

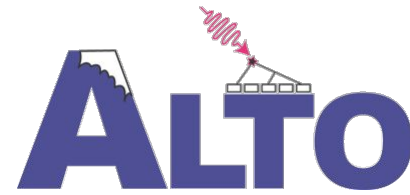
ALTO project:

Concept and Design Choices for a Wide-Field Southern VHE Gamma-Ray Observatory

Yvonne Becherini, Michael Punch, Satyendra Thoudam

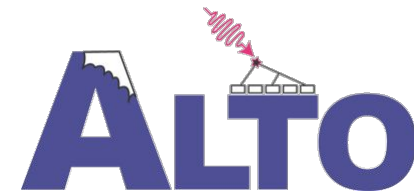
Partners:

LANL –	G. Sinnis, B. Dingus
CPPM –	J-P Ernenwein
CEA/Irfu –	B. Vallage, E Delagnes, J-F Glicenstein
NWU –	Markus Böttcher

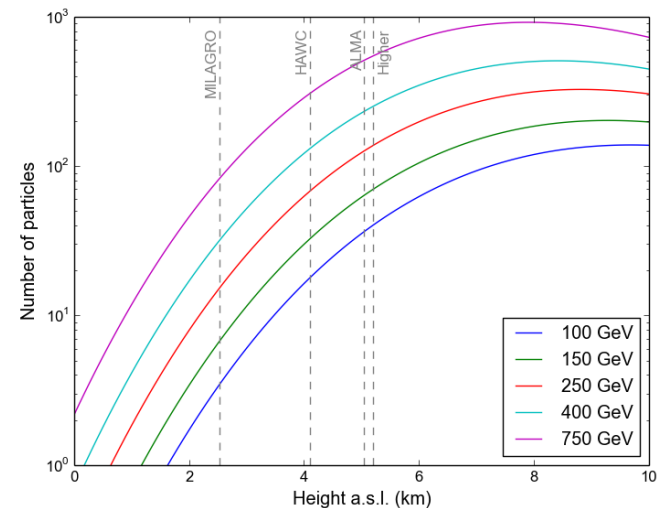
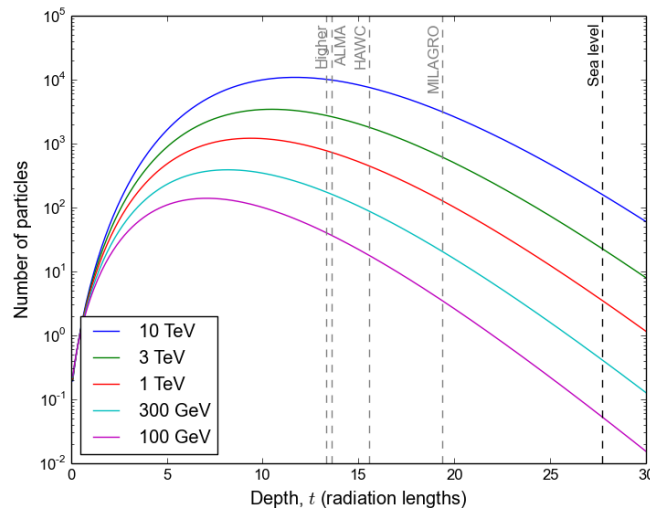


- Altitude ($> \sim 5\text{km}$):
 - For Physics goals, as a survey / alert instrument for transients
- Layout; Fine-grained array:
 - Smaller Water Cherenkov Detector (WCD) tanks than HAWC
 - Low dead-space
 - Improved angular resolution expected
- “Unit” design: Water Cherenkov tank + Muon-detector component
 - Liquid scintillator box (Scintillator Layer Detector, SLD)
 - Provides background rejection power
- Advanced electronics: NectarCam + White Rabbit
 - Trigger channel precisely time-stamped with “White Rabbit” system
 - Analogue memories + ADCs measure the waveform of the detector pulses
 - No cables from centre, only fibres

ALTO Altitude, Physics Considerations

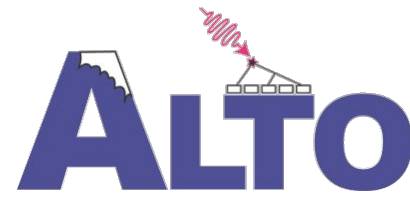


- Aim for $\sim 5.1\text{km}$ or higher
- Physics Goals:
 - Transients ... AGNs, GRBs, Grav. Waves...
 - Survey (esp. extragalactic)... Fermi Bubbles?



- \uparrow in altitude HAWC \rightarrow ALTO \Rightarrow $\times 2$ # particles
- Or for same # particles, energy \downarrow by 40%

ALTO Altitude, Technical Considerations



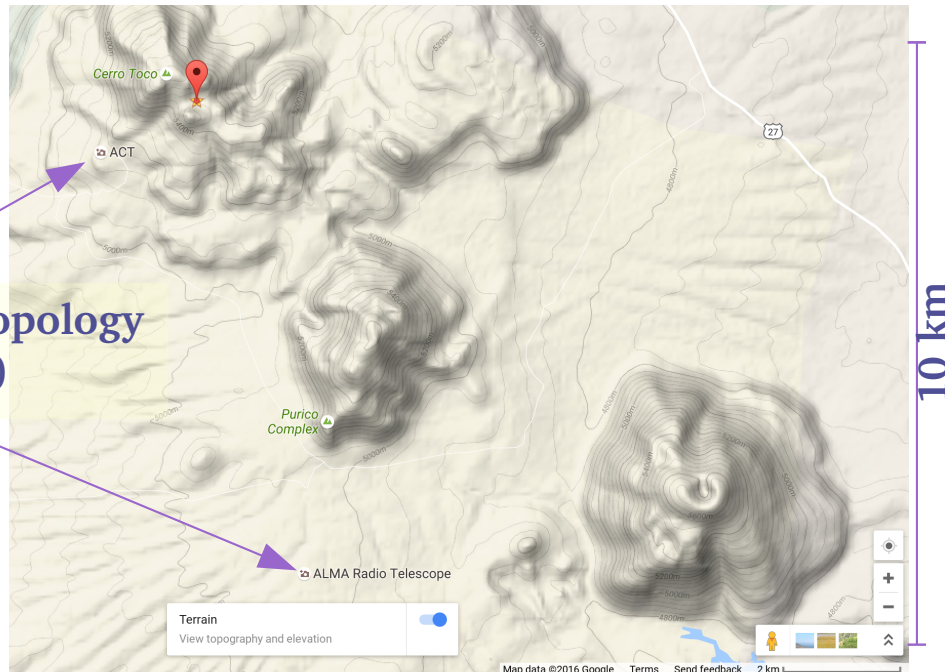
- High altitude \Rightarrow difficult working & operating conditions, so care about
 - Install-ability,
 - Reliability,
 - Maintainability
- Site characteristics
 - Plateau: a flattish space with diameter $\sim 160\text{m}$
 - Access: a nearby road needed
 - Water: source of water not too distant
 - Infrastructure:
 - Power grid or access to Diesel
 - Network access (data transfer, remote operation /monitoring)
 - Lower altitude accommodation for local crew
 - National interest and funding in host country

ALTO Site Candidates

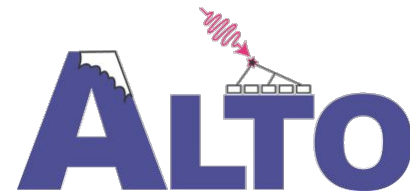


- Search “by google-hand” along the Andes
⇒ best current candidates
 - 4.8km at LLAMA (possible QUBIC) site
 - ~5km near Chacaltaya
 - 5.1km at ALMA/ACT site, perhaps up to 5.4km!

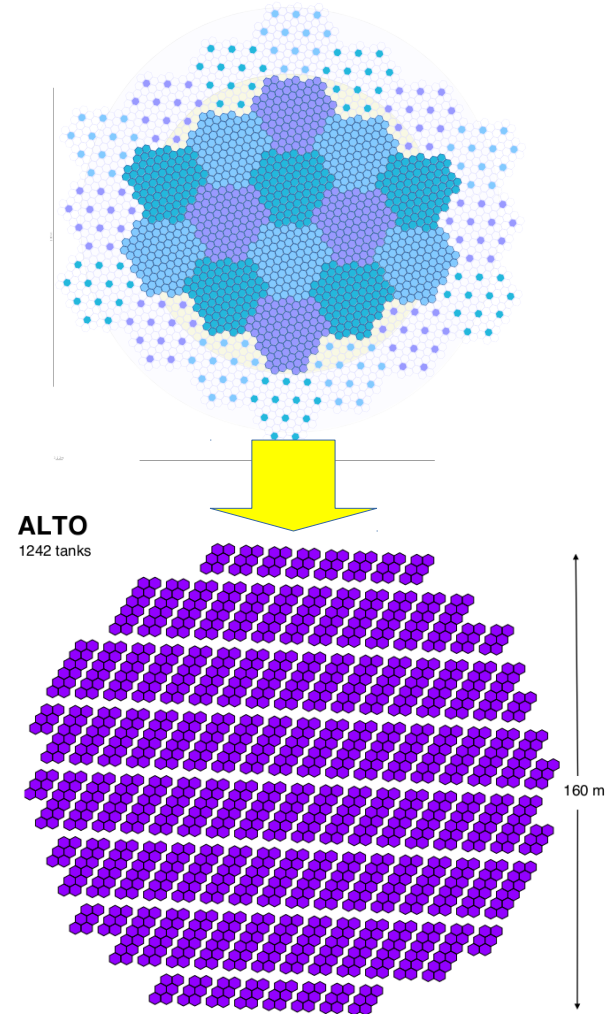
**ALMA/ACT Site Topology
(marker at 5.4km)**



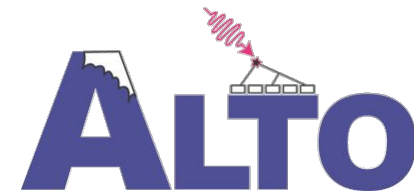
ALTO Layout Design Process: “The Simpler The Better”



- Initial design
 - All units contiguous, accessed from above
 - Scintillator box sunk into WCD tank
 - Optical Module (OM) with PMT up/down for WCD/SLD
 - ... probably unmaintainable
- Current design:
 - “Clusters” of 6 units
 - Scintillator boxes accessed from the sides
 - OM with single large PMT for WCD, separate smaller PMT for SLD
 - Access roads wide enough for truck, W-E



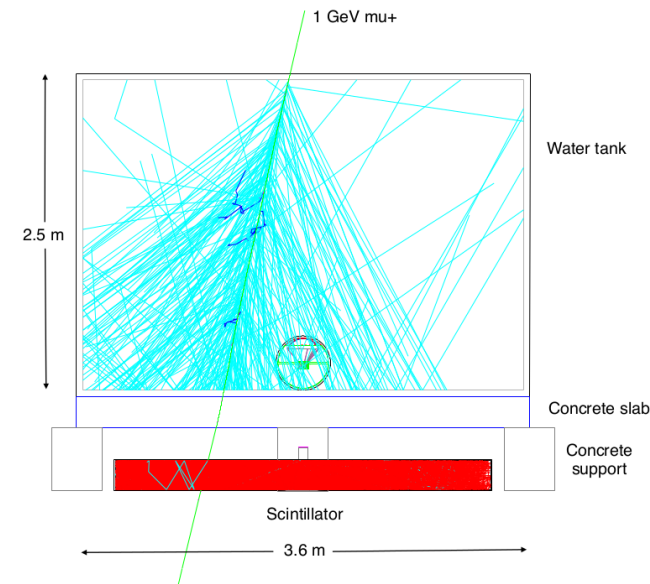
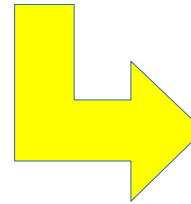
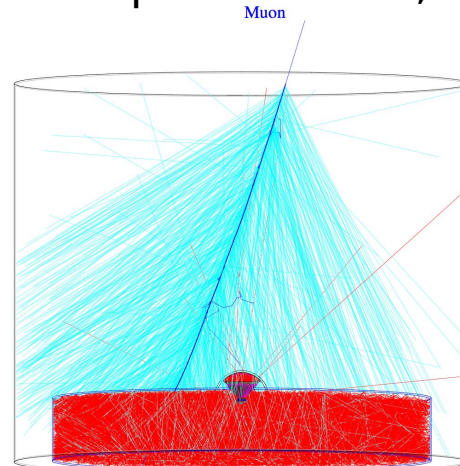
ALTO Unit Design Process



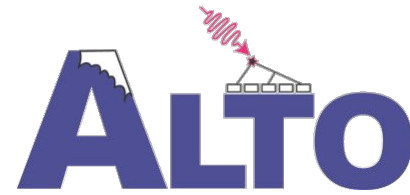
- Given diameter, WCD height imposed by geometry of Cherenkov angle
- Assuming Antares OM (see AMIGO presentation, F. Schüssler, tomorrow)

- Initial design
 - 1 central OM
 - Dismount Antares OM
 - Add 2nd smaller downwards PMT
 - \Rightarrow Too Complicated!!

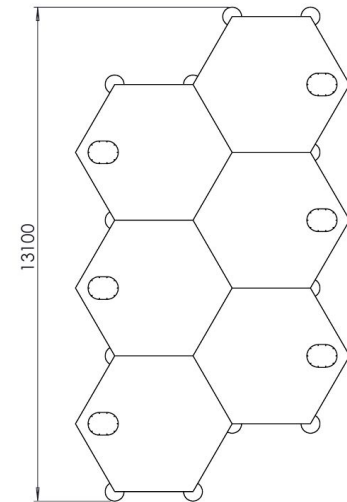
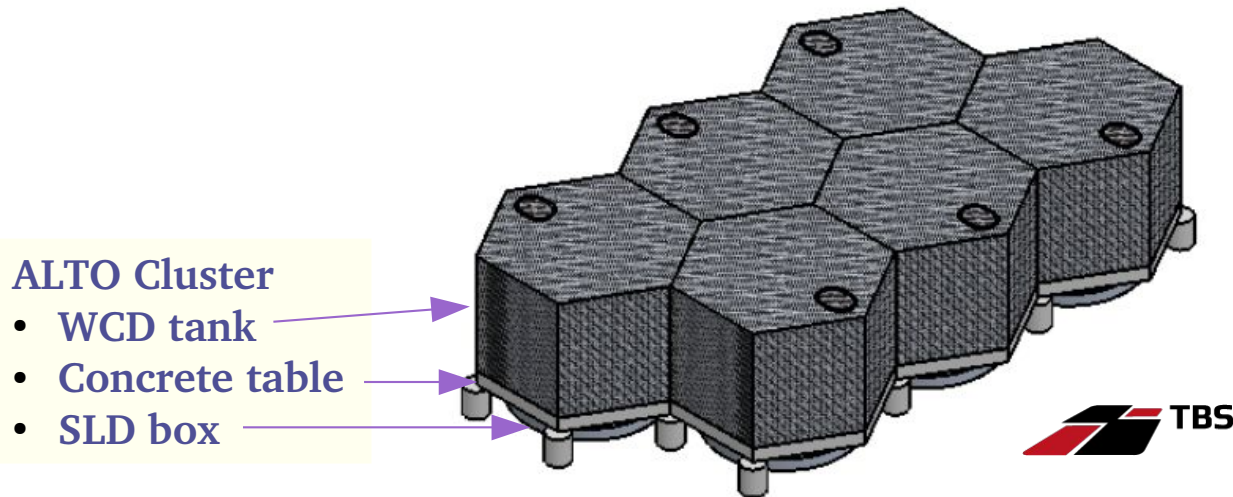
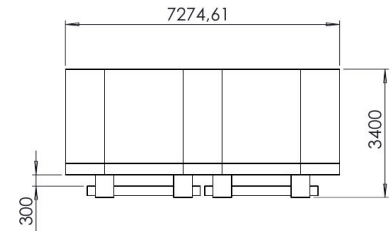
- Current design
 - Separate WCD and SLD
 - Keep Antares module intact!
 - Scintillator height lowered (from simulations)



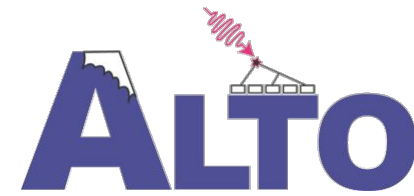
ALTO Cluster (6xUnits)



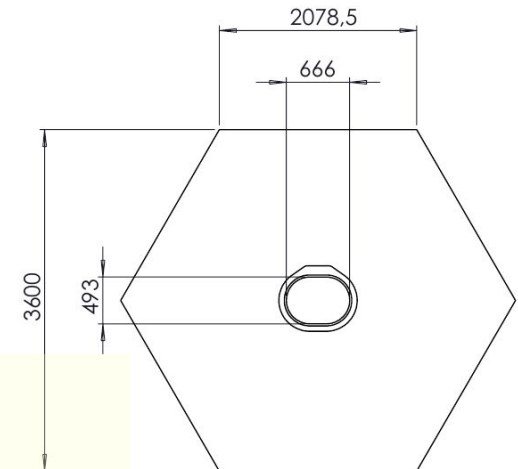
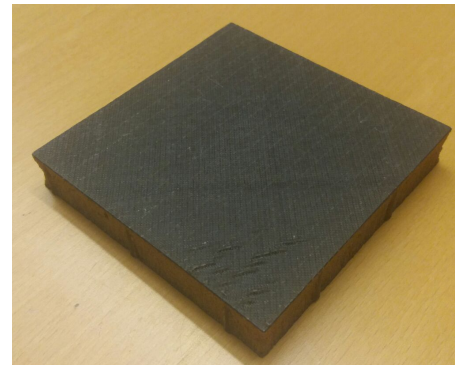
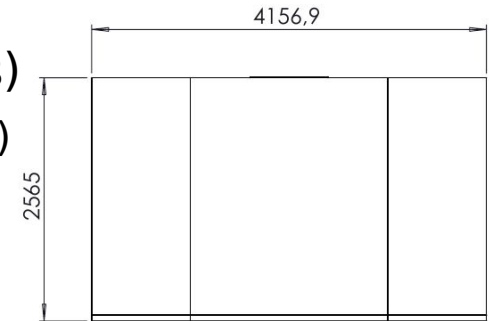
- Cluster = Group of 6 Units = 6 x (WCD + SLD)
 - WCDs on concrete “table” (1 pour for cluster)
 - SLDs below “table”, on telescopic rails
- Each cluster to have common:
 - Electronics readout unit
 - Solar panel + battery (TBD)
 - Communication to central control by fibre only



ALTO Unit Technical Details: WCD Tank



- WCD Tank: Carbon Fibre (TBS-Yard, Torsås... CF experts)
 - Strength to resist 28T water, but durable and light (<500kg)
 - Flat Pack (Ikea-type) for international transport
 - Glueing at lower altitude + UV-resistant paint applied
 - Transport by truck the final 10s of kms
 - Filling on-site from tanker trucks (via lower valve)



ALTO Tank

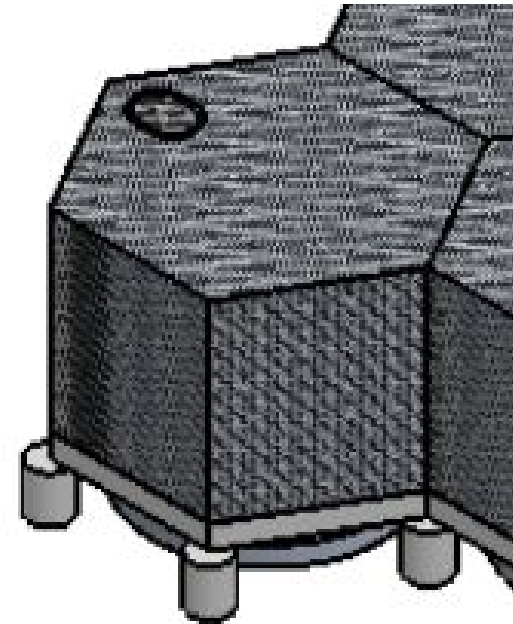
- Right: Views
- Above: Carbon Fibre sample
- Left: Strength calculations



ALTO Unit Technical Details: Table + SLD box



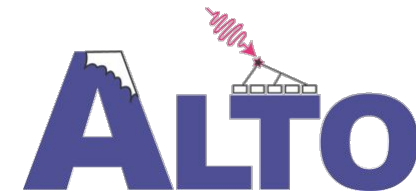
- Reinforced Concrete Table
 - 30cms is overkill for support, but...
 - Provides extra absorption of most e^\pm in showers
- Aluminium SLD box
 - Shallow cylinder, 3m diameter, 25cm height
 - Central protrusion for PMT housing
 - Box on telescopic rails hanging from Table \Rightarrow Ease of access
 - MP5 company in partnership with TBS yard
- Filled with liquid scintillator
(Collaboration with J-P Ernenwein):
 - Linear Alkyl Benzene (LAB) in which is diluted
 - 5 g/L PPO (Scintillation Grade: 2,5-Diphenyloxazole)
 - 0.02 g/L POPOP (1-4,bis-2-(5-Phenyloxazolyl)-benzene)
 - Non-toxic, stable liquid
 - Used for Antares surface array, SNOLAB, others...
- Plan to measure scintillator characteristics in lab (summer 2017)



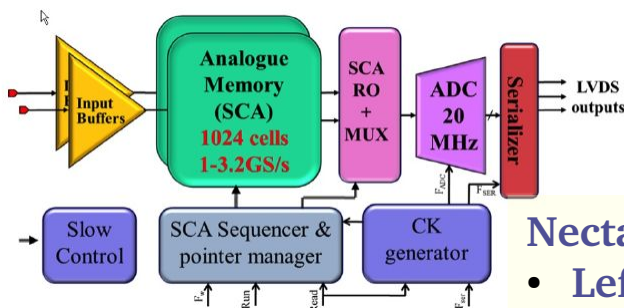
ALTO Unit

- WCD tank
- Concrete table
- SLD box

ALTO proposal for NectarCam electronics

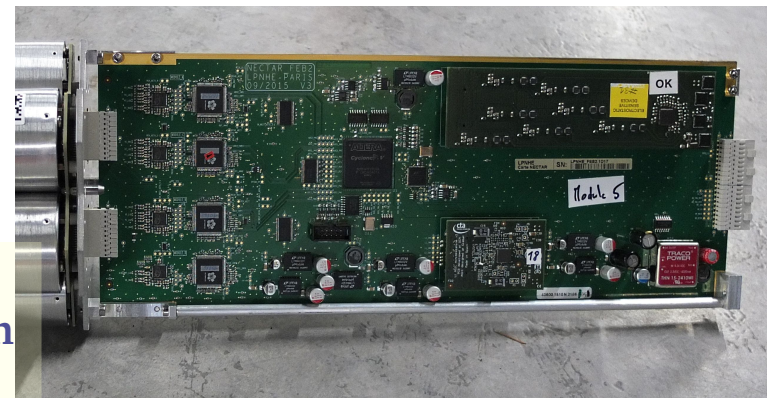


- Use of NectarCam front-end module allows a low-cost, low power read-out close to the detectors
- Main specifications of NectarCam module (module = 7 PMTs)
(adapted from J-F Glicenstein, Gamma2016)
 - Analogue bandwidth > 250 MHz (measured)
 - Analogue memory depth 1024 cells (differential)
 - Integration window 1-60 ns (cells @ 1 GHz)
 - Sampling frequency 0.5-2 GHz (nominal 1 GHz)
 - Dynamic range (<5% non linearity) 0.5-2000 photo-electrons (measured)
 - Charge resolution 32% (single p.e.), 2 % (2000 p.e)
 - Time resolution <1 ns (>5 p.e)
 - Dead time 1.93 – 6.52 μ s (for read-out of 16 – 64 samples)
 - Dead time fraction <3 % at 4.5 kHz trigger rate
 - Power consumption (per module) 18.5 W

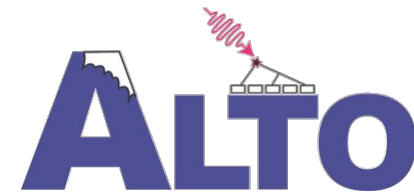


NectarCam

- Left: Nectar ASIC operation
- Right: NectarCam board

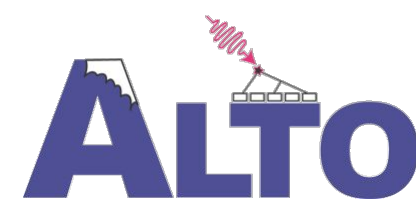


ALTO using NectarCam Electronics



- Modification of NectarCam electronics
 - Nectar Analogue memory (depth 1024) stores waveform (high and low-gain channels)
 - Each WCD/SLD self-triggers at a few p.e. (in standard NectarCam Pixel-Trigger channel)
 - Writing to Analogue memory stops, readout begins in defined window
 - Trigger signal sent to White Rabbit node for time-stamping (see next)
 - FPGA can provide integrated charge, timing information (max. time, width at half-max or above threshold ...)
 - Modify read-out to handle independent channel (Nectar ASIC) read-out?
- Design process will be driven by Simulations, initial tests
 - WCD Rate extrapolated from HAWC and others, expected $< 20\text{kHz}$
 - Use local coincidence of WCDs in cluster (L0) to lower rate, but affect on low-energy events?
 - Do we need GHz sampling? (note, Trigger time-stamping at ns is independent)
 - SLD rate extrapolated to be a few kHz ($< 2\text{ kHz?}$)
 - Cannot do local coincidence of SLDs, since looking for isolated muons!
 - But certainly can use $\leq 0.5\text{ GHz}$ sampling rate for Scintillation channel
 - If rates turn out too high for acceptable performance
 - explore DREAM (dead-time free) analogue memory

ALTO Electronics Trigger Time-stamping

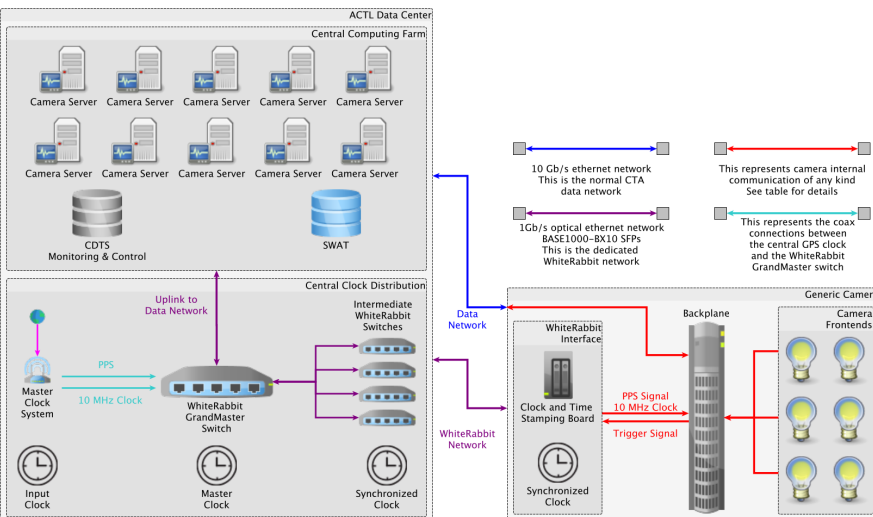
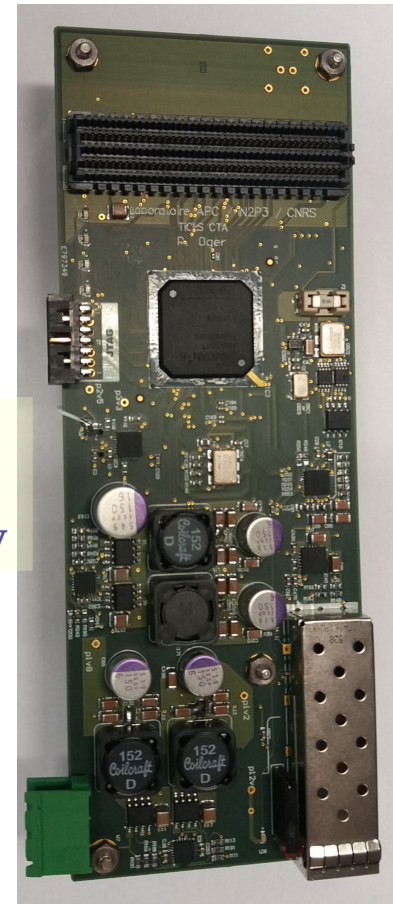


- Trigger for each WCD and SLD time stamped by White-Rabbit node (1-2 per cluster)
- Each WR node has its dedicated WR fibre, to send time-stamps to central point through hierarchy of WR switches at centre
- Software Array Trigger identifies coincidences
- In parallel, data sent to central server over standard 10Gbps fibres
- If no coincidence, non-coincident data are eliminated
- Coincident data are saved and treated
- Similar to the CTA Time-Stamping and Software Trigger strategy



White Rabbit

- **Right: SPEC/TiCkS node**
- **Left: CTA trigger strategy**

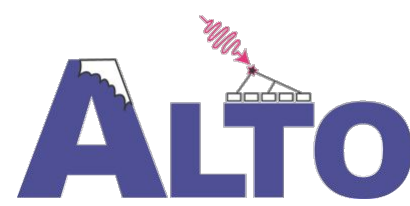


ALTO Background Rejection



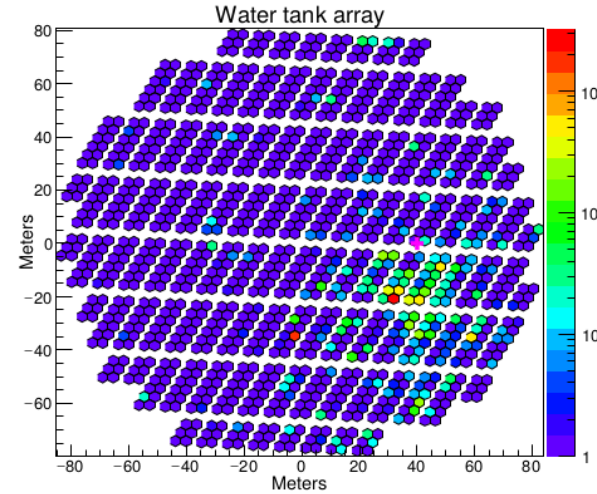
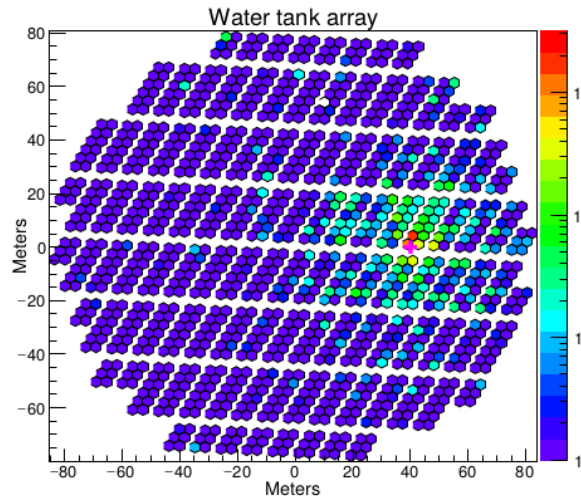
- Scintillator channel
 - \Rightarrow rejection from muon multiplicity
- Fine grained array
 - Better fit of the Lateral Distribution Function (LDF)
 - \Rightarrow rejection from parameters determined from fit
 - Good angular resolution expected
 - \Rightarrow background rejection for point sources
- See example showers on next slide,
and especially, next talk on ALTO Simulations, S. Thoudam

Gamma-ray and proton showers at same energy and core position

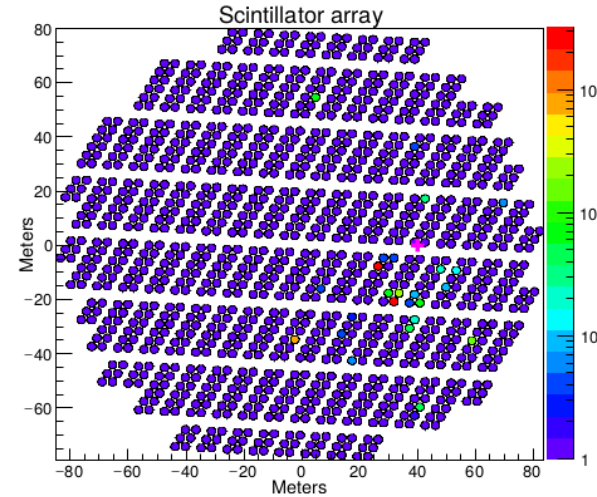
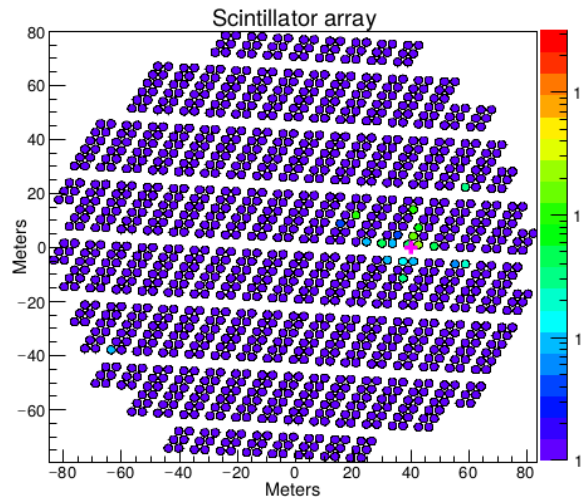


x, y 40, 0 m
E 1.5 TeV
Zen 12°

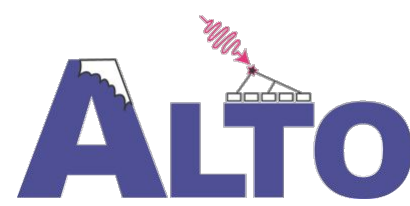
Gamma-ray



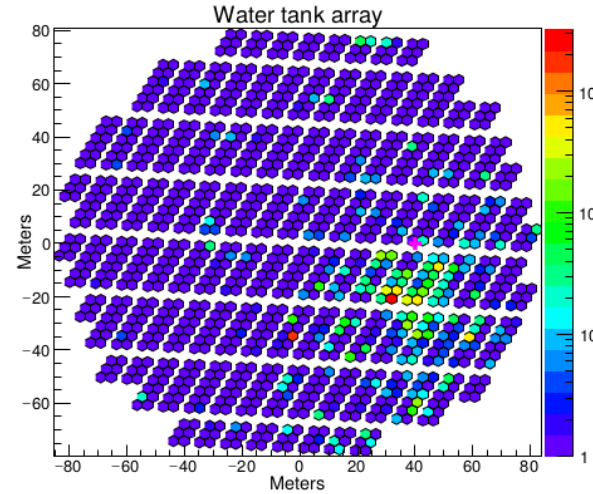
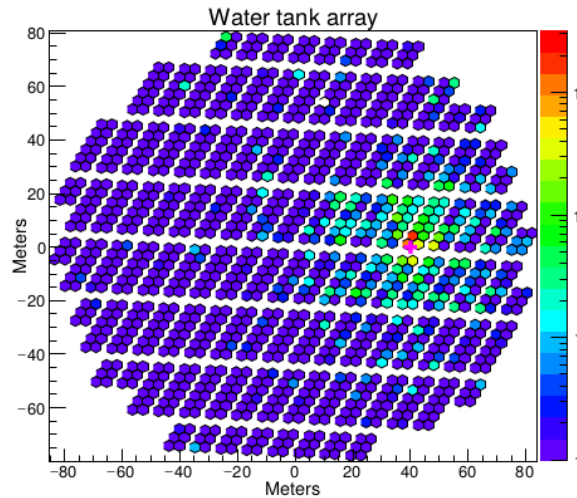
Proton



Gamma-ray and proton showers at same energy and core position



x, y 40, 0 m
E 1.5 TeV
Zen 12°



Gamma-ray

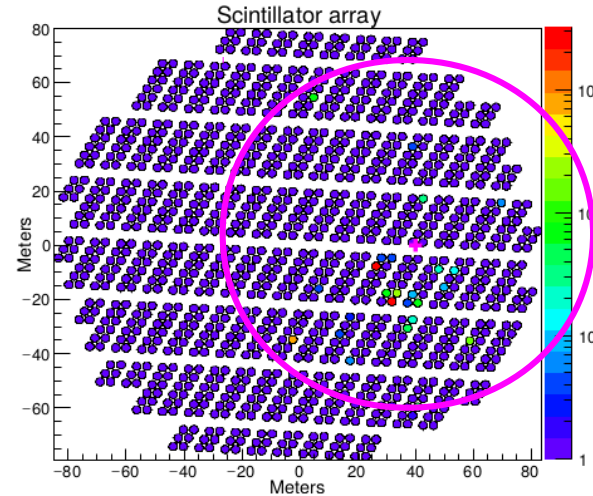
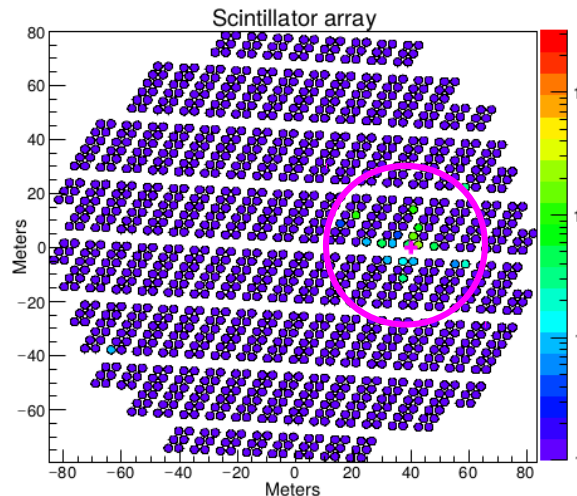
Scintillator signal
from e^\pm / γ 's.

Concentrated,
Correlated to core

Proton

Scintillator signal
from e^\pm / γ 's & μ 's.

Widely-Dispersed,
Not well-correlated
to core



ALTO Prototype Plans



- Prototype with 1 unit, fully funded
 - To be installed at LnU
 - May pour concrete table for ≥ 3 units (\Rightarrow expand as funding allows)



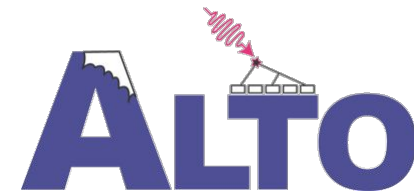
ALTO Prototype location

- Above: Aerial view
- Below: South / North view

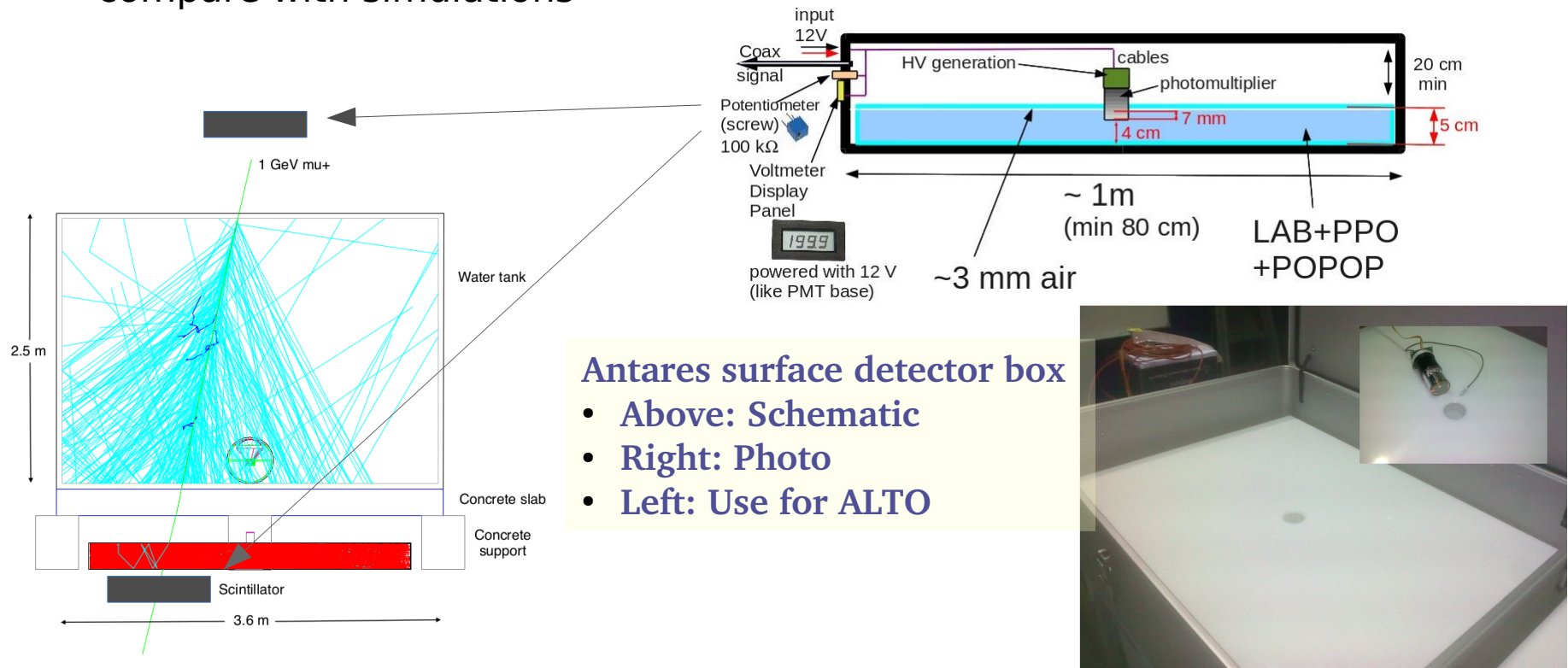


- Will provide
 - rates (WCD/SLD) to be compared with simulations for LnU altitude
 - Test if NectarCam cluster electronics can be used ~ “as-is”
 - Installation experience

ALTO Prototype Expected results



- Will test using “Muon boxes” from the Antares surface detector array (J-P Ernenwein)
- Can determine response to muons with \sim known trajectories, compare with simulations





- Design of ALTO is advancing
- Careful consideration of install-ability, reliability, and maintainability is influencing design
- Self-contained Electronics as close to detectors as possible
- Simulations are the next step to allow design decisions esp. for electronics, e.g.
 - Timing precision needed
 - Waveform precision and characteristics to save
 - Dynamic range needed
 - Coincidence strategy to be used
- Subsequent step is the prototyping of 1-3 units @ LnU
- Future step: Prototype installation on test site?