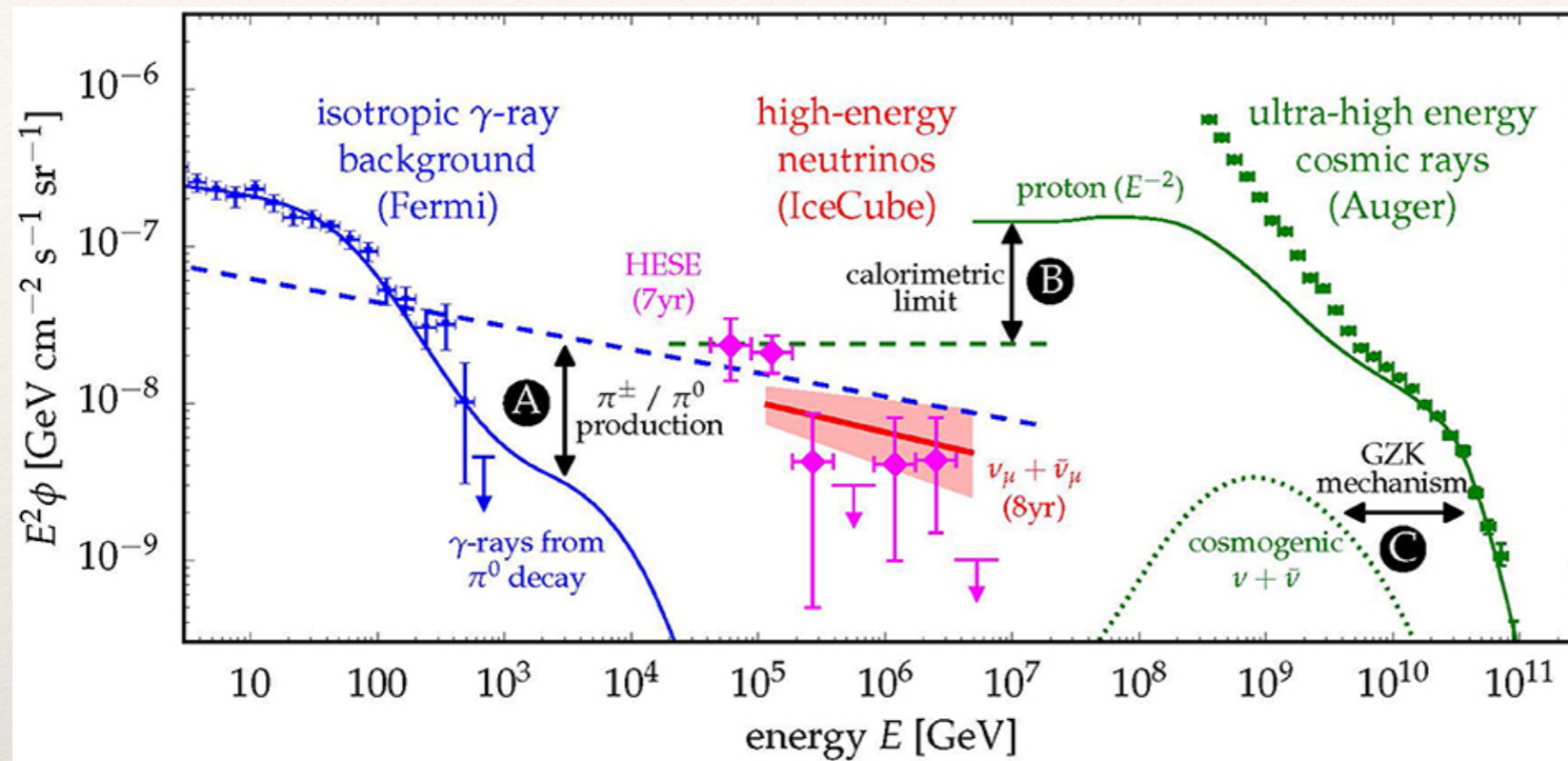


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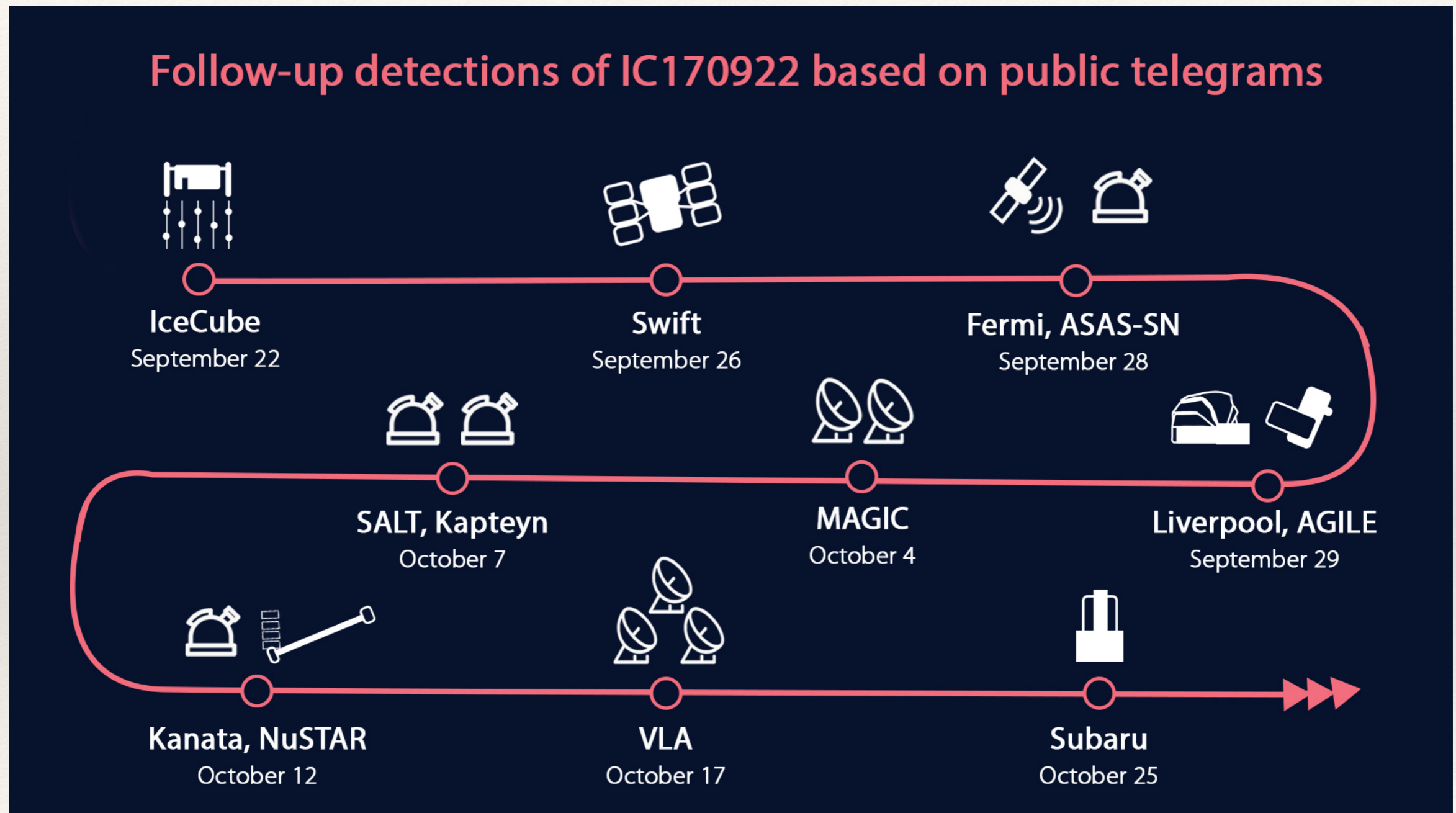
- ❖ Multimessenger interface **(B)** Cosmic ray emission models (solid green) of the most energetic cosmic rays imply a maximal flux (calorimetric limit) of neutrinos from the same sources (green dashed).
- ❖ Multimessenger interface **(C)** The same cosmic ray model predicts the emission of cosmogenic neutrinos from the collision with cosmic background photons (GZK mechanism).
- ❖ Neutrinos with higher energies are expected from the Greisen Ztsepin Kuzmin (GZK) effect, namely the interaction of ultrahigh-energy cosmic rays (UHECRs) with the cosmic microwave background (CMB) and the extragalactic background light (EBL), but have not yet been detected

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- ❖ Gravitational Waves: <https://youtu.be/hhbMpe17fzA>
- ❖ Cosmic Rays: <https://svs.gsfc.nasa.gov/cgi-bin/details.cgi?aid=11209>
- ❖ Must read for all that is discussed today and more: <https://www.frontiersin.org/articles/10.3389/fspas.2019.00032/full>

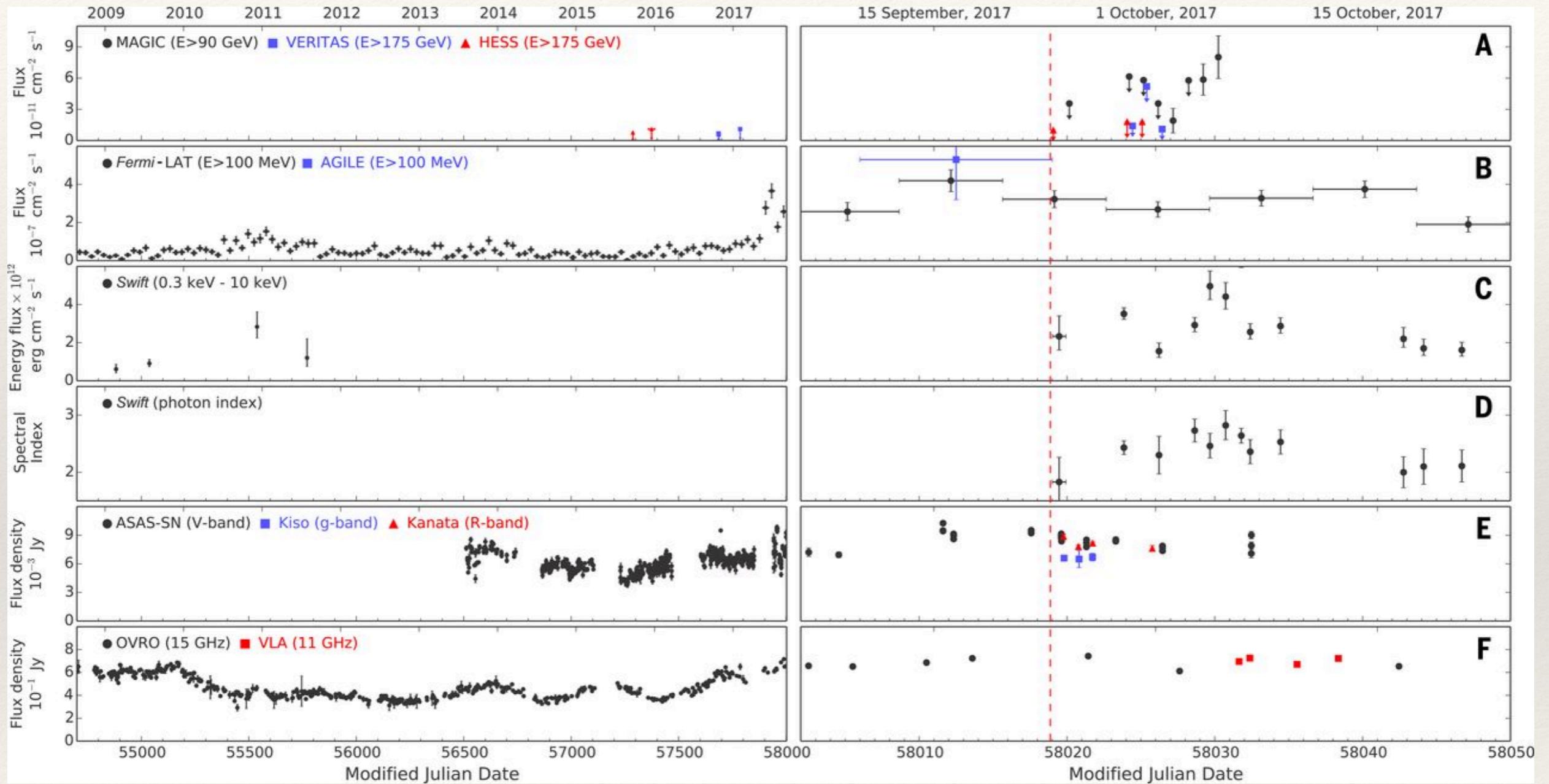
Neutrinos and Photons

- ❖ Or Why TXS0506+056 is so famous?

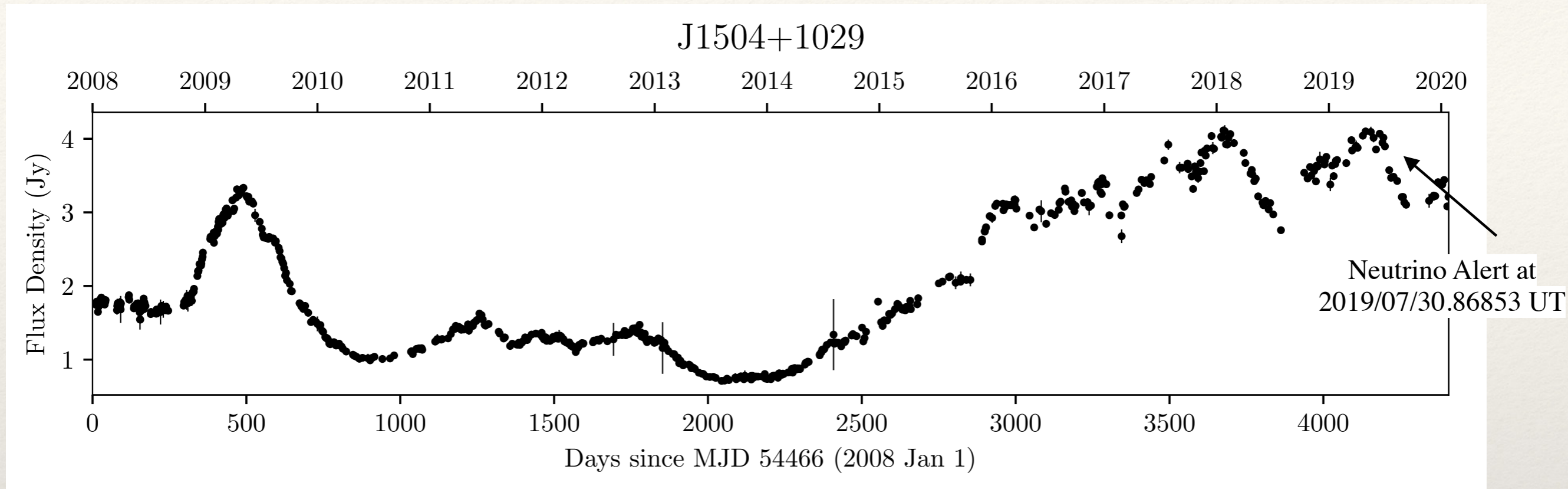


Neutrinos And Photons: TXS 0506+056

❖ Neutrino Alert: IC-170922A (Red dashed line)



Exciting Research Work In Progress



Data taken from the OVRO 40-m monitoring program (Richards, J. L. et al. 2011, ApJS, 194, 29) [Link](#)

- ❖ The Blazar PKS1502+106 was found to be possibly correlated with an Icecube alert ([IC190730A](#))
- ❖ At the time of the alert, the radio observations of the FSRQ were seen reaching an all time peak flux of 4 Jy (S. Kiehlmann et.al. [The Astronomer's Telegram 12996, 1 \(2019\).](#)) (See Below)
- ❖ Moreover, positive correlations would help us better understand the neutrino production processes in AGN (as also discussed in [Plavin et al 2020](#))

Gravitational Waves

**FIRST COSMIC EVENT OBSERVED
IN GRAVITATIONAL WAVES AND LIGHT**
Colliding Neutron Stars Mark New Beginning of Discoveries

Collision creates light across the entire electromagnetic spectrum. Joint observations independently confirm Einstein's General Theory of Relativity, help measure the age of the Universe, and provide clues to the origins of heavy elements like gold and platinum

Gravitational wave lasted over 100 seconds

On August 17, 2017, 12:41 UTC, LIGO (US) and Virgo (Europe) detect gravitational waves from the merger of two neutron stars, each around 1.5 times the mass of our Sun. This is the first detection of spacetime ripples from neutron stars.

Within two seconds, NASA's Fermi Gamma-ray Space Telescope detects a short gamma-ray burst from a region of the sky overlapping the LIGO/Virgo position. Optical telescope observations pinpoint the origin of this signal to NGC 4993, a galaxy located 130 million light years distant.

LIGO Georgia Tech Center for Relativistic Astrophysics

Gravitational Waves and Photons

- ❖ Advertisement for Raamis's talk regarding search for high-energy neutrino emission from merger candidates reported by the LIGO-Virgo Collaboration (LVC) throughout its first three observing runs.

Multimessenger With IceCube

- ❖ Powerful real time follow-up program at IceCube targets the detection of transient sources.
- ❖ This multimessenger program sends alerts of single and clusters of high-energy neutrino events (multiplets), typically within one minute of the event detection.
- ❖ In collaboration with other observatories, IceCube aims to identify the electromagnetic counterpart of a rapidly fading source or coincident gravitational waves. Single event alerts are distributed publicly as [GCN alerts](#), while [multiple alerts](#) are distributed through individual agreements with optical, X-ray, and gamma-ray observatories.
- ❖ Searches for bursts of low-energy neutrinos from [nearby supernovas](#) are performed, and above threshold detection is announced rapidly within the SNEWS network.

Taken from the IceCube Science Highlights Page

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- ❖ IceCube Multimessenger Highlights: https://icecube.wisc.edu/science/highlights/neutrino_astronomy

Questions?
or
Comments!