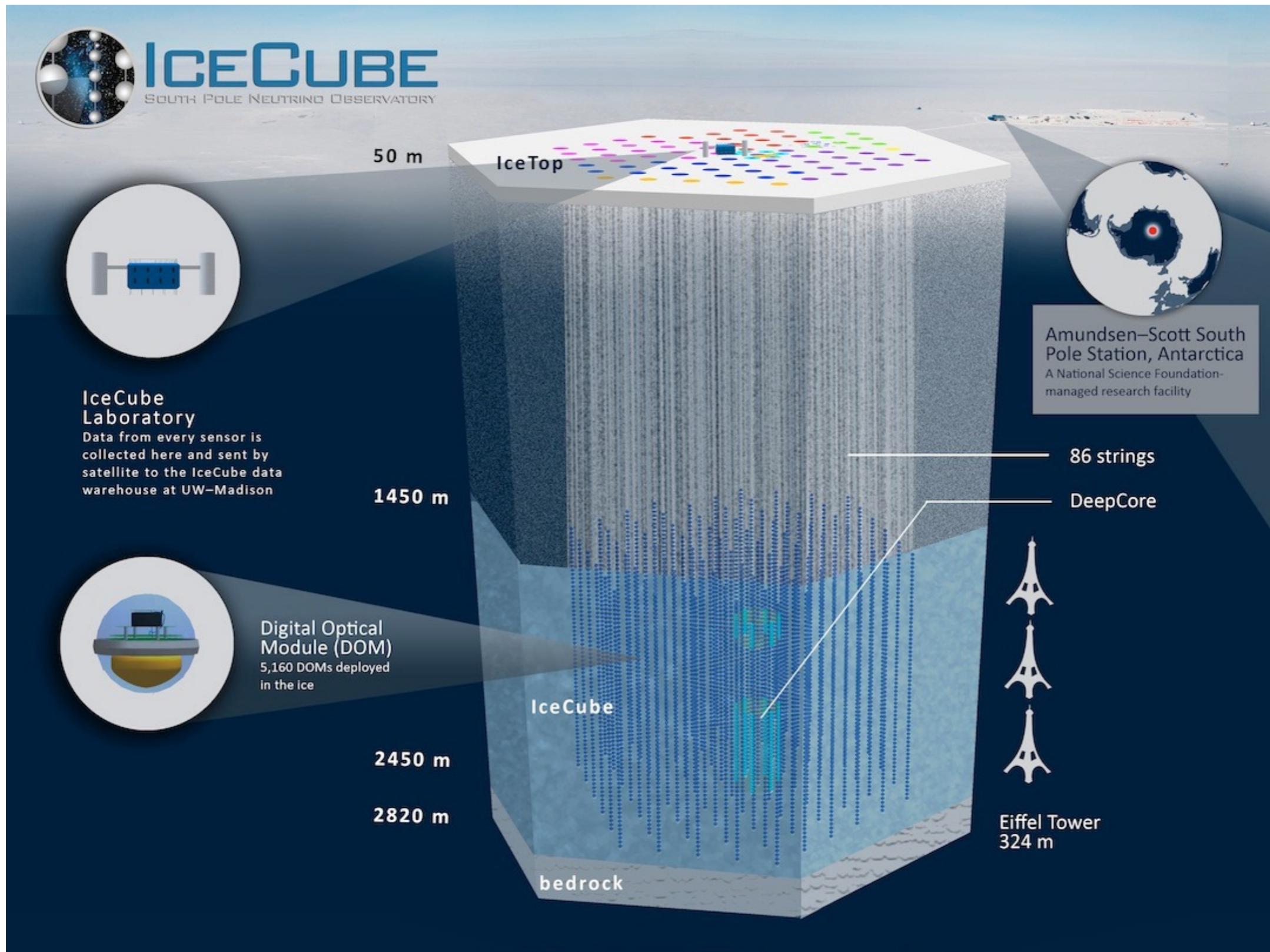


i³SIM

IceCube Bootcamp 2017

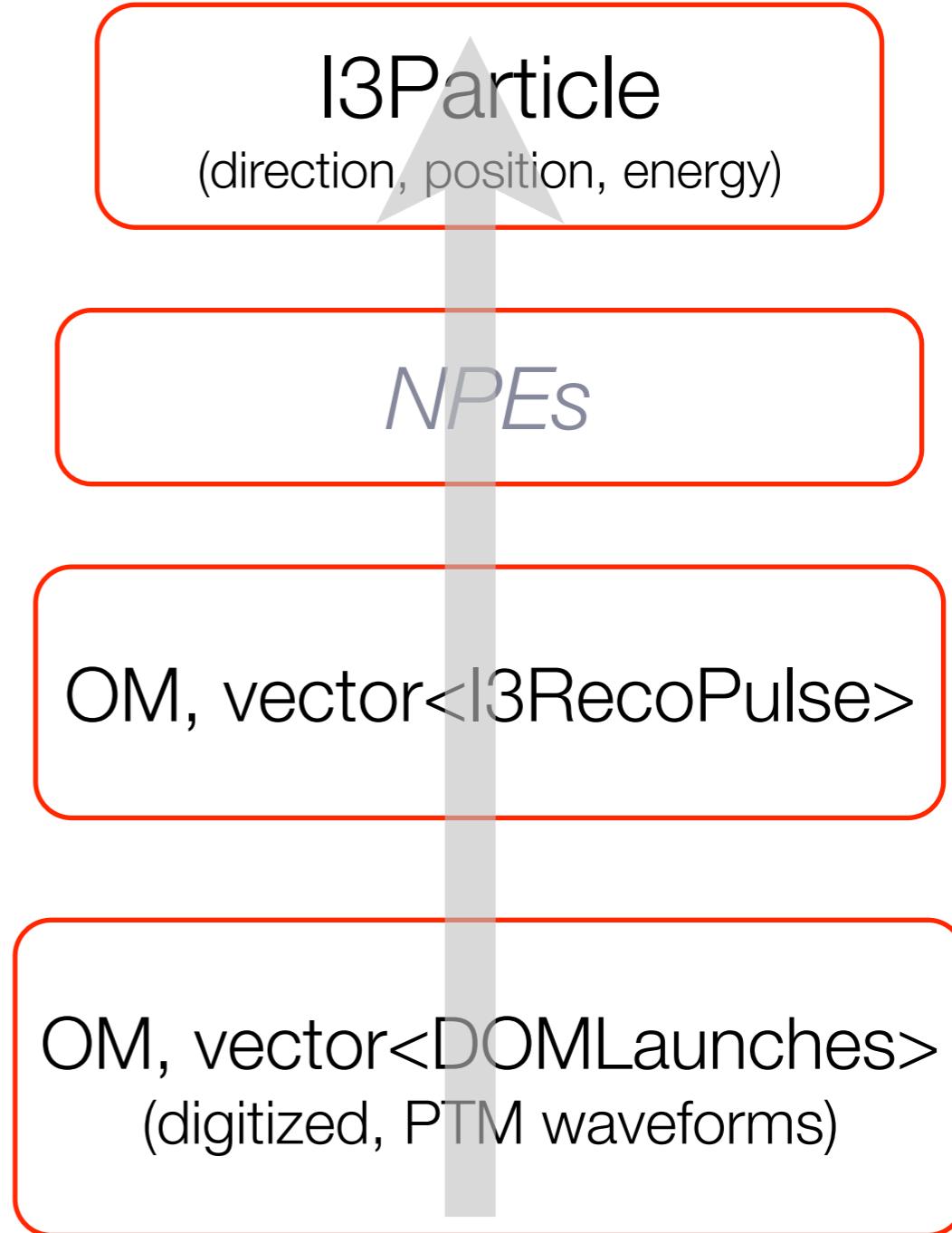
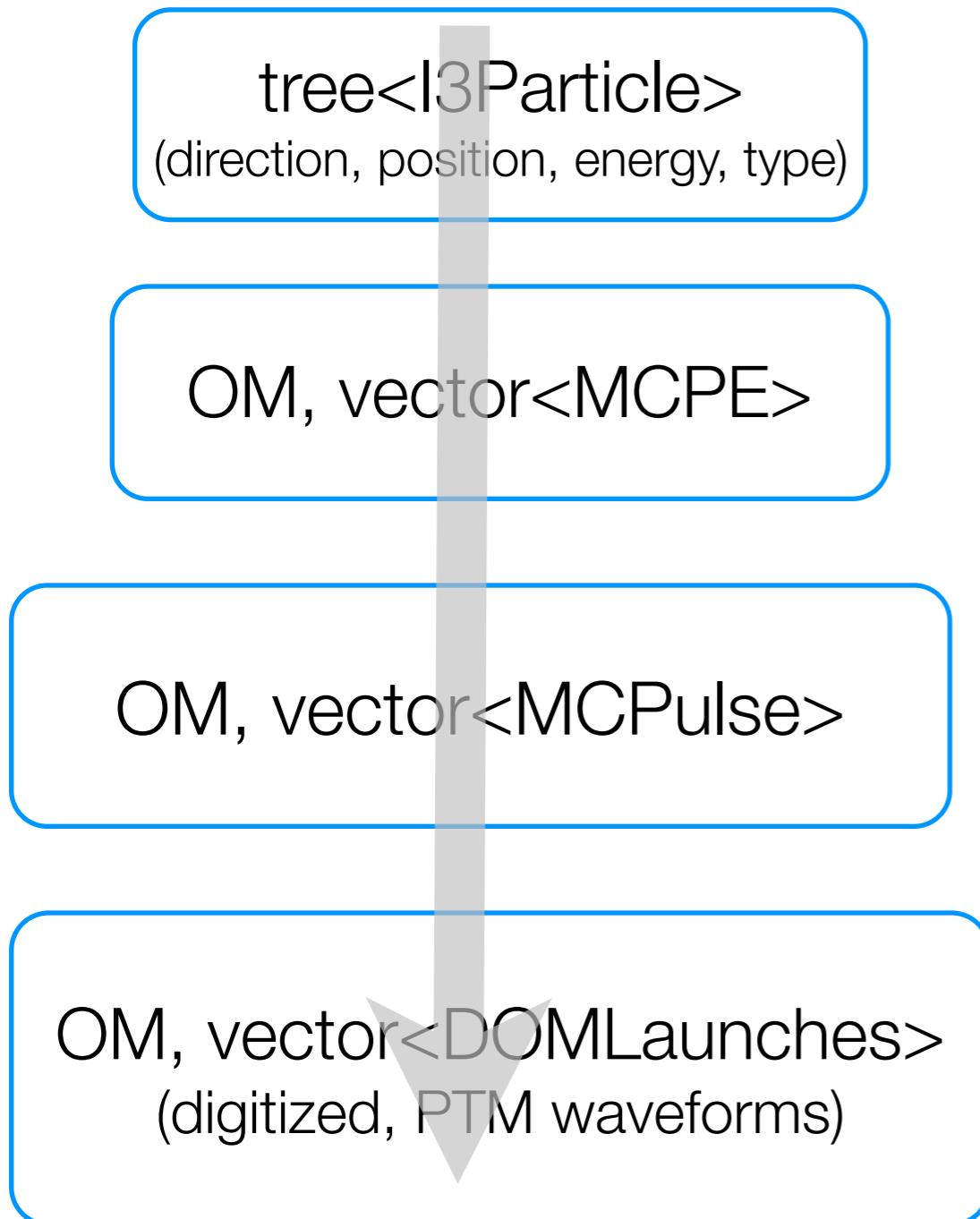
Juan Carlos Díaz Vélez

The IceCube Observatory



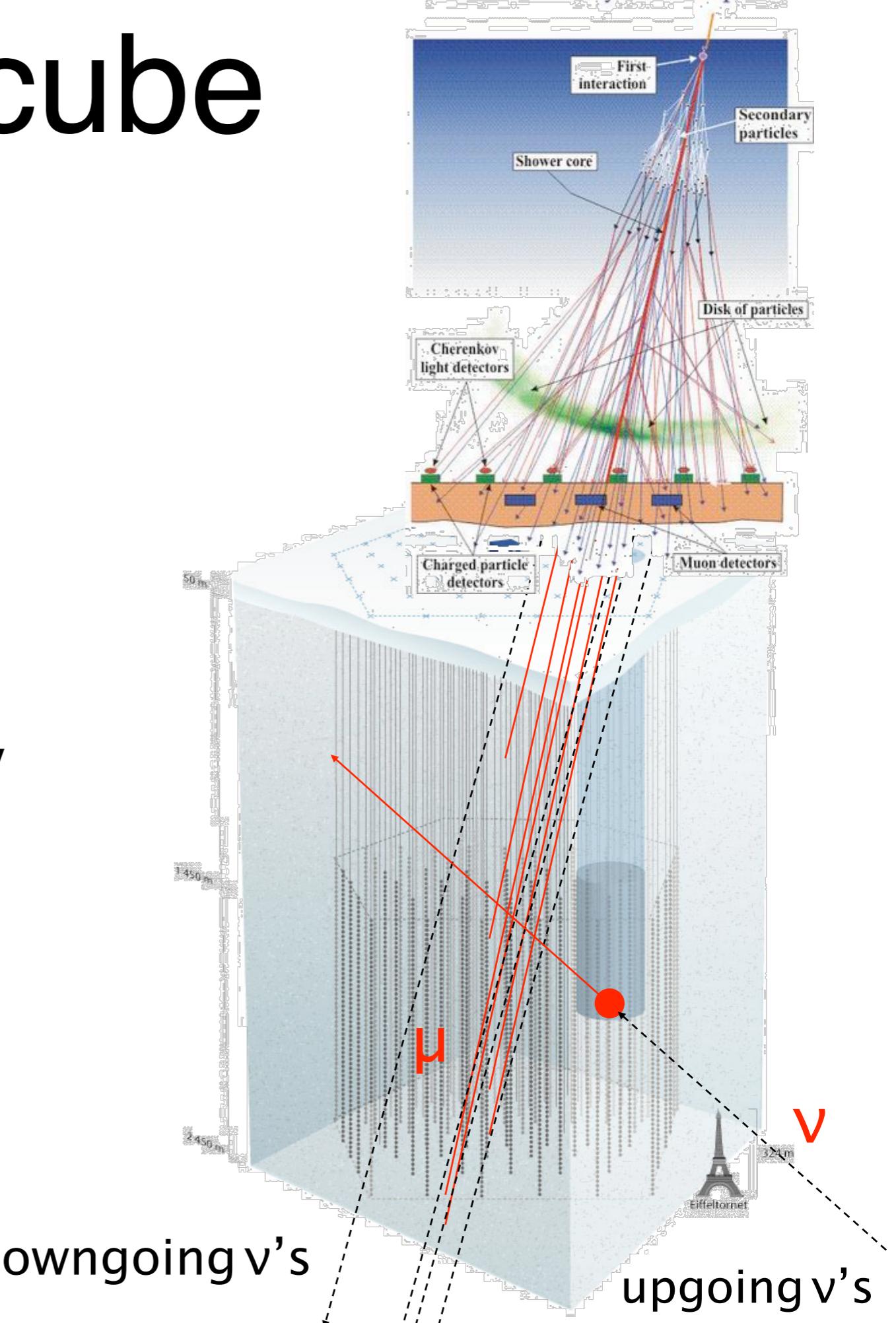
Simulation

Reconstruction



Events in icecube

- Air shower detection @ surface
- Penetrating muon detection in deep ice
- Events dominated by cosmic ray muons : $10^6 \mu$ for every ν that interacts in IceCube
- Atmospheric ν 's



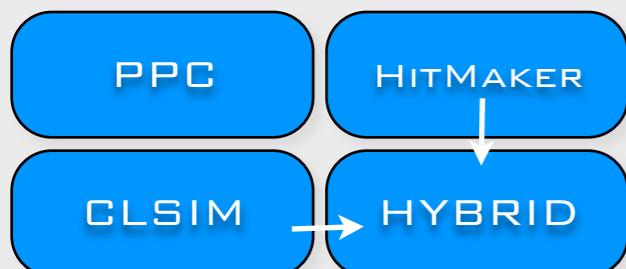
GENERATION



PROPAGATION



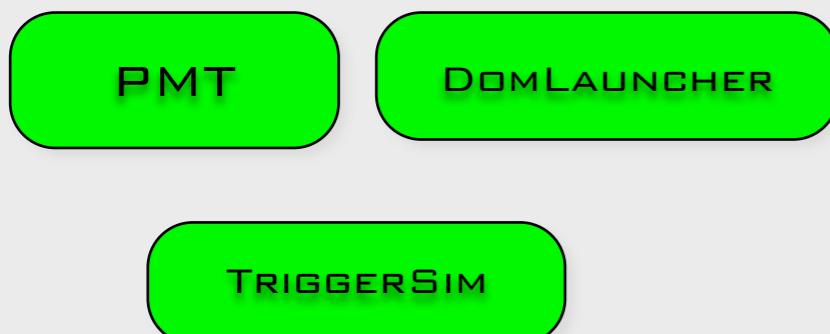
PHOTON PROPAGATION



BGND NOISE/MUONS

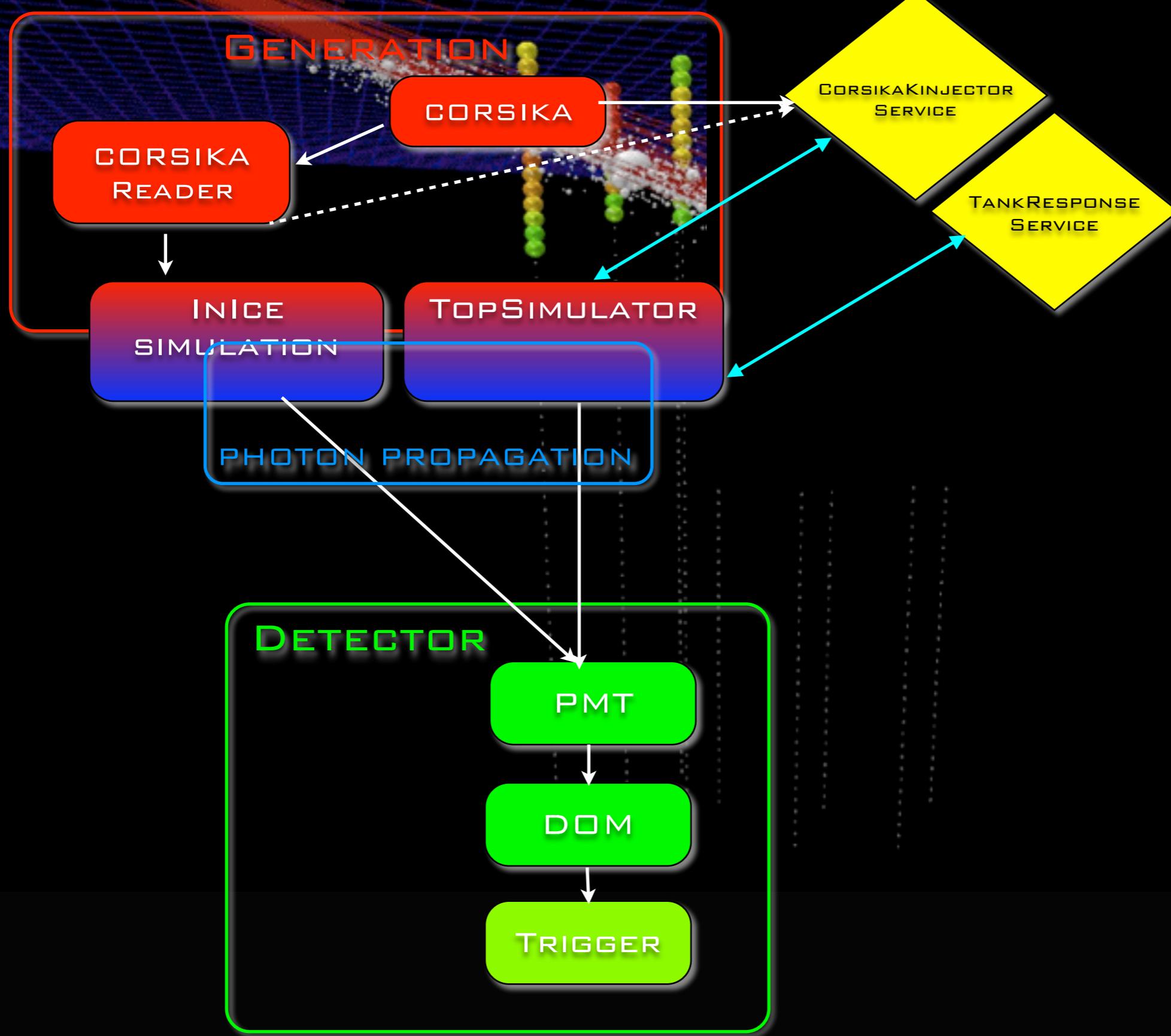


DETECTOR



i³SIM

simulaton chain (IT)



Generators

- ▶ Cosmic-ray Air Showers:
 - ▶ **CORSIKA** (FORTRAN stand-alone)
 - ▶ **corsika-reader**: IceTray reader for standard format
 - ▶ **CorsikaInjectorService** (IceTop)
- ▶ Muons:
 - ▶ **MuonGun**: parametrization of flux of atm. muons under the ice.
- ▶ Neutrinos:
 - ▶ **neutrino-generator**: injects neutrinos, propagates them through Earth, forces interaction in detector volume.
 - ▶ **genie-icetray**: detailed simulation of neutrino interactions with GENIE.
(Used for low-energy simulations)
 - ▶ **LeptonInjector / NuFSGen** (not yet available): weighted leptons+weights to account for flux models, interaction models, in-earth propagation, etc.

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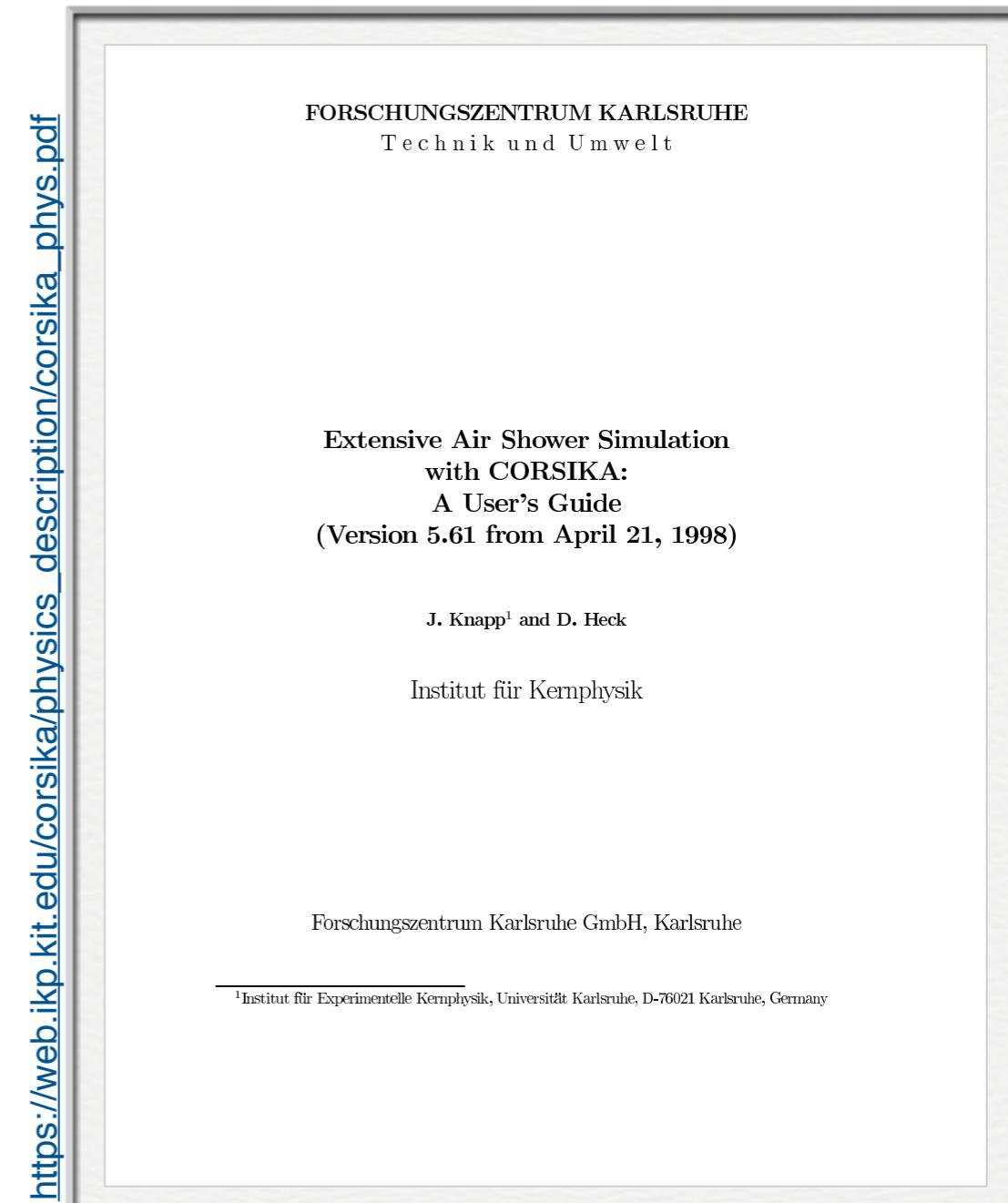
Generators (cont.)

- ▶ Other:
 - ▶ **wimpsim-reader**: IceTray interface for WimpSim (FORTRAN stand-alone)

Generators : CORSIKA

(**C**Osmic **R**ay **S**Imulations for **K**Ascade)

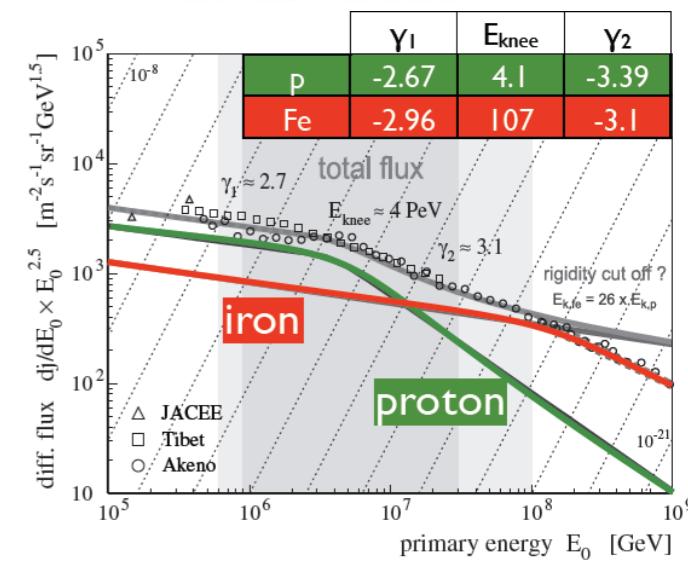
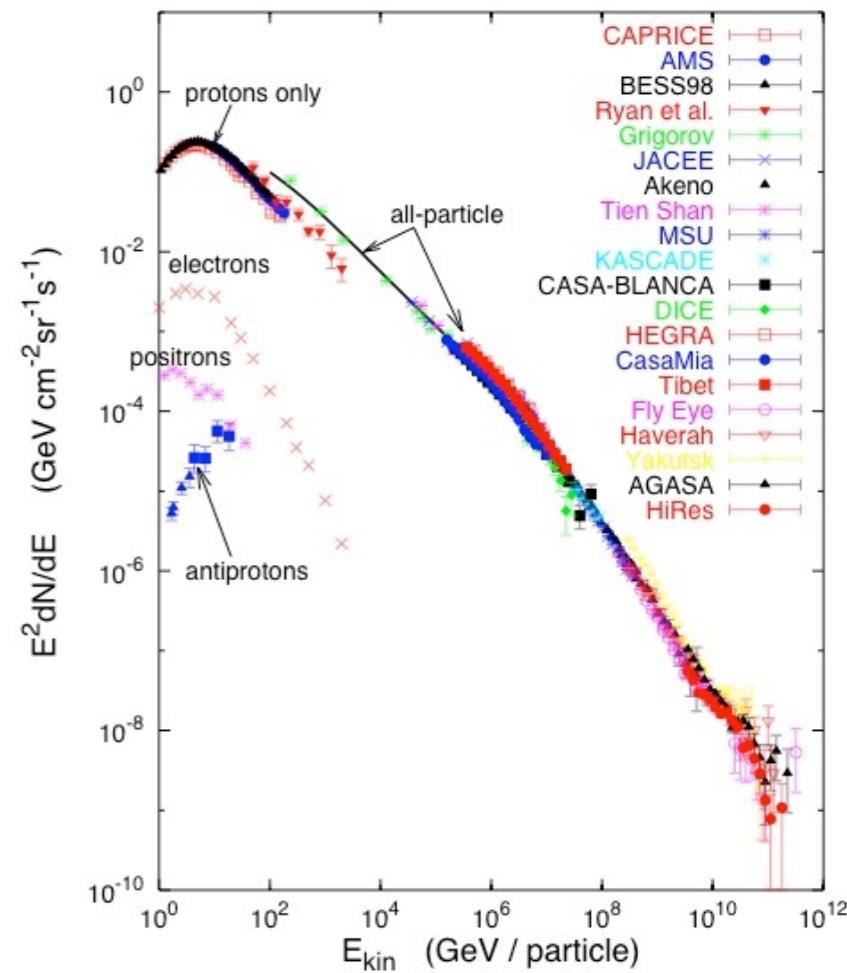
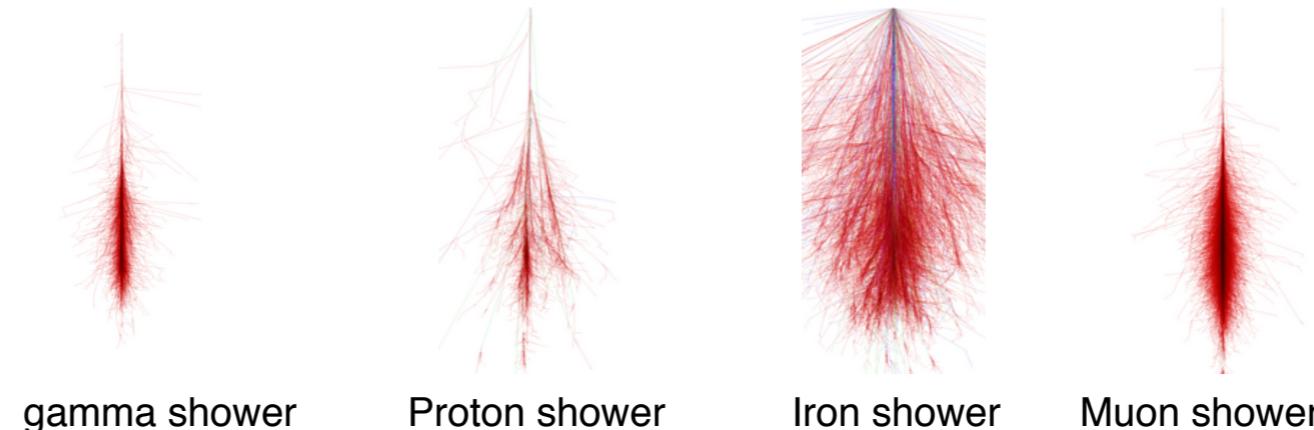
- Particles are tracked through the atmosphere until they undergo reactions with the air nuclei or - in the case of unstable secondaries - decay.
- The hadronic interactions at high energies may be described by several reaction models alternatively:
 - *VENUS*, *QGSJET*, and *DPMJET* (Gribov-Regge theory),
 - *SIBYLL* (minijet model).
 - *neXus*, *EPOS* (combination of *QGSJET* and *VENUS*).
 - *HDPM* (Dual Parton Model).
- Hadronic interactions at lower energies:
 - *GHEISHA*, *FLUKA* , or *UrQMD* models.
- For electromagnetic interactions
 - Tailored version of *EGS4*.
 - Analytical *NKG* formulas.



Generators : CORSIKA

(COsmic Ray SImulations for KAscade)

- ▶ weighted events : artificially flat spectrum
 - ▶ better livetime efficiency @ 10 TeV but poor efficiency @ TeV
 - ▶ energy-targeted generation of (H,He,CNO,Mg,Fe) with $E^{-1(2)}$



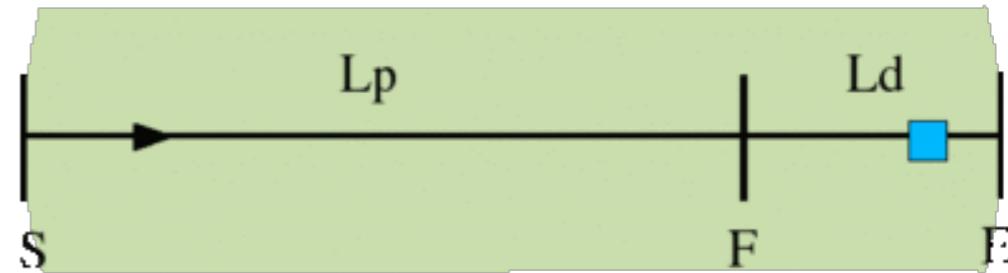
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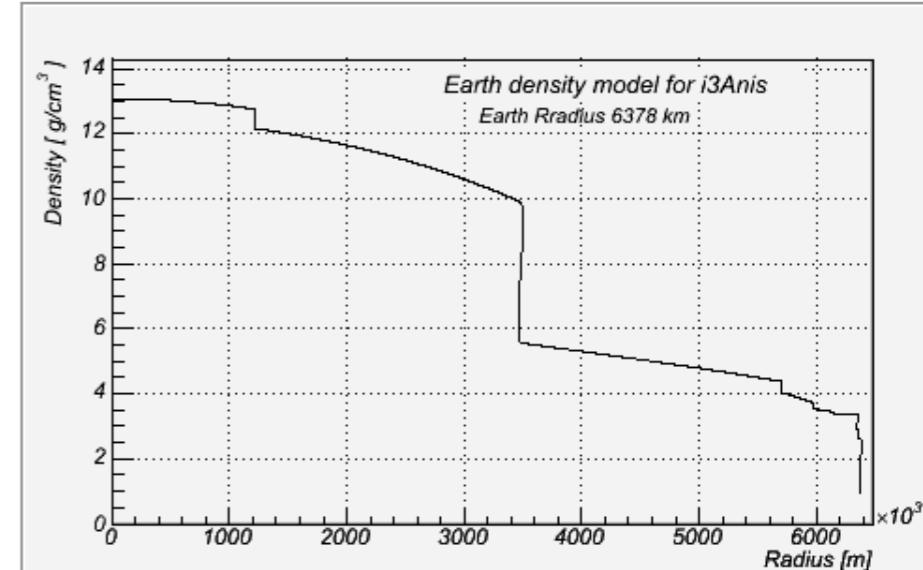
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neutrino-generator



1. Calculate total path length inside the Earth using injected neutrino geometry.
 - a. Separate the total path length into propagation area (SF) and
 - b. detection volume (FE).
2. Define a step length $dx[m]$ using propagation area and step number.
3. For each step:
 - a. Calculate a column depth and Earth's density at the step point.
 - b. Calculate a total cross section at the step point.
 - c. Calculate a probability that the injected neutrino interacts within the step. Try Monte-Carlo, and decide whether an interaction happened within the step.
 - d. If interaction occurred: choose interaction randomly.
 - i. If CC-interaction is selected with injection particle NuMu or NuE, break (event is killed).
 - ii. else, generate secondaries and continue to next step.
 - e. If nothing happens, continue next step.
4. Finish propagation when injected neutrino + secondaries reach surface of detection volume (point F), then process a weighted interaction.



- produce a $E^{-\gamma}$ ν_μ , ν_e , ν_τ with
 - ▶ PRELIM Earth's density model
- A log-log plot of cross sections σ [mb] versus $\log_{10}(\text{Energy [GeV]})$. The y-axis ranges from 10^{-12} to 10^4 , and the x-axis ranges from 0 to 12. Four curves are shown: CC nu + N (black), CC nubar + N (red), NC nu + N (green), and NC nubar + N (blue). All curves show an increase in cross section with energy, with CC interactions being significantly higher than NC interactions.

 - ▶ parton distribution functions
 - ▶ prop & interaction of neutrinos into a weight

neutrino-generator

- Calculates the propagation probability (i.e. that the neutrino will reach the detector)
- It forces an interaction within a volume around the detector and computes the probability of this interaction

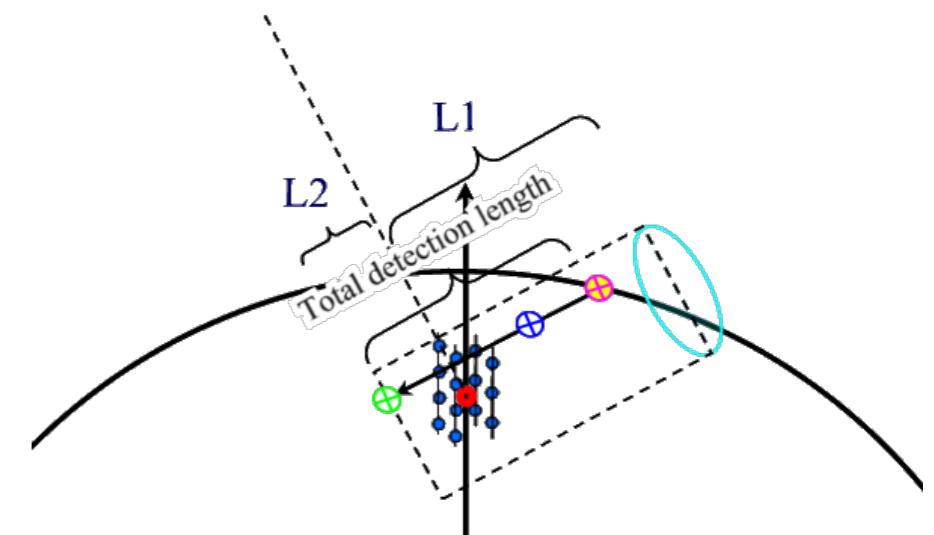
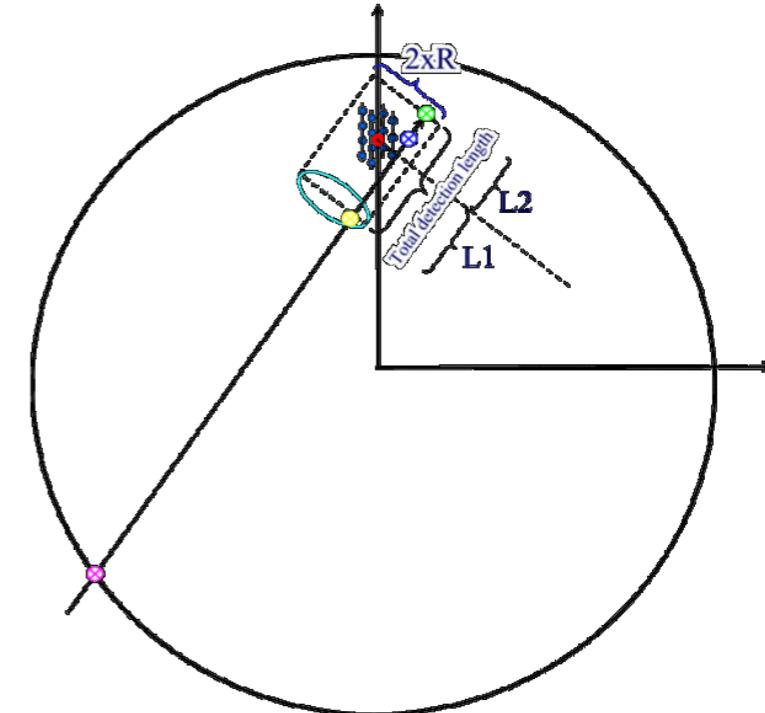
$$OneWeight = \left(\frac{P_{int}}{E^{-\gamma}} \right) \cdot \int_{E_{min}}^{E_{max}} E^{-\gamma} dE \cdot Area \cdot \Omega \cdot T [GeV \cdot cm^2 \cdot sec \cdot sr]$$

where P_{int} = $TotalInteractionProbabilityWeight$, $E^{-\gamma}$ is the neutrino generation energy spectrum shape, E_{min} and E_{max} is the minimum and maximum generation energy of neutrinos, $Area$ is the generation surface, Ω the generation solid angle and $T = 1sec$ is the timescale.

- The weight corresponding to a given theoretically motivated neutrino flux is

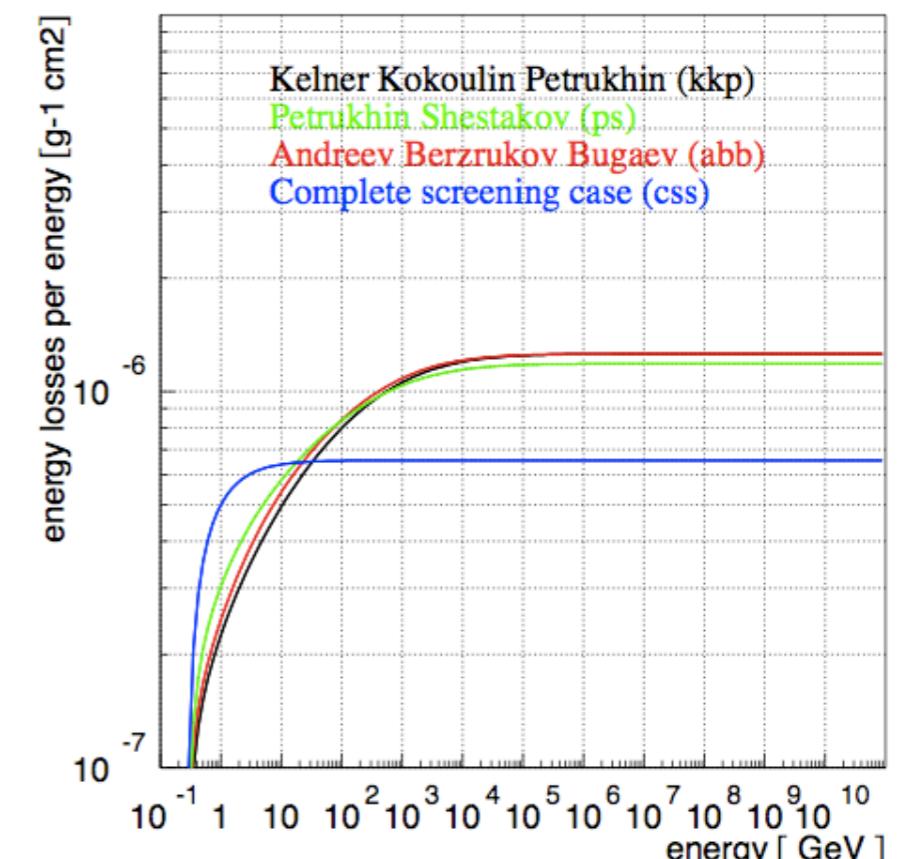
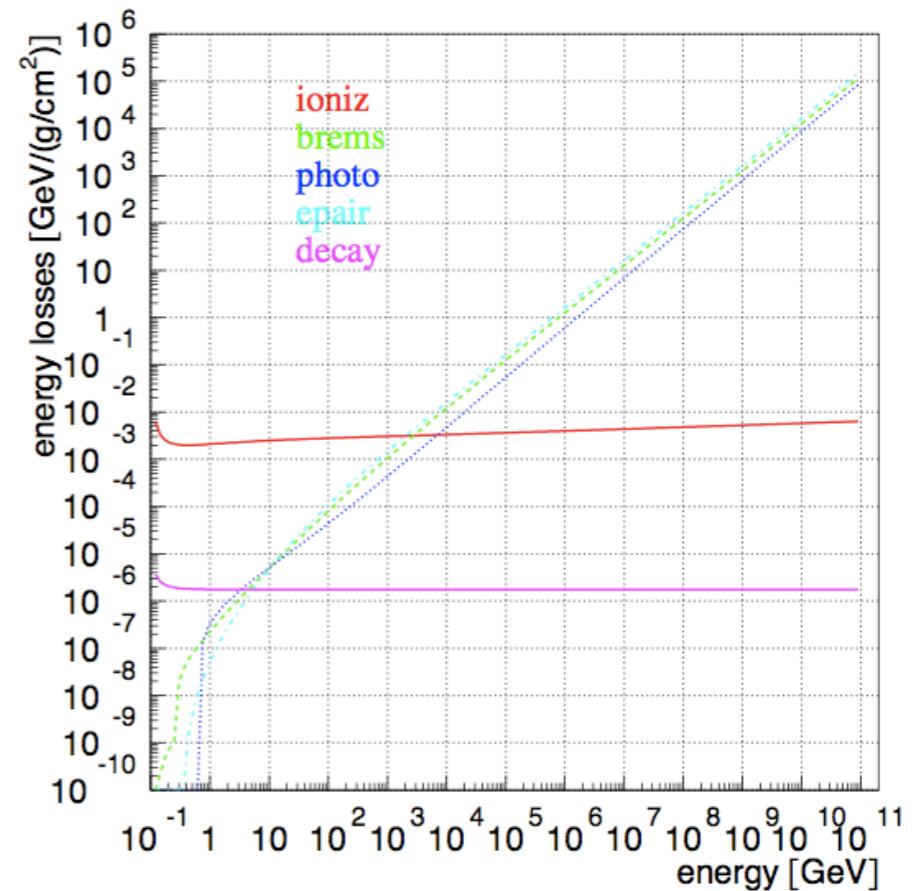
$$w_i = \frac{OneWeight_i}{NEvents} \times \frac{d\Phi_\nu(E_\nu)}{dE_\nu}$$

- For more details on how to use OneWeight see:



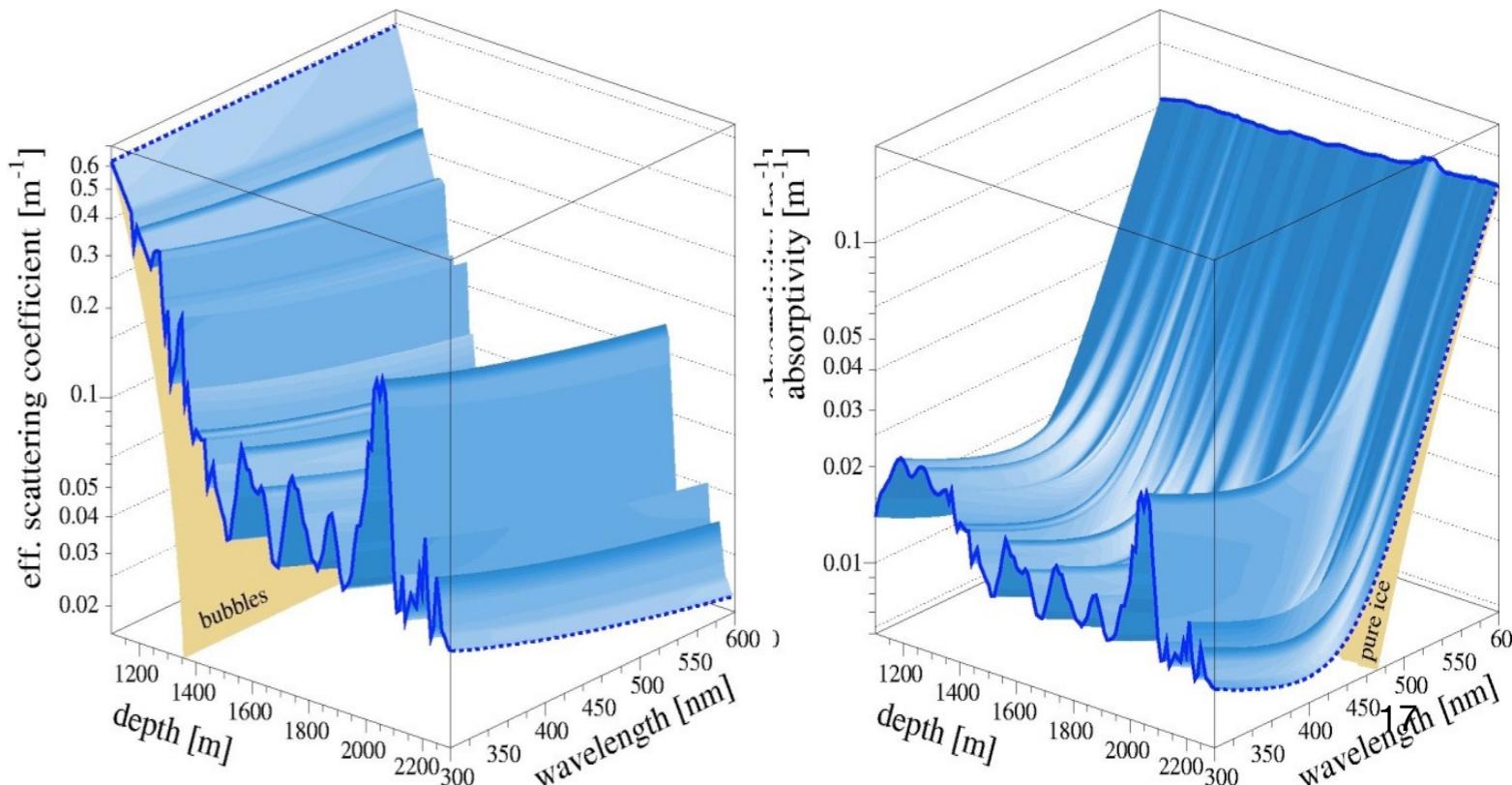
Lepton propagation

- ▶ PROPOSAL: parametrized interactions with the medium.
 - ▶ Stochastic energy losses include:
 - ▶ ionization
 - ▶ electron-pair production
 - ▶ bremsstrahlung
 - ▶ photo-nuclear interaction
 - ▶ decay
- ▶ GEANT4: Detailed particle propagation in media.
 - ▶ 3rd-party G4 library used by CLSim to propagate leptons for low-energy simulations (CPU-intensive).



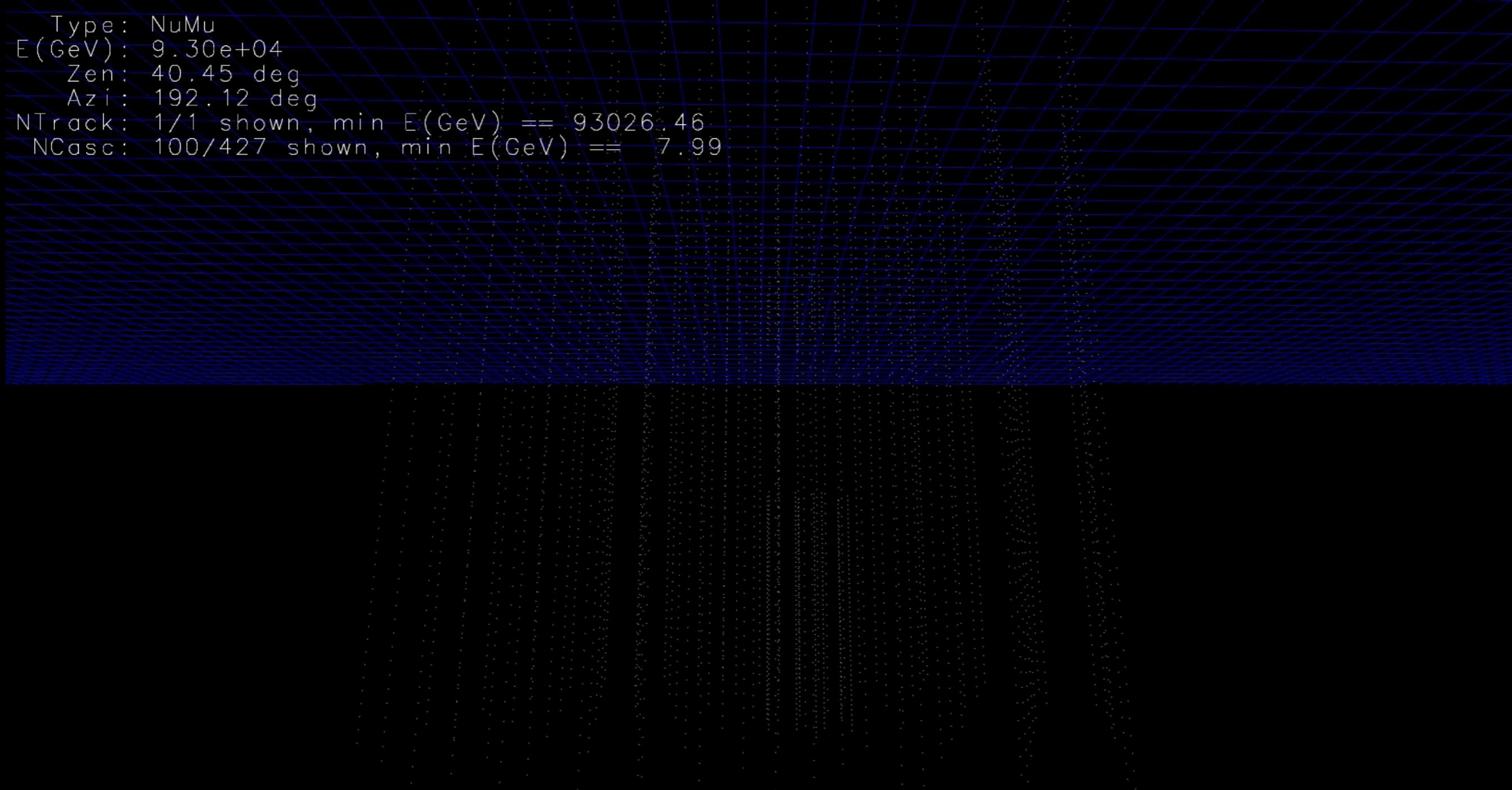
Photon Propagation

- μ energy lost + cascades \rightarrow photons \rightarrow p.e.
 - Photon propagation : ice properties + PMT response + DOM glass/gel
 - Pre-generated lookup splined table :
 - I3PhotonicsHitMaker
 - Amplitude and time distribution
 - Direct photon tracking
 - PPC
 - CLSim
 - Hybrid photon tracking
 - HitMaker + CLSim



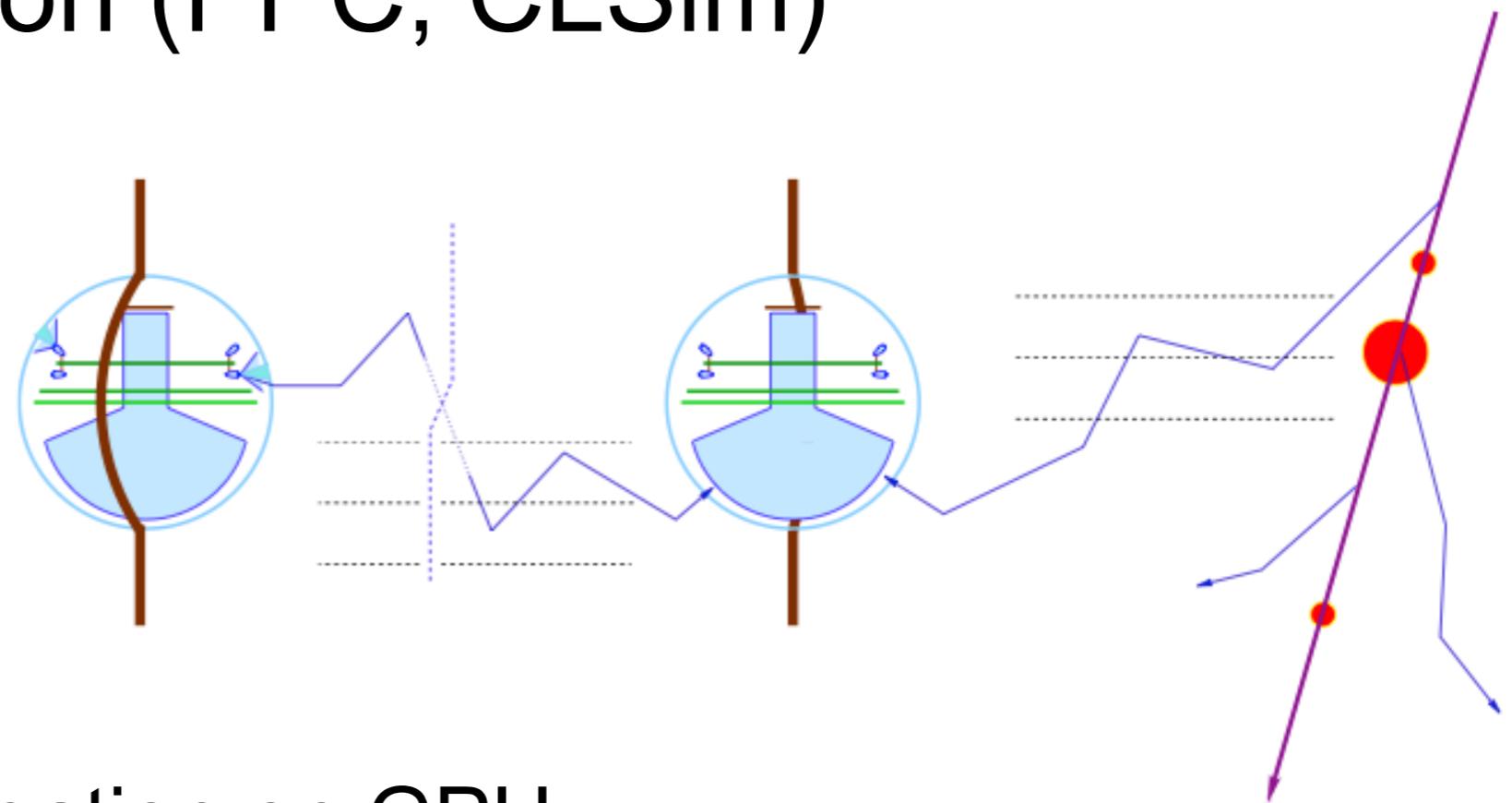
Photon Propagation (PPC, CLSim)

Type: NuMu
E(GeV): 9.30e+04
Zen: 40.45 deg
Azi: 192.12 deg
NTrack: 1/1 shown, min E(GeV) == 93026.46
NCasc: 100/427 shown, min E(GeV) == 7.99

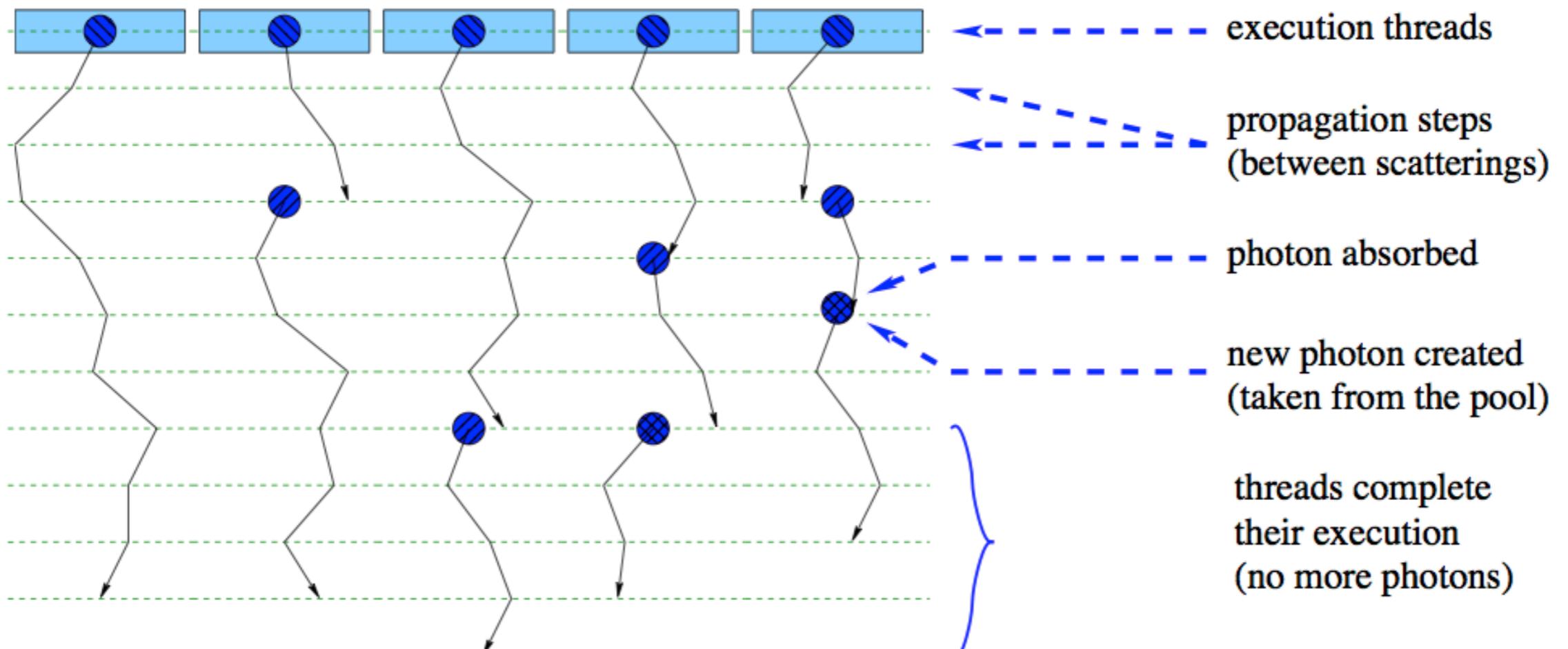


http://icecube.wisc.edu/~ckopper/muon_with_photons.mov

Photon Propagation (PPC, CLSim)



Direct photon propagation on GPU



Polyplopia

(from gr., πολύς - polús, “many”, and ὄψ-ops , "vision")

Coincident atmospheric shower events in IceCube

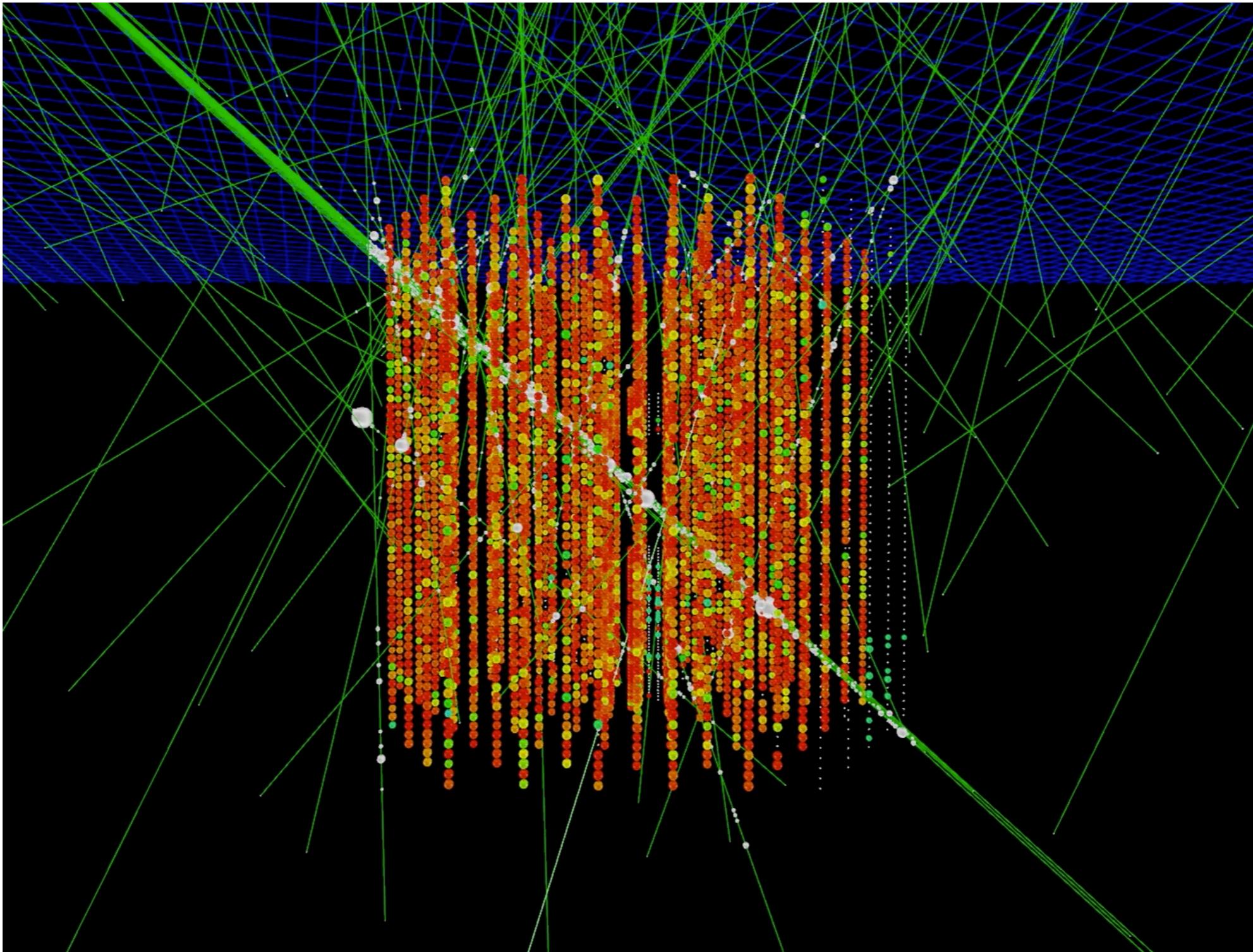


- **polyplopia::PoissonMerger**
 - Injects background event read from a separate file on top of primary events in the chain by sampling from a Poisson distribution over a time window Δt .
 - Also makes use of a *CoincidentEventService* that could be drop-in replaced with other event services such as a MuonGun-based service.
 - Writes a separate I3MCTree with background particles.
 - Writes a combined I3MCPE map for signal and background.
- **polyplopia::MPHitFilter**
 - Removes events that don't produce light in the detector and removes branches of I3MCTrees whose particles don't produce enough PEs in the detector,
 - Reduces the storage requirements.
- It is then up to the trigger-sim to split up Q-frames into P-frames events based on triggers.

Polyplopia

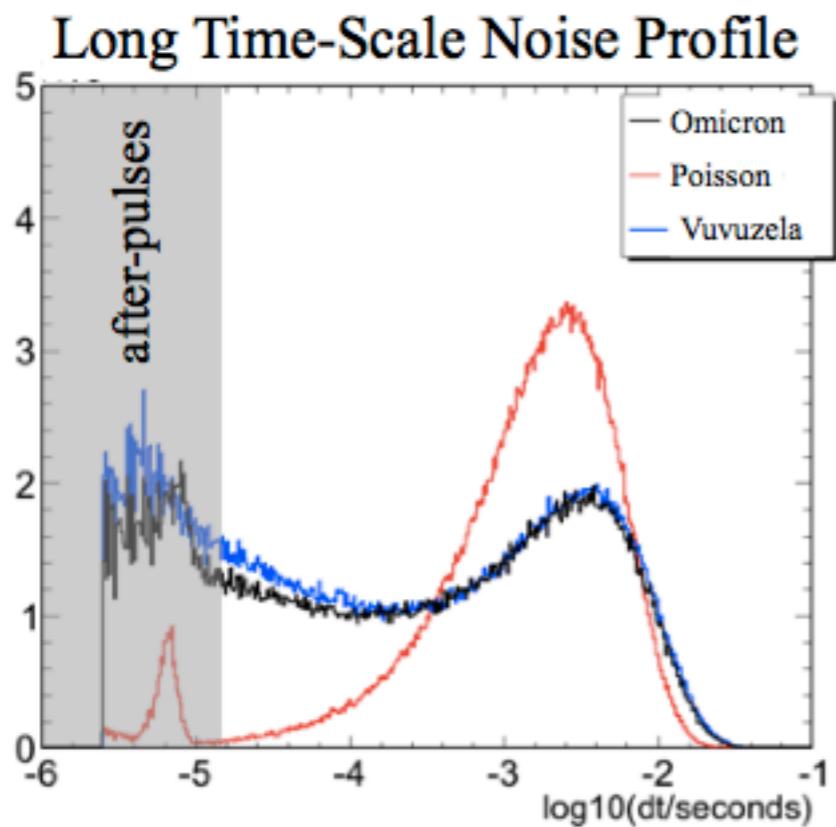
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Coincident atmospheric shower events in IceCube

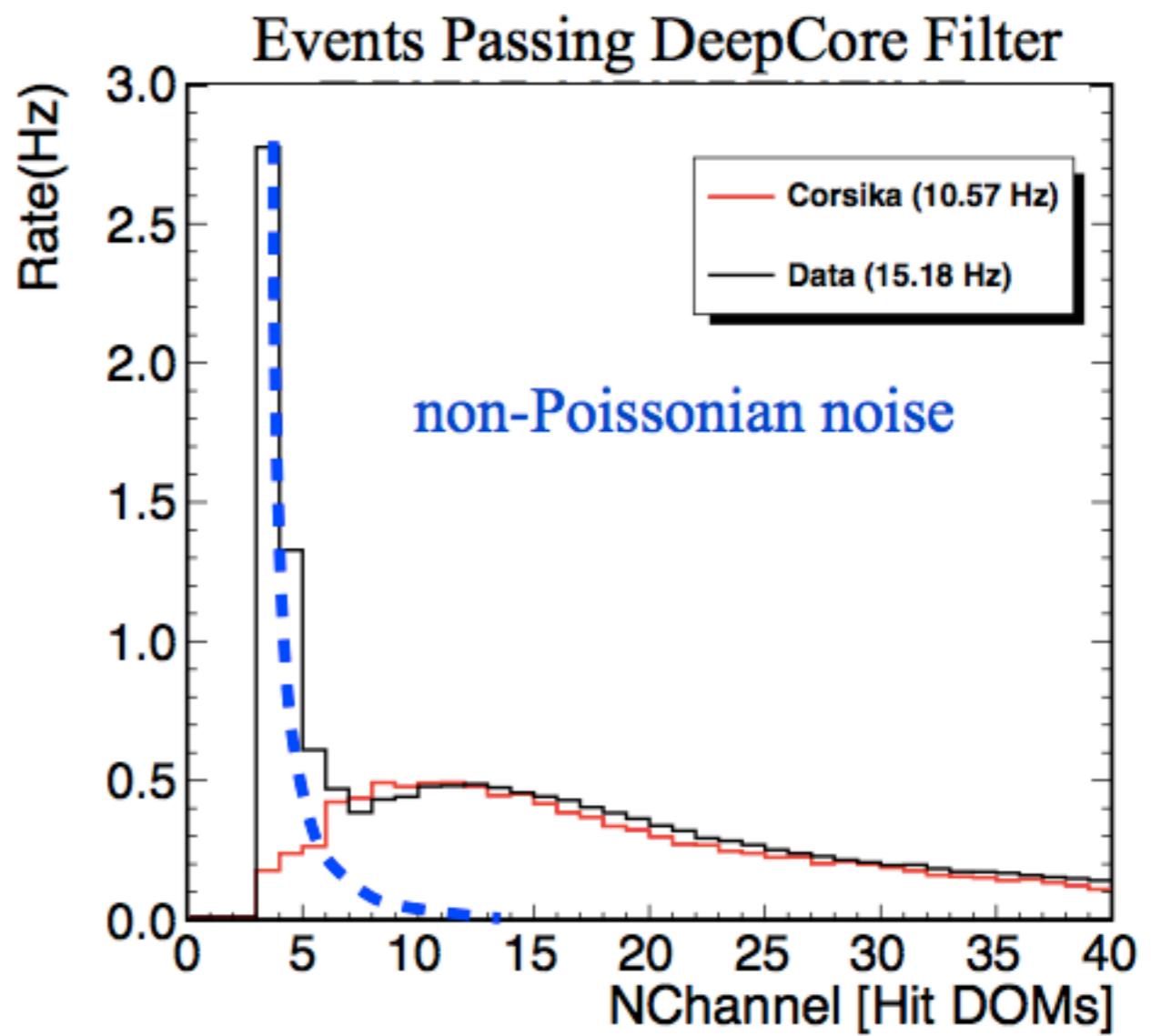


Noise Generation

Previous simulation used simplified Poissonian model. Vuvuzela uses exponential for **thermal and radioactive decays** and log-normal for **scintillation**.



*Courtesy of M.Larson (U.Alabama)



*Courtesy of J.Koskinen (PSU)

Noise Generation

→ (MCPes)

Noise Model

Thermal Noise (~few Hz)
[Poisson process]

~ ms Timescales

+

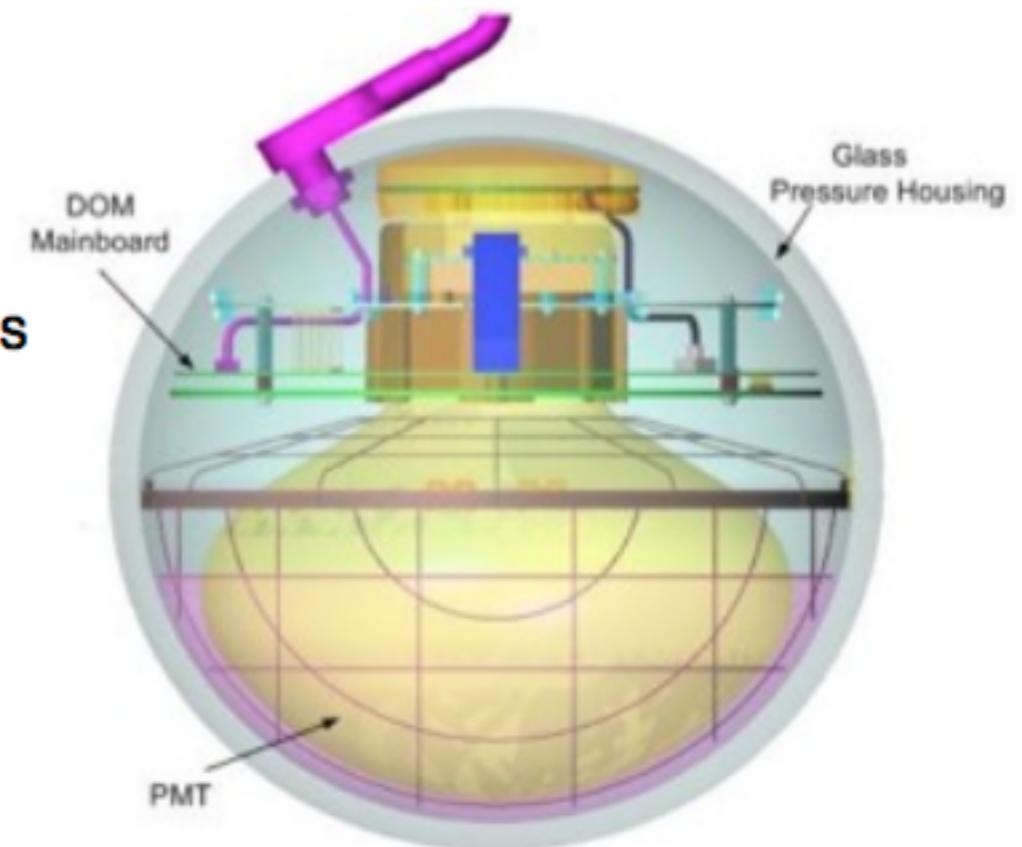
Radioactive Decay in Glass
[Poisson process]

~ ms Timescales

Energy deposited in glass

Glass scintillates/fluoresces
over long timescale
[Log-normal]

≤ 500 μ s Timescales



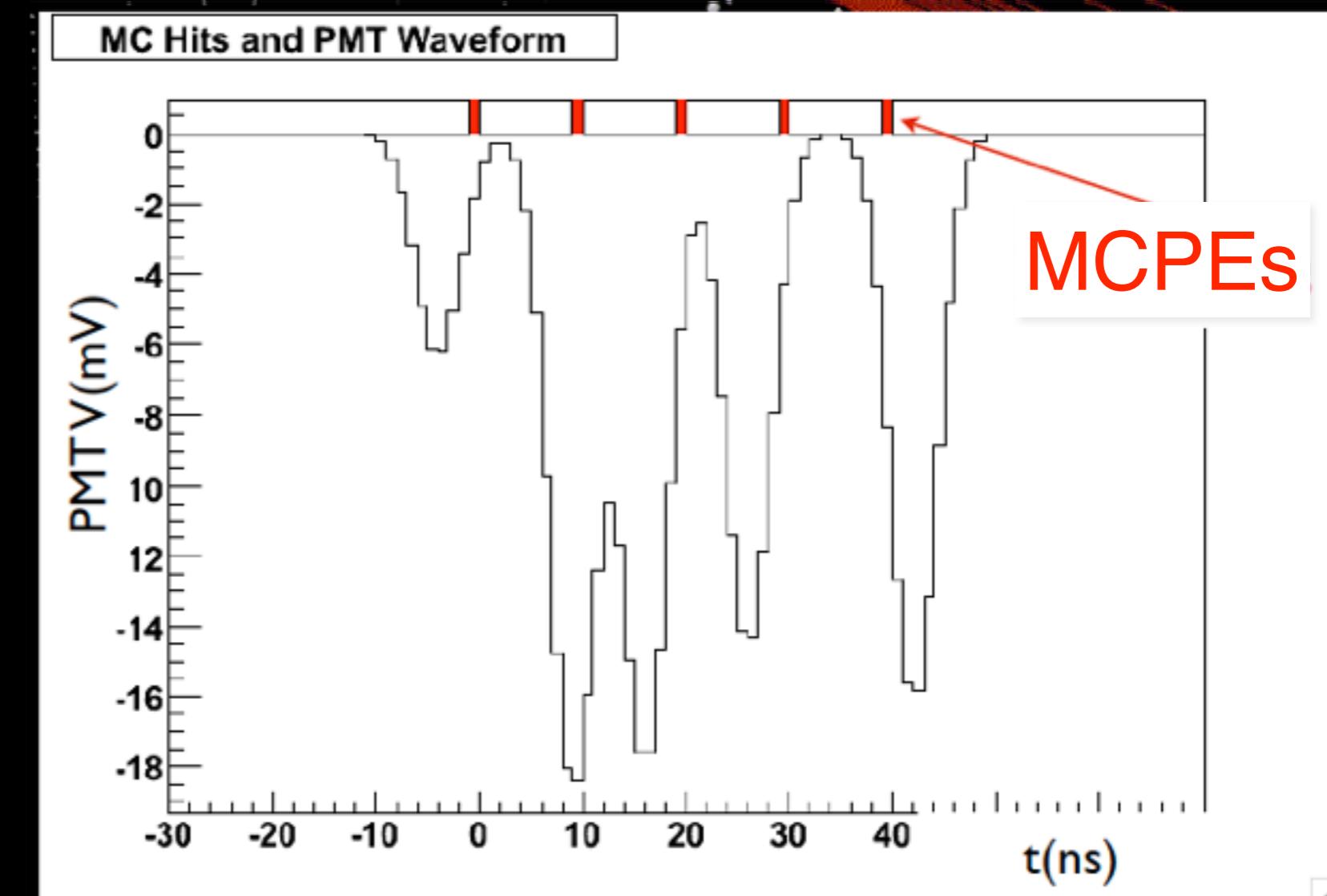
DOMLauncher:: PMTResponseSimulator

PMT

Generates PMT Waveform

From distribution of
(combined) MCPEs.

Outputs I3MCPulseSeries
for each DOM.

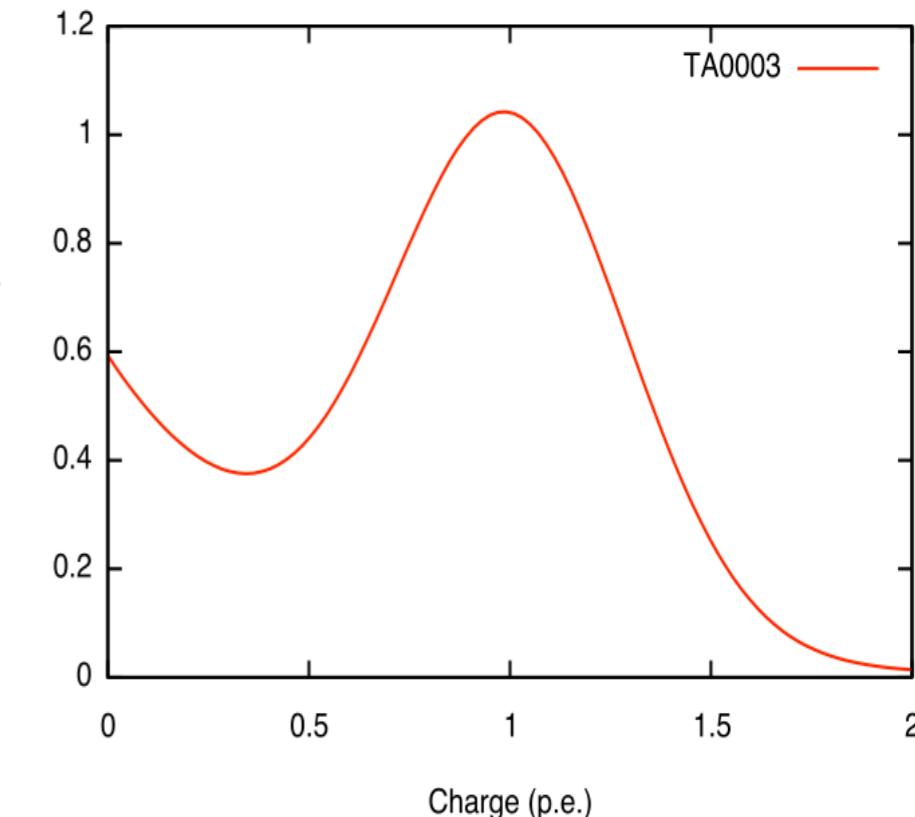


PMTResponseSimulator

Input: I3MCPEs
Output: I3MCPulses

Processing MCPEs :

- Give each MCPE a weight → corresponding to the pulse charge that photon would yield.
- Generate prepulses, late pulses and after pulses.
- Apply time jitter.
- Simulate the effect of saturation.



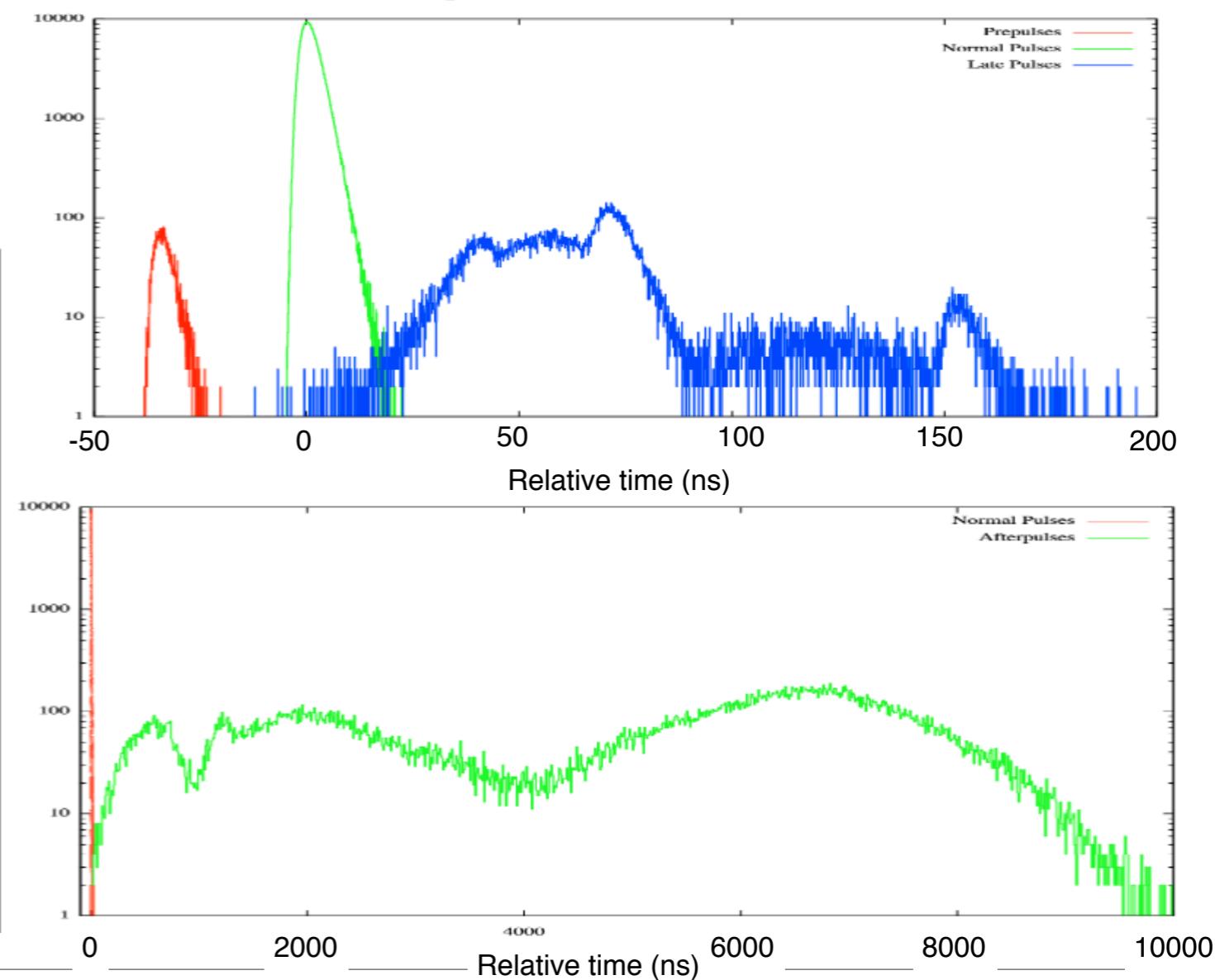
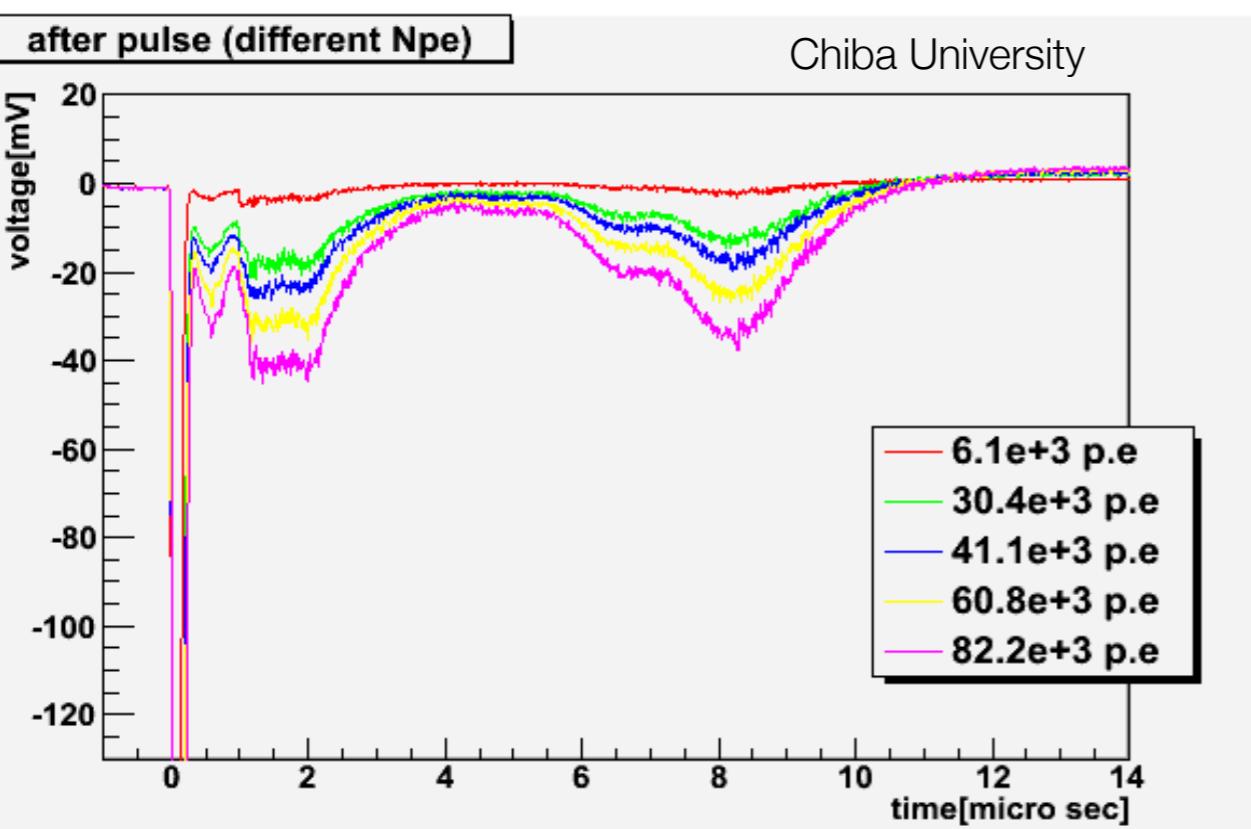
Weights from SPE Charge Distribution

PAL pulses

Pre-pulses: photoelectrons ejected from the first dynode,

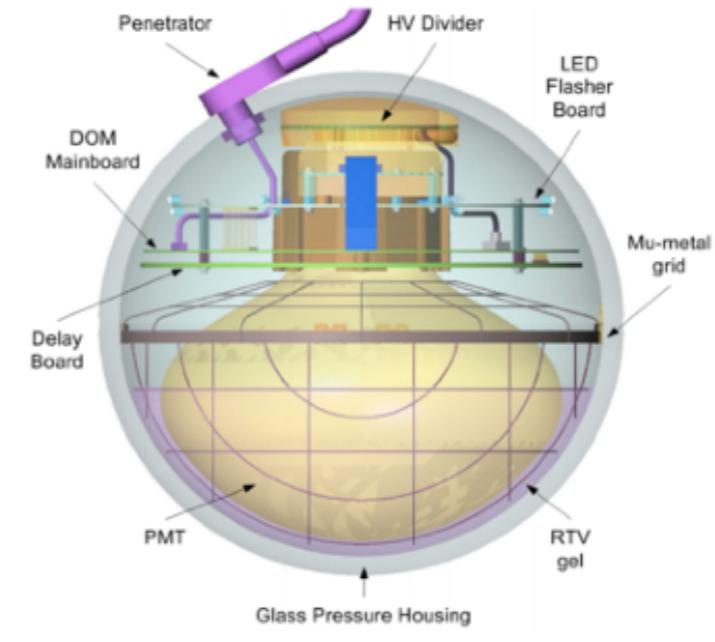
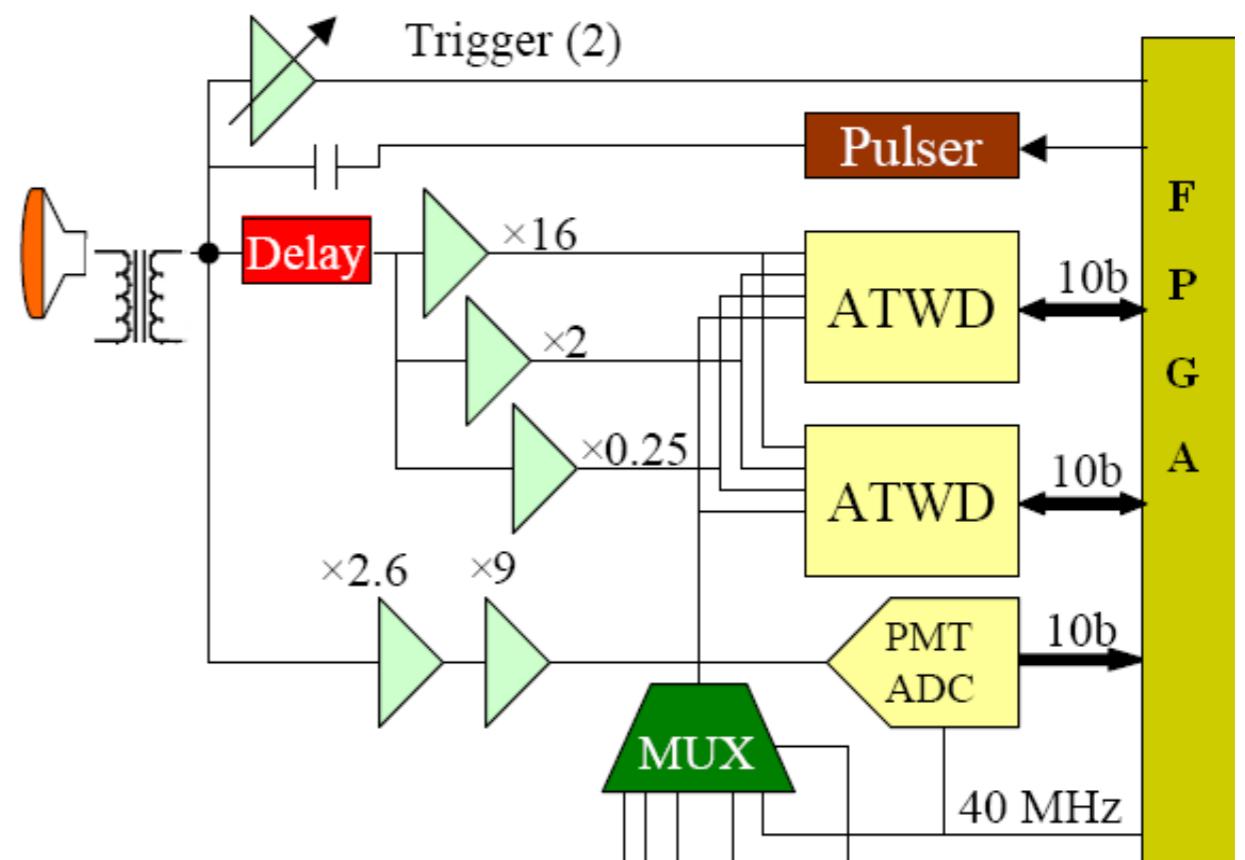
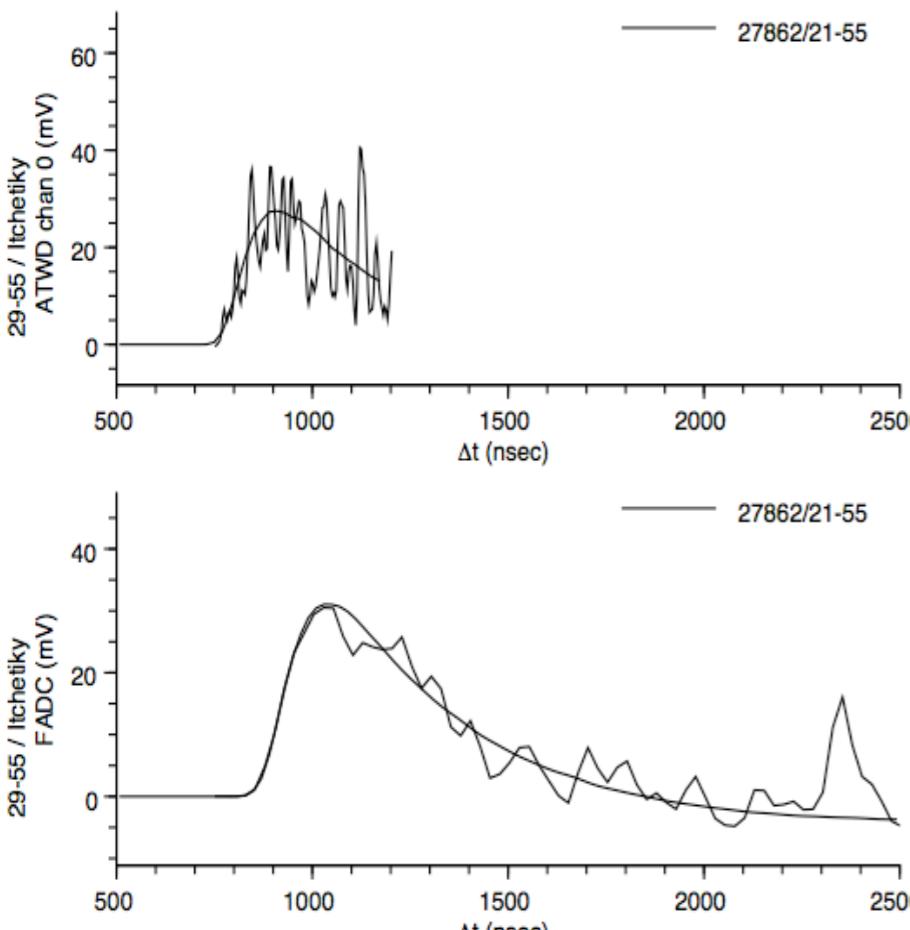
Late pulses: electrons backscatter from dynode to cathode.

After-pulses: ionization of residual gases by electrons accelerated in the space between dynode.



DOMLauncher: DOM mother board simulation

- Discriminator
- LC-logic
- Digitization
- Simulated effects
 - Electronic noise in the digitizers
 - Beacon launches (CPU triggered launches)
 - The FPGA Clock phase
 - RAPcal time uncertainty



Trigger

trigger-sim

InIce/Icetop

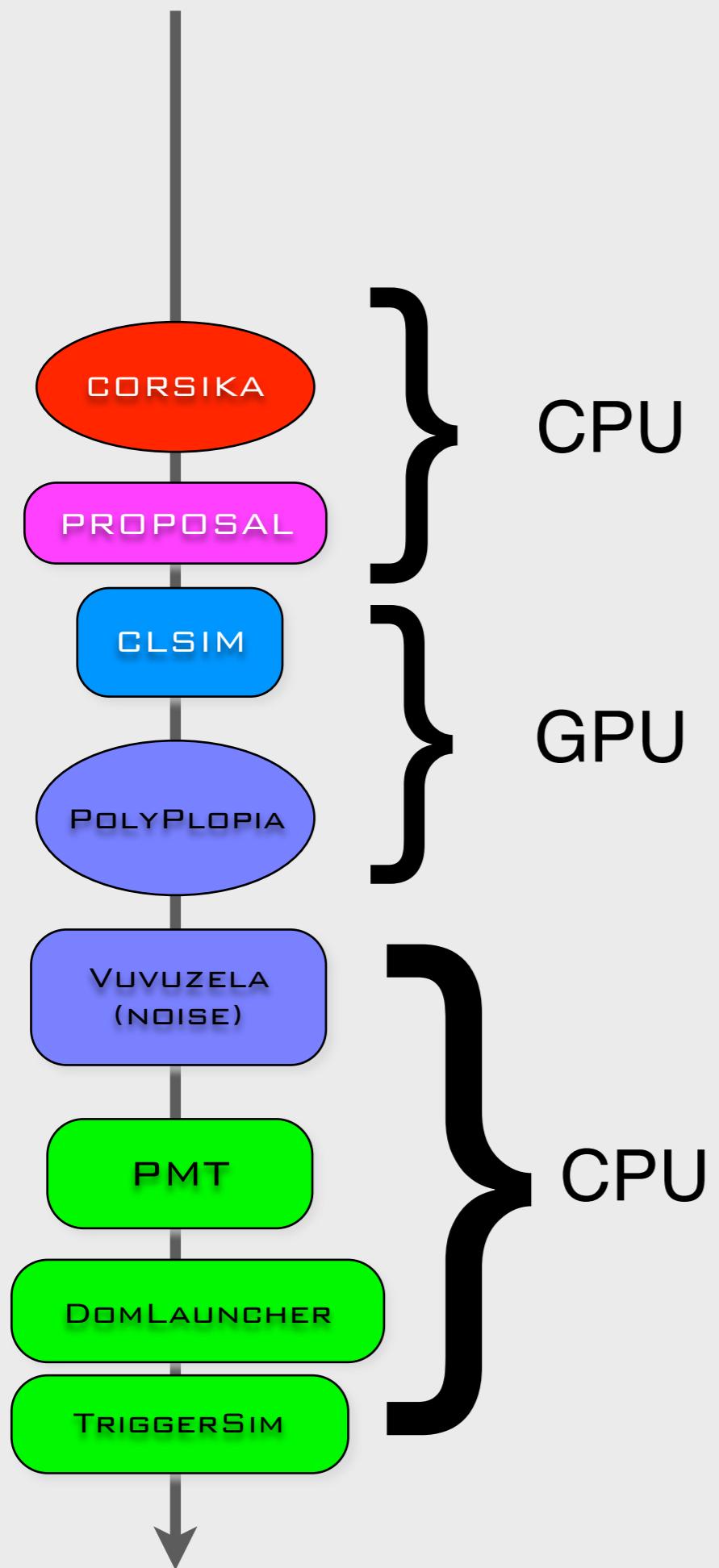
- Simple Majority Trigger
 - SMT8
 - SMT3 (DC)
- Cluster Trigger
- Volume Trigger
- SlowMP

Global Trigger
I3TimeShifter

Trigger Types

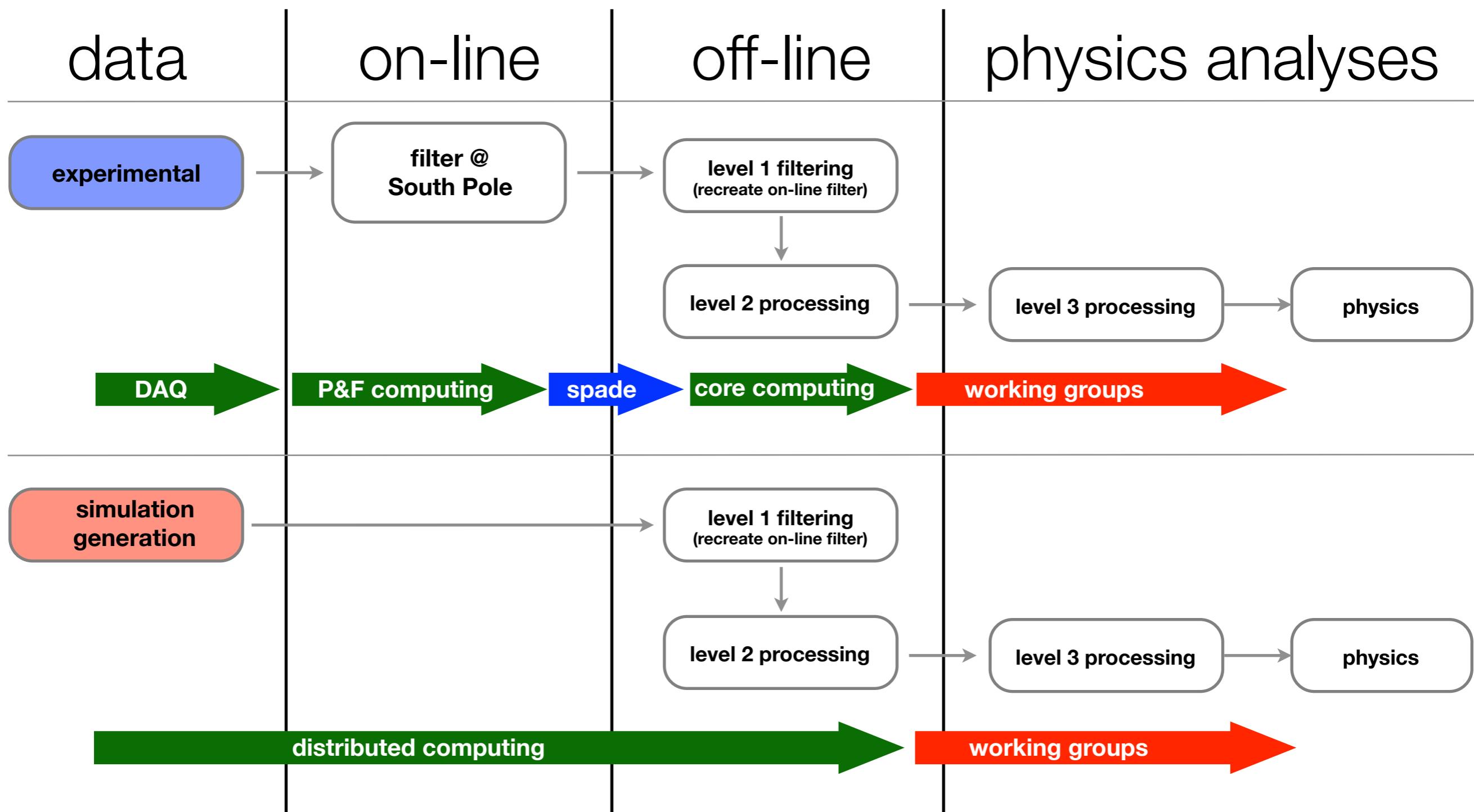
- **Simple Multiplicity Trigger (SMT)**
 - N HLC hits or more in a time window
 - Example: InIce SMT8 with $N_{\text{hits}} \geq 8$ in $5 \mu\text{s}$
 - readout window around this captures early and late hits ($-4 \mu\text{s}, +6 \mu\text{s}$)
- **String trigger** (a.k.a. Cluster trigger in DAQ-land)
 - N HLC hits out of M DOMs on a string in a time window
 - Example: 5 hits from a run of 7 adjacent DOMs in a time window of 1500 ns
- **Volume trigger** (a.k.a Cylinder trigger in DAQ-land)
 - simple majority of HLC hits (SMT4) with volume element including one layer of strings around a center string
 - cylinder height is 5 DOM-layers (2 up and down from the selected DOM).
- **Slow Particle trigger (SLOP)**
 - slow-moving hits along a track
 - lengths of the order of $500 \mu\text{s}$ and extending up to milliseconds
- ~~Fixed Rate trigger, Minimum Bias trigger, Calibration trigger~~

J. Kelley - DAQ

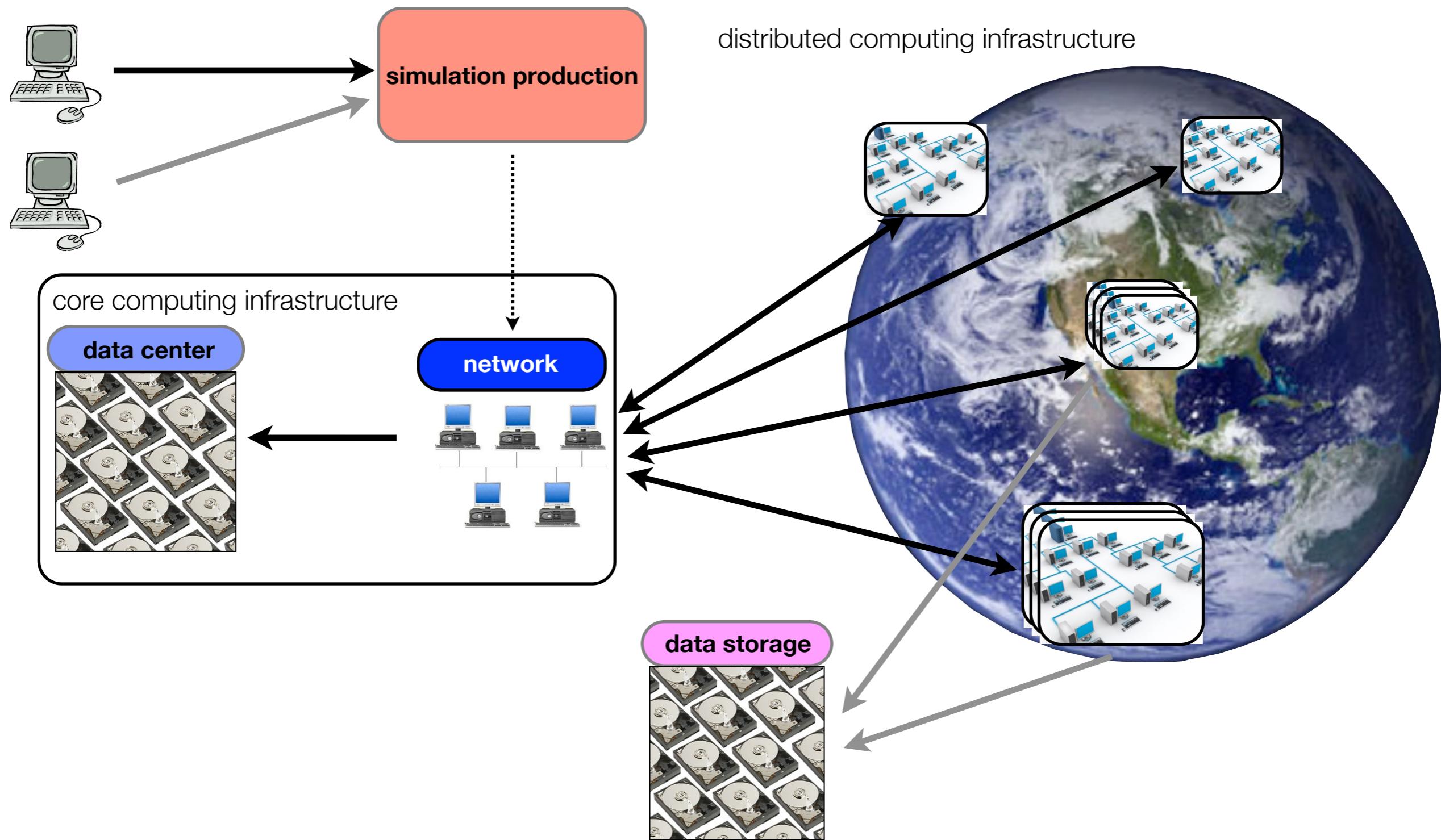


i3SIM

flow of experimental and simulation data



Simulation Production



This project is a collection of scripts, tray segments and IceProd modules used in simulation production. The aim is to provide a central place with standard segments for running simulation in both production and privately.

- **Tray Segments:** IceTray meta-modules that contain several I3Modules with default parameters.
- **IceProd modules:** basic wrappers around tray segments that provide an interface for IceProd.
- **Scripts:** collection of python scripts used in simulation production
- **Examples:** The directory simprod-scripts/resources/examples contains a collection of example scripts for running IPModules
- **Tests:** are run on the build-bots to check that the different parts of the simulation are not broken with each commit to the software repository.

Simprod-Scripts

[http://software.icecube.wisc.edu/documentation/projects/
simprod_scripts/index.html](http://software.icecube.wisc.edu/documentation/projects/simprod_scripts/index.html)

IceProd Modules

\$I3_SRC/simprod-scripts/python/modules

Corsika

MuonGunGenerator

NuGen

GENIE

CLSim

PPC

Detectors

IceTop

Simprod-Scripts

[http://software.icecube.wisc.edu/documentation/projects/
simprod_scripts/index.html](http://software.icecube.wisc.edu/documentation/projects/simprod_scripts/index.html)

Tray Segments

\$I3_SRC/simprod-scripts/python/segments

Calibration

GenerateIceTopShowers

DetectorSim

GenerateNeutrinos

GenerateAirShowers

GenerateNoiseTriggers

GenerateCosmicRayMuons

HybridPhotonicsCLSim

GenerateFlashers

Polyplopia

GenerateIceTopShowers

PropagateMuons

simprod-scripts

http://software.icecube.wisc.edu/documentation/projects/simprod_scripts/examples/fullSimulation.html

Examples: \$I3_SRC/simprod-scripts/resources/examples/fullSimulation.py

```
./fullSimulation.py -n 10 --seed=43682 --datasetnumber=1 --runnumber=1 --no-hybrid --icemodel=SpiceLea \
--detector=IC86 --unshadowed-fraction=0.99 --flavor=NuTau --outfile=taus.i3 --skip-calibration \
--from-energy=1000 --to-energy=10000000 --include-gcd-in-outfile
```

-n 10 Simulate 10 events (the output file might contain fewer events because not everything will trigger).

--seed=43682 A random number generator seed

--datasetnumber=1 --runnumber=1 A combination of “dataset” and “run” numbers. You can keep the same seed for all of them, each one will give a distinct independent set of events. (So you can think of the combination of seed, dataset and run as the actual random number seed. Or, put another way, there is no need to have different seeds for different runs.)

--no-hybrid Do not use hybrid simulation mode, i.e. propagate everything using direct simulation. There will be no photon tables and things like ice anisotropy/SpiceLea will work.

--icemodel=SpiceLea The other two models are “Spice1” and “SpiceMie”. Including things like WHAM would be trivial if you need it.

--detector=IC86 This will select a GCD file from \$I3_PORTS automatically. Currently works for IC86 and IC79.

--unshadowed-fraction=0.99 This is the “DOMEfficiency”, currently named like this for compatibility with other tools that use the same name.

--flavor=NuTau You can set this also to “NuE” and “NuMu”.

--outfile=... The name of your final .i3 file you want to generate.

--from-energy=1000 --to-energy=10000000 The energy range in GeV. Currently not implemented for MuonGun which uses fixed 36 energy ranges. Should be fixed soon.

--include-gcd-in-outfile Use this option if you want to generate output files with GCD frames in them. It makes them much easier to use, but of course wastes some space..

simprod-scripts

Scripts: \$I3_SRC/simprod-scripts/resources/scripts
(run the individual pieces as broken down by production tasks)

```
$ python nugen.py -h

        Usage: nugen.py [options]

        Options:
          -h, --help            show this help message and exit
          --no-execute          boolean condition to execute
          --gcdfile=GCDFILE    GeoCalibDetStatus filename
          --outputfile=OUTPUTFILE
                                Output filename
          --summaryfile=SUMMARYFILE
                                XMLSummary filename
          --mjd=MJD             MJD for the GCD file
          --seed=RNGSEED         RNG seed
          --procnum=RNGSTREAM
```

...

simprod-scripts

Exercise: Running scripts:

```
icecube@M16:~$ ssh submitter
[submitter]$
[submitter]$ condor_submit interactive_gpu.condor -interactive
Submitting job(s).
1 job(s) submitted to cluster 120263704.
Waiting for job to start...
Welcome to slot1@gtx-00.icecube.wisc.edu!

[gtx-00]$ cd ${CONDOR_SCRATCH_DIR}
[gtx-00]$ /cvmfs/icecube.opensciencegrid.org/py2-v2/icetray-env simulation/v05-01-01
*****
*                                         *
*          W E L C O M E   t o   I C E T R A Y   *
*          Version combo.trunk      r147415       *
*                                         *
*          You are welcome to visit our Web site   *
*          http://icecube.umd.edu                  *
*****
[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/nugen.py \
    --gcdfile=/cvmfs/icecube.opensciencegrid.org/data/GCD/GeoCalibDetectorStatus_2016.57531_V0.i3.gz \
    --outputfile=nutau.i3 --nevents=100 \
    --seed=123 --procnum=0 --nproc=1 \
    --FromEnergy=1e5 --ToEnergy=1e6 --NuFlavor=NuTau

[gtx-00]$ dataio-pyshovel nutau.i3
```

simprod-scripts

Exercise: Running scripts:

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icecube@M16:~$ ssh submitter
[submitter]$
[submitter]$ condor_submit interactive_gpu.condor -interactive
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Waiting for job to start...
Welcome to slot1@gtx-00.icecube.wisc.edu!

[gtx-00]$ cd ${CONDOR_SCRATCH_DIR}
[gtx-00]$ /cvmfs/icecube.opensciencegrid.org/py2-v2/icetray-env simulation/v05-01-01
*****
*                                         *
*          W E L C O M E   t o   I C E T R A Y           *
*          Version combo.trunk      r147415             *
*                                         *
*          You are welcome to visit our Web site        *
*          http://icecube.umd.edu                      *
*****
[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/nugen.py \
    --gcdfile=/cvmfs/icecube.opensciencegrid.org/data/GCD/GeoCalibDetectorStatus_2016.57531_V0.i3.gz \
    --outputfile=nutau.i3 --nevents=100 \
    --seed=123 --procnum=0 --nproc=1 \
    --FromEnergy=1e5 --ToEnergy=1e6 --NuFlavor=NuTau

[gtx-00]$ dataio-pyshovel nutau.i3
```

simprod-scripts

Exercise: Running scripts:

```
[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/clsim.py \
    --gcdfile=/cvmfs/icecube.opensciencegrid.org/data/GCD/
GeoCalibDetectorStatus_2016.57531_v0.i3.gz \
    --inputfilelist=nutau.i3 --outputfile=mcpes.i3 \
    --seed=123 --procnum=0 --nproc=1 --no-RunMPHitFilter \
    --UseGPU
```

```
[gtx-00]$ dataio-pyshovel mcPes.i3
```

```
[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/detector_ic86.py \
    --gcdfile=/cvmfs/icecube.opensciencegrid.org/data/GCD/
GeoCalibDetectorStatus_2016.57531_v0.i3.gz \
    --inputfile=mcPes.i3 --outputfile=det.i3 \
    --seed=123 --procnum=0 --nproc=1 --RunID=123
```

```
[gtx-00]$ dataio-pyshovel det.i3
```

simprod-scripts

Exercise: Running scripts:

```
[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/corsika.py \
          --nshowers=10000 --outputfile=corsika_bg.i3 --seed=1234 \
          --procnum=0 --nproc=1 --CORSIKAseed=123 --ranpri=2 \
          --corsikaVersion=v6960-5comp \
          --corsikaName=dcorsika --skipoptions=compress

[gtx-00]$ dataio-pyshovel corsika_bg.i3

[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/polyplopia.py \
          \
          --gcdfile=/cvmfs/icecube.opensciencegrid.org/data/GCD/
GeoCalibDetectorStatus_2016.57531_v0.i3.gz \
          --infile=mcpes.i3 --outfile=merged_pes.i3 \
          --seed=1234 --procnum=0 --nproc=1 --
backgroundfile=corsika_bg.i3 --mctype=NuTau

[gtx-00]$ python $I3_BUILD/simprod-scripts/resources/scripts/
detector_ic86.py \
          \
          --gcdfile=/cvmfs/icecube.opensciencegrid.org/data/GCD/
GeoCalibDetectorStatus_2016.57531_v0.i3.gz \
          --infile=merged_pes.i3 --outfile=det_wcoinc.i3 \
          --seed=123 --procnum=0 --nproc=1 --RunID=123

[gtx-00]$ dataio-pyshovel det_wcoinc.i3
```

Simprod-Scripts

frame objects

I3 Data Shovel			Press '?' for help
Name	Type	Bytes	
BackgroundI3MCPESeriesMap	I3Map<OMKey, vector<I3MCPE> >	41	
BackgroundI3MCTree	TreeBase::Tree<I3Particle, I3ParticleID, __g...	32	
BackgroundI3MCTree_preMu... BackgroundMMCTrackList	TreeBase::Tree<I3Particle, I3ParticleID, __g...	794	
BeaconLaunches	I3Vector<I3MMCTrack>	40	
I3EventHeader	I3Map<OMKey, vector<I3DOMLaunch> >	46	
I3MCPESeriesMap	I3EventHeader	99	
I3MCPESeriesMapWithoutNoise	I3Map<OMKey, vector<I3MCPE> >	435781	
I3MCPulseSeriesMap	I3Map<OMKey, vector<I3MCPulse> >	432958	
I3MCPulseSeriesMapParticl...	I3Map<OMKey, map<I3ParticleID, vector<unsign...	140992	
I3MCTree	TreeBase::Tree<I3Particle, I3ParticleID, __g...	72963	
I3MCTree_preMuonProp	TreeBase::Tree<I3Particle, I3ParticleID, __g...	4646	
I3MCWeightDict	TreeBase::Tree<I3Particle, I3ParticleID, __g...	422	
I3TriggerHierarchy	I3Map<string, double>	1236	
I3Triggers	I3Tree<I3Trigger>	792	
IceTopRawData	I3Tree<I3Trigger>	414	
InIceRawData	I3Map<OMKey, vector<I3DOMLaunch> >	46	
MMCTrackList	I3Map<OMKey, vector<I3DOMLaunch> >	56823	
NuGPrimary	I3Vector<I3MMCTrack>	2864	
PolyplopiaPrimary	I3Particle	150	
RNGState	I3Particle	150	
SignalI3MCPEs	I3SPRNGRandomServiceState	73	
TimeShift	I3Map<OMKey, vector<I3MCPE> >	432958	
	I3PODHolder<double>	36	
Key: 1/23	StartTime: 2015-05-18 17:04:01 UTC		
Frame: 3/12 (25%)	Duration: 24831.9 ns		
Stop: DAQ			
Run/Event: 123/0			
SubEvent: (n/a)			

More on simulation

1. http://wiki.icecube.wisc.edu/index.php/Simulation_Documentation_Wiki
2. Weighting tutorial: <https://grid.icecube.wisc.edu/simulation/weighting/example.htm>
3. <http://software.icecube.wisc.edu/documentation/>
4. http://wiki.icecube.wisc.edu/index.php/Simulation_Production
5. <http://grid.icecube.wisc.edu/simulation/table>
6. SLACK: #simulation