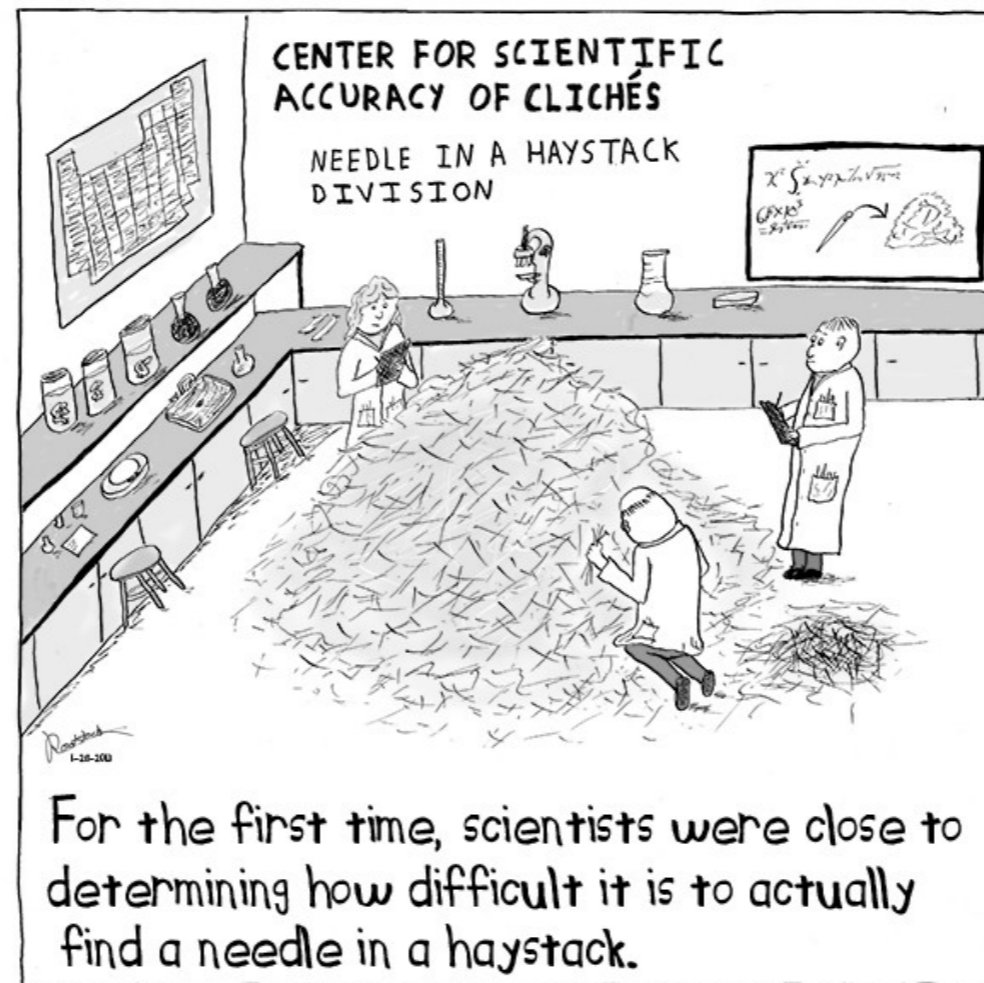


Point Source Likelihood Technique

“Finding needles in haystacks”



Slides by Josh Wood

The Problem

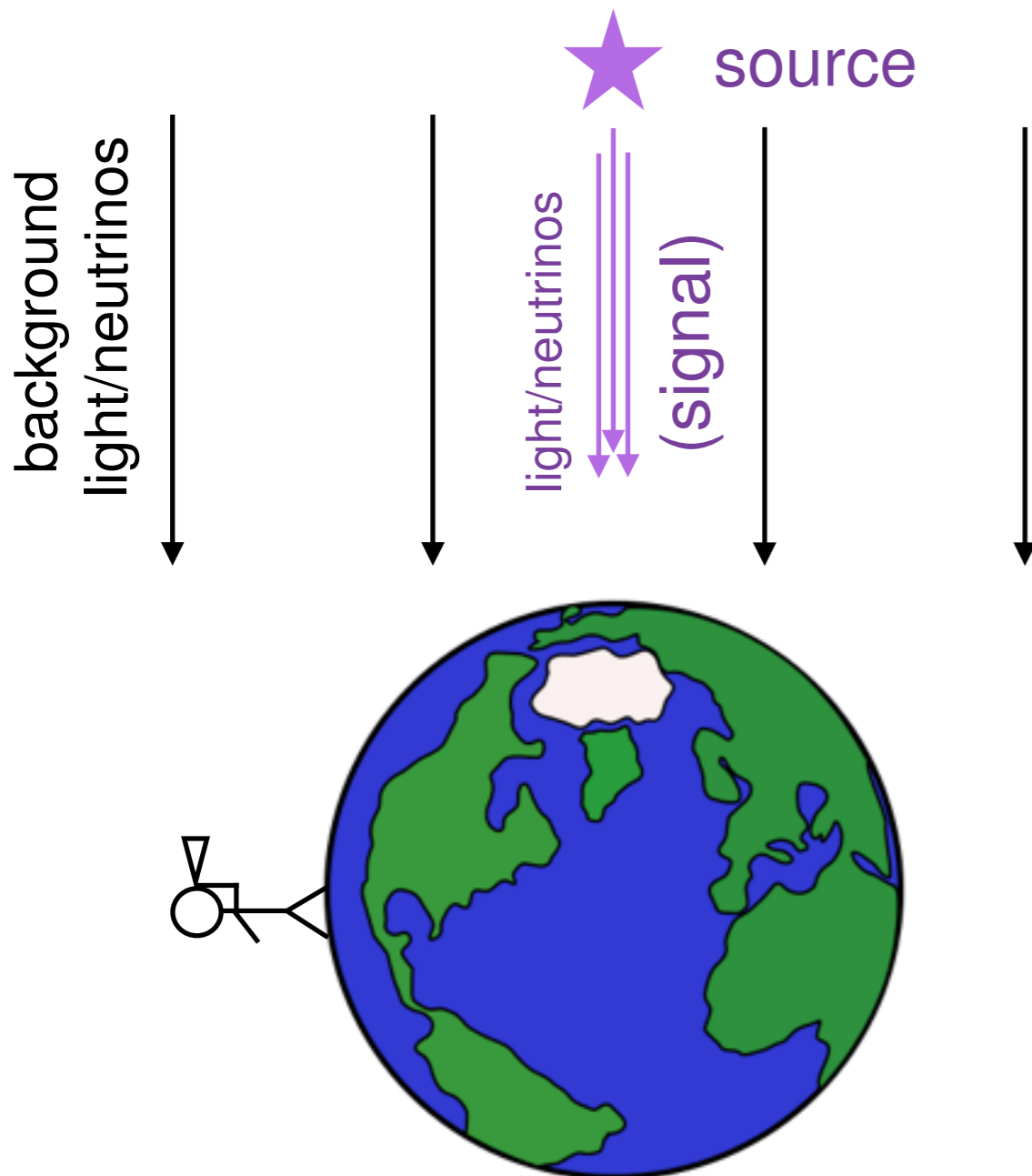
- Astronomy is easy when you don't have background



- For some messengers (high energy photons, neutrinos) we can't turn backgrounds off, but we still want to find sources. How to find sources on top of background?

What to do? Think about it!

Imagine you're an astronomer looking for a point source in the presence of uniform background.



def: **signal** is a particle that came from the source you're looking for

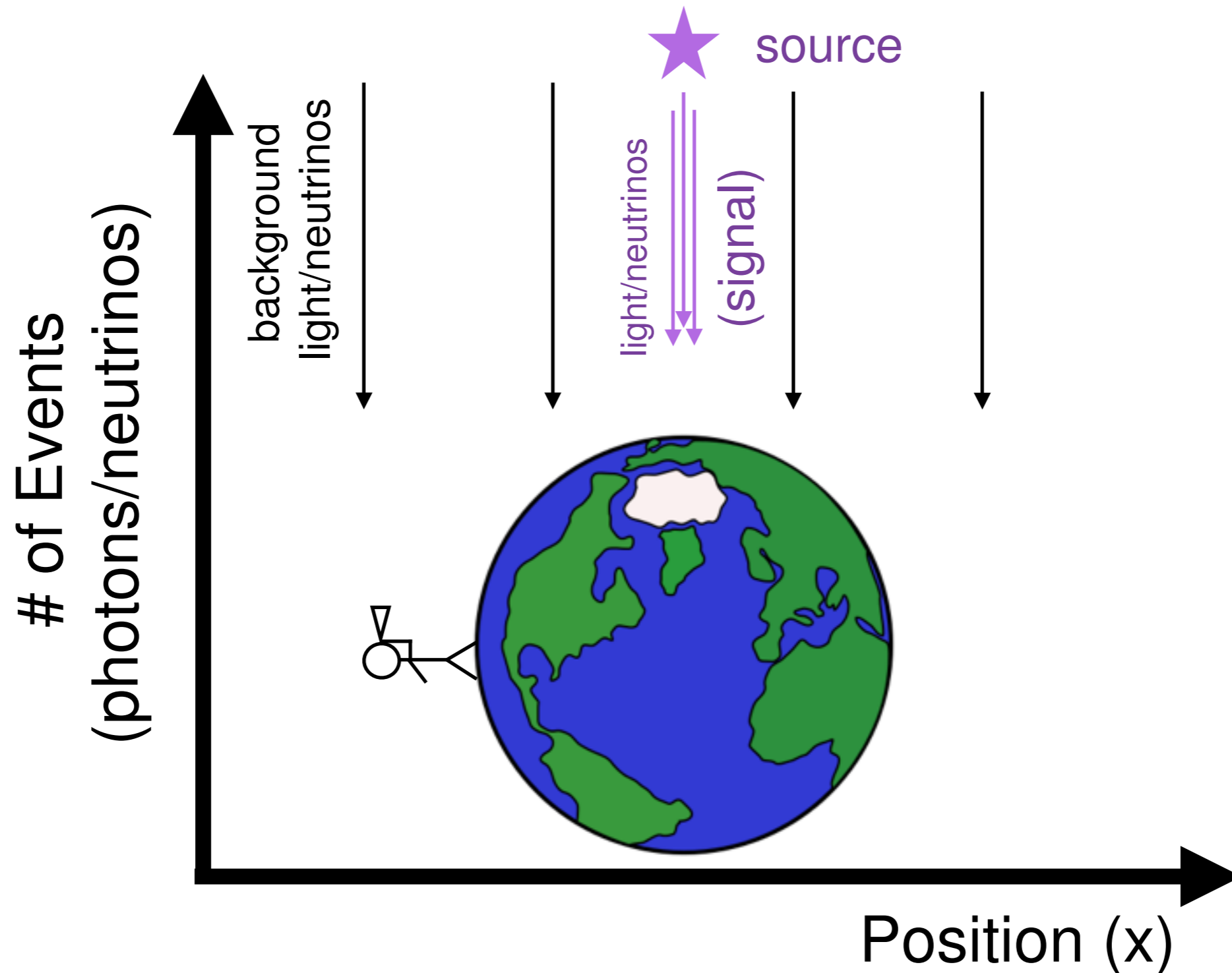
def: **background** is a particle that did not come from the source but looks identical to a particle emitted by the source

Ex. photon/neutrino with same energy as one from the source

def. **event** a detected particle.
Can be photon, neutrino etc.

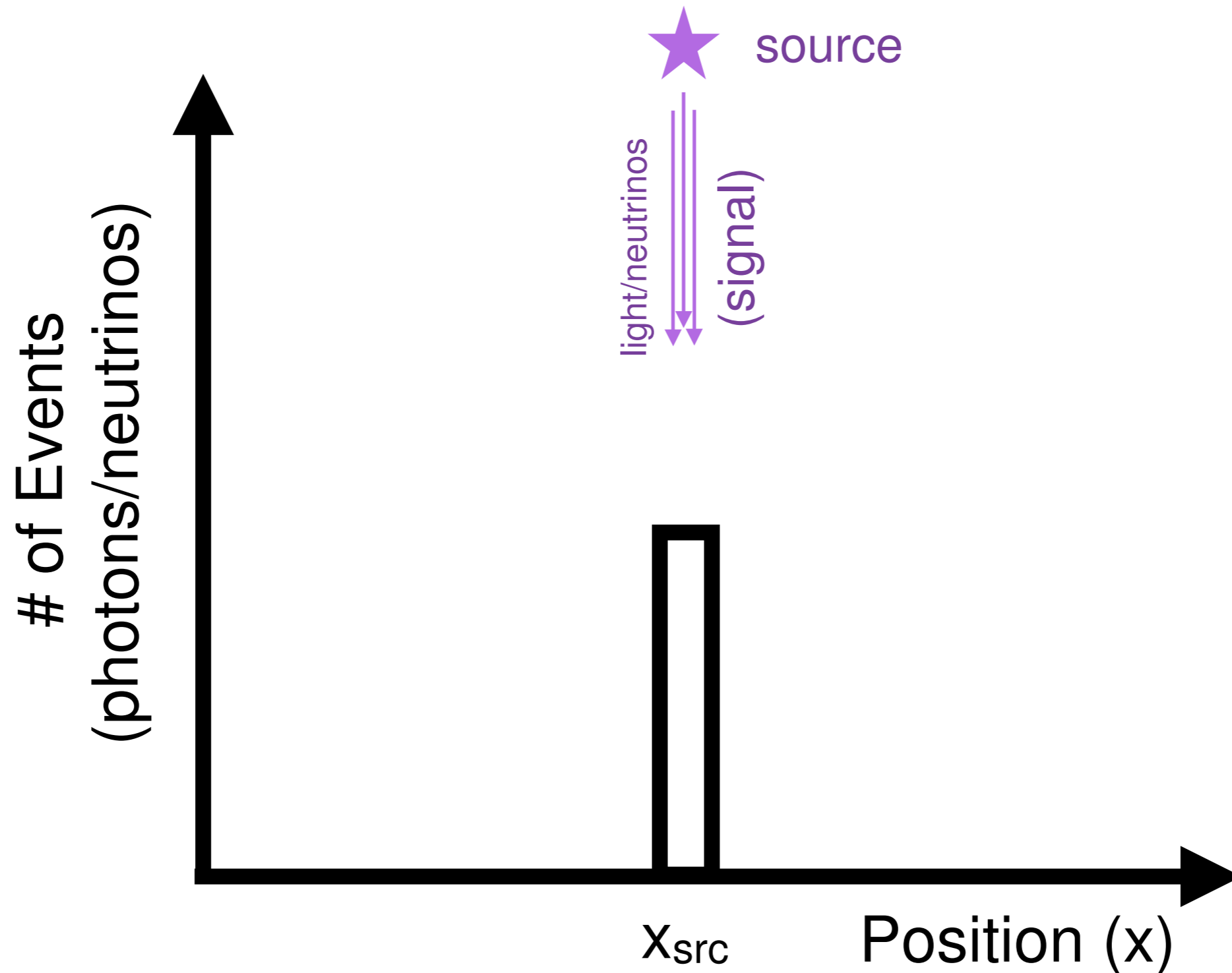
Formalism Part I: Spatial Distributions

Let's start by adding some axes to our example



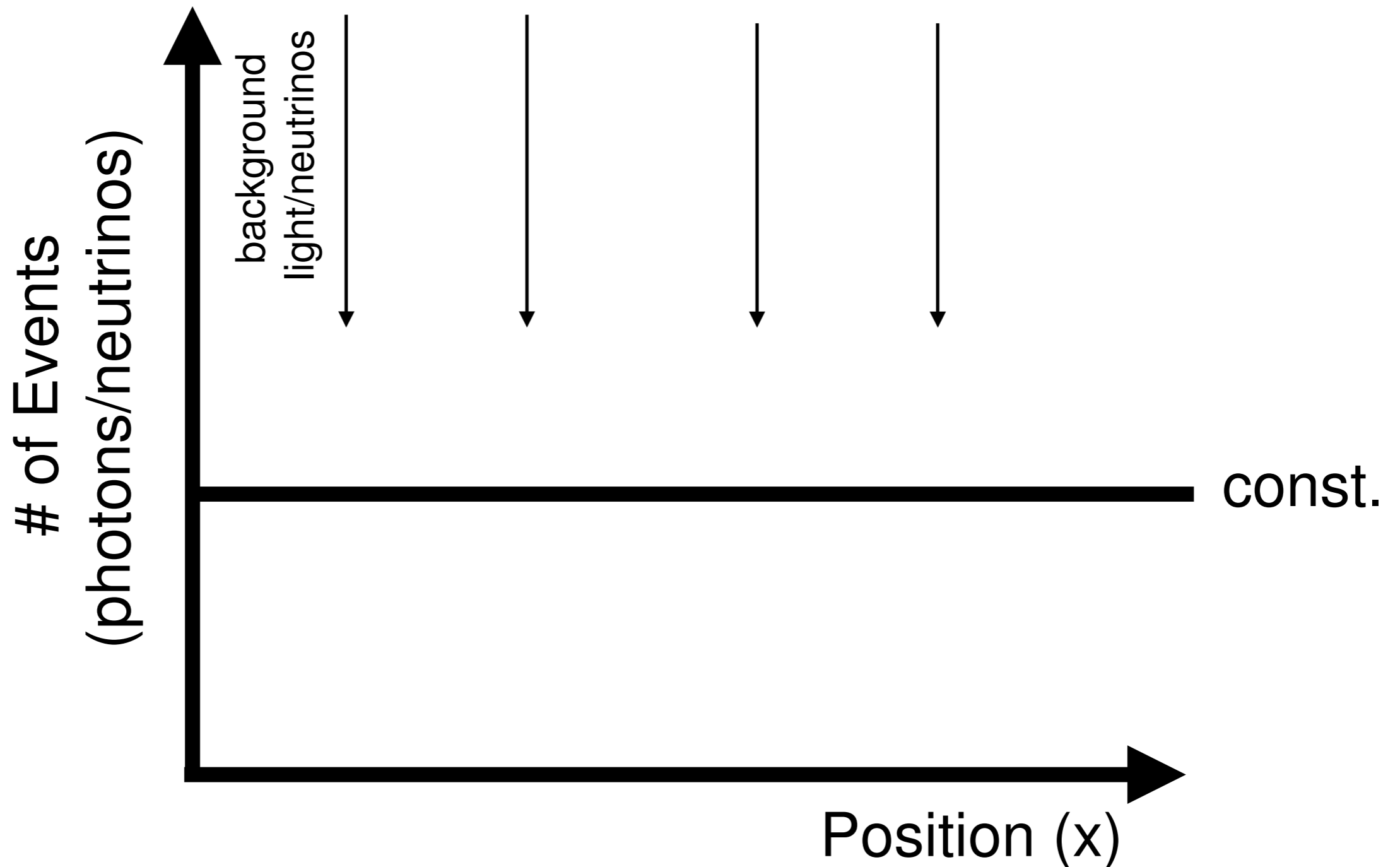
Formalism Part I: Spatial Distributions

Let's start by adding some axes to our example

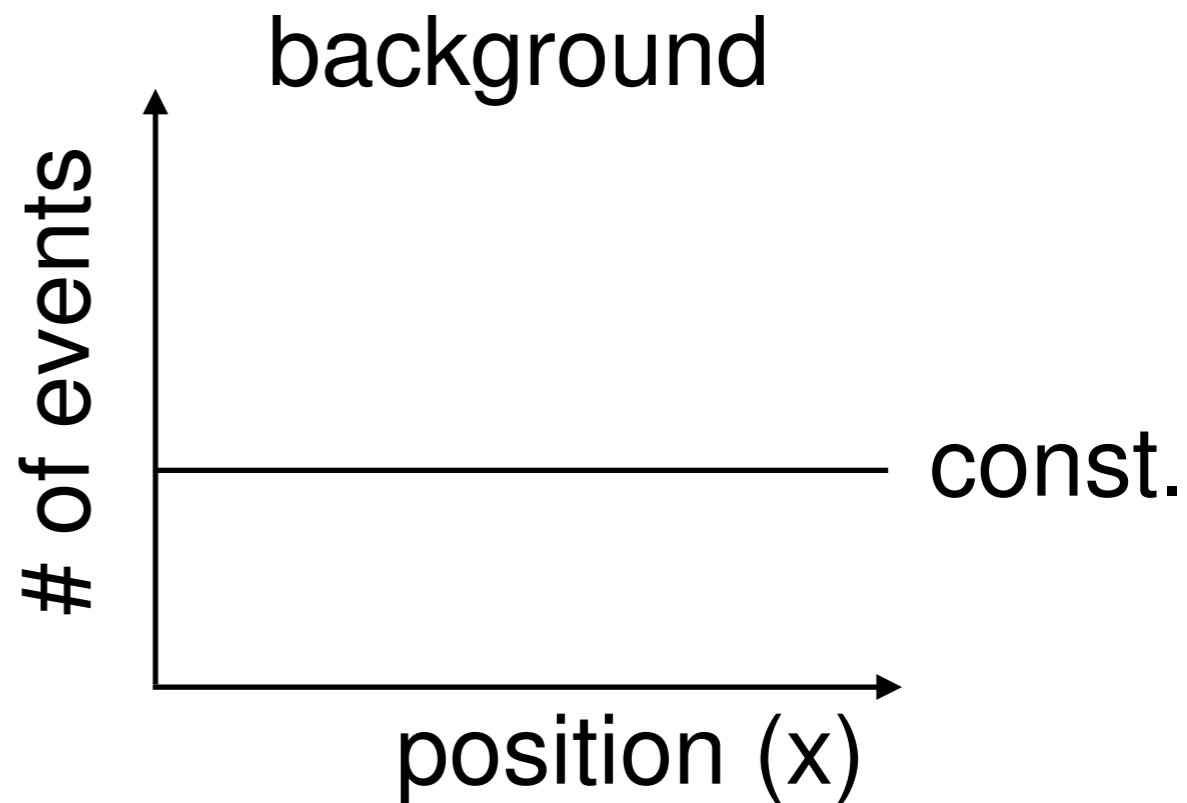
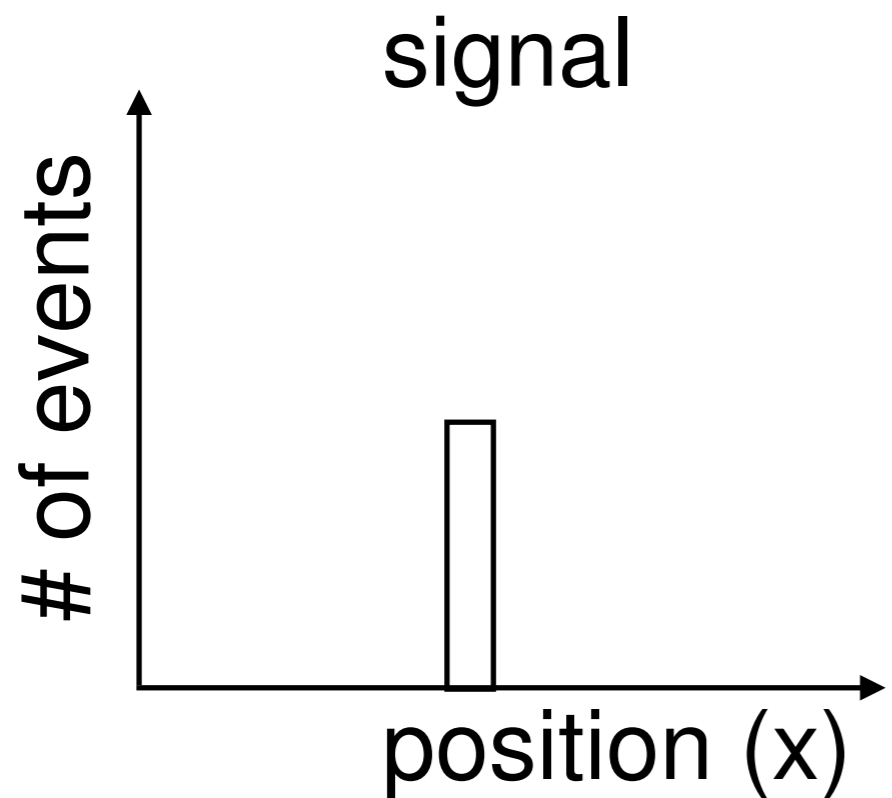


Formalism Part I: Spatial Distributions

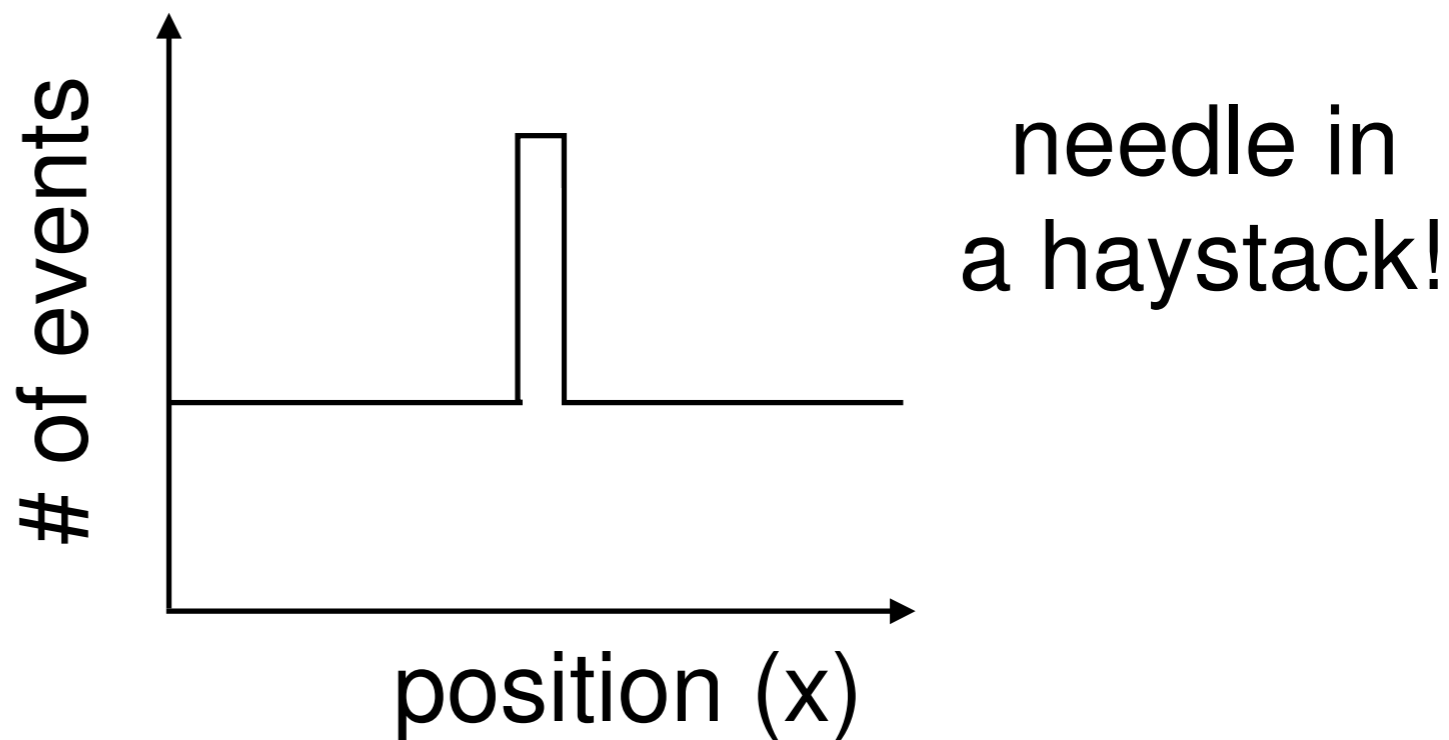
Let's start by adding some axes to our example



Formalism Part I: Spatial Distributions



data = signal +
background



Formalism Part II: Probabilities

Now that we know what signal & background distributions look like, we can formulate them in terms of **probabilities**

def. **probability** is the chance of getting a given result out of the total number of outcomes.

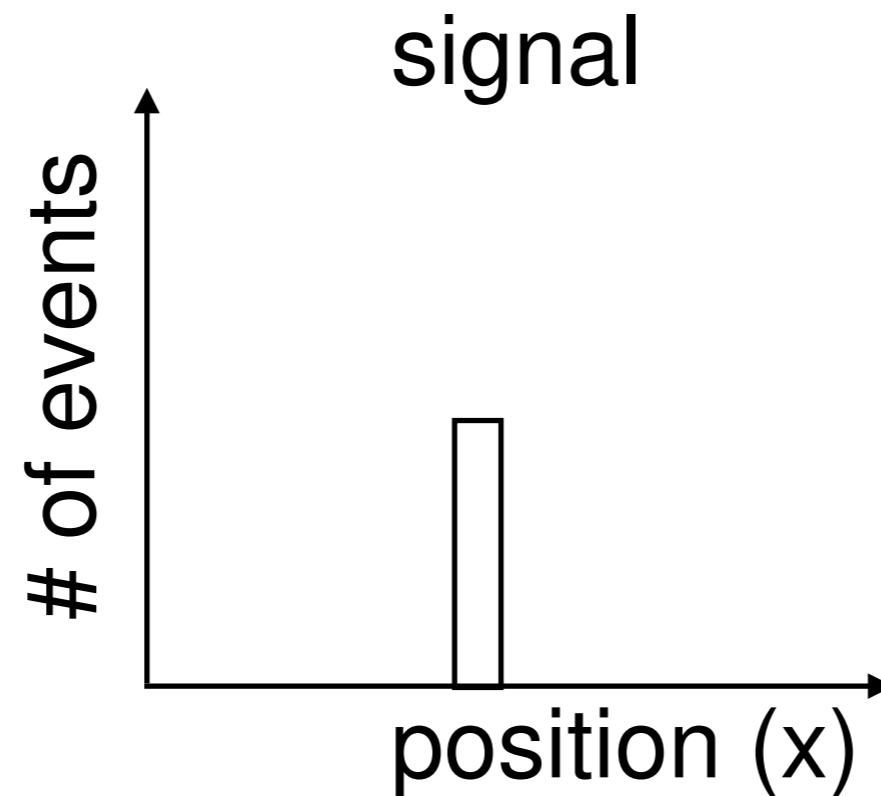
- > ranges 0 to 1 (never to always)
- > sum of all outcomes must be 1

This way we can ask the question: what is the probability that our data are consistent with background + signal versus the case of background only?

- > quantify if a point source is present in data

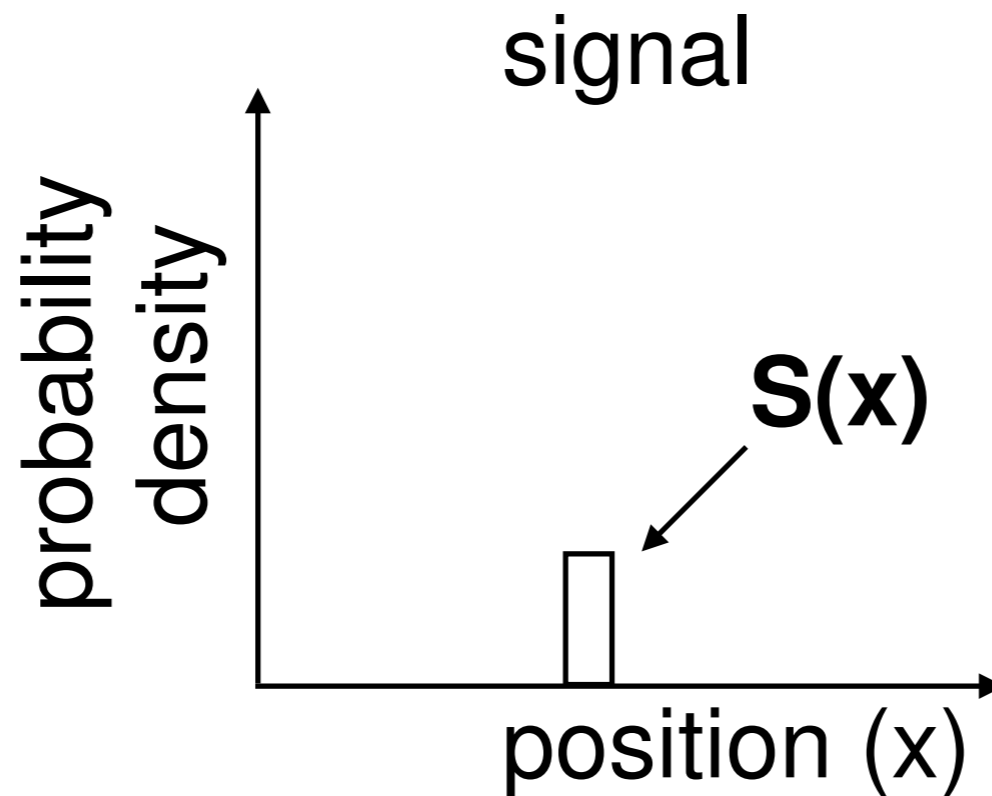
Formalism Part II: Probabilities

Ok, let's turn our distributions of events into probability densities \rightarrow scale such that integral of distribution is 1



Formalism Part II: Probabilities

Ok, let's turn our distributions of events into probability densities \rightarrow scale such that integral of distribution is 1

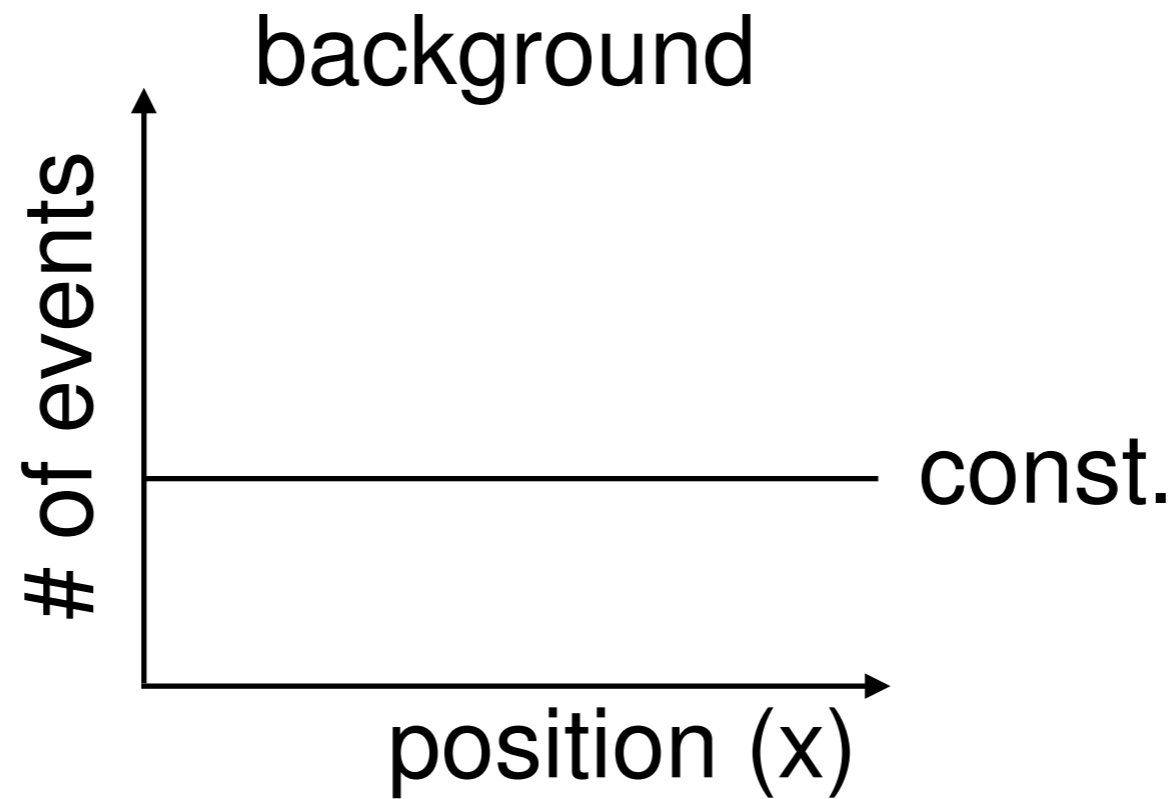


$S(x)$ = probability density of finding signal at x

$S(x) dx$ = probability of finding signal within dx of x

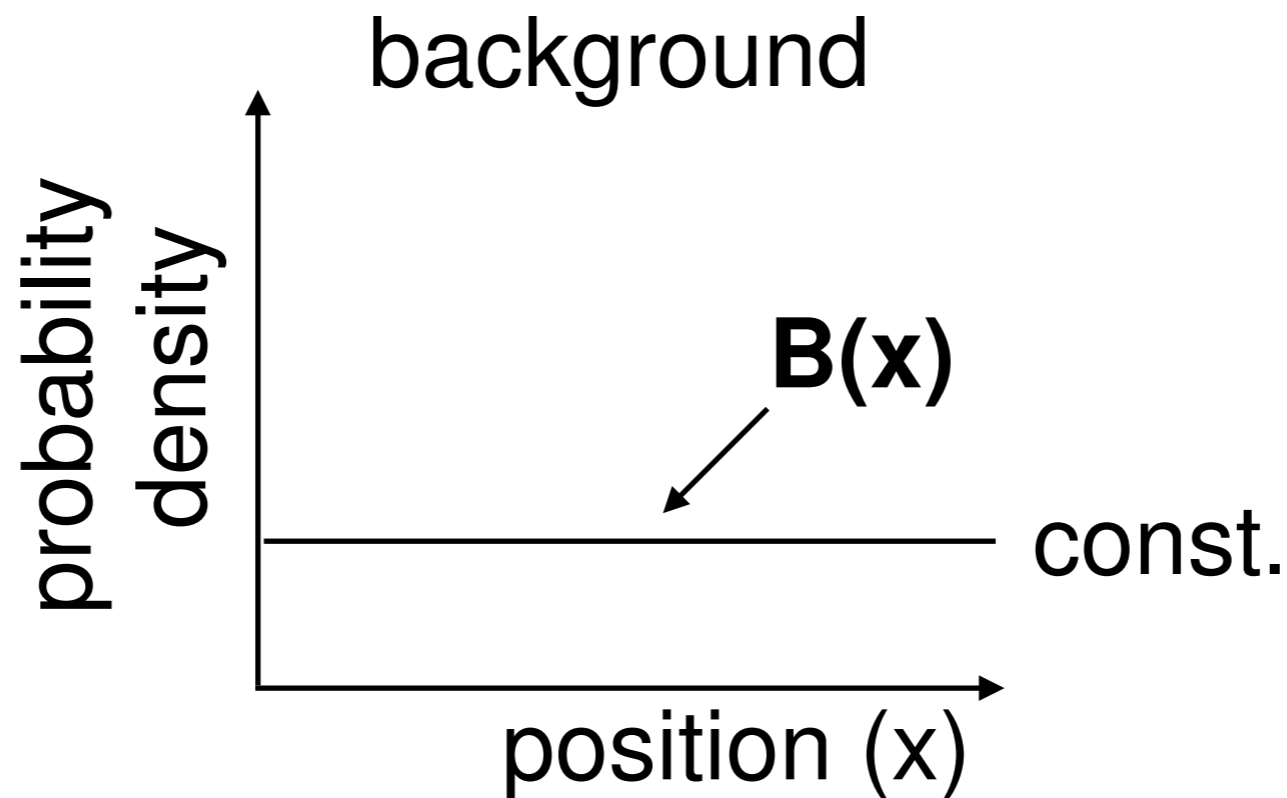
Formalism Part II: Probabilities

Ok, let's turn our distributions of events into probability densities \rightarrow scale such that integral of distribution is 1



Formalism Part II: Probabilities

Ok, let's turn our distributions of events into probability densities \rightarrow scale such that integral of distribution is 1



$\mathbf{B(x)}$ = probability density of finding background at x

In astronomy, we typically work on surface of a sphere \rightarrow uniform $\mathbf{B(x)} = 1/4\pi$

Formalism Part II: Probabilities

For a dataset with:

→ **N** total events

→ **n_s** signal events

→ **x_i** is the position where we detect the *i*th event ($i \in [1, N]$)

total *i*th signal prob.

$$\frac{n_s}{N} * \mathbf{S}(\mathbf{x}_i)$$

↑
probability density
of signal at **x_i**

probability *i*th event
is a signal event

+

total *i*th background prob.

$$\left(1 - \frac{n_s}{N}\right) * \mathbf{B}(\mathbf{x}_i)$$

↑
probability density
of background at **x_i**

probability *i*th event
is a background event

Formalism Part II: Probabilities

For a dataset with:

→ **N** total events

→ **n_s** signal events

→ **x_i** is the position where we detect the *i*th event (*i* ∈ [1, N])

$$L(\mathbf{n}_s) = \prod_{i=1}^N \left(\frac{n_s}{N} * \mathbf{S}(\mathbf{x}_i) + \left(1 - \frac{n_s}{N}\right) * \mathbf{B}(\mathbf{x}_i) \right)$$

L(n_s) is the total probability of the dataset containing **n_s** signal events. It is called a **likelihood function**.

The best estimate for the true value of **n_s** is the value which **maximizes the likelihood function**.

now for some math voodoo...

Formalism Part III: Voodoo

Working with **ratios of likelihoods** has some nice statistical properties.

Also, addition is easier and faster than multiplication so we define a **test statistic (TS)**:

$$\mathbf{TS} = 2 \log(L(n_s) / L(n_s = 0))$$

$$\mathbf{TS} = 2 \sum_{i=1}^N \log \left[\frac{n_s}{N} * \frac{\mathbf{S}(x_i)}{\mathbf{B}(x_i)} + \left(1 - \frac{n_s}{N} \right) \right]$$

Finding the value of n_s which maximizes TS is equivalent to maximizing the likelihood. TS = 0 means consistent with background only. TS ~ 25 typically proof of a point source.

