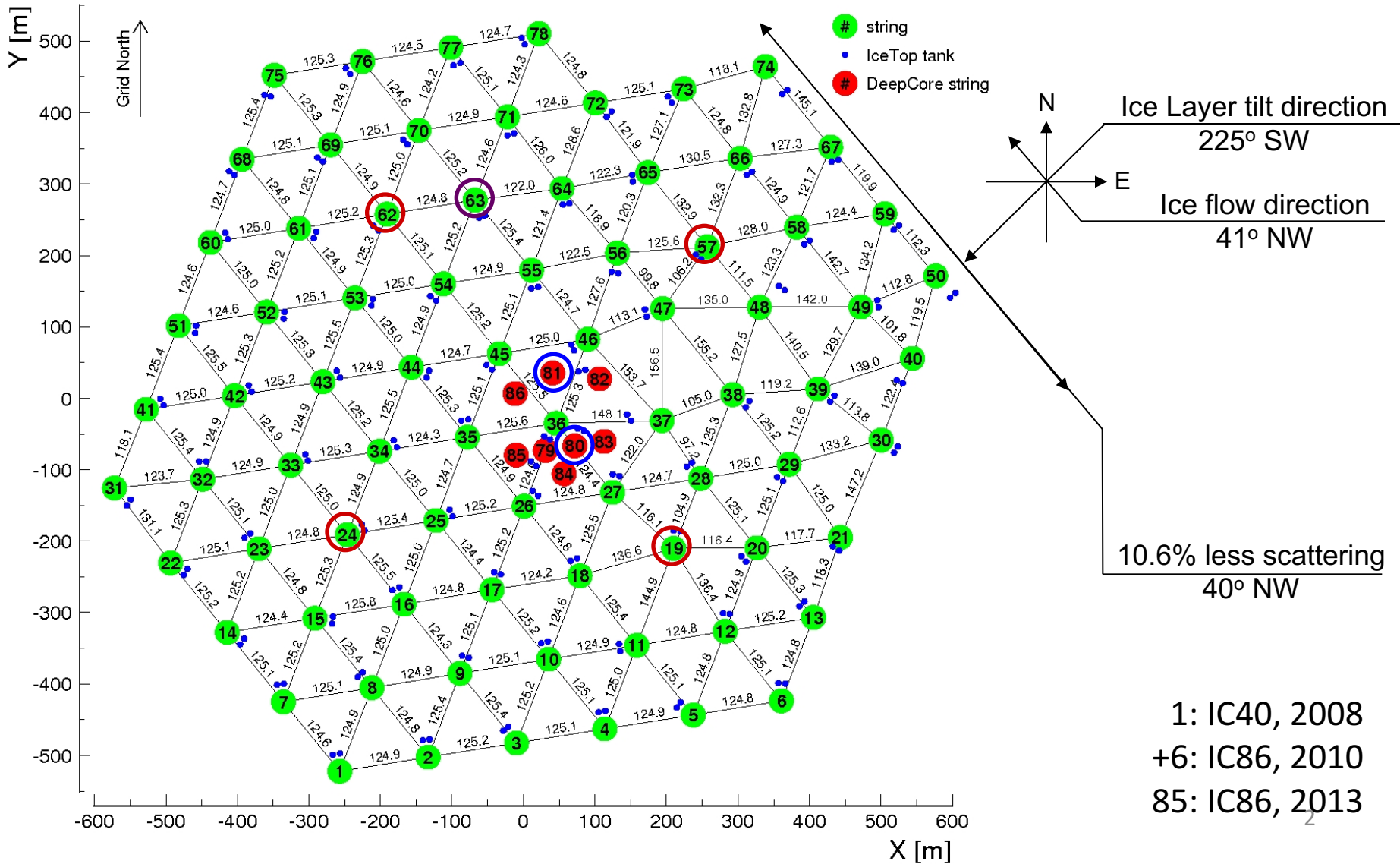


Ice model with flasher unfolding

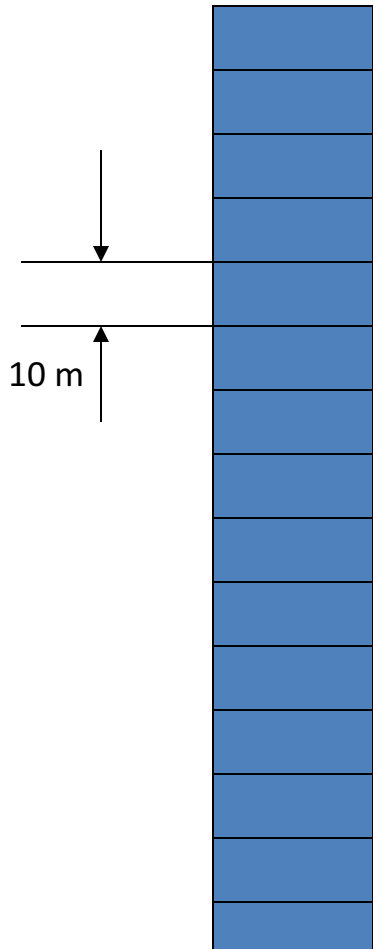


Flasher data-derived ice models

IceCube-86 (78+8) interstring (surface) distances

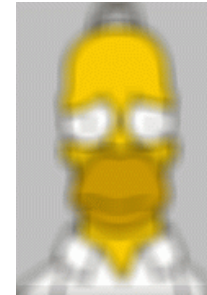


Ice layer parametrization

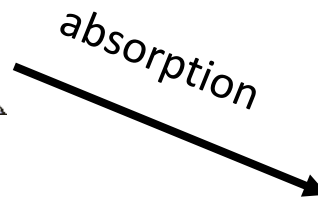


In each 10-meter layer define:

- scattering
- absorption



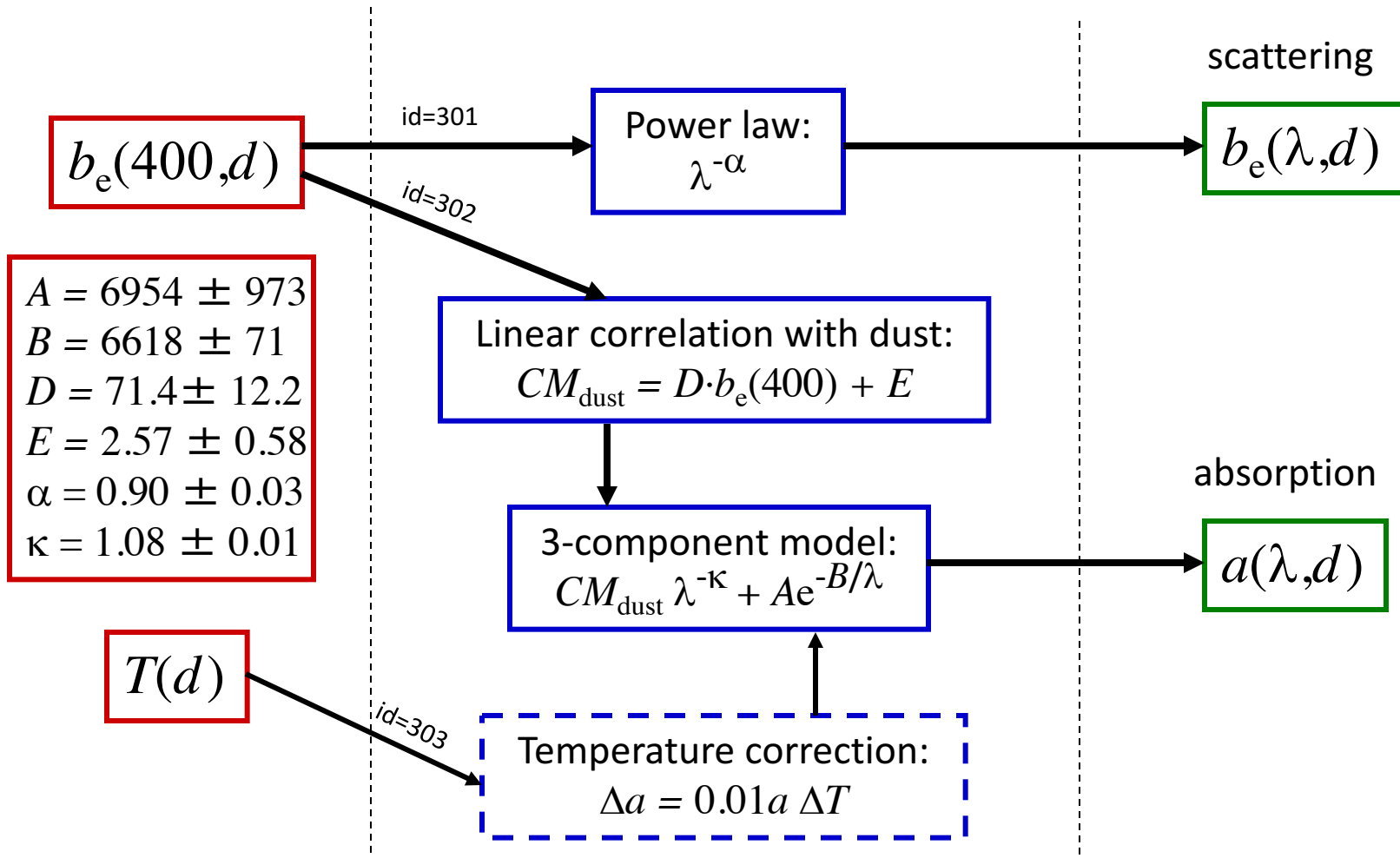
Source is blurred



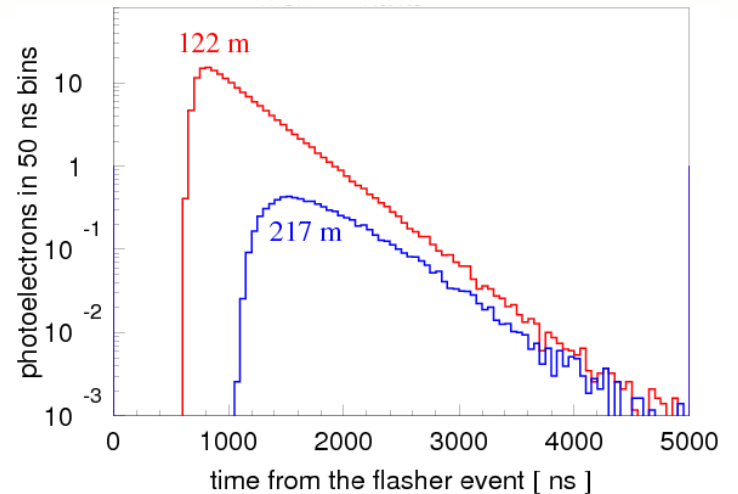
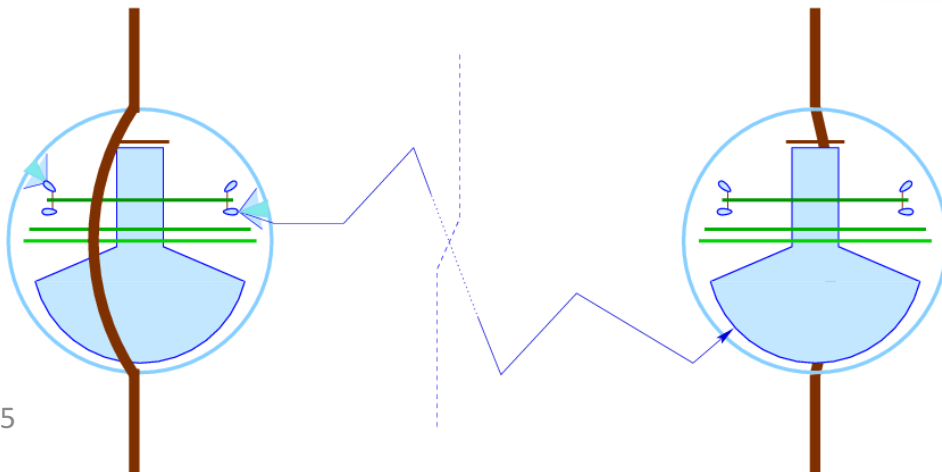
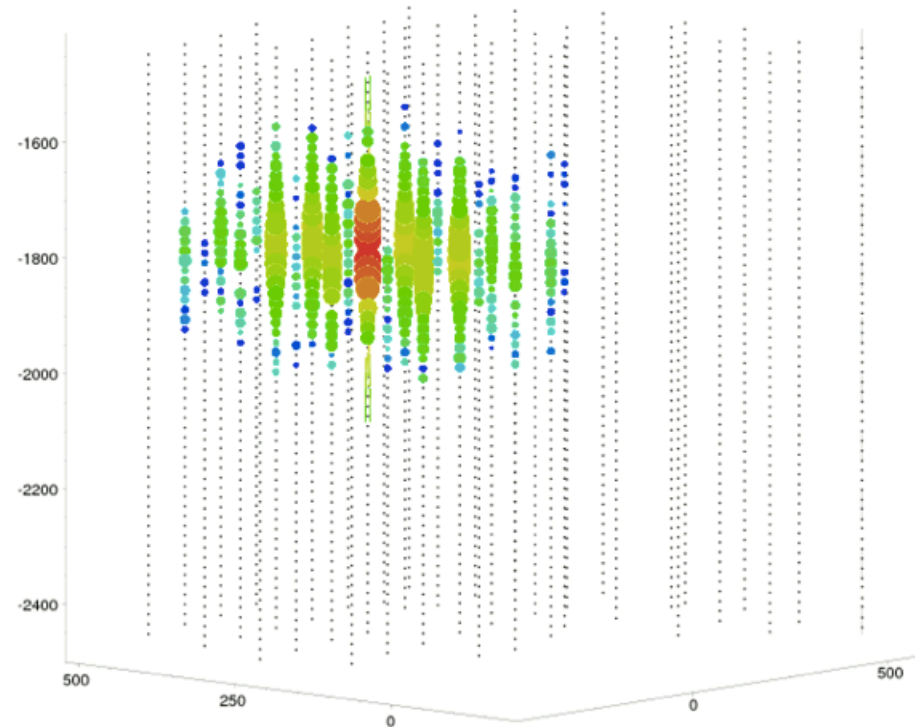
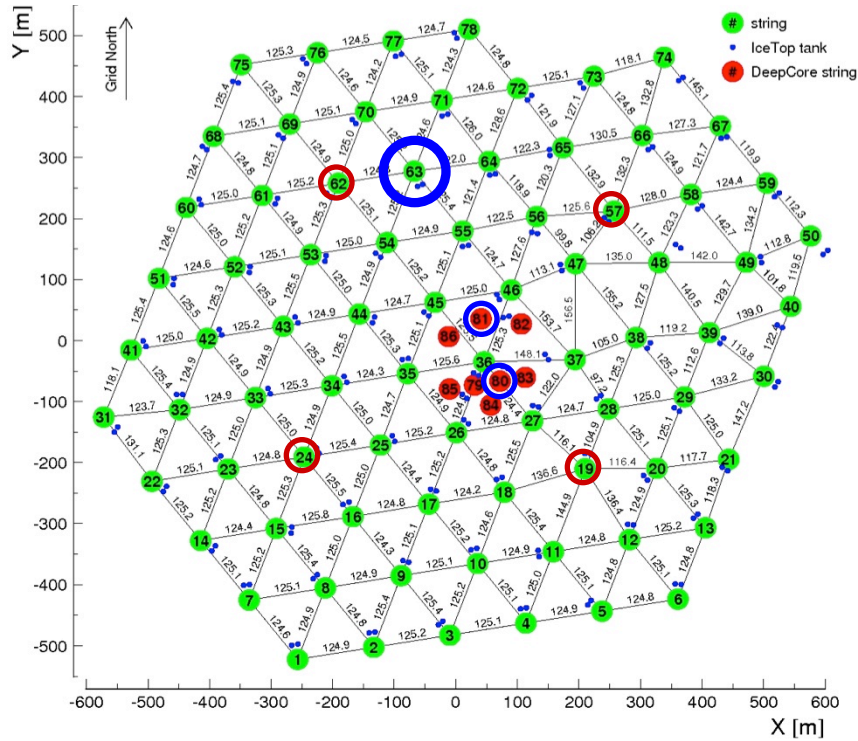
Source is dimmer

$a = \text{inverse absorption length } (1/\lambda_{\text{abs}})$
 $b = \text{inverse scattering length } (1/\lambda_{\text{sca}})$

A 6-parameter Plug-n-Play Ice Model



SPICE model fit



Timeline

AMANDA ice models:

bulk, f125, mam, mamint, stdkurt, sudkurt, kgm, ...
millennium (published 2006) → AHA (2007)

model error

55%

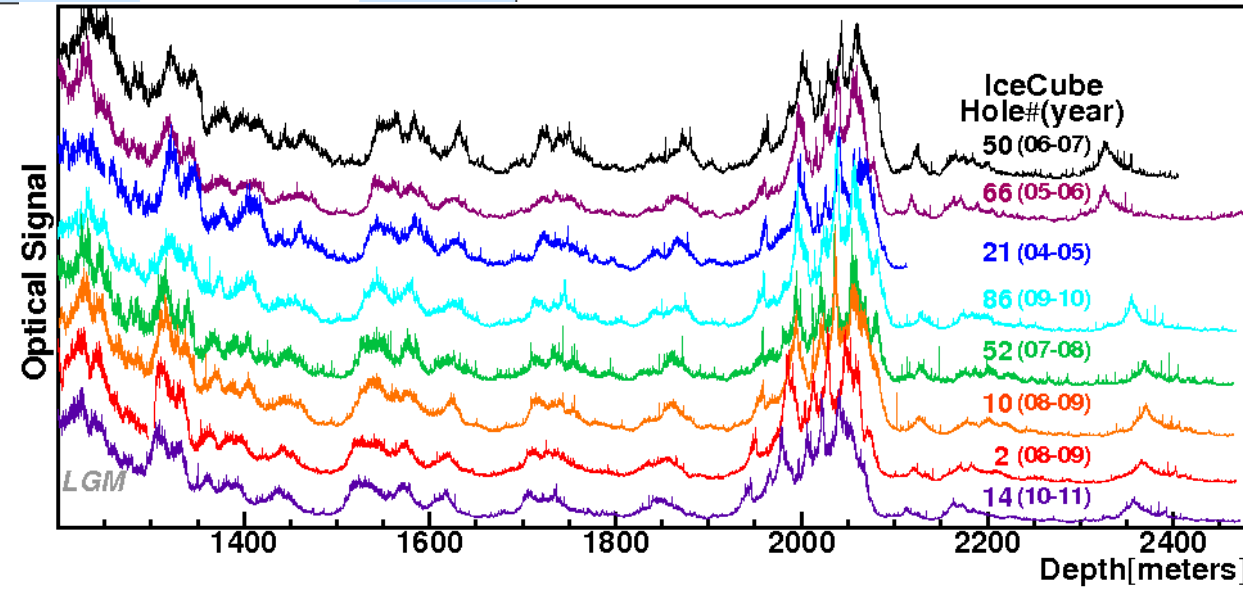
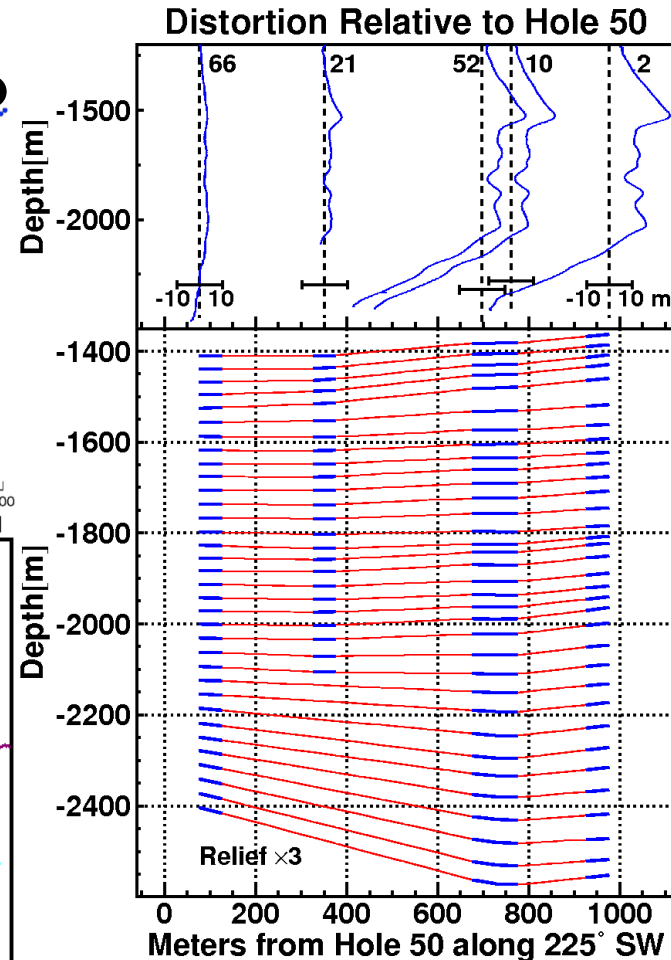
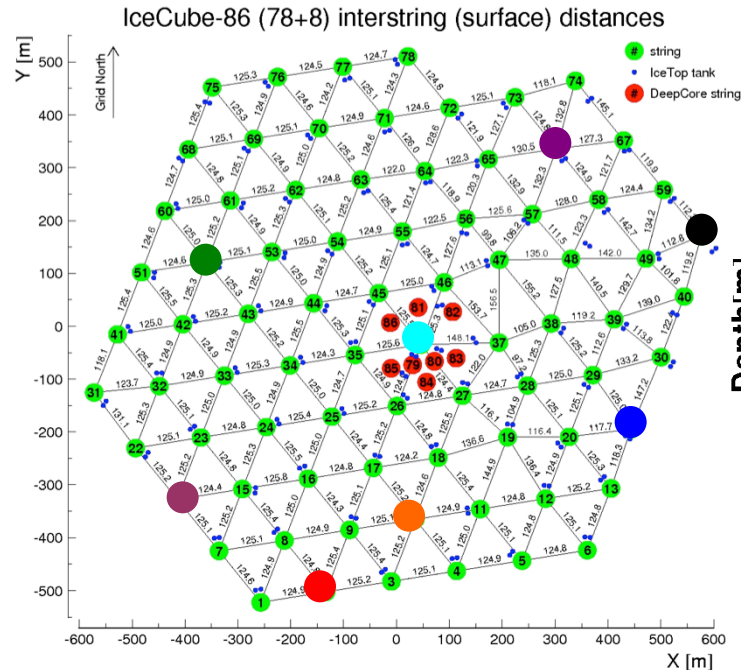
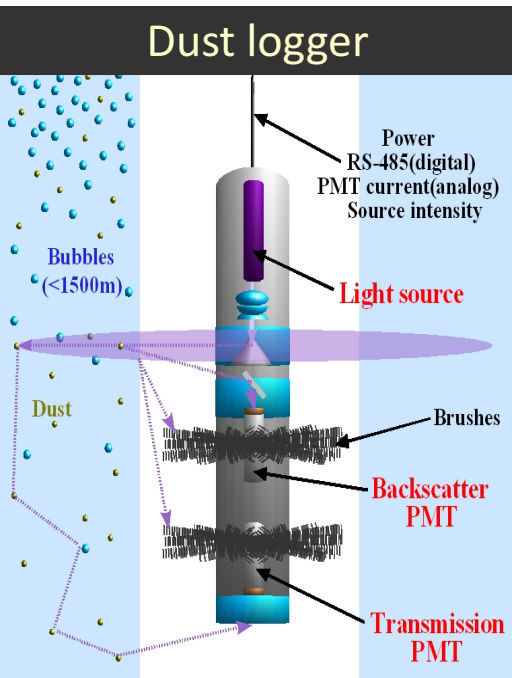
IceCube ice models:

WHAM	(2011)		42%
SPICE 1	(2009)		29%
SPICE 2, 2+, 2x, 2y	(2010)	added ice layer tilt	
SPICE Mie	(2010)	fit to scattering function	29%
SPICE Lea	(2012)	fit to scattering anisotropy	20%
SPICE (Munich)	(2013)	7-string, LED unfolding	17%
SPICE ³ (CUBE)	(2014)	llh fixes, DOM sensitivity fits	11%
SPICE 3.0	(2015)	improved RDE, ang. sens. fits	10%
SPICE 3.1, 3.2	(2016)	85-string, correlated model fit	<10%
SPICE HD, ...		direct HI and DOM sens., cable, DOM tilt	<9%

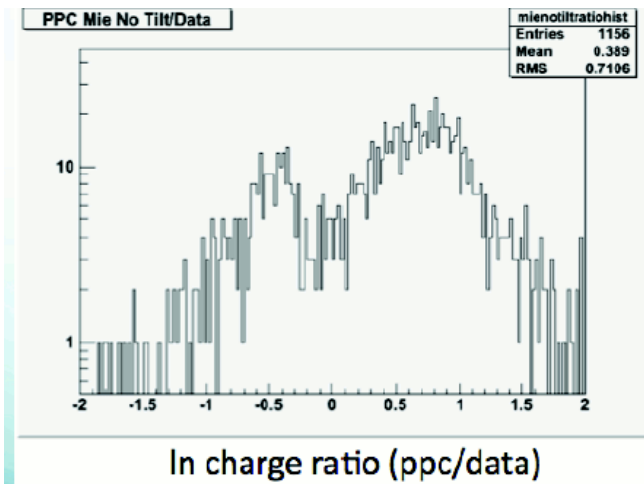
Model error (precision in charge prediction): <10%
Extrapolation uncertainty: 13% (sca) / 15% (abs)
Linearity: < 2% in range 0.1 ... 500 p.e.

(models available in ice-models module)

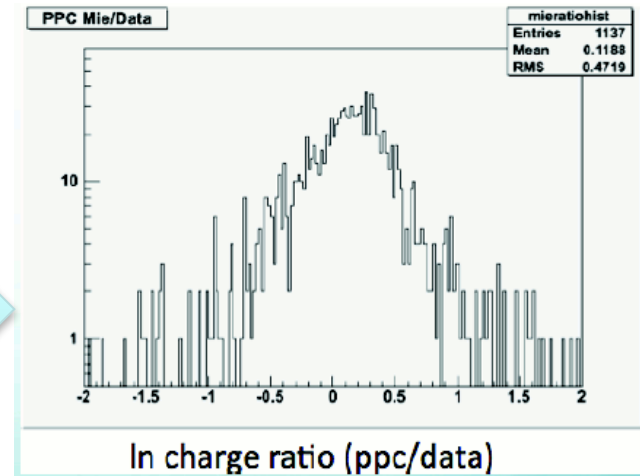
Dust logger discovers ice tilt



Tilt is important for flasher data!



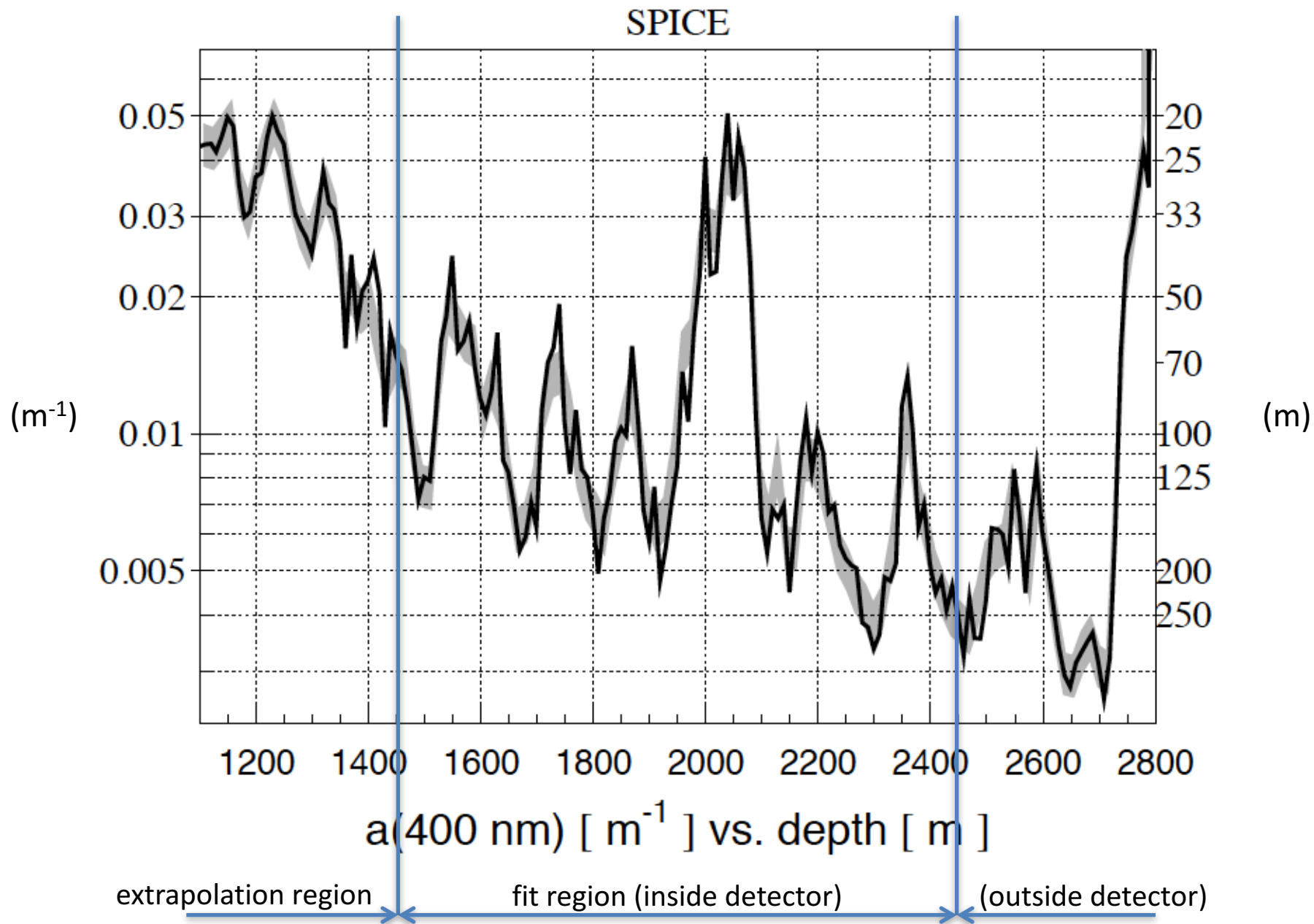
Comparison of LED flasher simulation to data without tilt



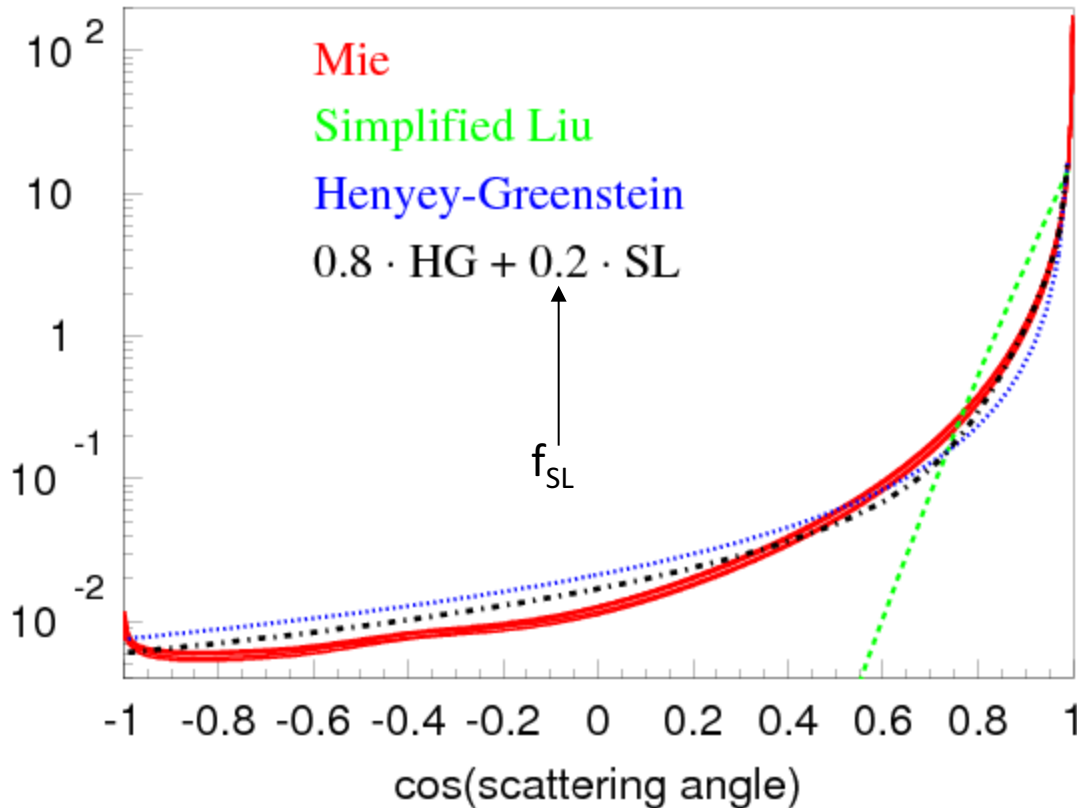
Comparison of LED flasher simulation to data with tilt

black line: fit to flasher data

gray band: scaled merged dust log



Approximation to Mie scattering



Simplified Liu:

$$p(\cos \theta) \sim (1 + \cos \theta)^\alpha, \quad \text{with} \quad \alpha = \frac{2g}{1-g}$$

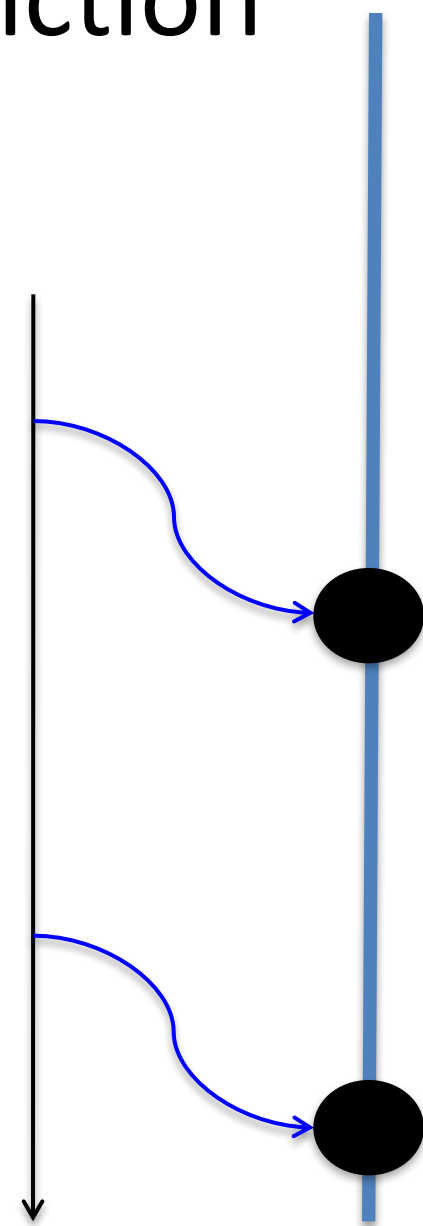
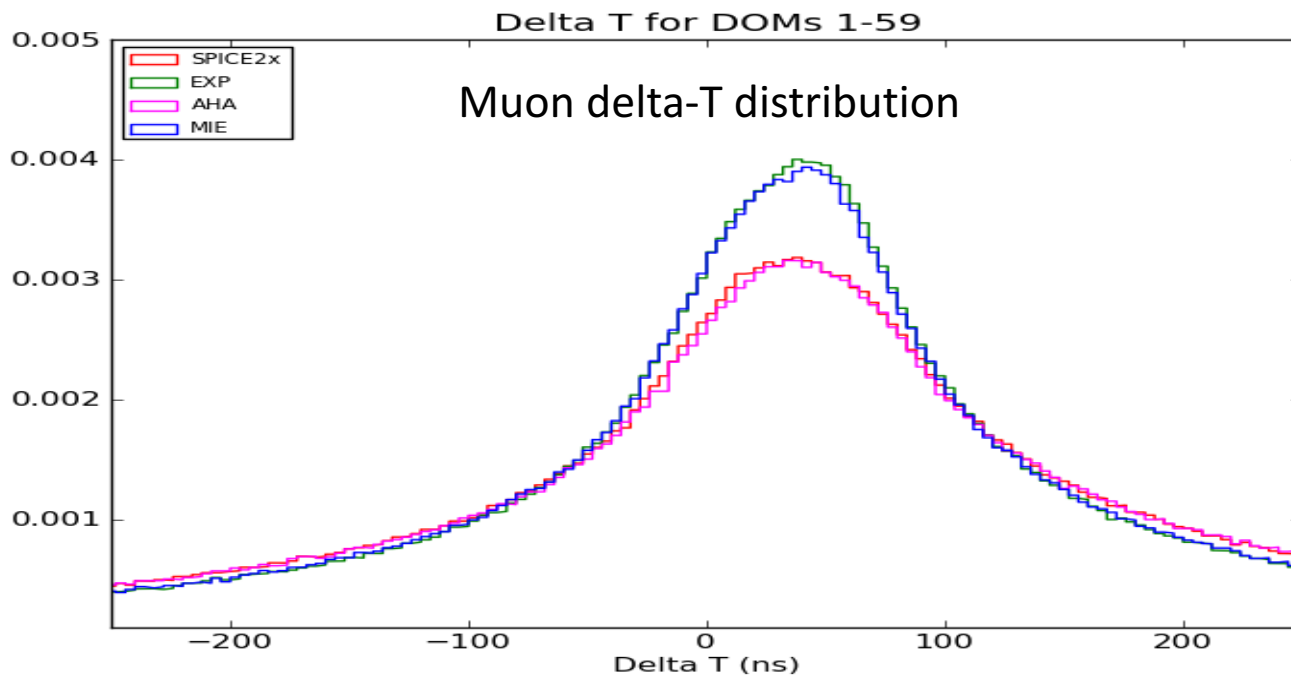
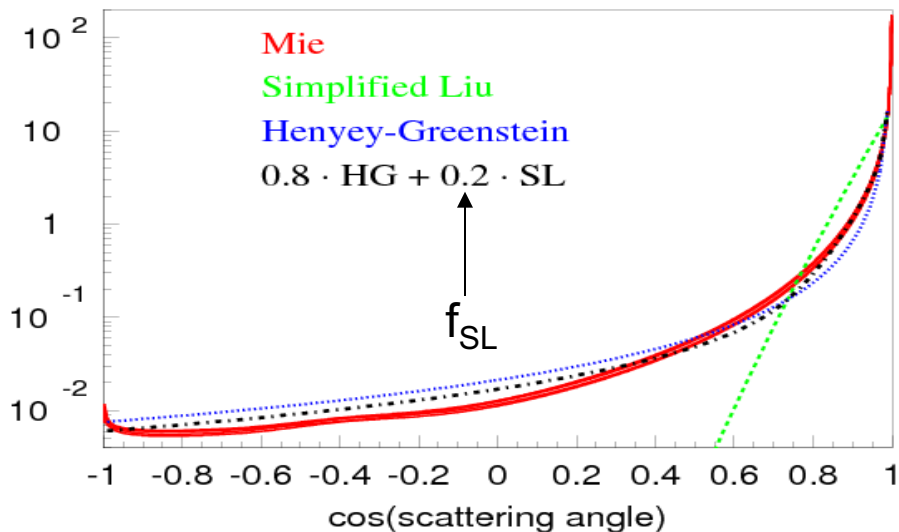
Henyey-Greenstein:

$$p(\cos \theta) = \frac{1}{2} \frac{1 - g^2}{[1 + g^2 - 2g \cdot \cos \theta]^{3/2}}$$

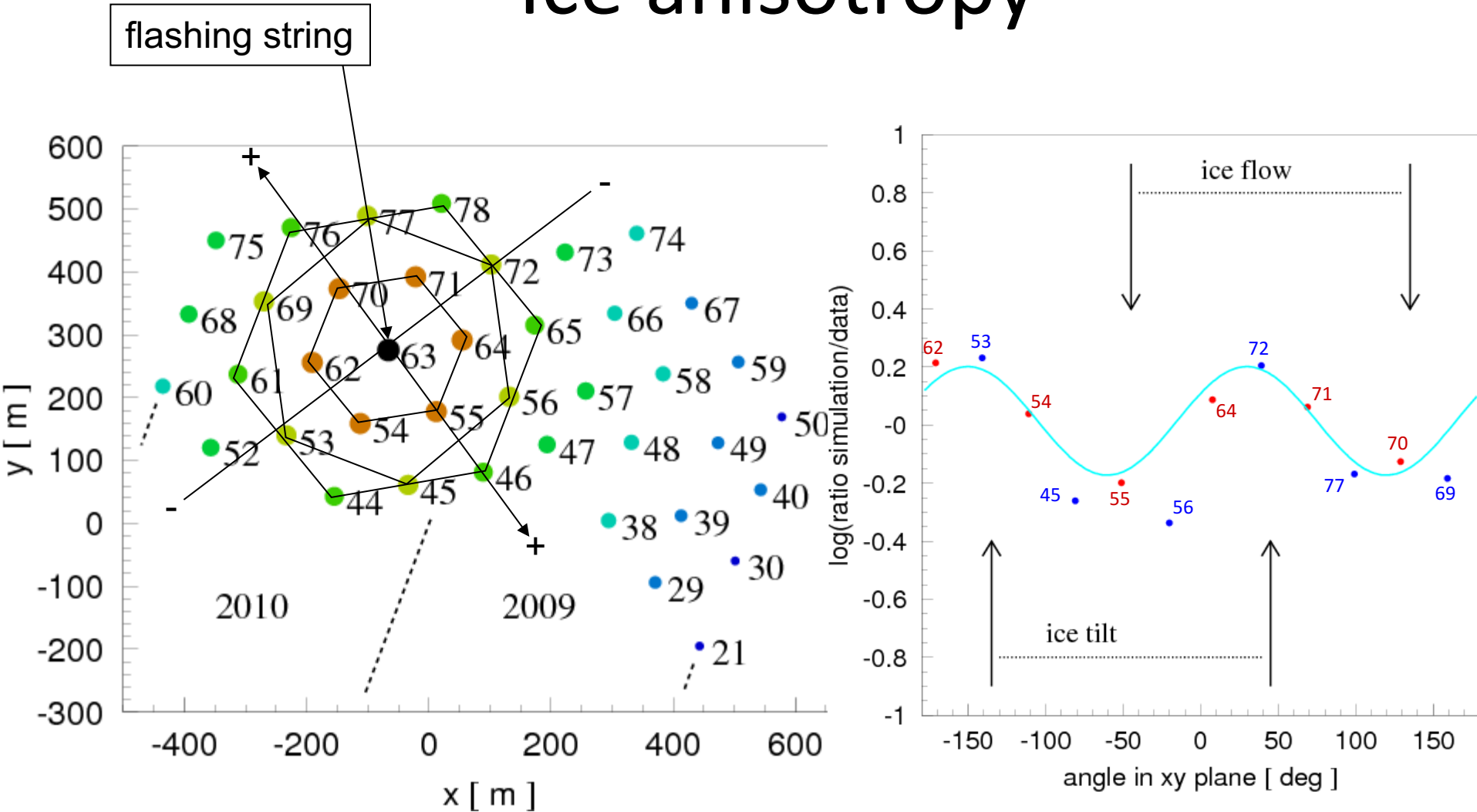
Mie:

Describes scattering on acid, mineral, salt, and soot with concentrations and radii at SP

Fitting the scattering function

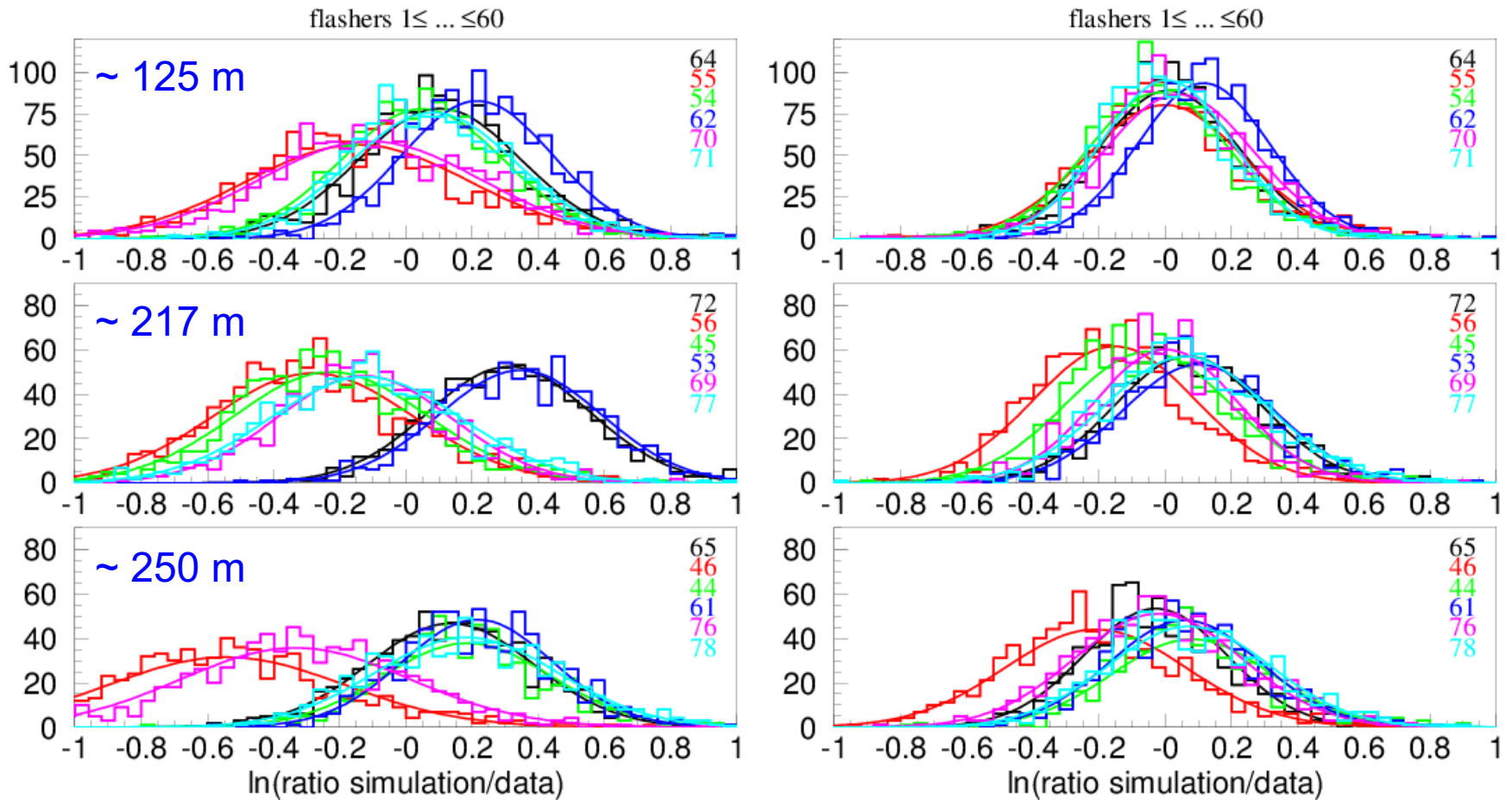


Ice anisotropy



10-20% per 100 m azimuth modulation in charge observed!

Charge variation vs. distance

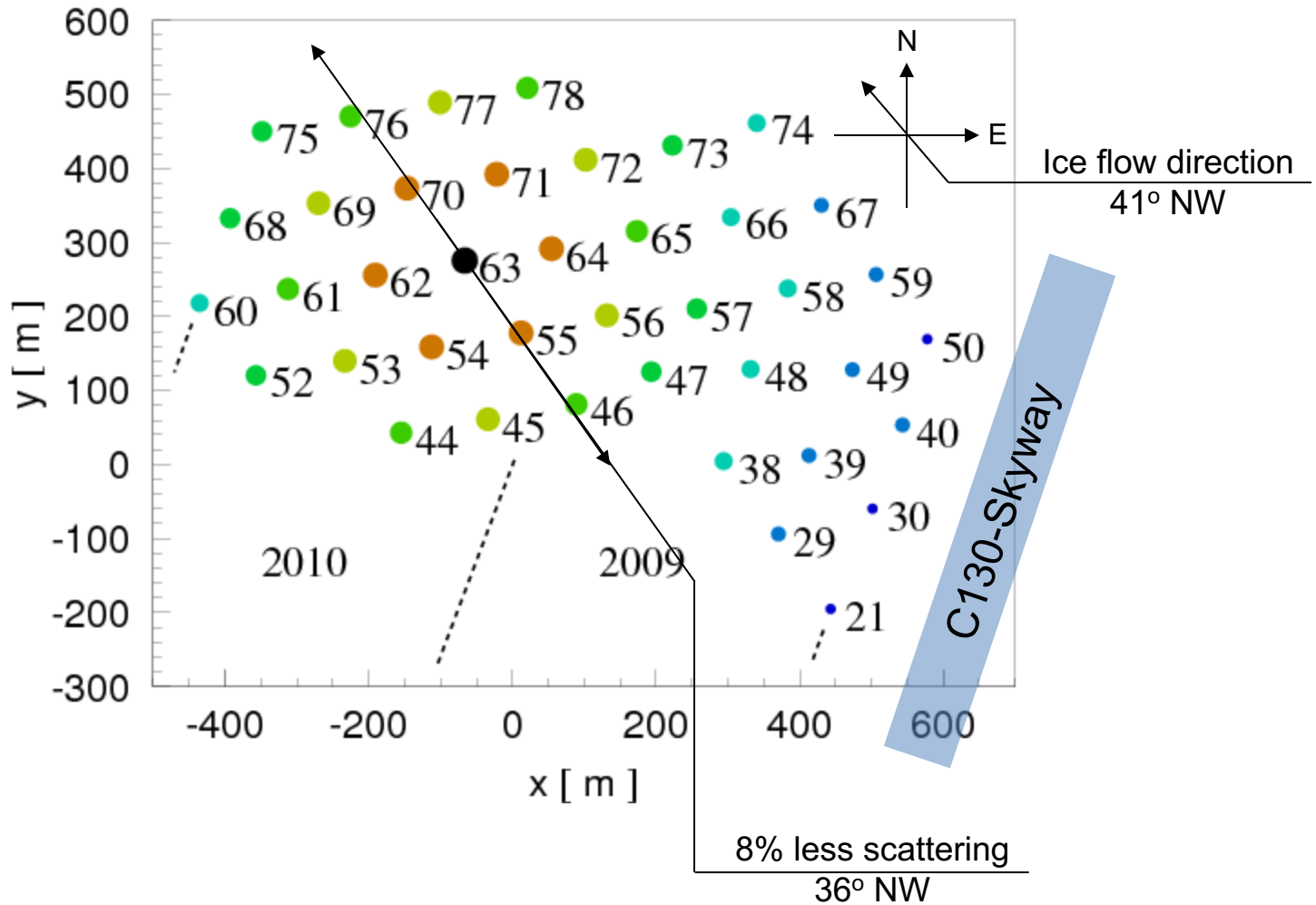


SPICE Mie [SPICE Paper]



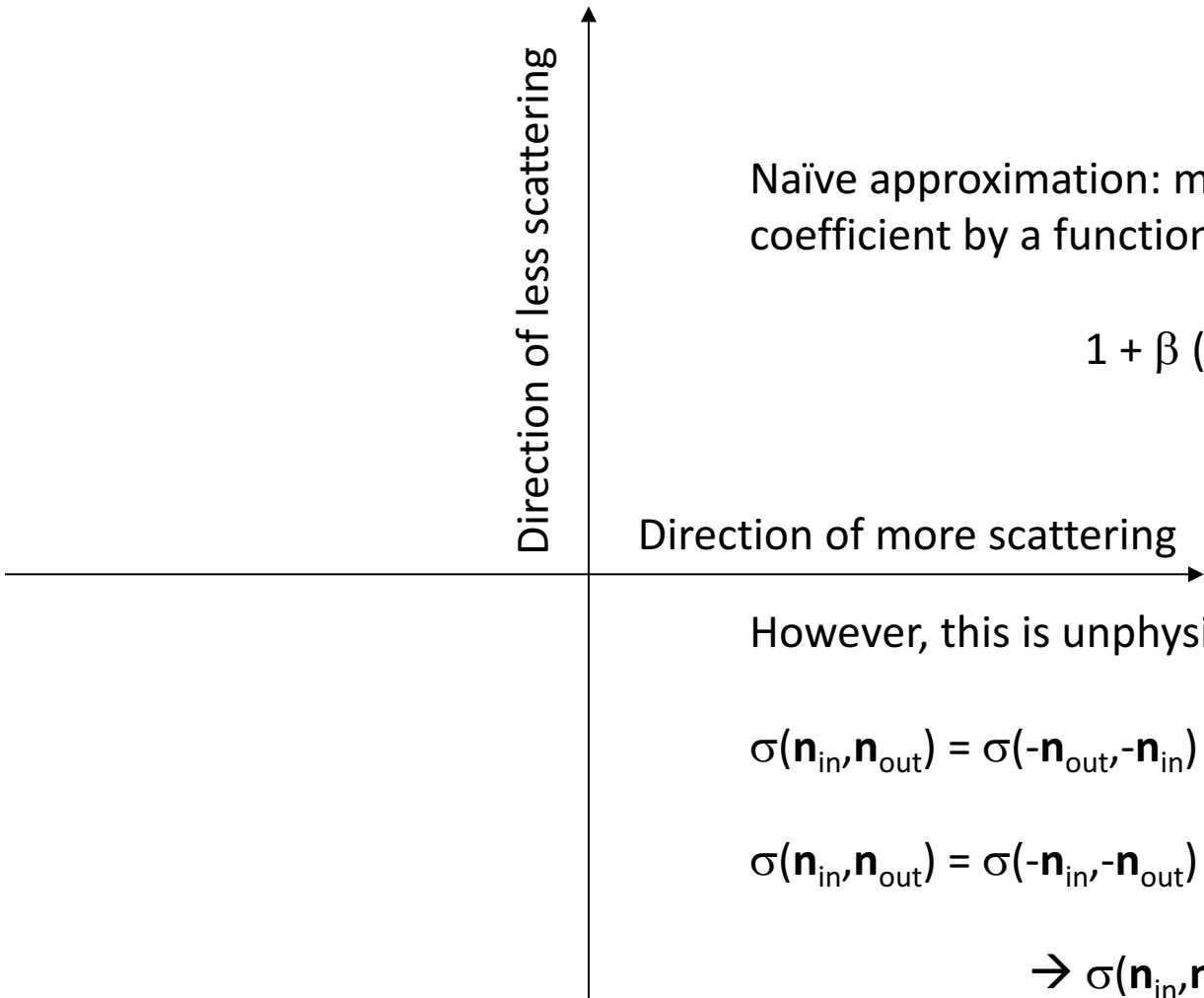
SPICE Lea

Fit to the anisotropy hypothesis



What is Ice anisotropy

Direction of less scattering



Naïve approximation: multiply the scattering coefficient by a function of photon direction, e.g., by

$$1 + \beta (\cos^2\theta - 1/3)$$

Direction of more scattering

However, this is unphysical:

$$\sigma(\mathbf{n}_{in}, \mathbf{n}_{out}) = \sigma(-\mathbf{n}_{out}, -\mathbf{n}_{in}) \text{ (time-reversal symmetry)}$$

$$\sigma(\mathbf{n}_{in}, \mathbf{n}_{out}) = \sigma(-\mathbf{n}_{in}, -\mathbf{n}_{out}) \text{ (symmetry of ice)}$$

$$\rightarrow \sigma(\mathbf{n}_{in}, \mathbf{n}_{out}) = \sigma(\mathbf{n}_{out}, \mathbf{n}_{in})$$

A possible parameterization

The scattering function we use is $f(\cos \theta)$, a combination of HG and SL.

How about this extension: $f(\vec{n}_i \cdot \vec{n}_o) \rightarrow f(\vec{k}_i \cdot \vec{k}_o)$, $\vec{k}_{i,o} = \frac{A\vec{n}_{i,o}}{|A\vec{n}_{i,o}|}$

$A = \begin{pmatrix} \alpha & 0 & 0 \\ 0 & \beta & 0 \\ 0 & 0 & \gamma \end{pmatrix}$ the basis of the 2 scattering axes and z (α, β are, e.g., 1.05).

However, function $f(\cos \theta)$ is well-defined for only $\cos \theta$ between -1 and 1.

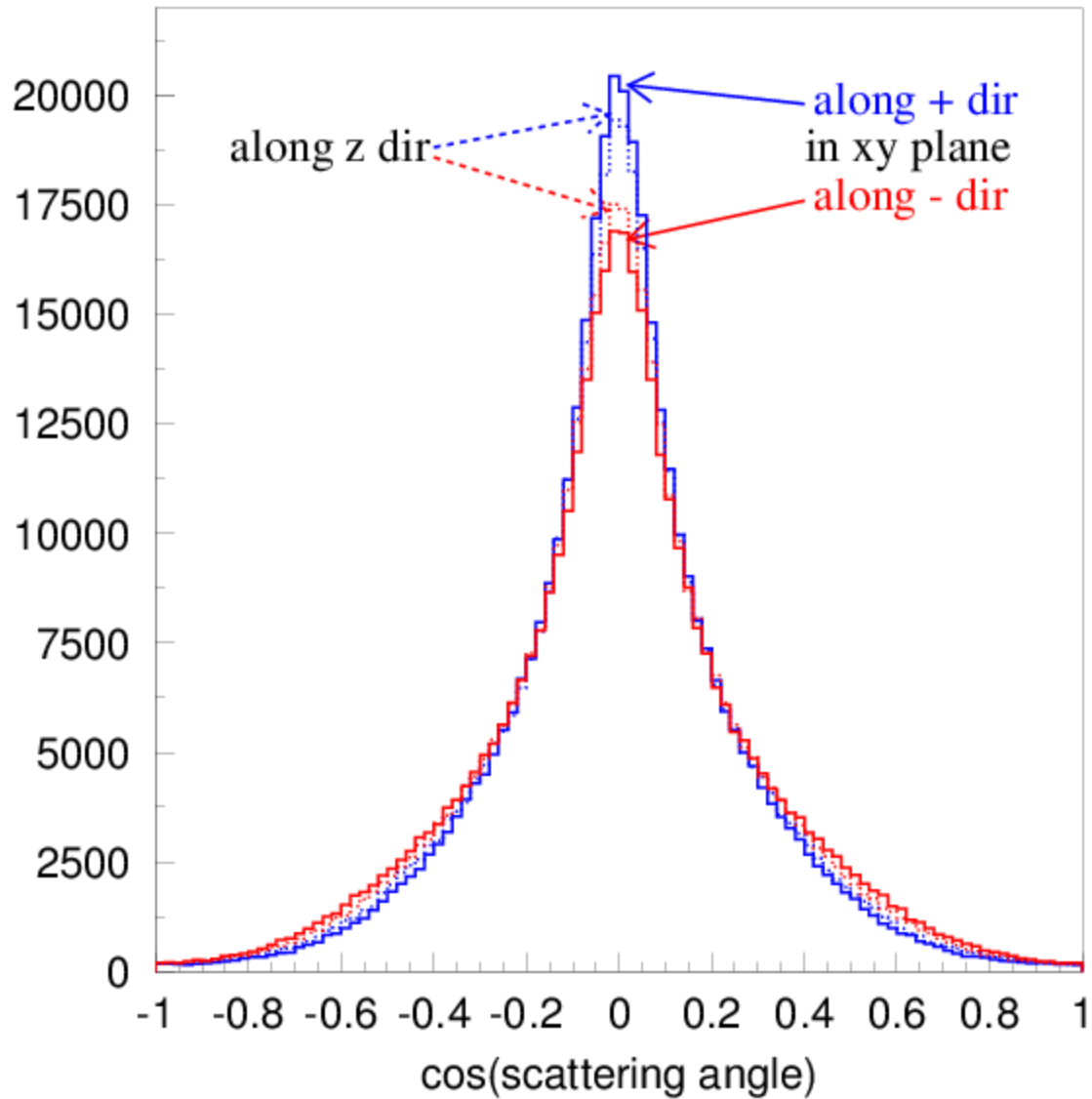
A possible modification is $\mathbf{n}_{in} \rightarrow A\mathbf{n}_{in}/|A\mathbf{n}_{in}| \rightarrow \mathbf{n}_{out} \rightarrow A^{-1}\mathbf{n}_{out}/|A^{-1}\mathbf{n}_{out}|$.

This introduces two extra parameters: α, β (in addition to the direction of scattering preference).

The geometric scattering coefficient is constant with azimuth. However, the effective scattering coefficient receives azimuthal dependence as:

$$1 - \langle \cos \theta \rangle = (1 - g) \cdot \frac{1}{2} \cdot (B_{in} B_{in} - n^i B_{in} n^j B_{jn}) \cdot |A\vec{n}|^2 \quad B = A^{-1}$$

Scattering example (5% anisotropy)



Anisotropy coefficient

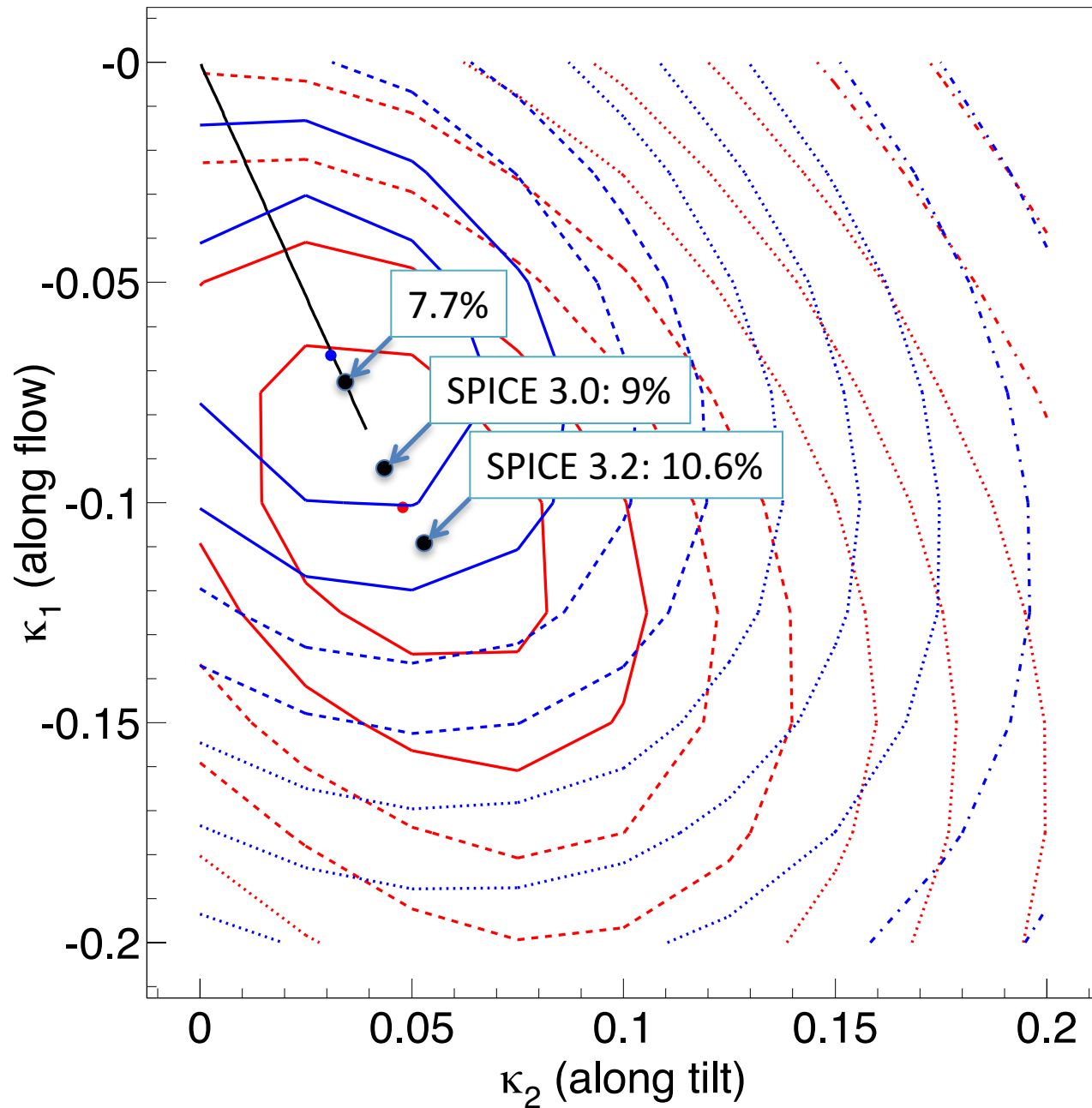
Two values are currently derived from the same data:

10.6% (SPICE 3.2) – fit uses full flasher light profile unfolding and RDE table

7.7% (SPICE HD) – fit uses cylindrically-symmetrical (or otherwise fixed) emitted light pattern and nominal DOM efficiencies

The unfolded flasher light pattern has an overall shape that is correlated with the direction of (main axis of) ice anisotropy – this is not good

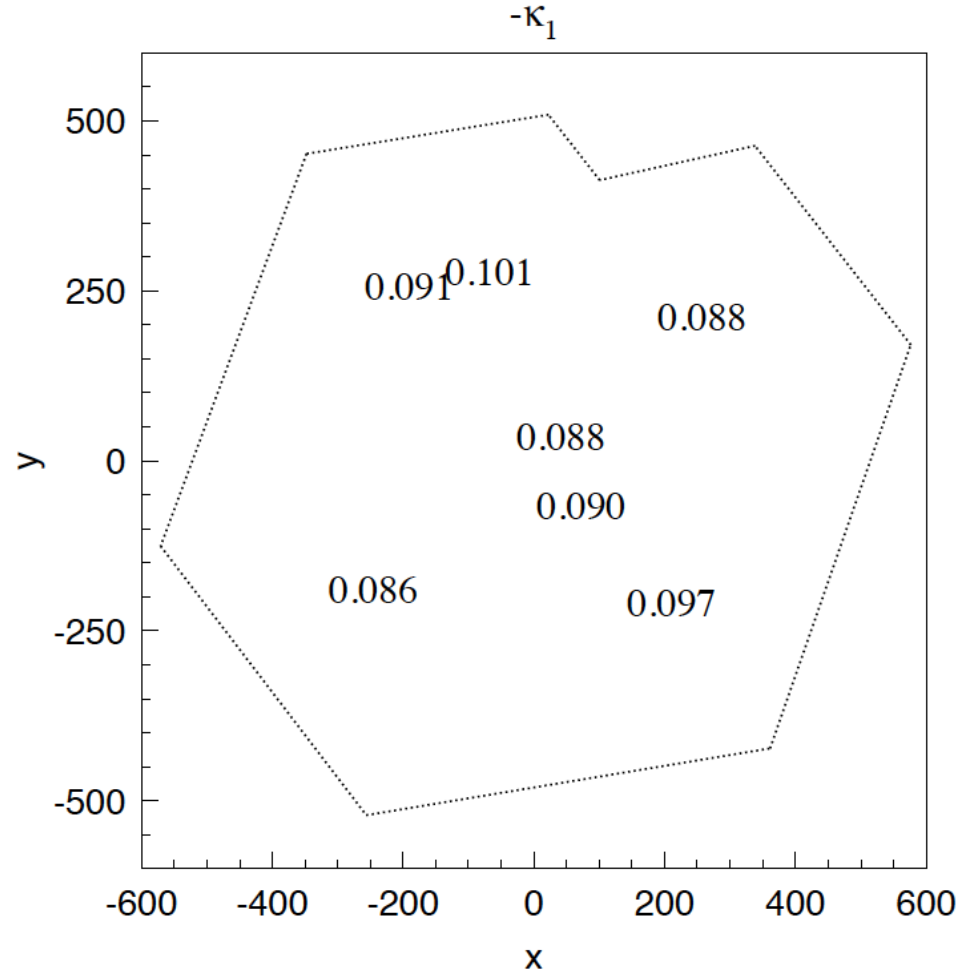
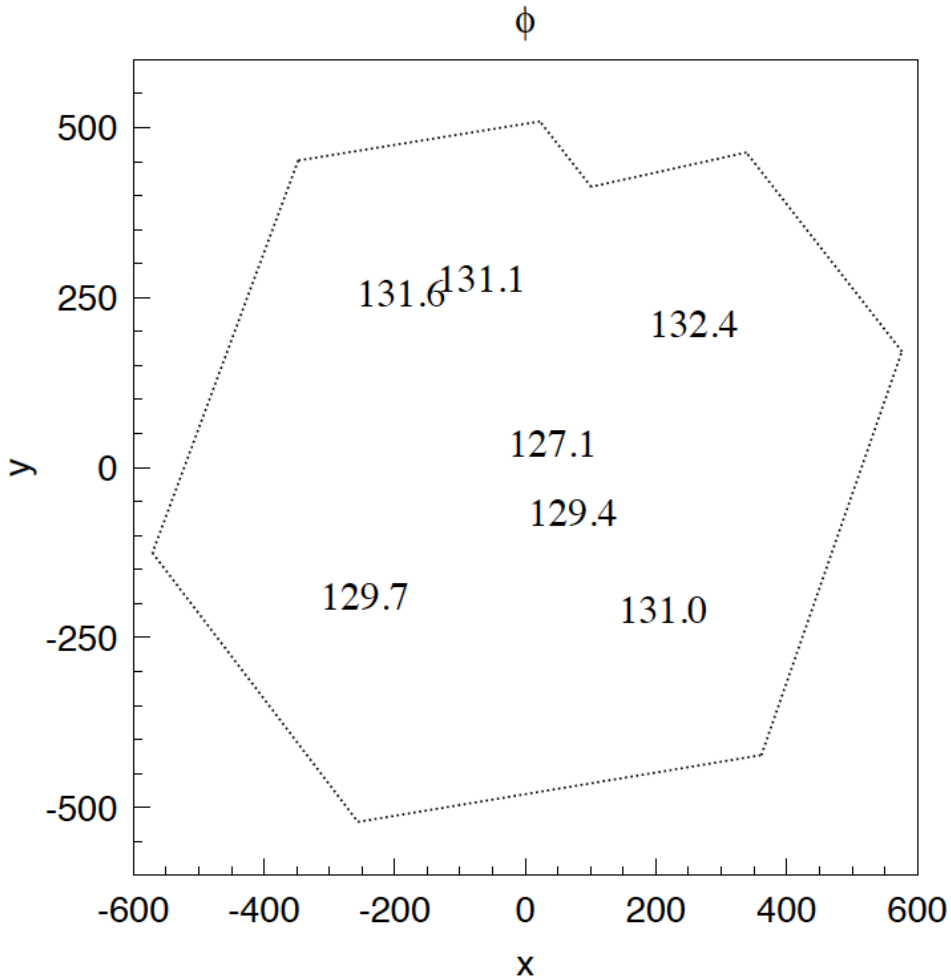
Also, Martin R. found evidence for depth-dependence of anisotropy by fitting the anisotropy coefficient at different depths



From SPICE Lea
ICRC report:

Red: charge-only
Blue: time-binned

ϕ/κ_1 inhomogeneity



SPICE 3.x

SPICE 3.0 used flashers on 7 strings, updated tilt maps and extrapolation data, fitted individual DOM sensitivities and flasher DOM LED pattern unfolding

SPICE 3.1 updated the model using flashers on 85 strings, full likelihood fit to the DOM sensitivities with regularizations around nominal values (1 and 1.35)
Also: unfolded angular sensitivity to the same set of flasher data

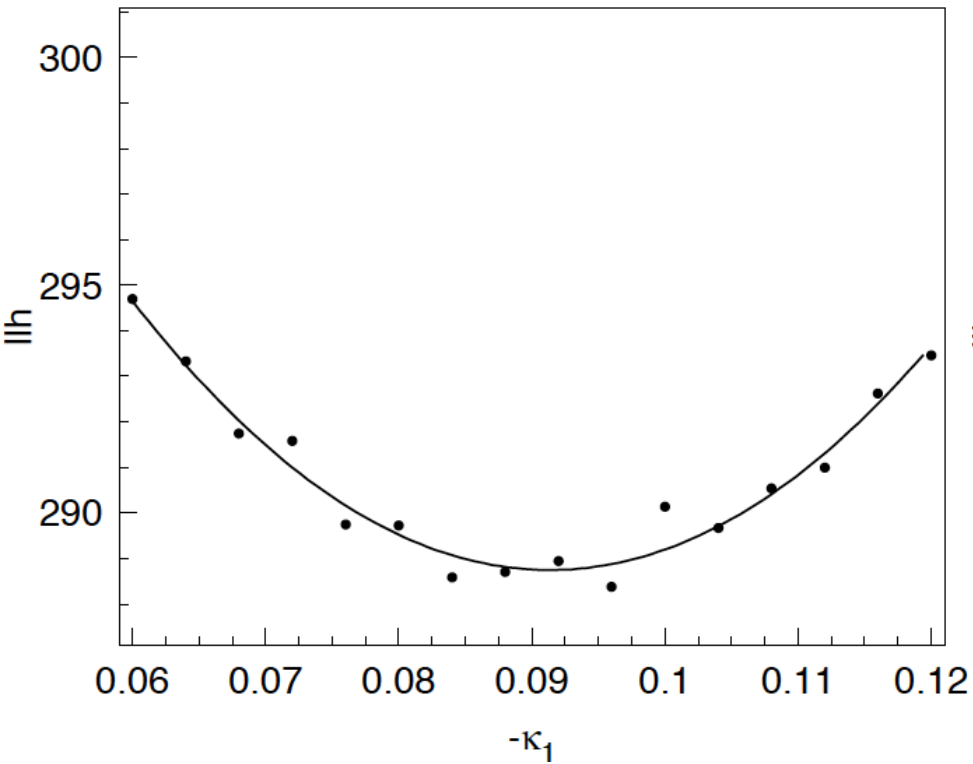
SPICE 3.2 employed a special (correlation model) refining method for adding small corrections to the table of scattering and absorption coefficients

(available in ice-models module)

Selected ice model parameters

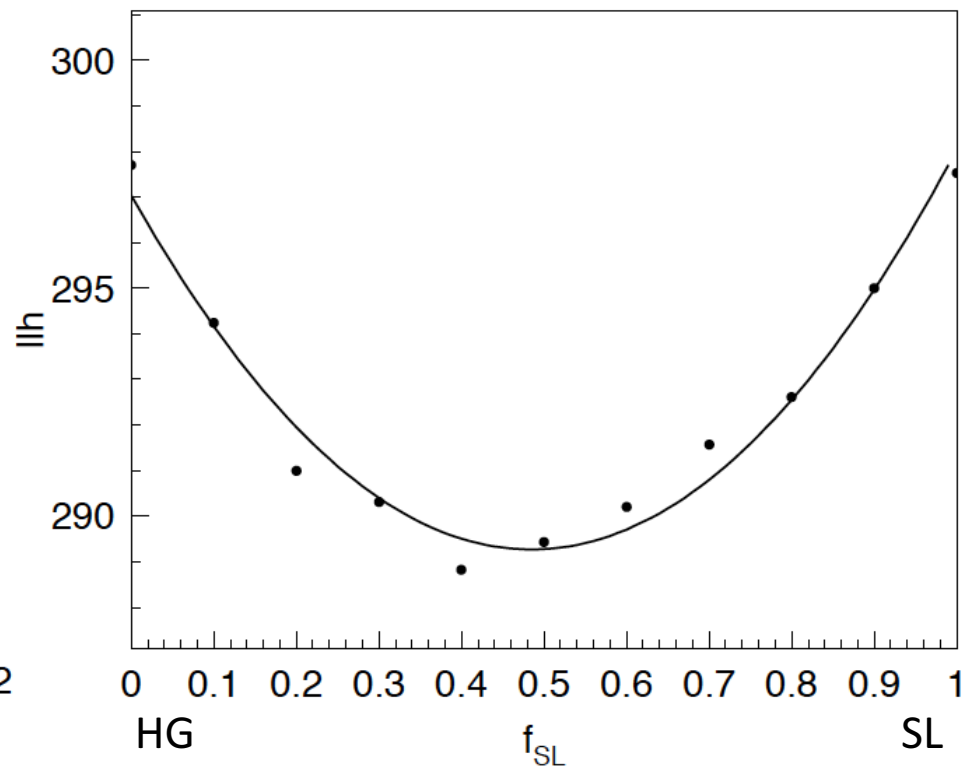
SPICE 3.0

min at 0.0914



SPICE Lea: 0.08

min at 0.48

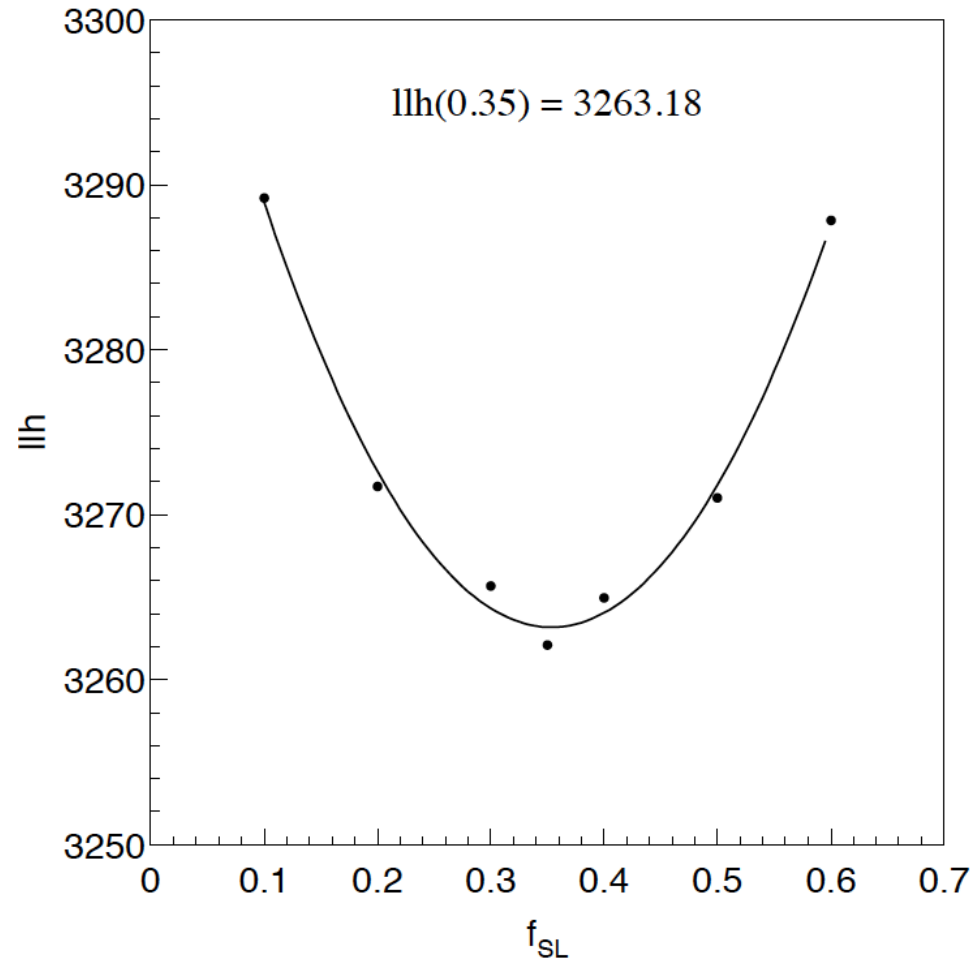
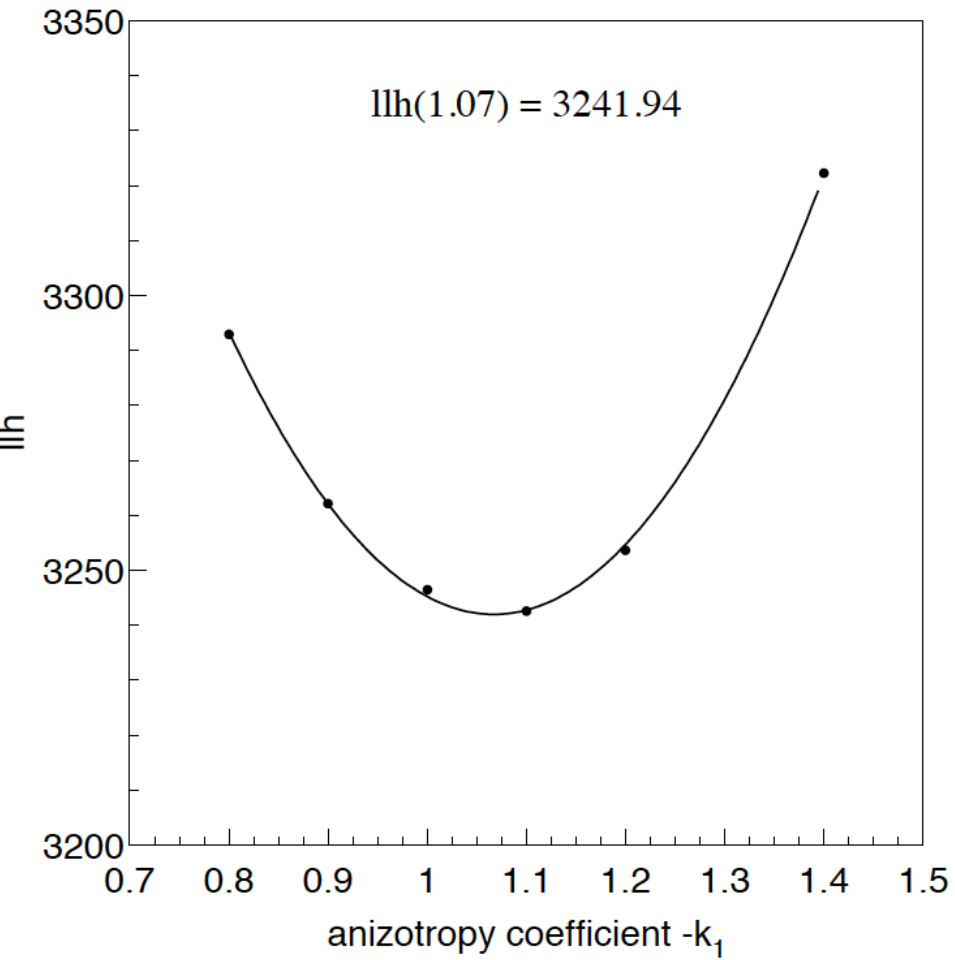


SPICE Mie: 0.45

SPICE Lea: 0.41

Selected ice model parameters

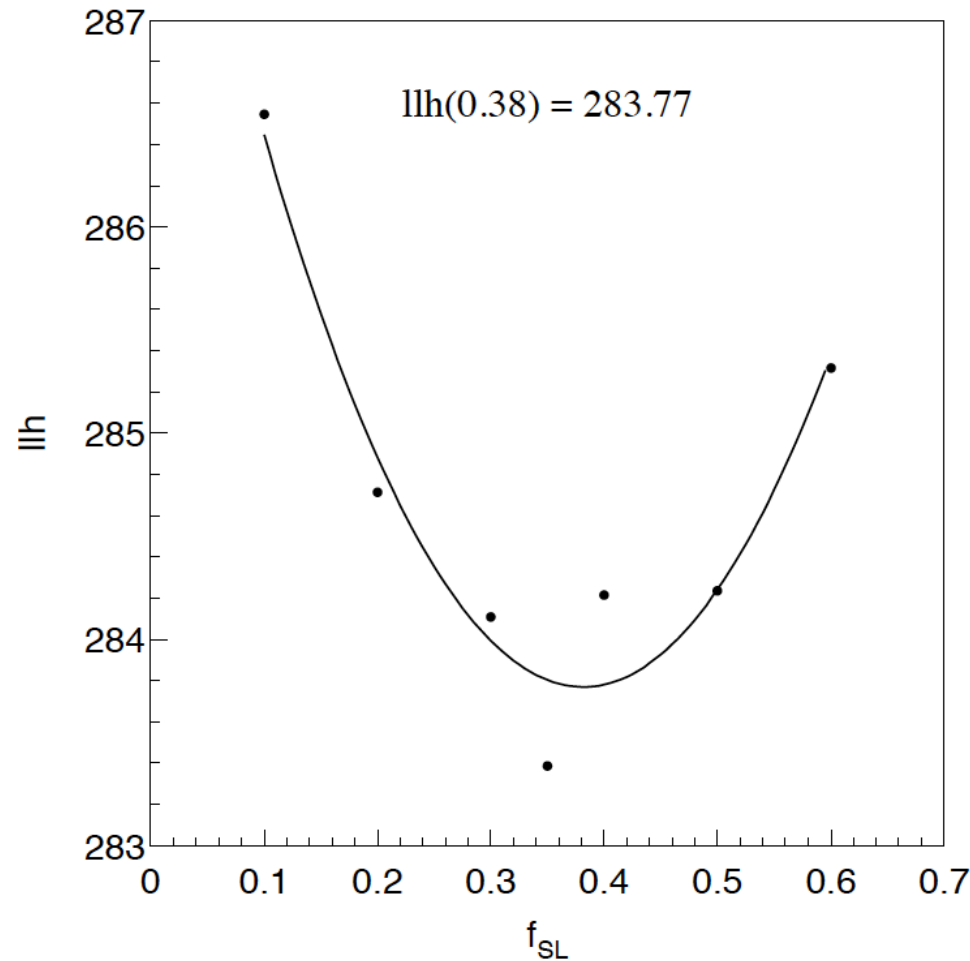
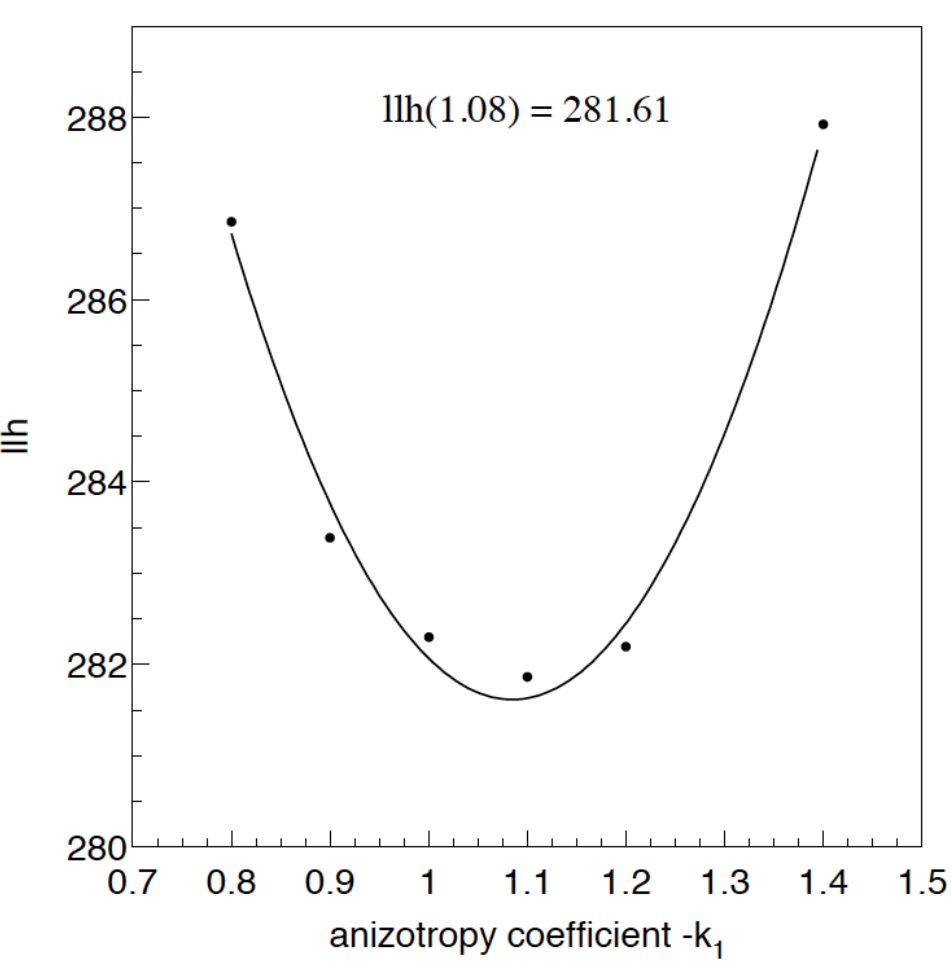
SPICE 3.1



85 strings of flasher data

Selected ice model parameters

SPICE 3.1



Applied to only 7 strings of flasher data

Ice model comparison

Ice models	SPICE Mie	SPICE Lea	SPICE 3	SPICE 3.1	SPICE 3.2
Strings fit	1	1	7	85	85
Anizotropy	0	8%	9%	10.8%	10.6%
f_{SL} (scat.)	0.45	0.41	0.5	0.35	0.35
Model error	29%	20%	10.7%	9.8%	9.8%
g.o.f (x10)				1.2115	1.1799

Other improvements in SPICE 3:

- updated tilt maps and extrapolation (relies on dust logger/EDML data)
- DOM sensitivity fits and flasher DOM LED pattern unfolding

Main difference to Mie and Lea: 4% less absorption on average

(all of the above models are available in ice-models module)

SPICE HD

Ice model developed by Martin R

shares the ice tables with SPICE 3.2

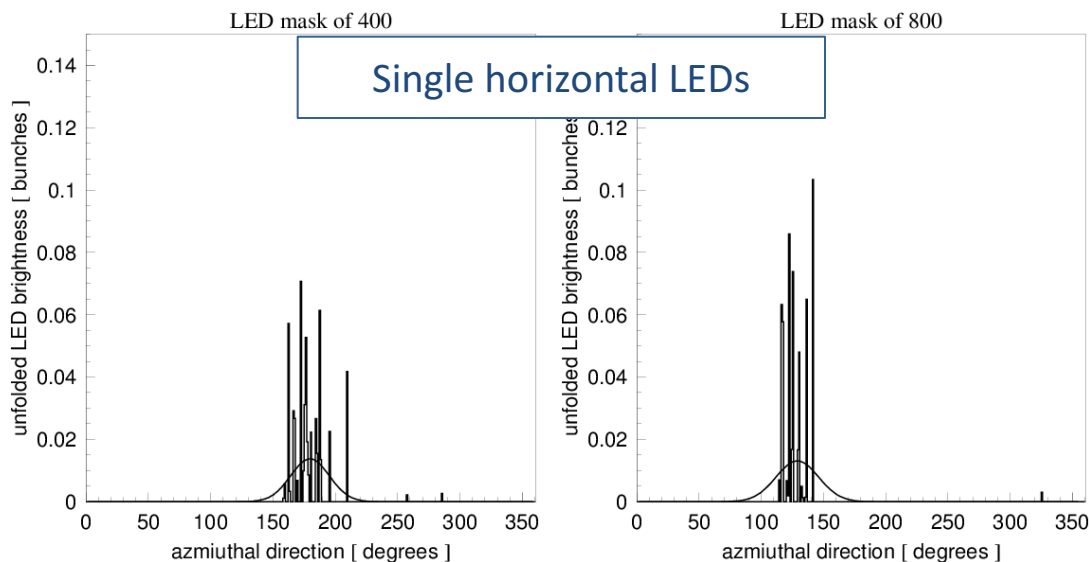
(except with smaller anisotropy coefficient, 7.7%)

(new: with depth-dependent anisotropy)

(new: with different anisotropy model)

Unfolding of flasher LEDs

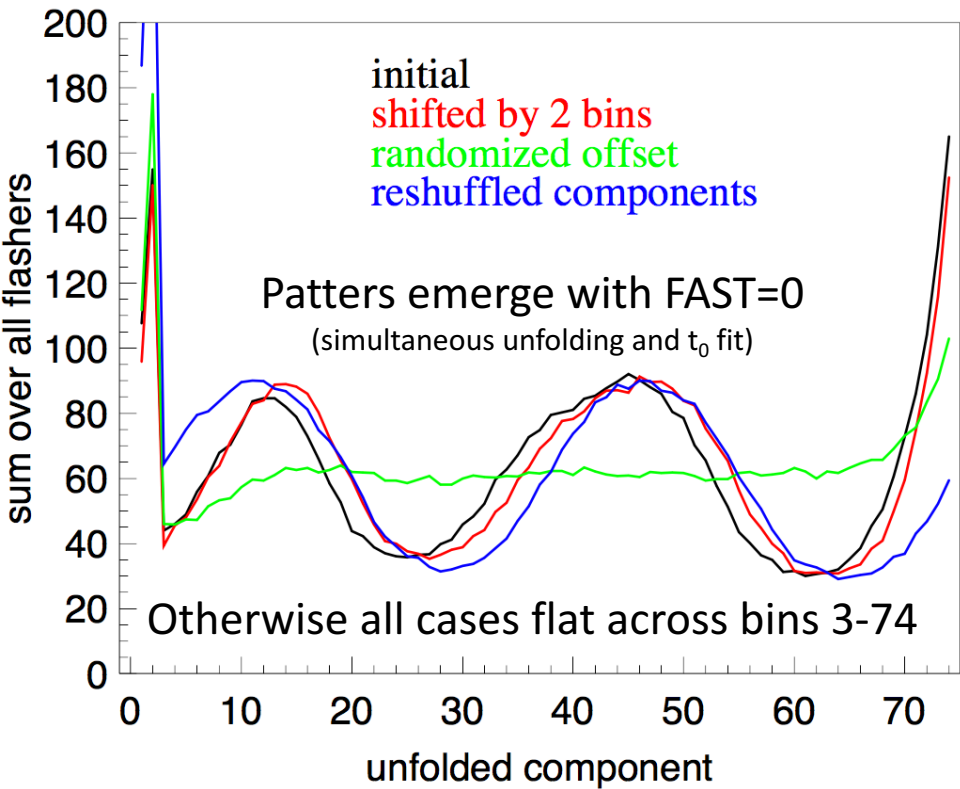
- Simulate LED light (2d gaussian with 9.7 degree rms) every 5 degrees in azimuthal direction from 0 to 355 degrees, also up and down components
- Create a [azimuth x charge_in_DOM] matrix, and unfold to charge_in_DOM in data
- The unfolded pattern is re-simulated and llh is calculated



This adds 30340 ($410 \cdot 74$) nuisance parameters

Additional parameters optionally introduced at a later step: individual DOM efficiencies (~ 3400 values)

Correlated unfolded pattern



components 1,2 are up/down
components 3-74 correspond to
0-355 degrees spaced by 5 degrees

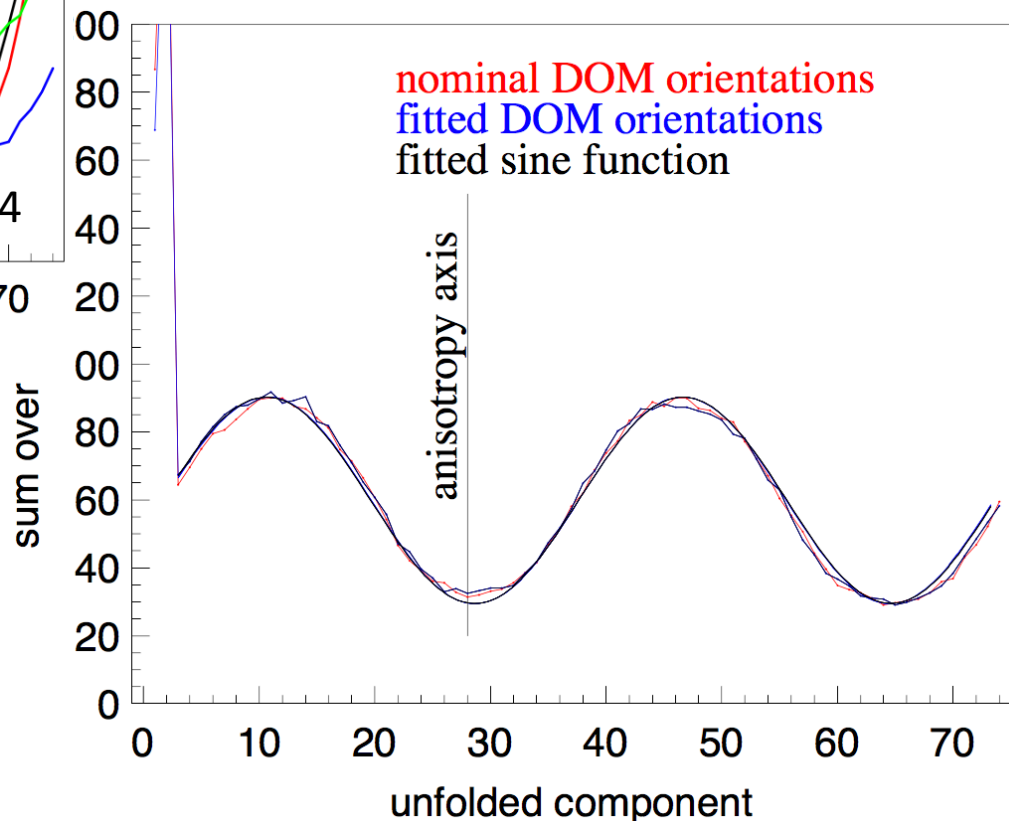
Unfolding favors higher components

Starting (initial) LLH=3189.17
Reshuffled (fixed) LLH=3097.57

SREP=10

Model error: 11.2% → 10.9%

Fixed pattern is now sine-like



One ice table for everything

Nominal DOM efficiencies, h2-50cm hole ice model, 1x simulation (10 events)

Ice model comparison: GOF

Nominal DOM efficiencies, h2-50cm hole ice model, 1x simulation (10 events)

With flasher LED pattern unfolding, spice 3.2: 3656.51 (over 4746 flashers)

No unfolding:

anizotropy	✓	✗	✗
tilt	✓	✓	✗
SPICE Mie	4664.05	5735.30	6649.60
SPICE Lea	4415.63	5746.48	6504.19
SPICE 3.2	4330.60	5303.99	6277.33

By removing anizotropy and, if necessary tilt, from SPICE 3.2, the resulting ice model is improved over SPICE Mie and SPICE Lea

Ice model comparison: model error

Nominal DOM efficiencies, h2-50cm hole ice model, 1x simulation (10 events)

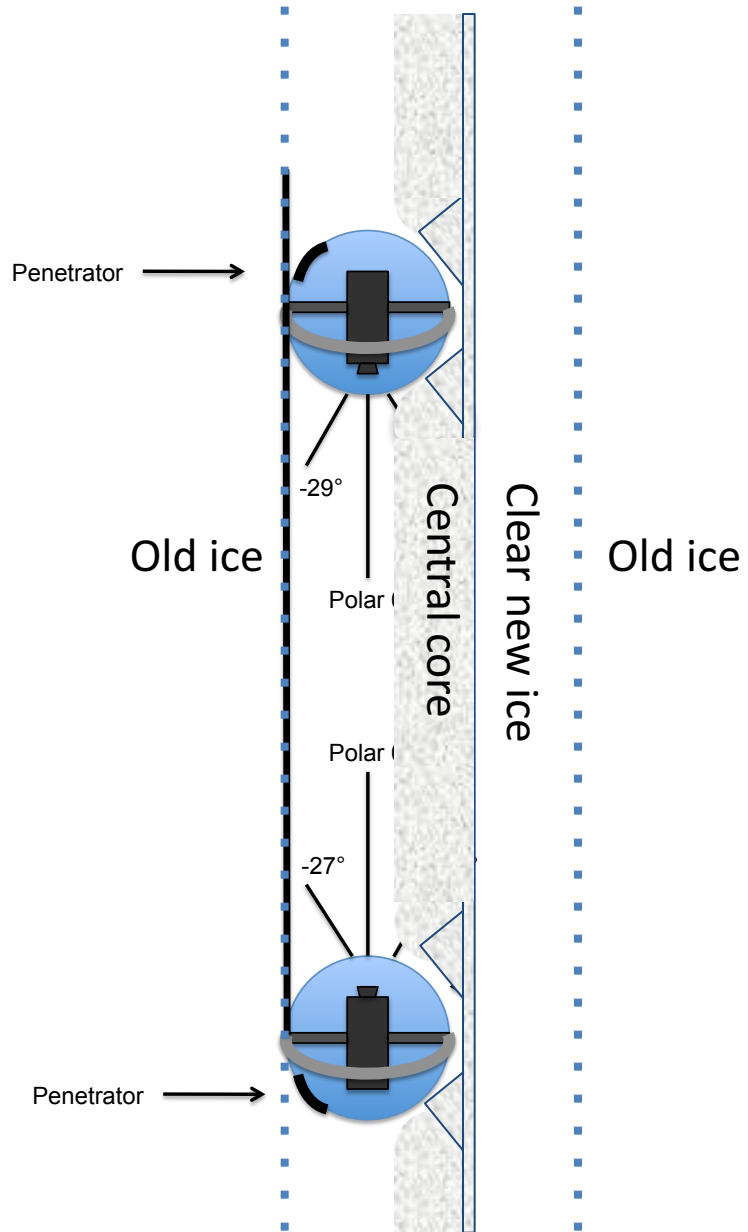
With flasher LED pattern unfolding, spice 3.2: 14.7%

No unfolding:

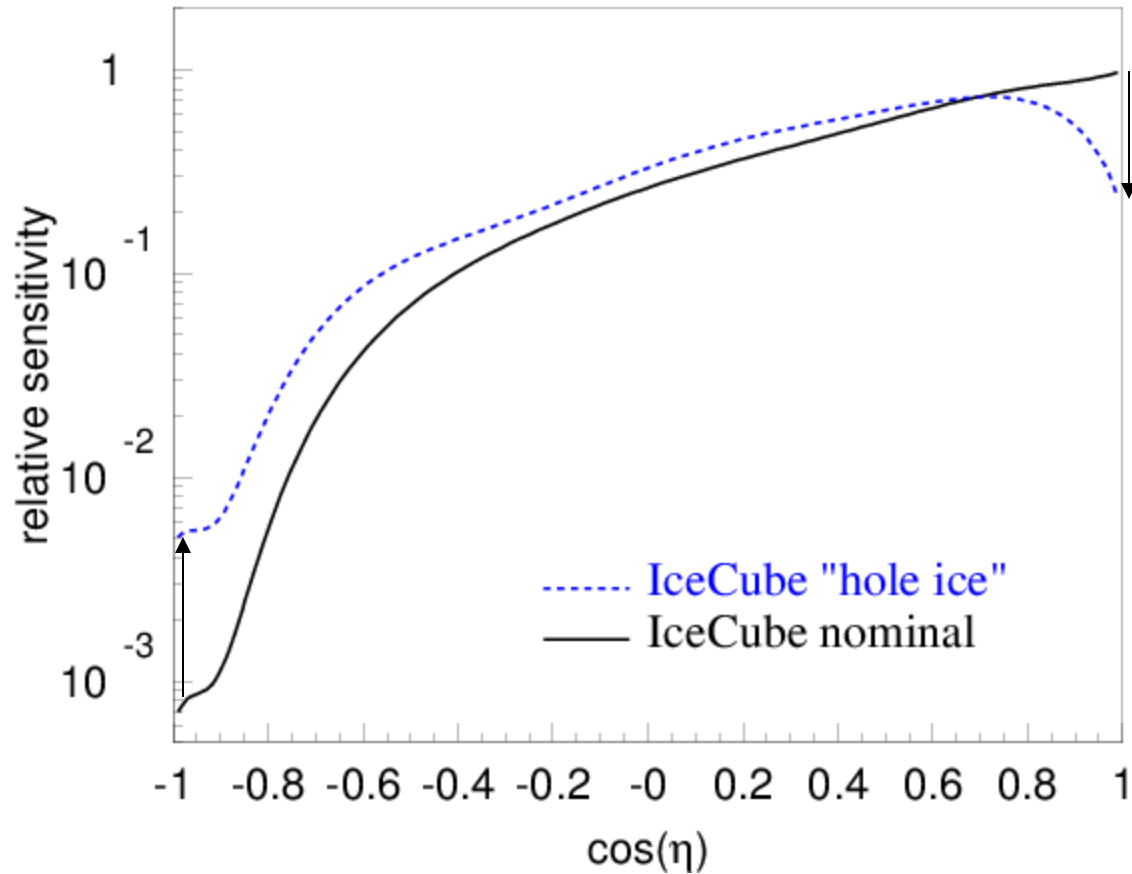
anizotropy	✓	✗	✗
tilt	✓	✓	✗
SPICE Mie	17.4%	26.6%	29.6%
SPICE Lea	18.6%	27.5%	30.0%
SPICE 3.2	16.5%	25.4%	28.6%

By removing anizotropy and, if necessary tilt, from SPICE 3.2, the resulting ice model is improved over SPICE Mie and SPICE Lea

Hole ice



Traditional “hole ice” angular sensitivity

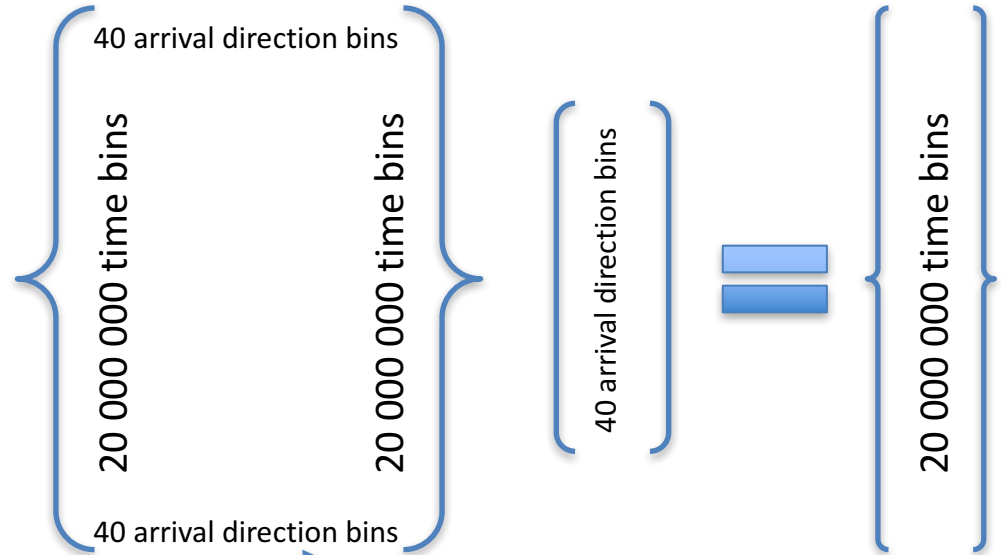
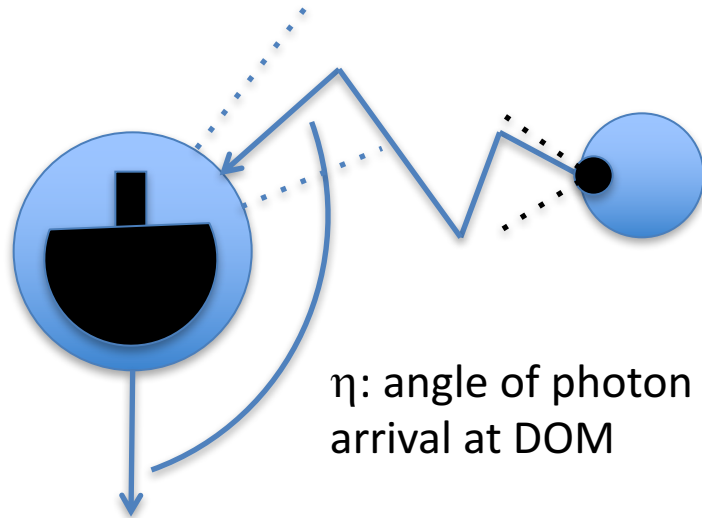


Unfolding the angular sensitivity

Same likelihood construction as in SPICE³ fit to flasher data

Nuisance parameters:

5046 receiver DOM efficiencies
(72+2)*4746 flasher parameters

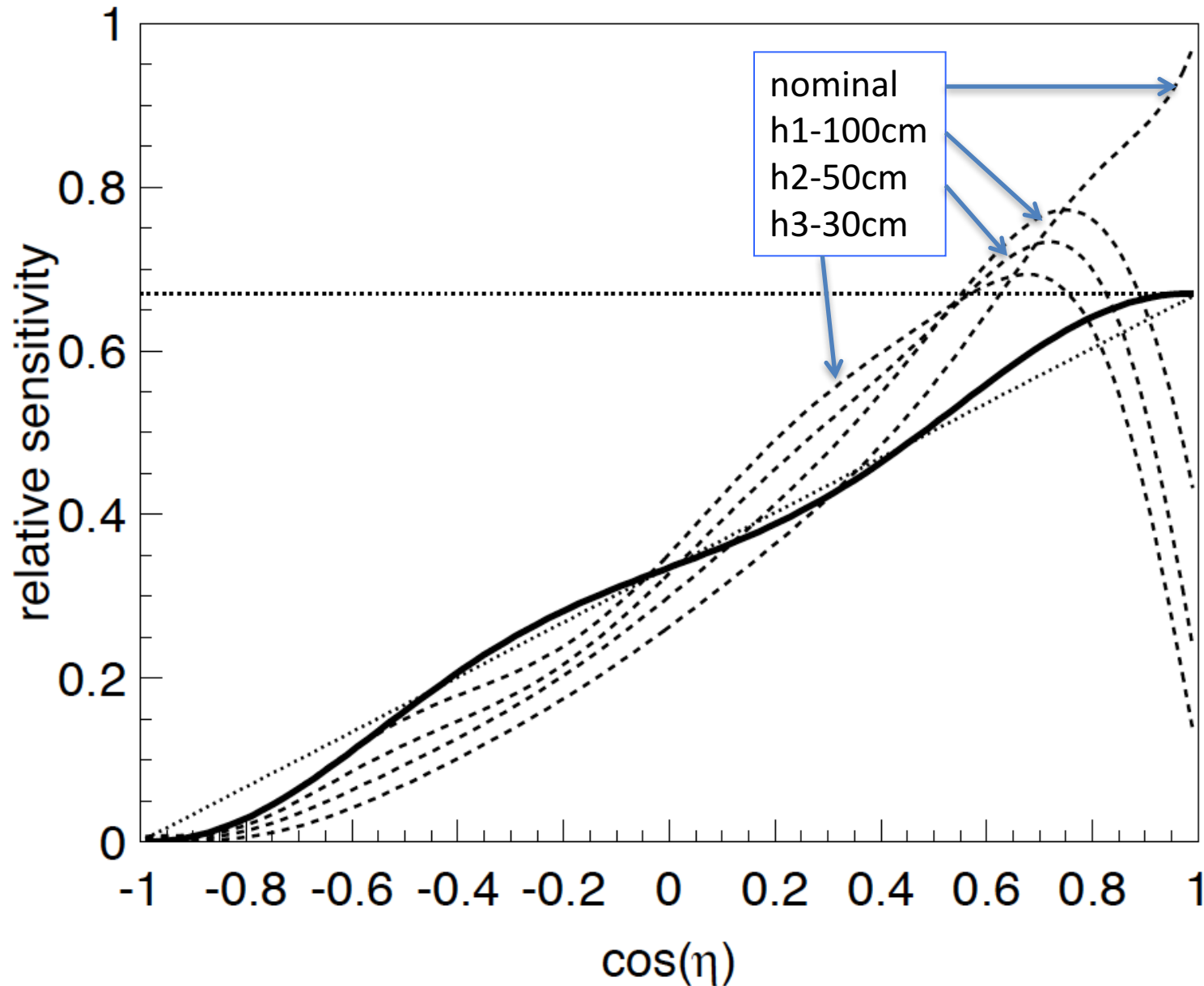


Unfolding matrix (from simulation):

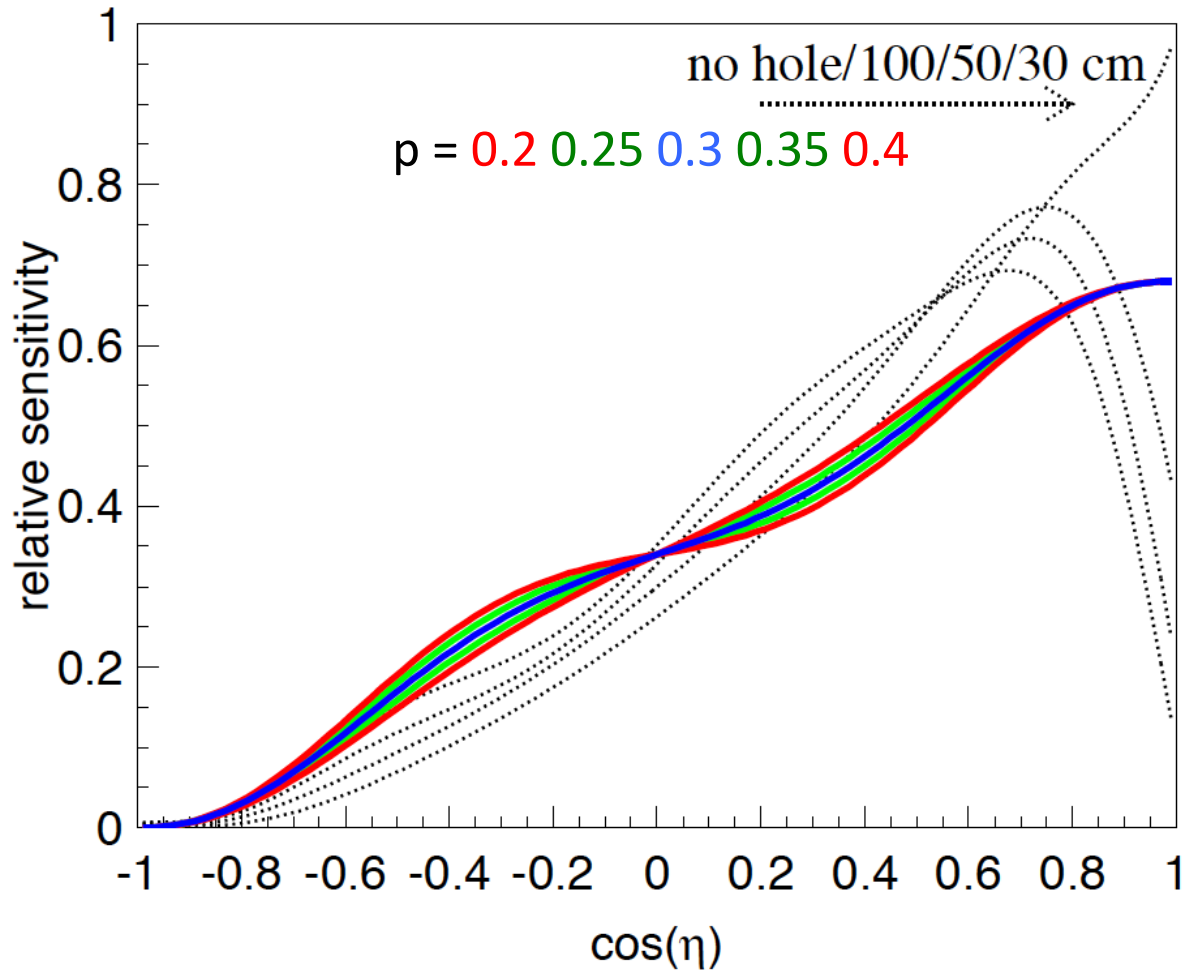
40 $\cos(\eta)$ unfolding bins

20 000 000 charge values of emitter-receiver time bin pairs (4746 flashers)

Best fit to all-string flasher set



Fitted shape parameter: $p=0.3\pm 0.1$



An unfolded solution as fitted to the all-purpose flasher data

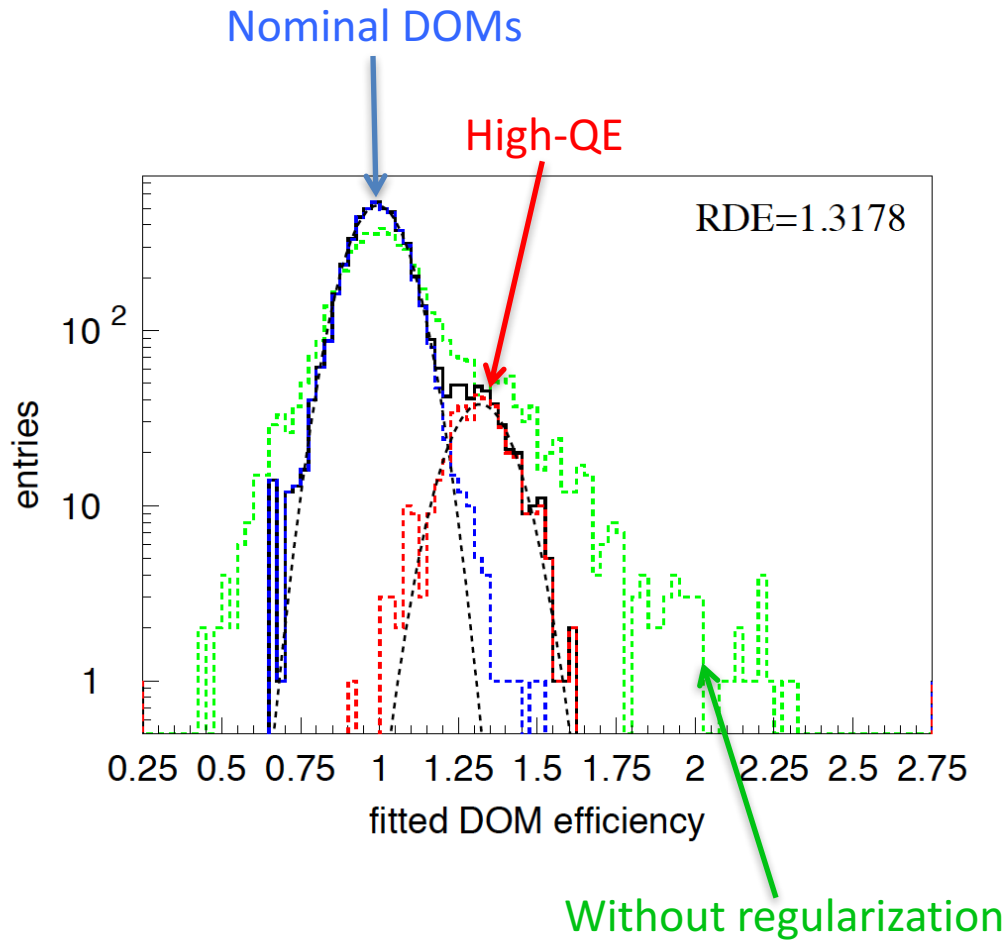
Hole ice

Low-energy group developed an extension that uses an additional shape parameter (p_2) to describe the peak

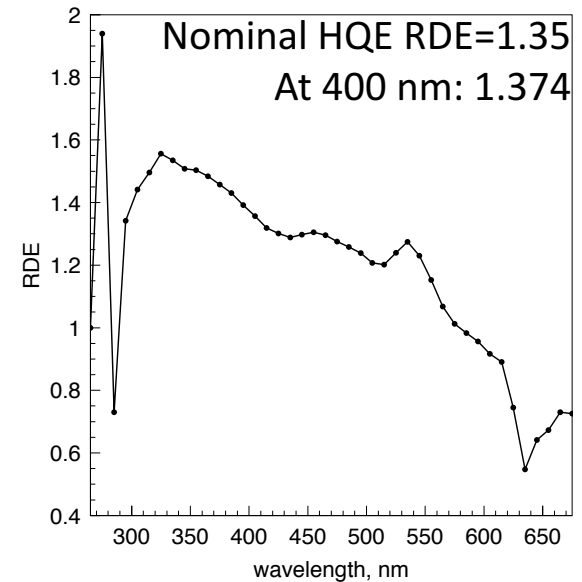
Martin R. has DARD-derived and lab measurement models, with an ice model extension, SPICE HD

Relative DOM efficiencies (RDE)

DOM efficiencies



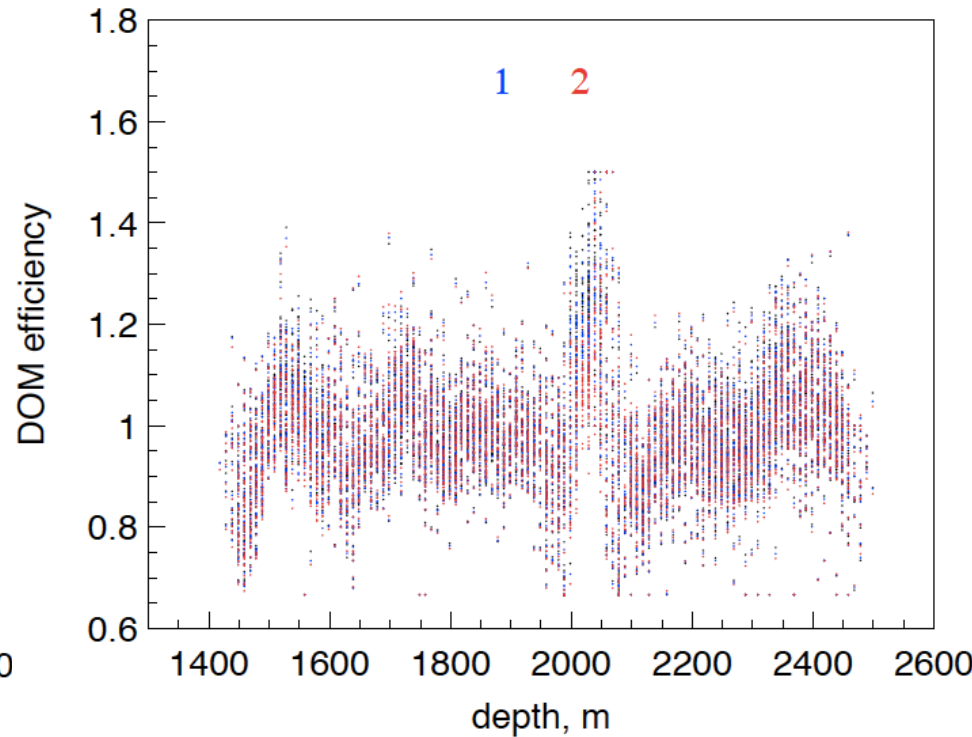
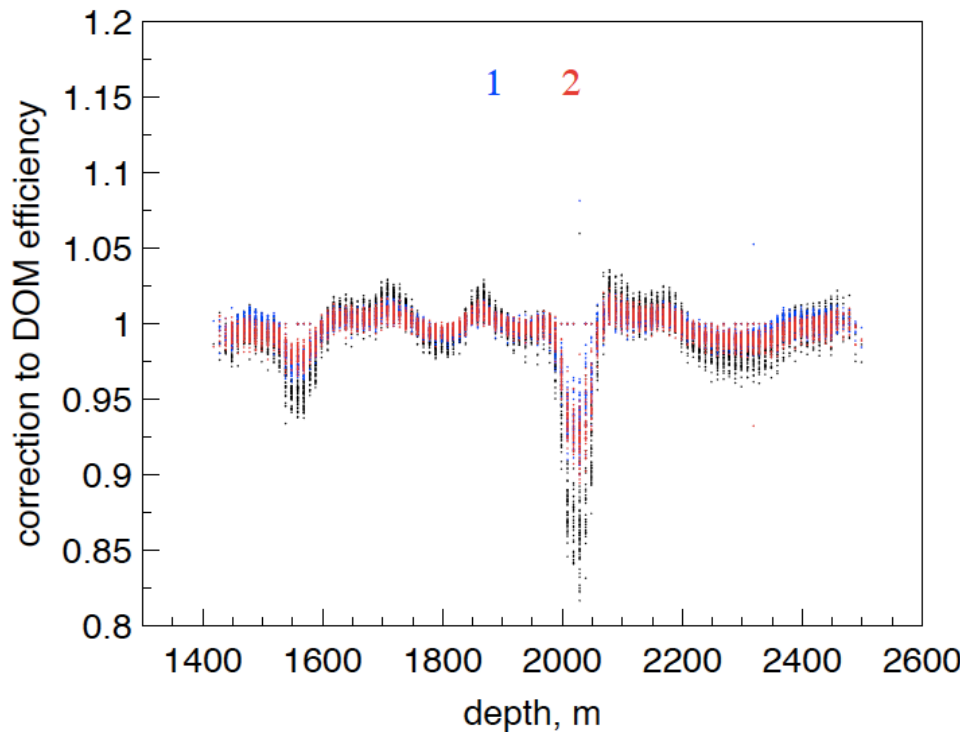
Use fitted RDE table and two wavelength acceptance curves (nominal and high-QE) simultaneously



(available in ice-models module, directory resources/models/spice-latest-full)

Derived DOM efficiencies

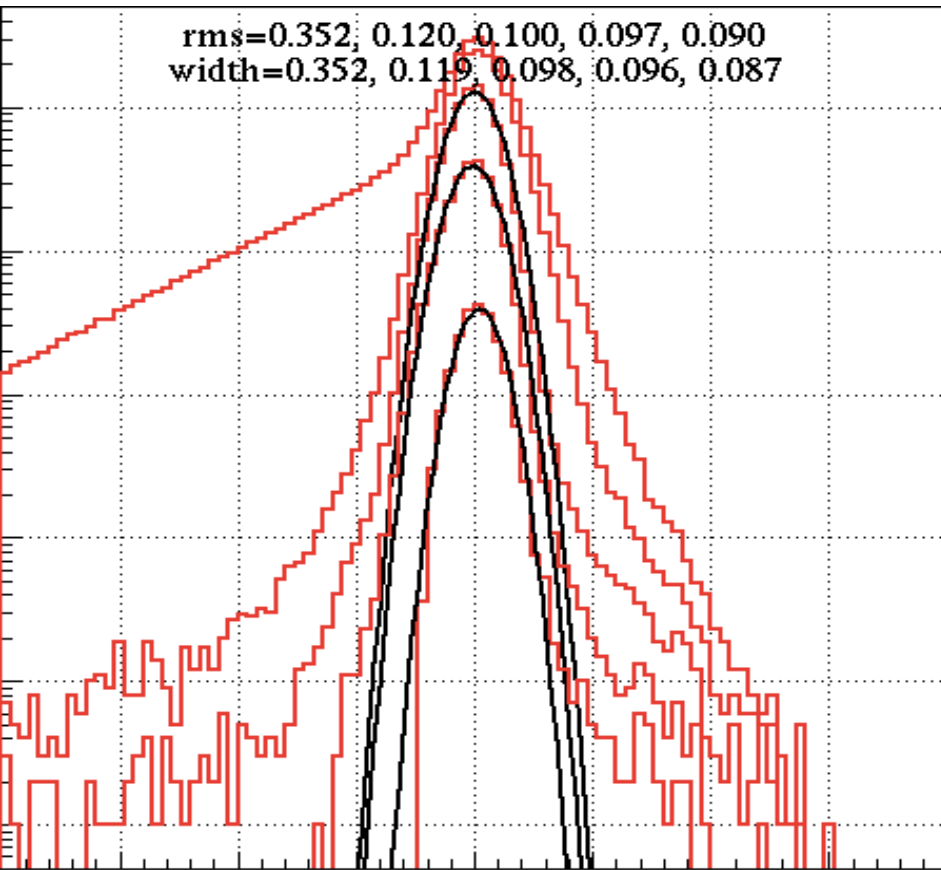
Shown are nominal DOMs only



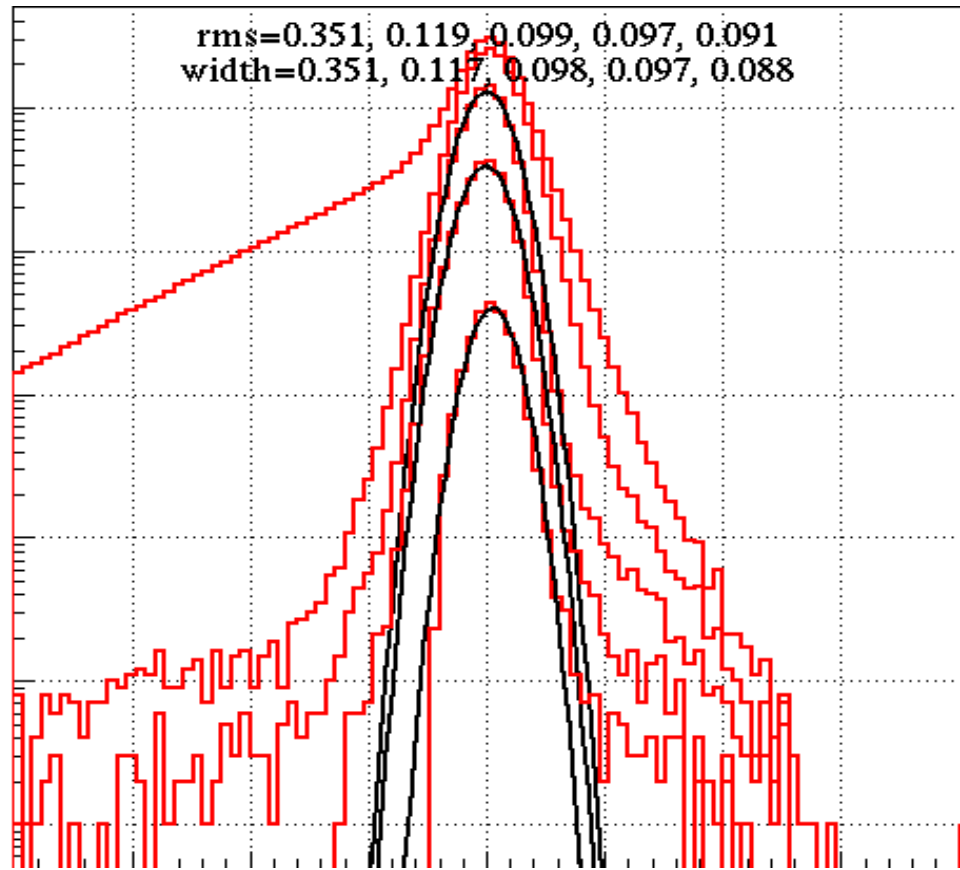
Cumulative change shown in black (in plot on the left)
(available as part of ice-models module in [resources/models/spice-latest-full/eff-f2k](#))

Other

Model error



SPICE 3.1



SPICE 3.2

100 sim. events (10x statistics)

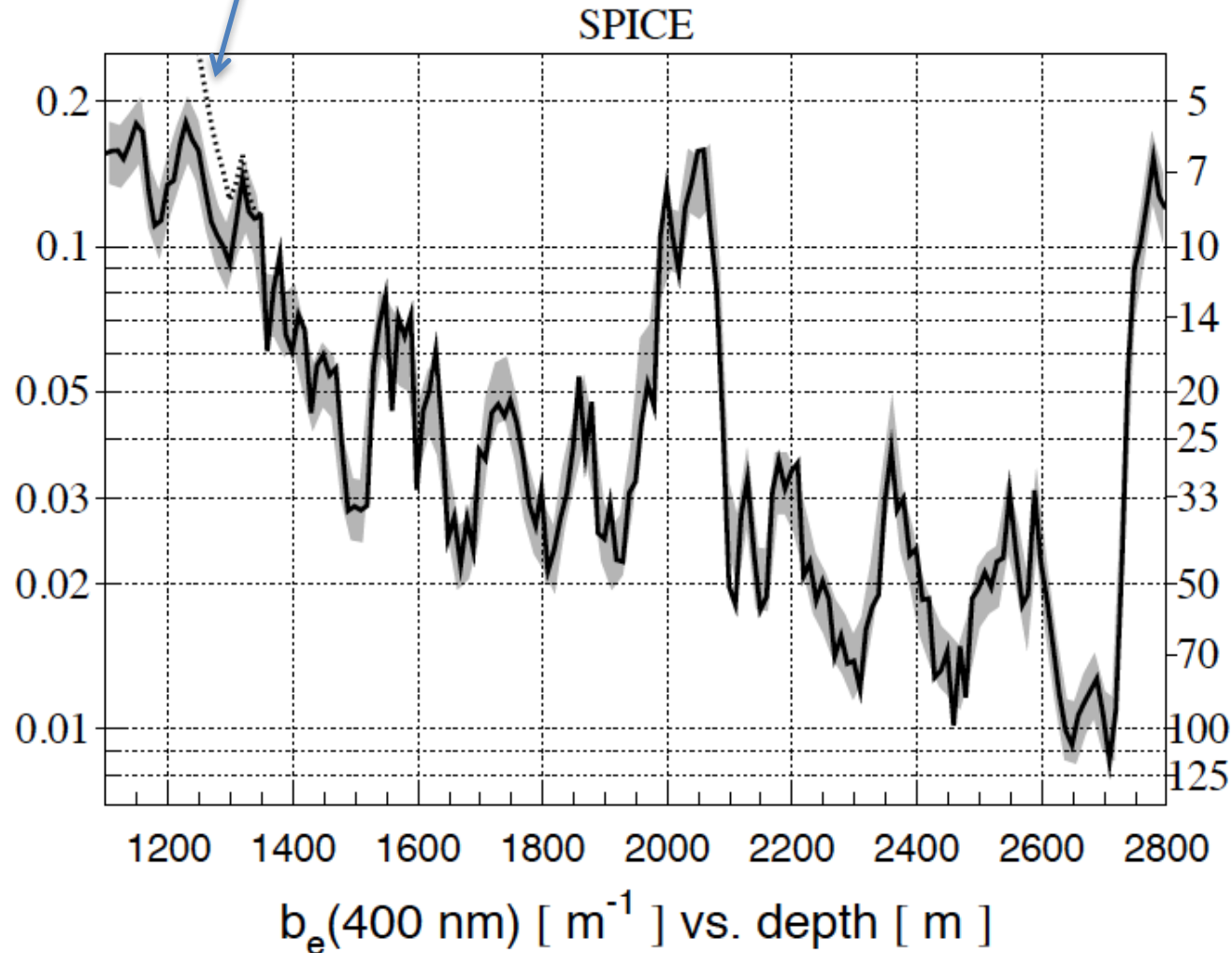
benchmark value unchanged at 9.8%

Very slight improvement for smaller charge: 11.9 → 11.7

Air bubbles at shallow depths

Parameterized in file icemodel.bbl

Air bubble contribution (from WHAM!):
 $11.7 * (1 - d/1350) * (1 - d/1400)$



Backup: anisotropy coefficient

$$\kappa_1 = 10.6\%$$

0: all directions

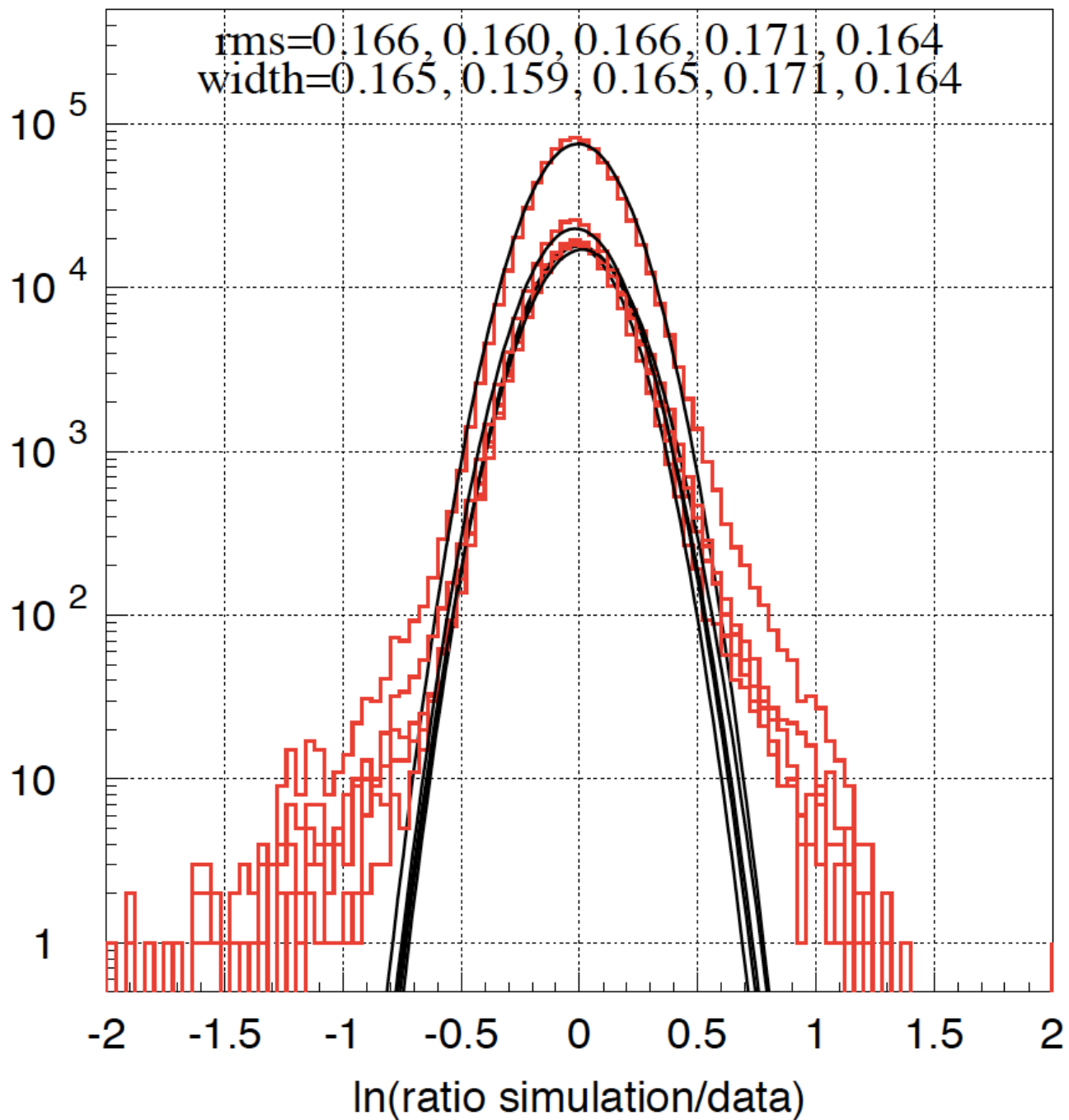
1,3: flasher-receiver directions within 45 degrees of major anisotropy axis

2,4: flasher-receiver directions within 45 degrees of minor anisotropy axis

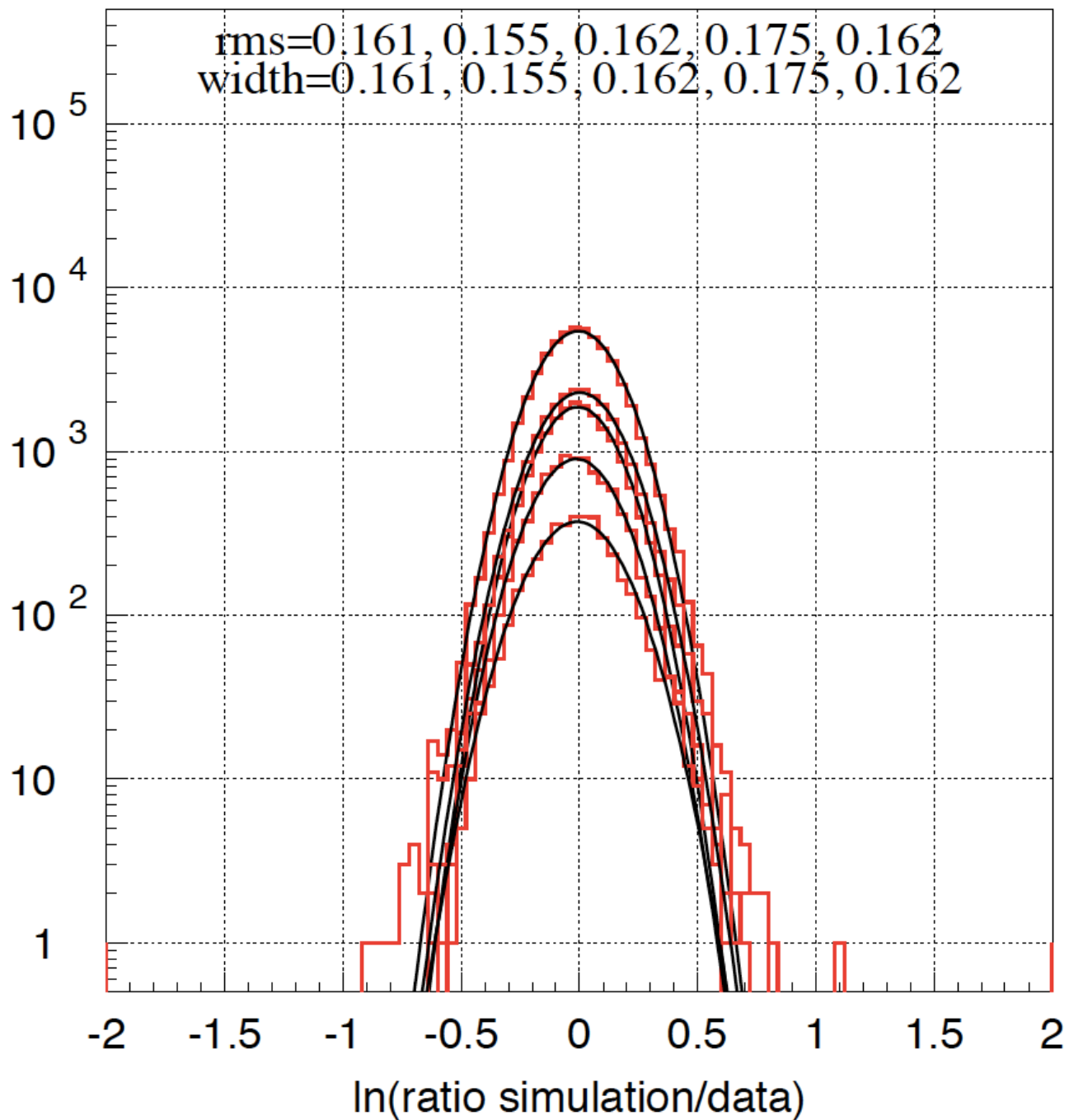
No unfolding, nominal DOM efficiencies, h2-50cm hole ice

GOF=4330.6

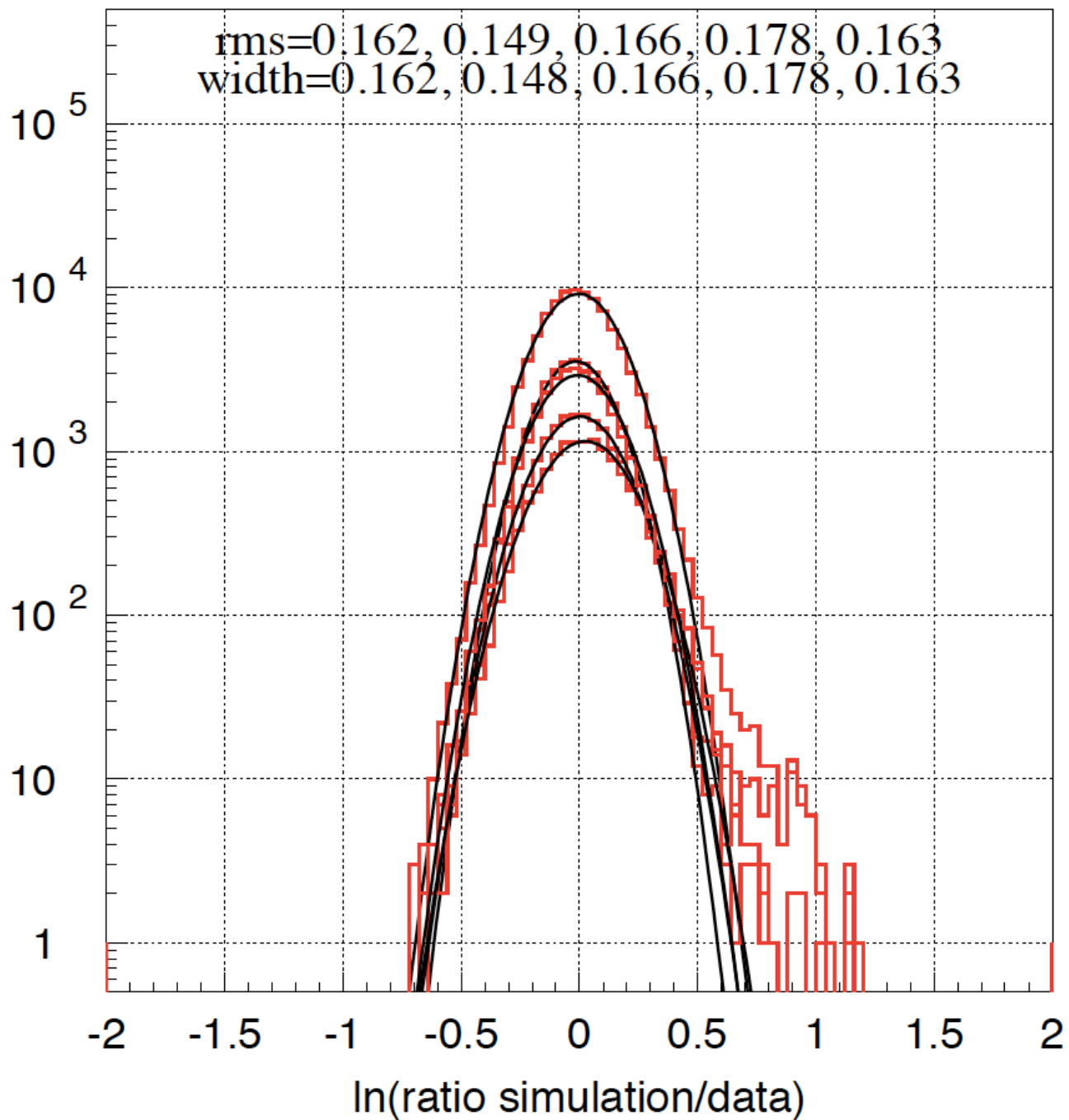
All flashers



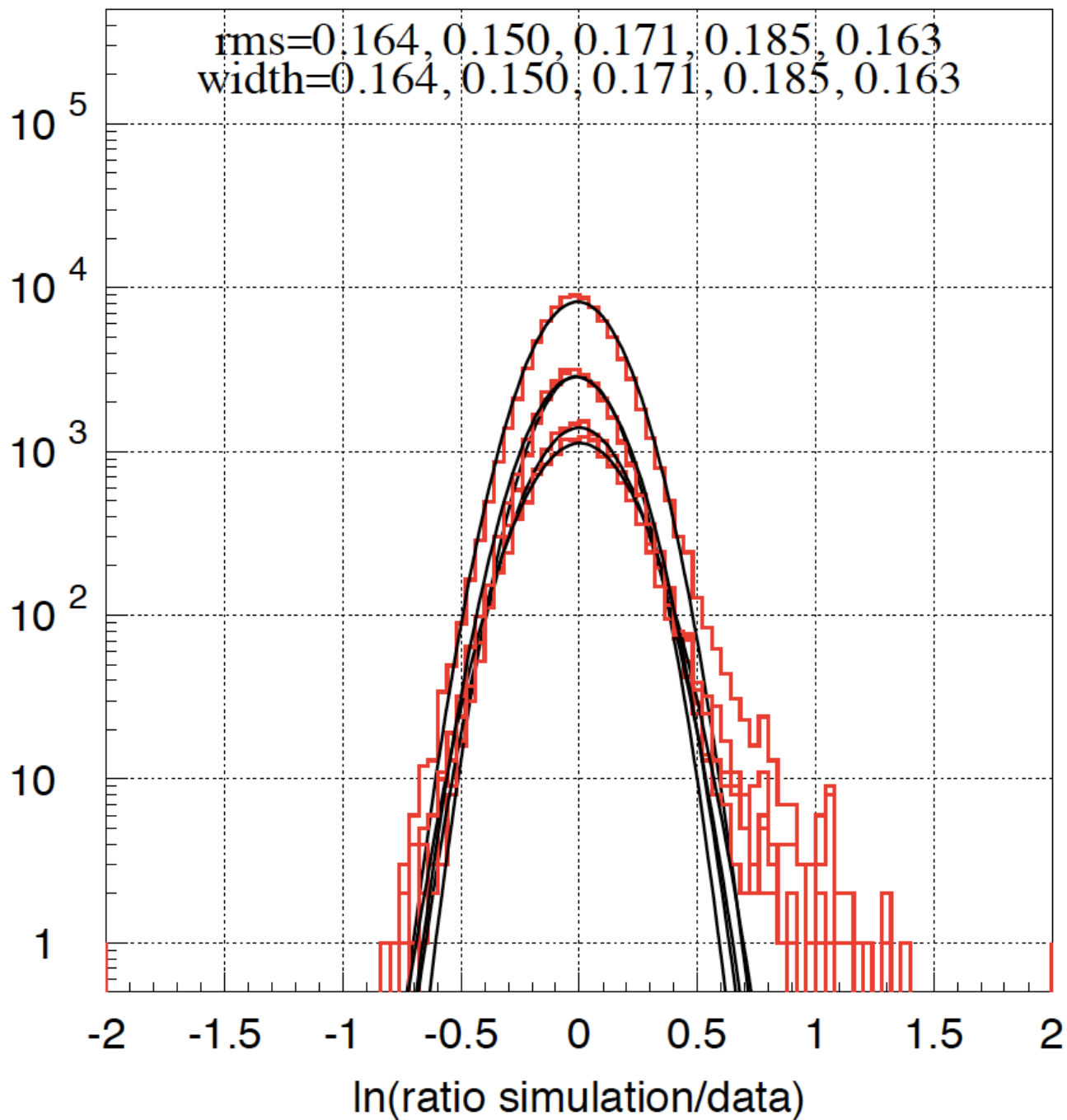
Flashes 1...10



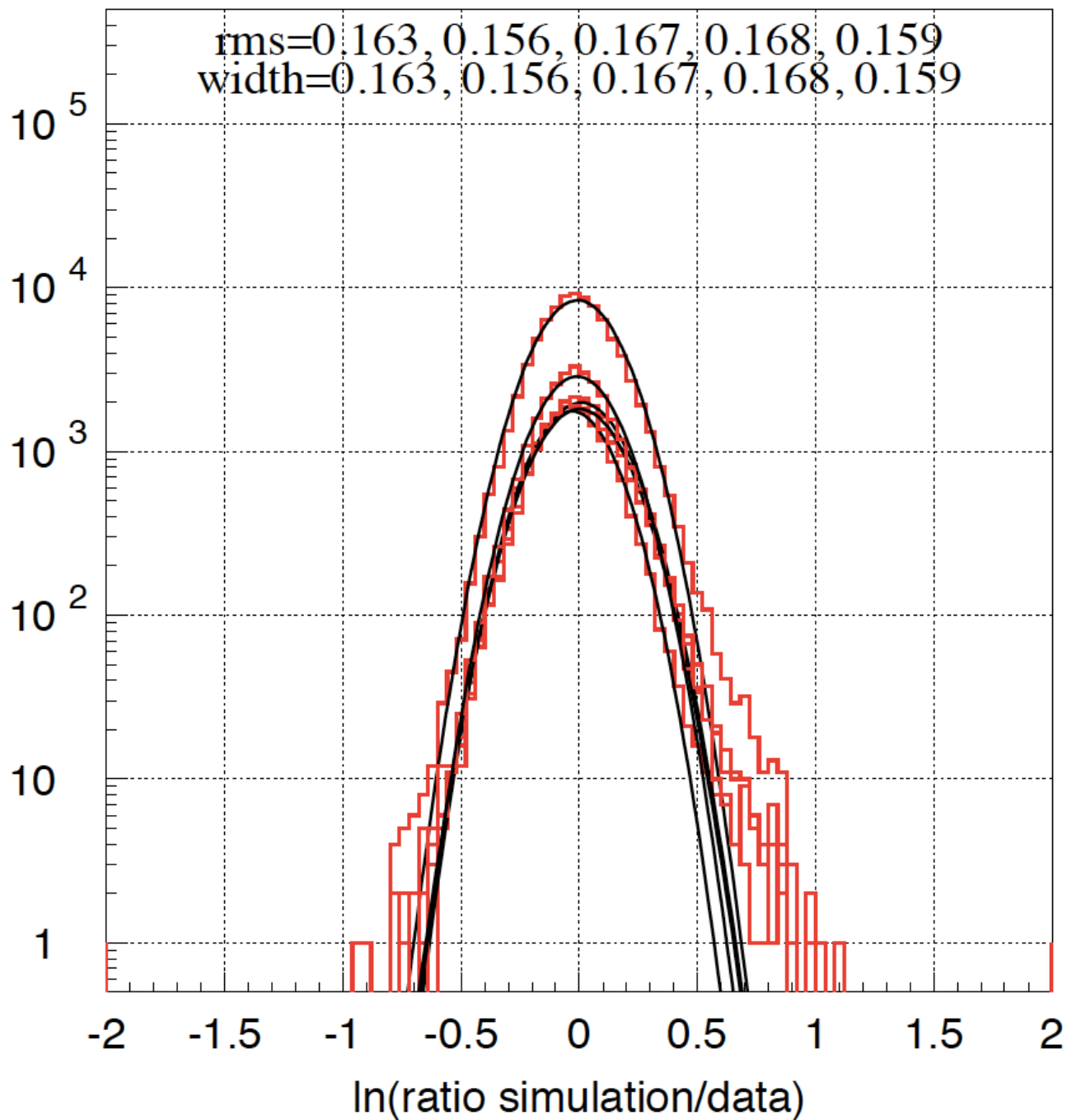
Flashes 11...20



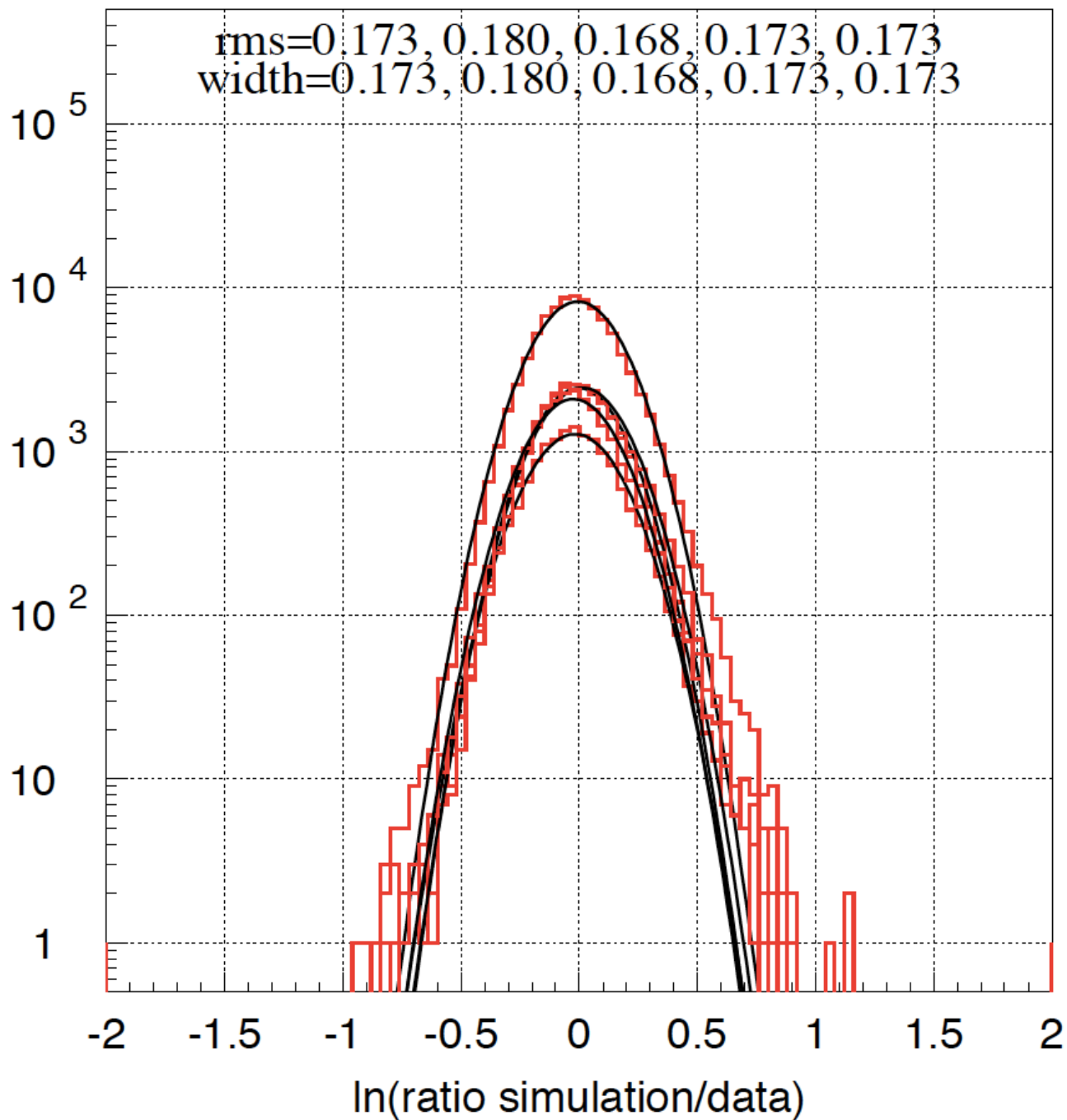
Flashers 21...30



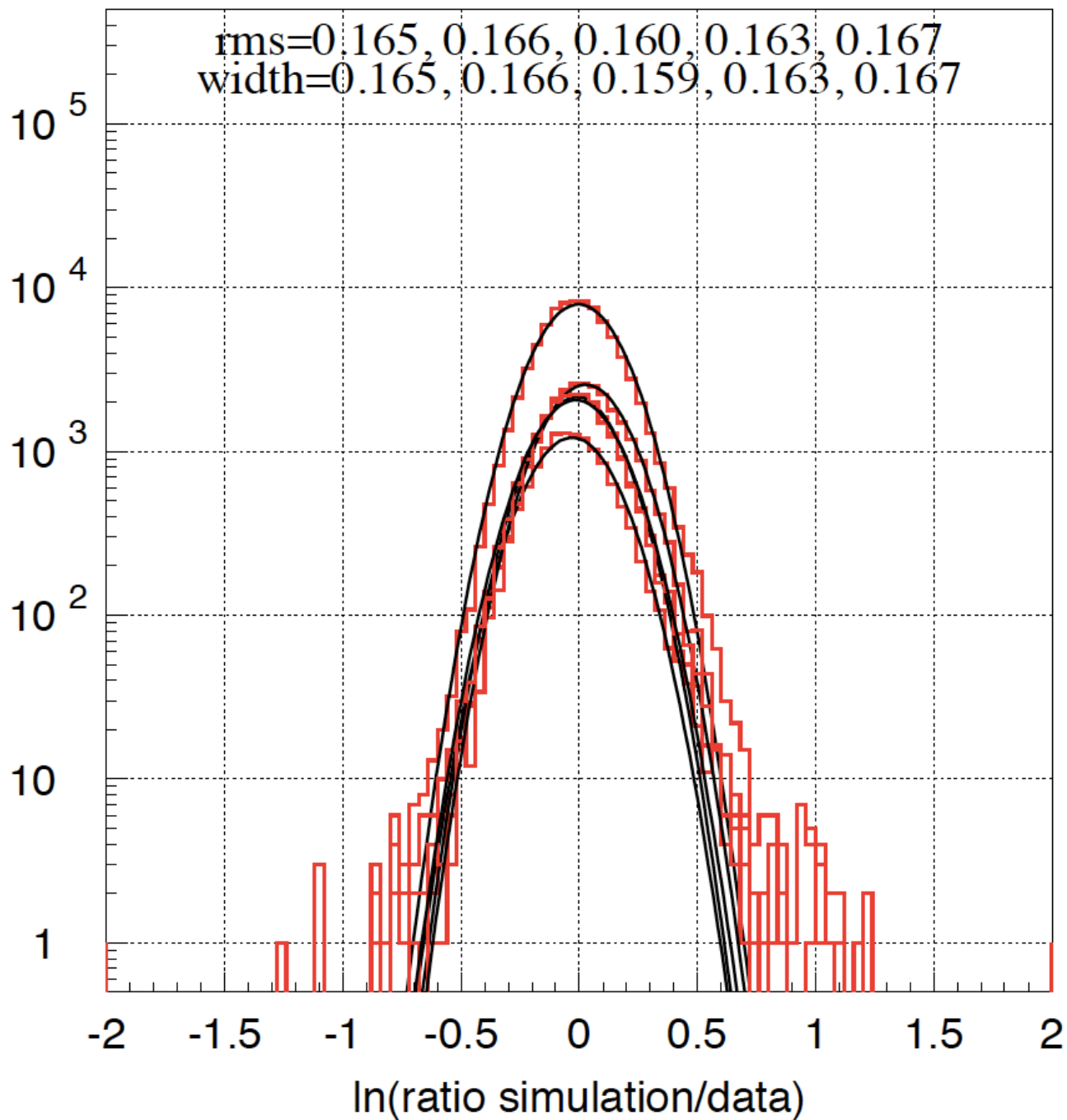
Flashes 31...40



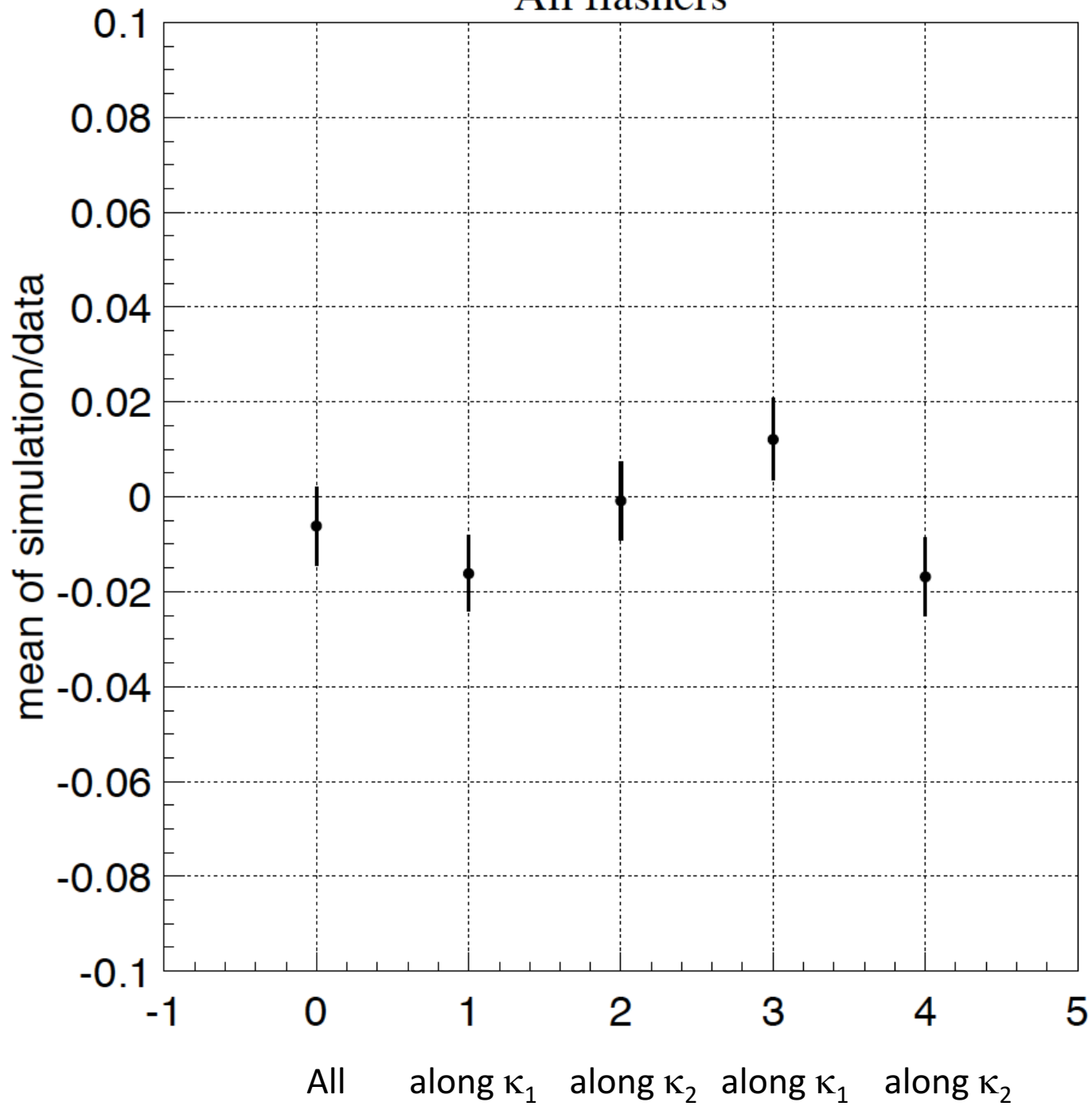
Flashes 41...50



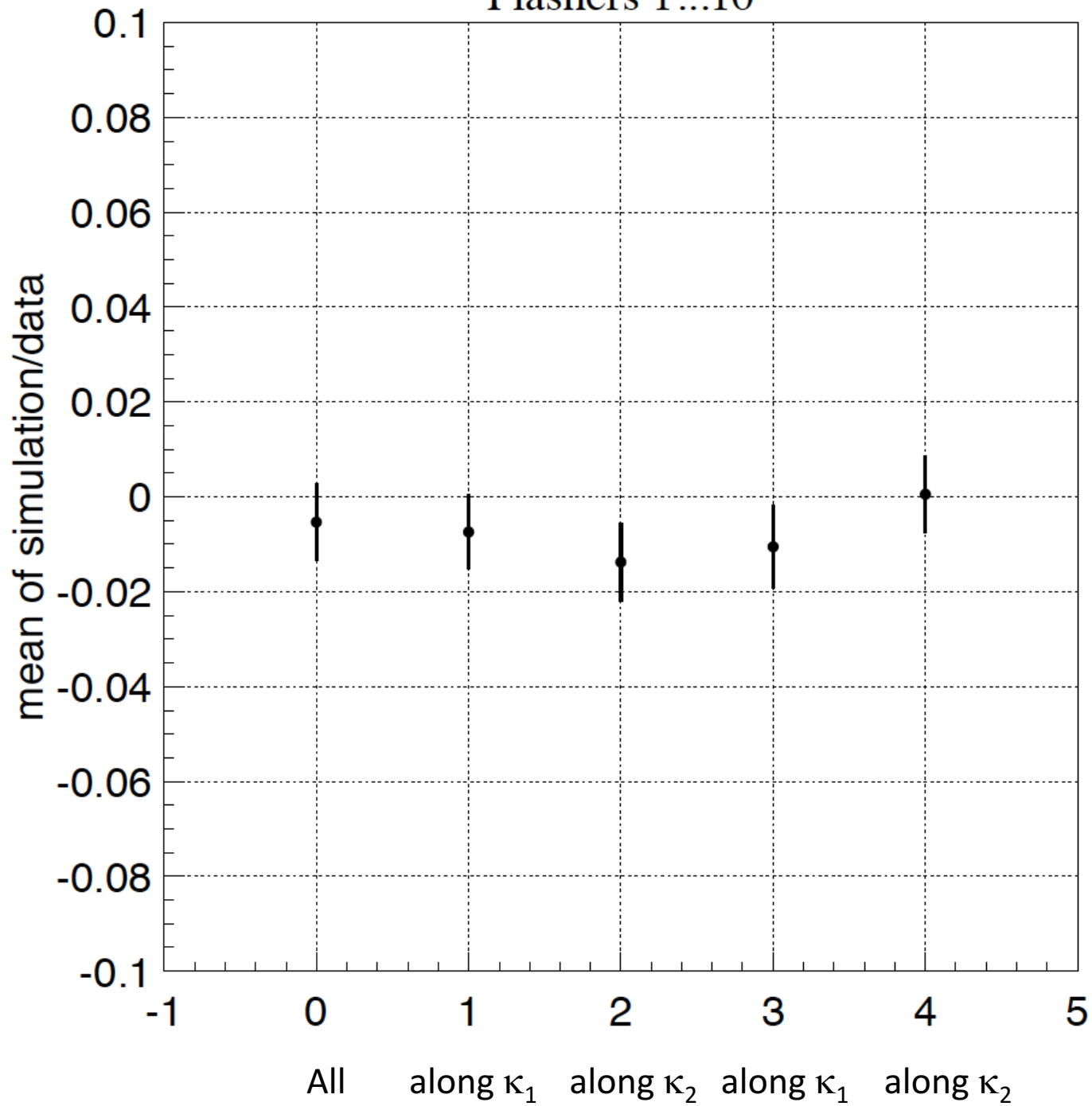
Flashes 51...60



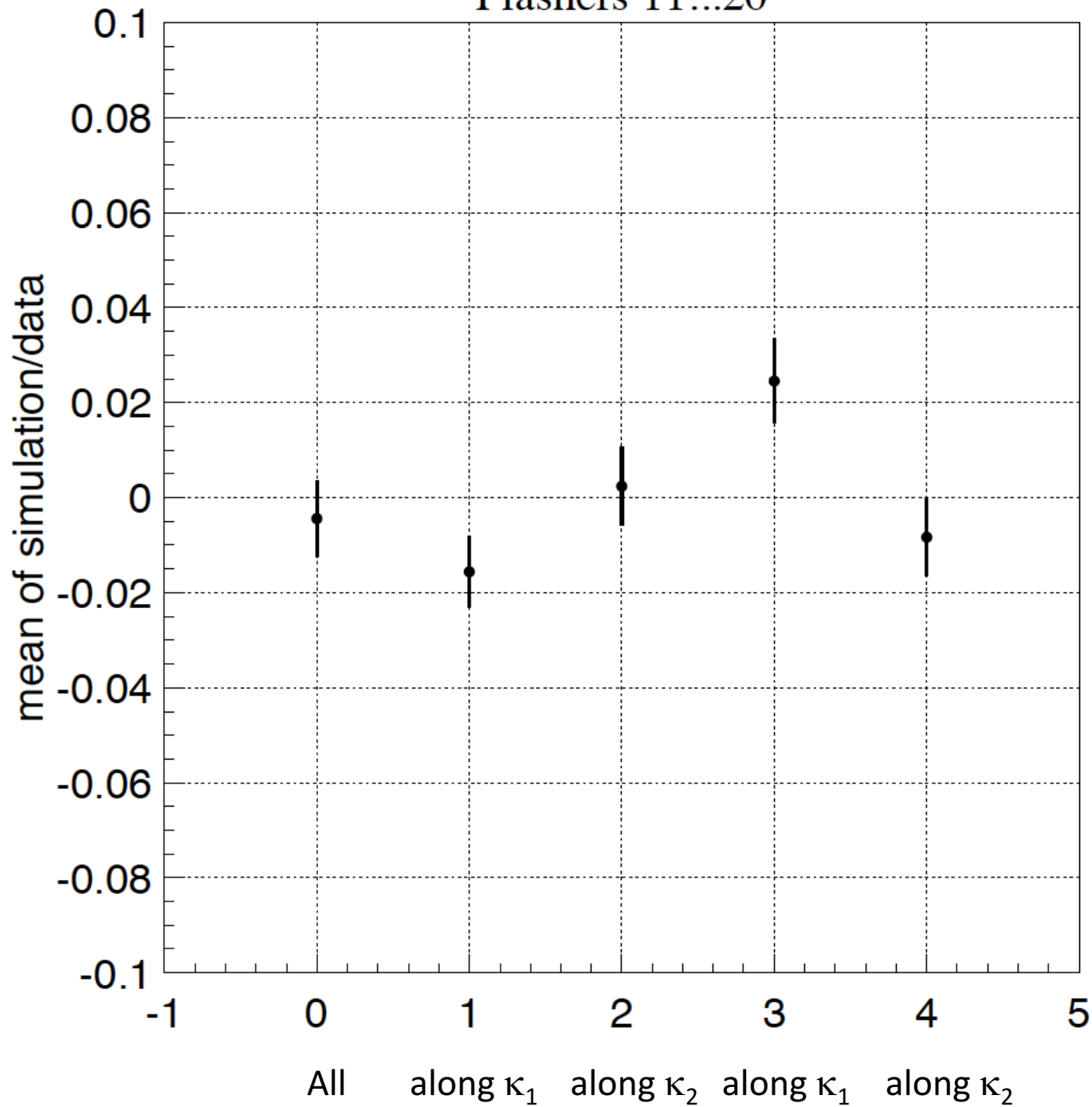
All flashers



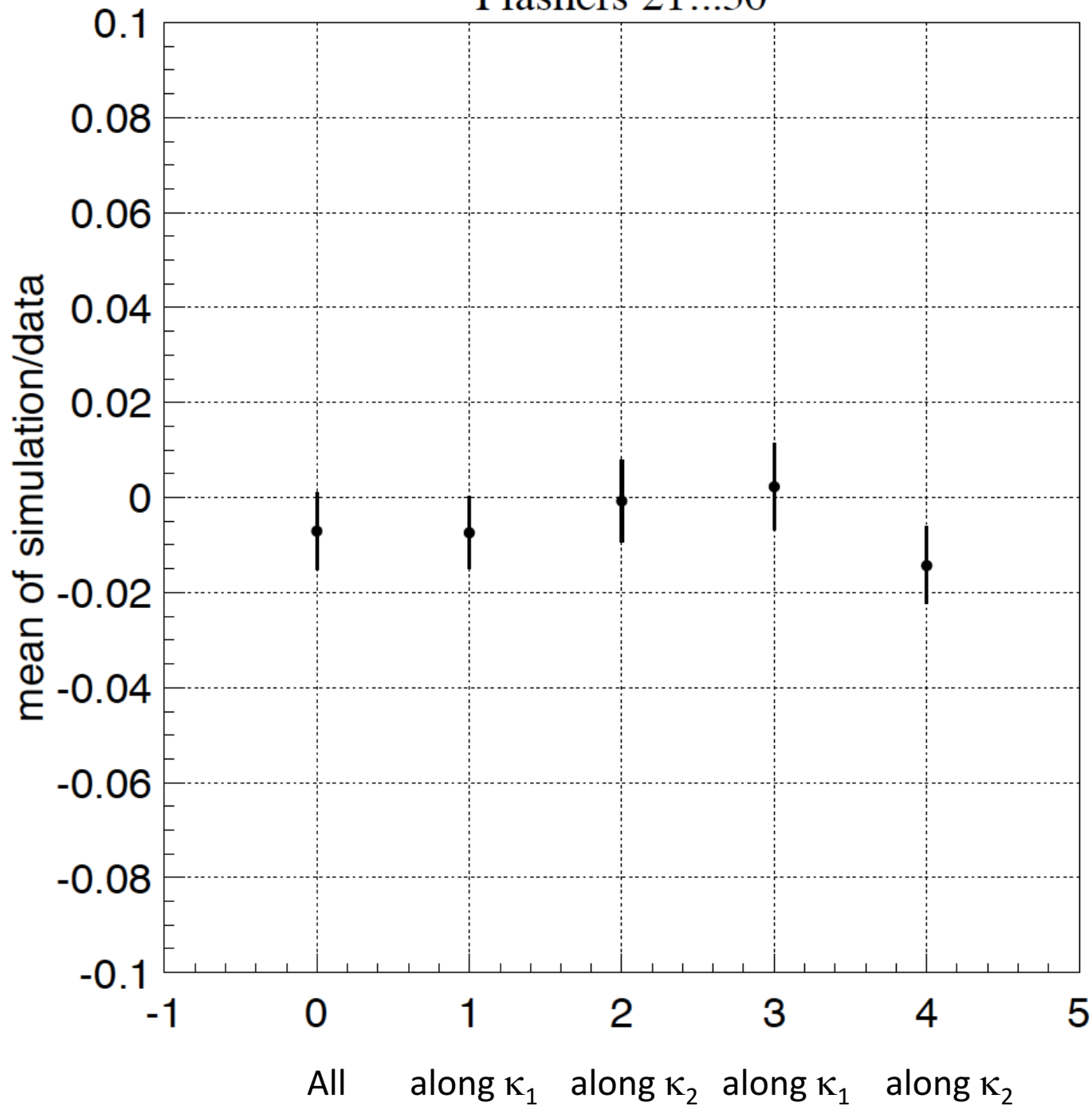
Flashers 1...10



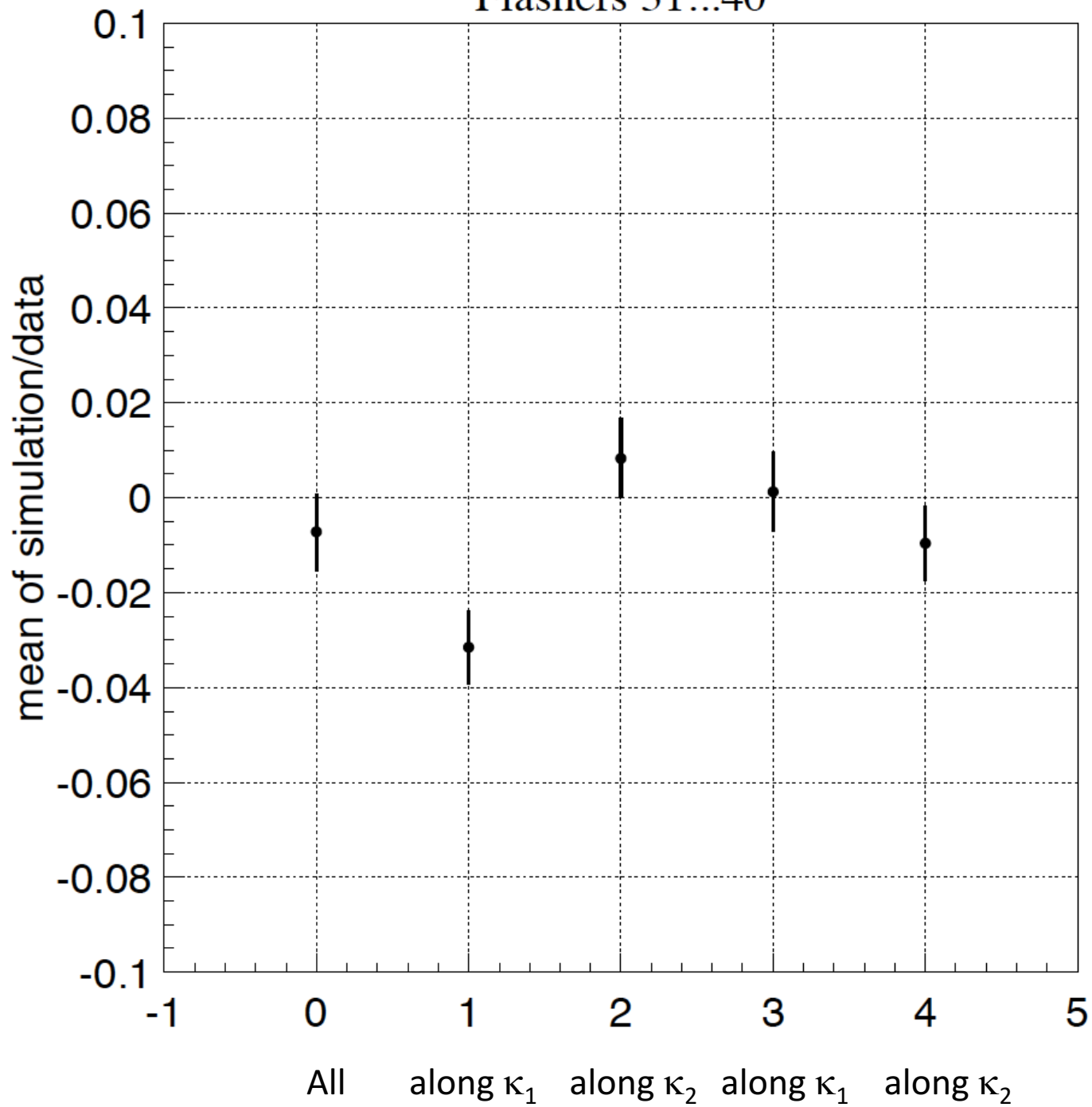
Flashers 11...20



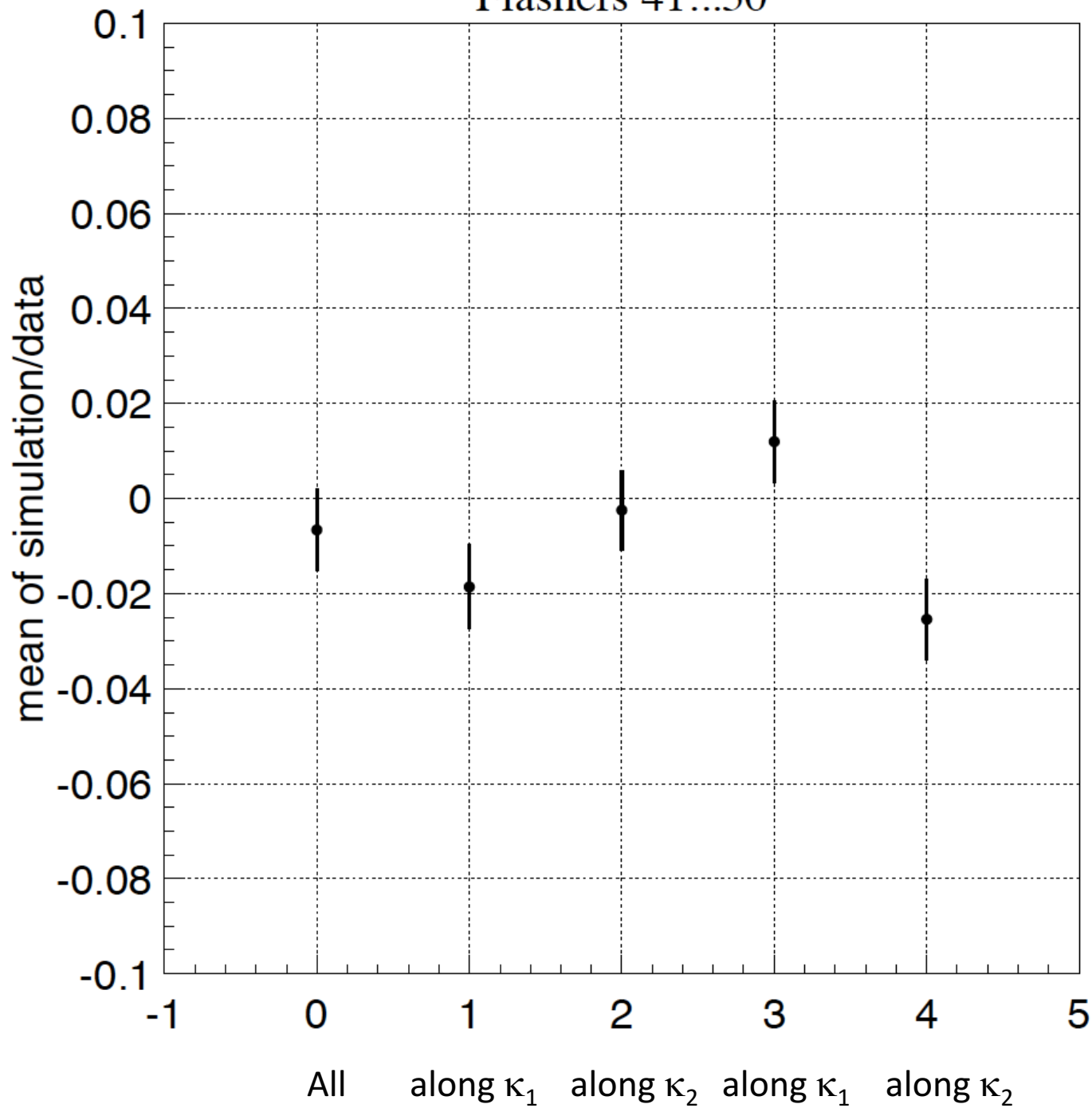
Flashers 21...30



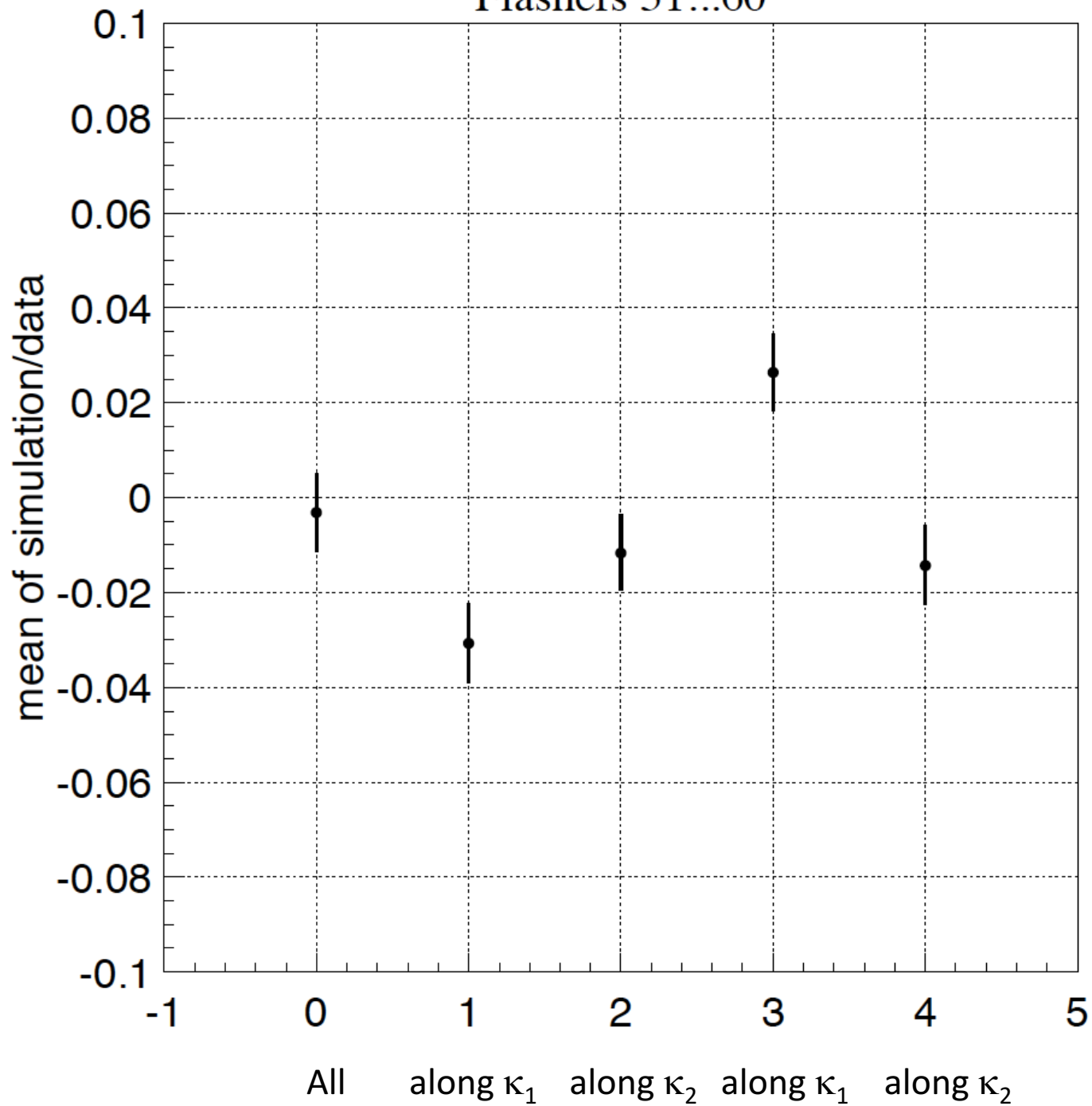
Flashers 31...40



Flashers 41...50



Flashers 51...60



$$\kappa_1 = 7.7\%$$

0: all directions

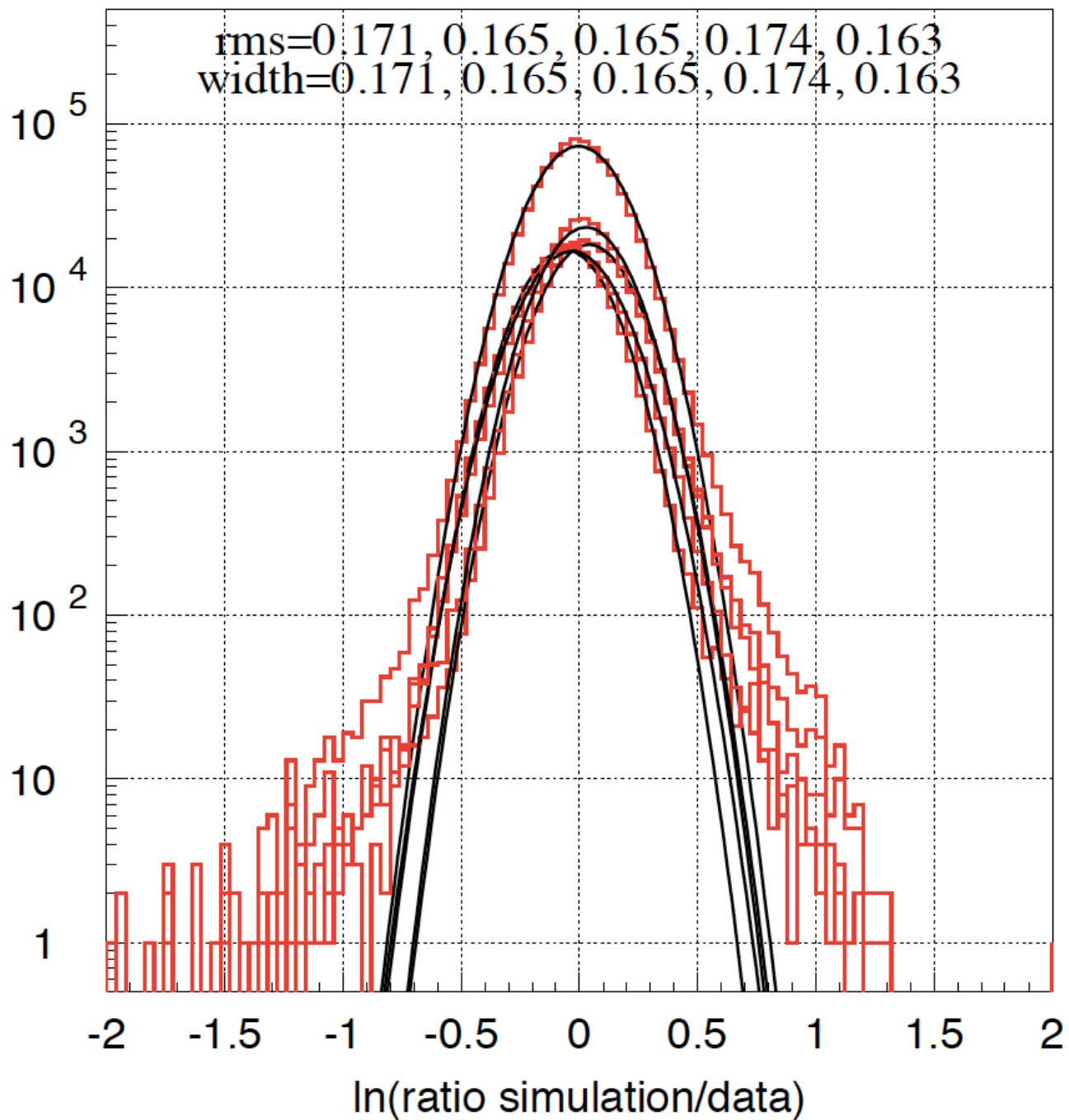
1,3: flasher-receiver directions within 45 degrees of major anisotropy axis

2,4: flasher-receiver directions within 45 degrees of minor anisotropy axis

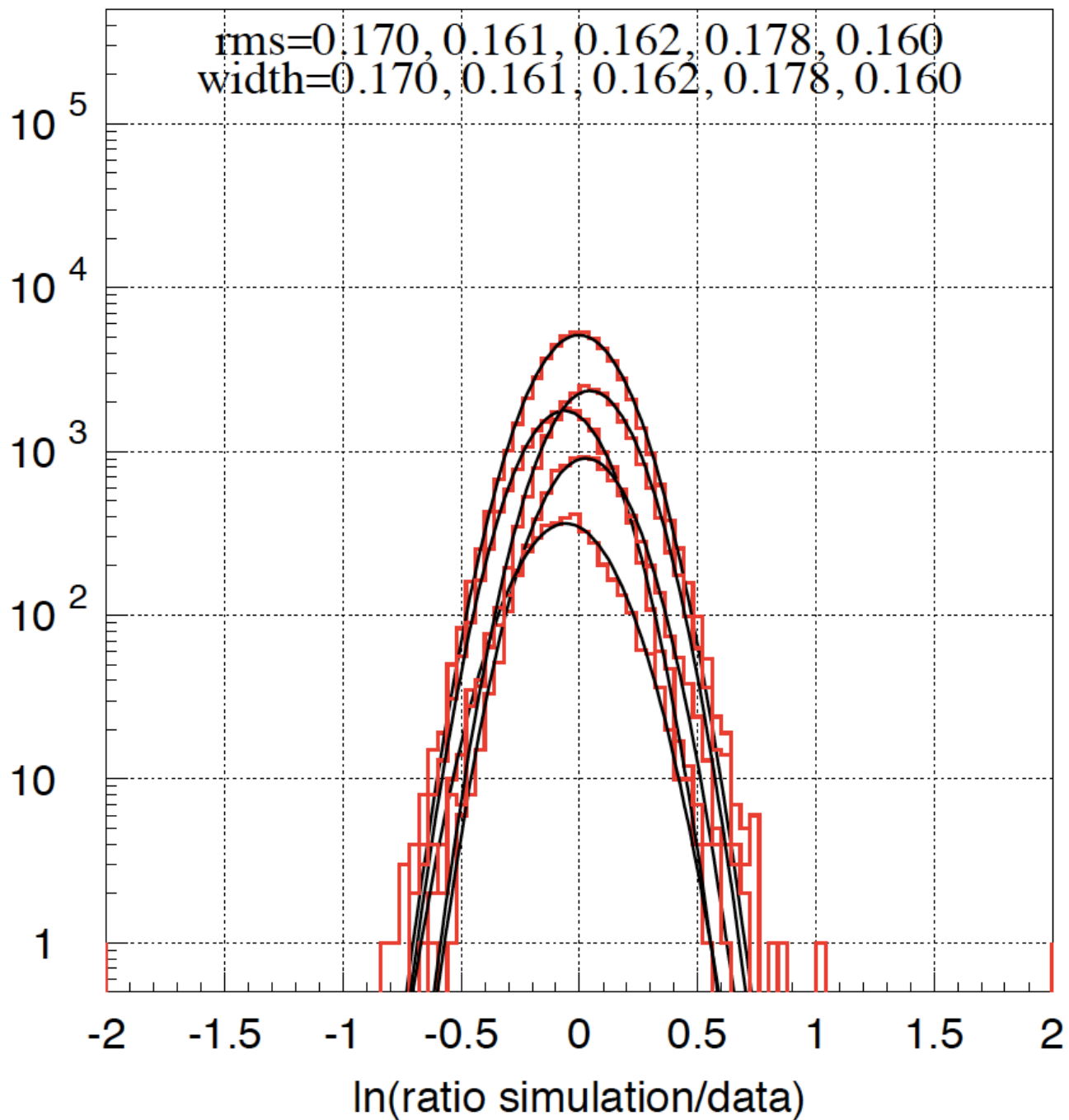
No unfolding, nominal DOM efficiencies, h2-50cm hole ice

GOF=4163.65

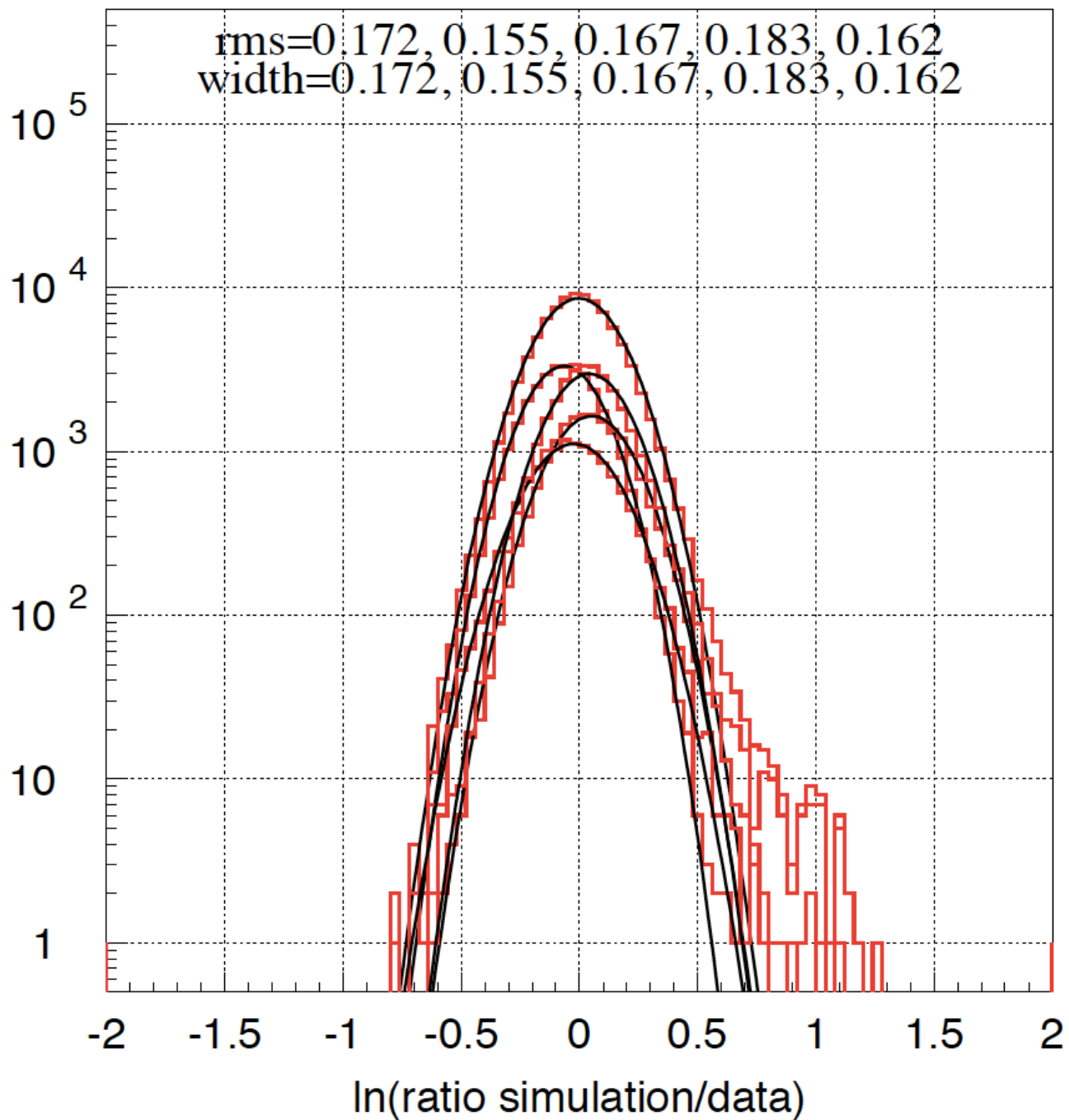
All flashers



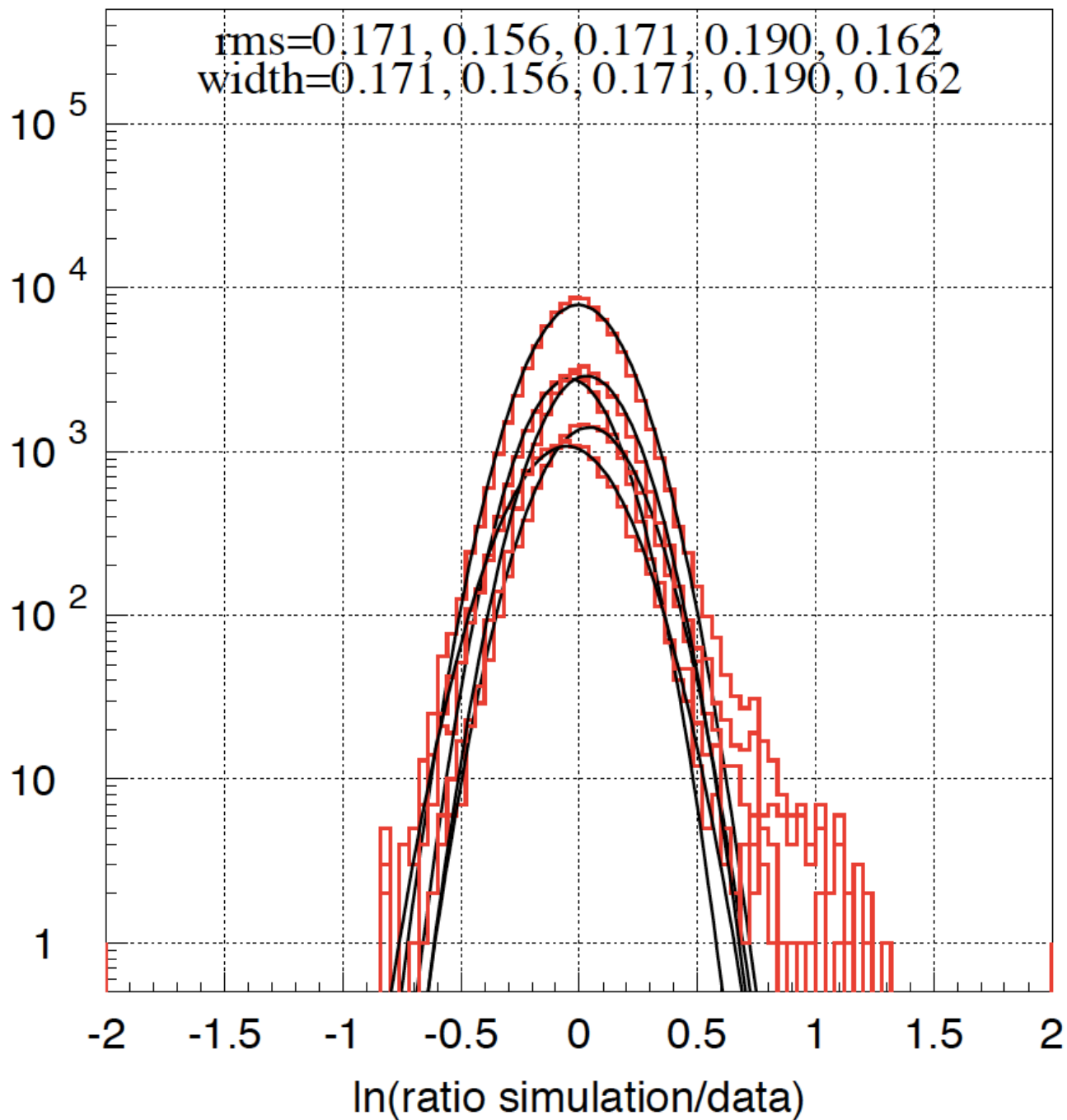
Flashes 1...10



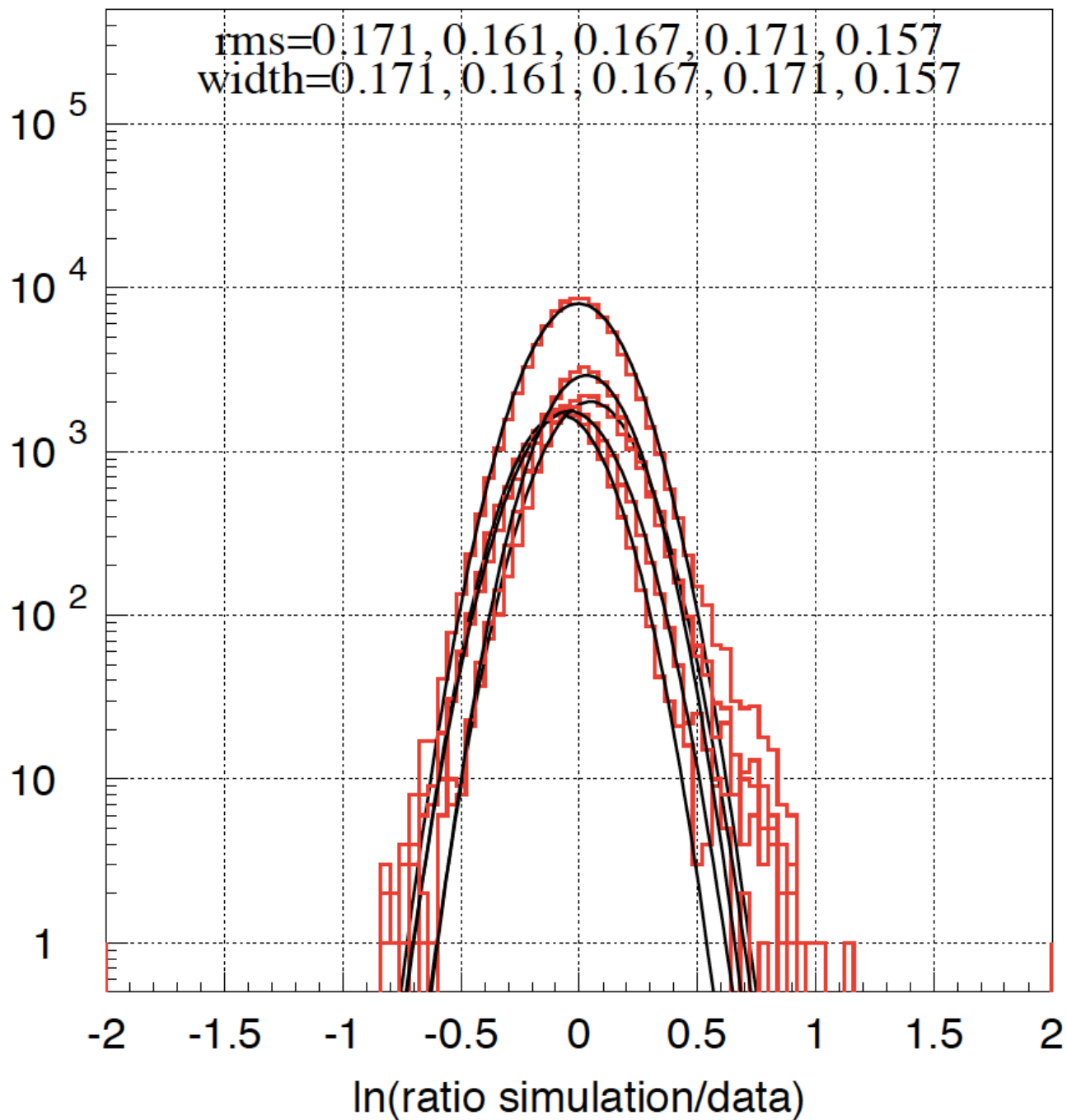
Flashers 11...20



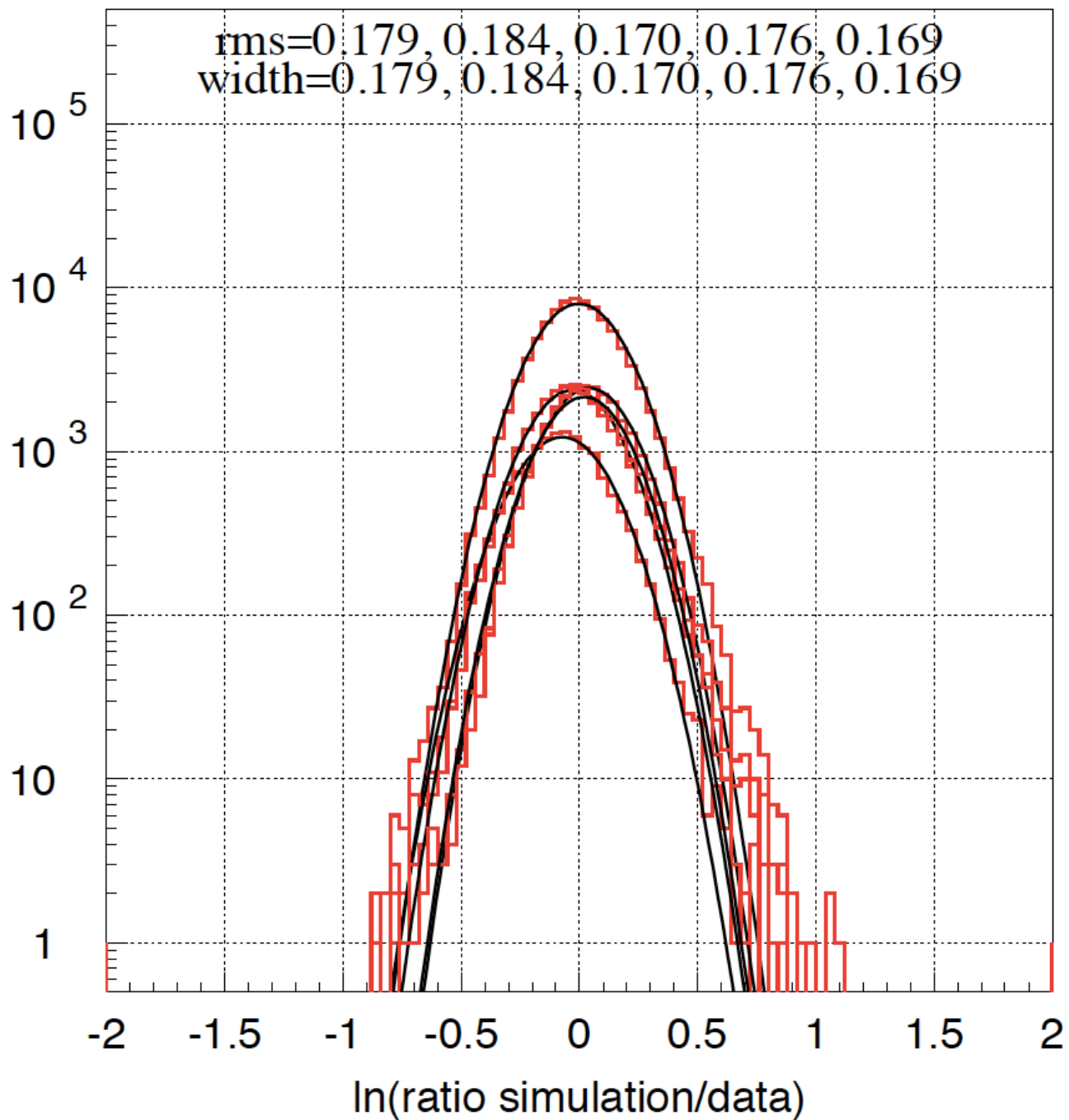
Flashers 21...30



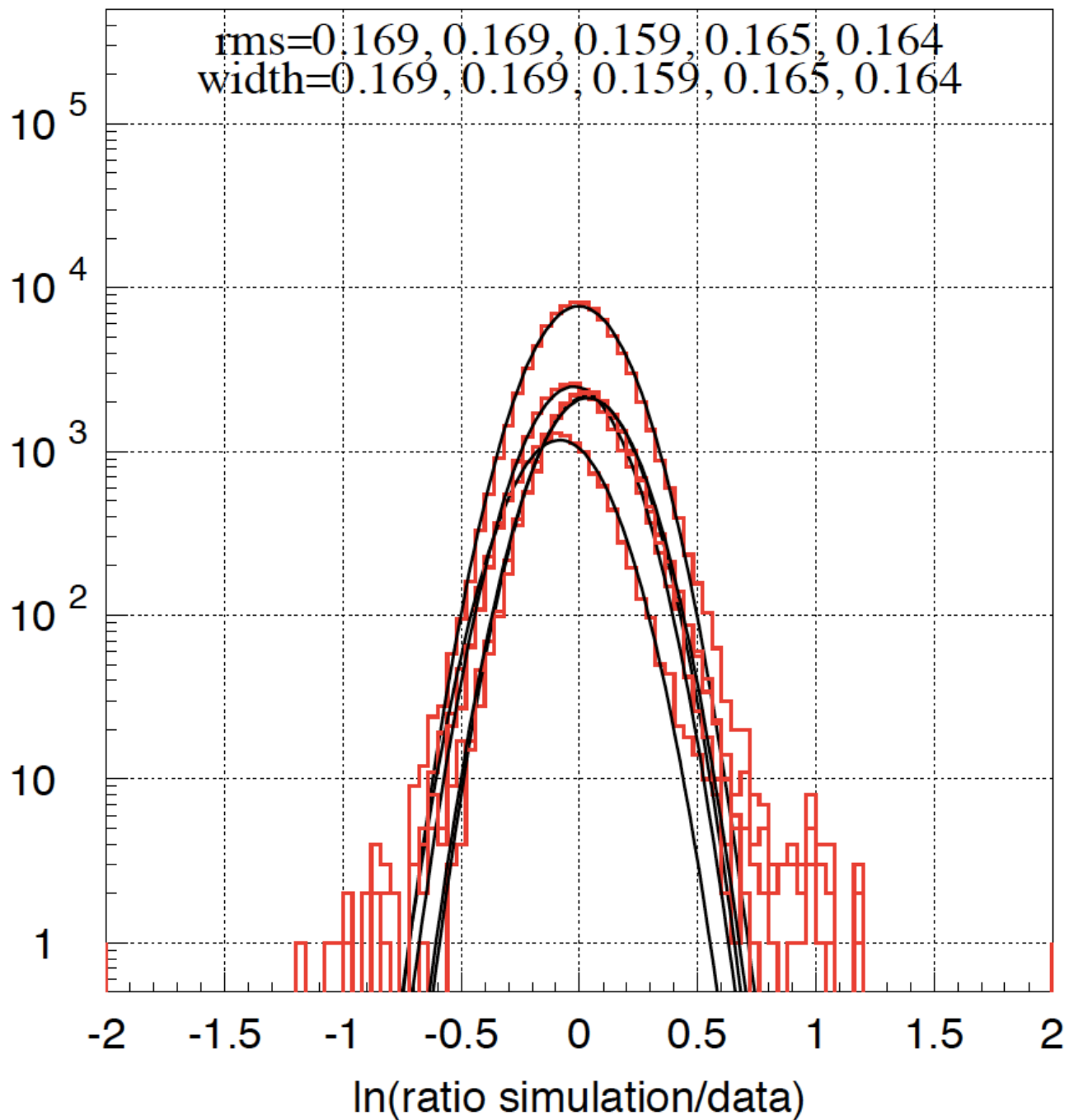
Flashers 31...40



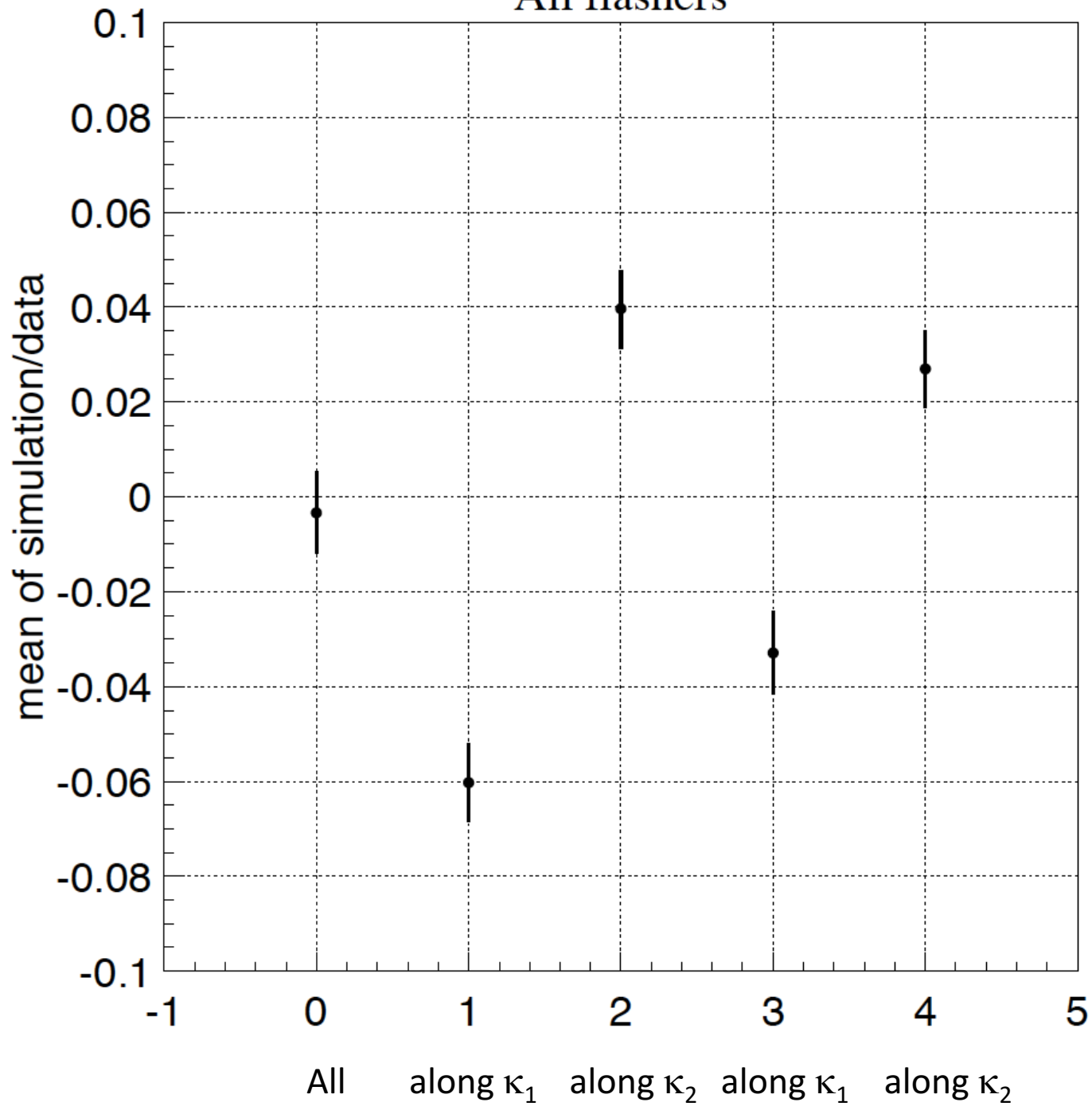
Flashers 41...50



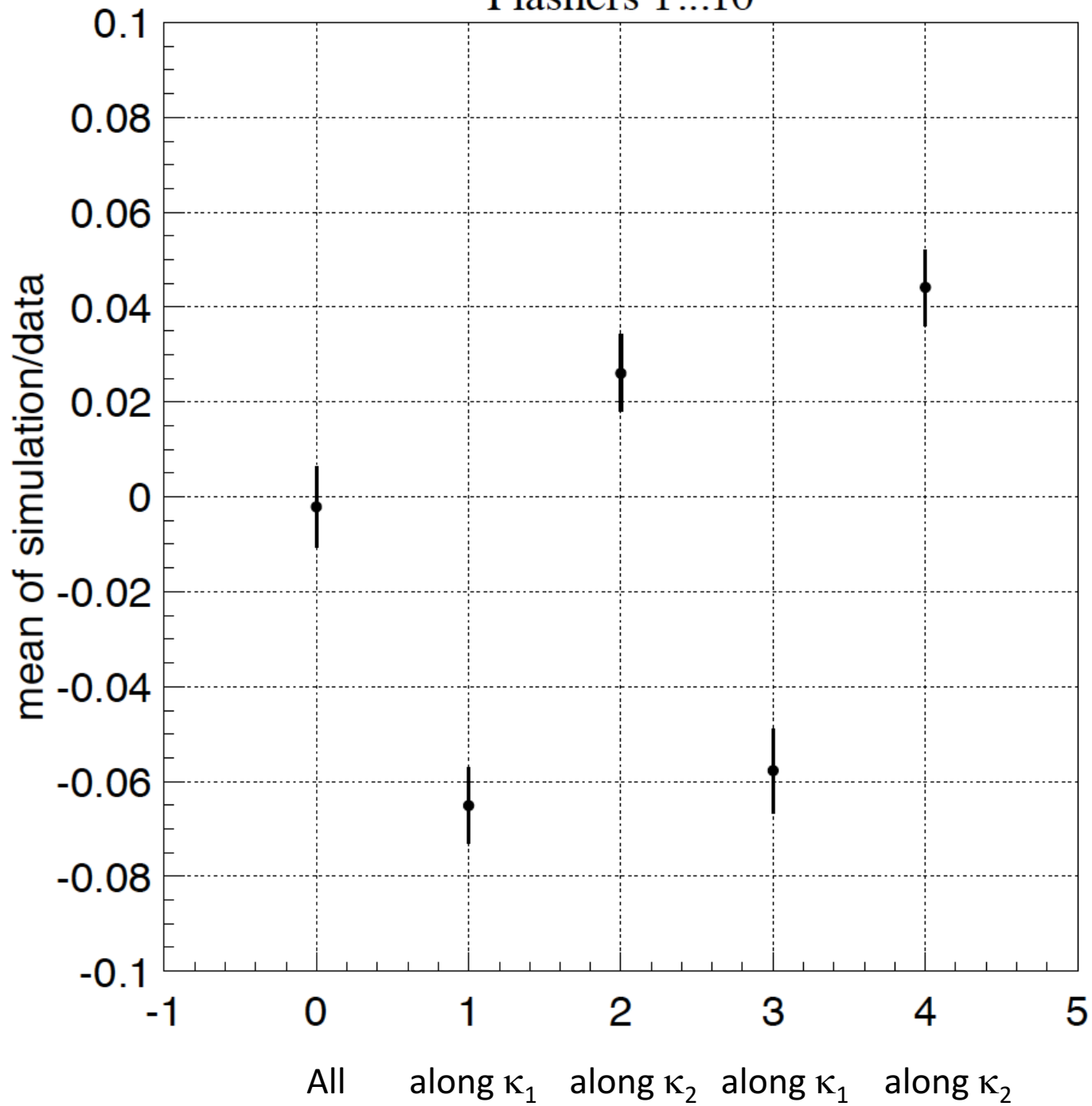
Flashers 51...60



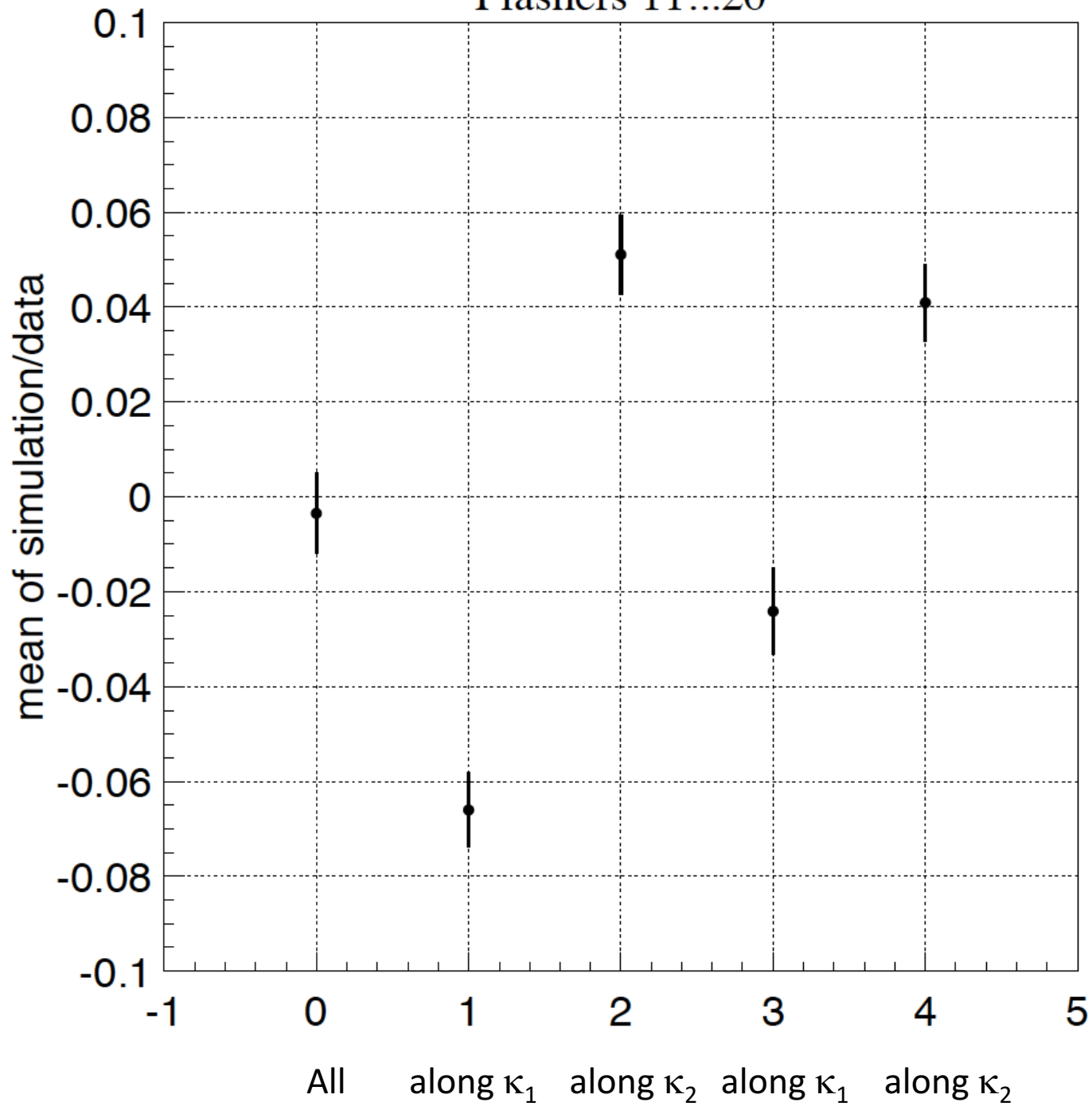
All flashers



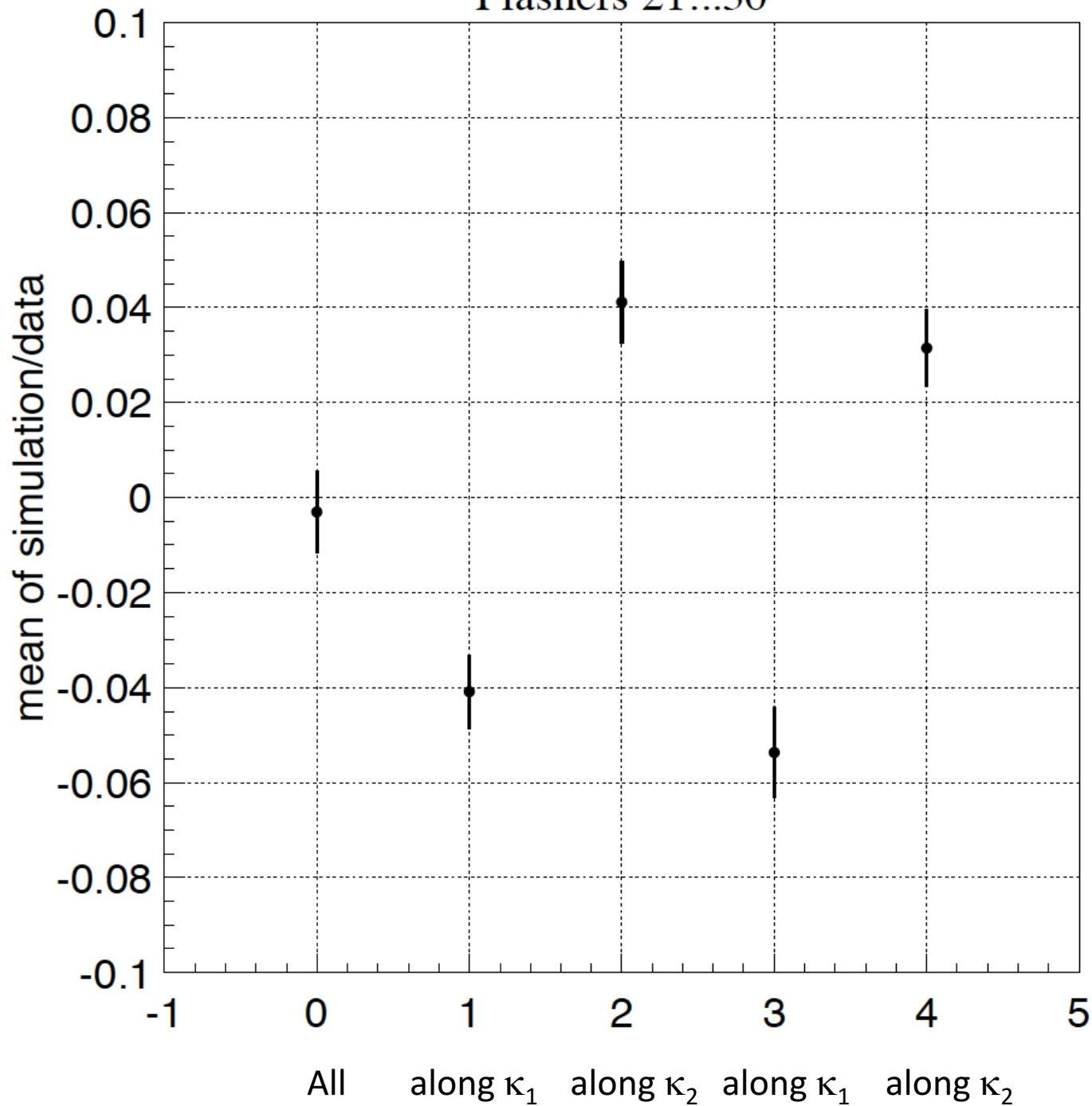
Flashers 1...10



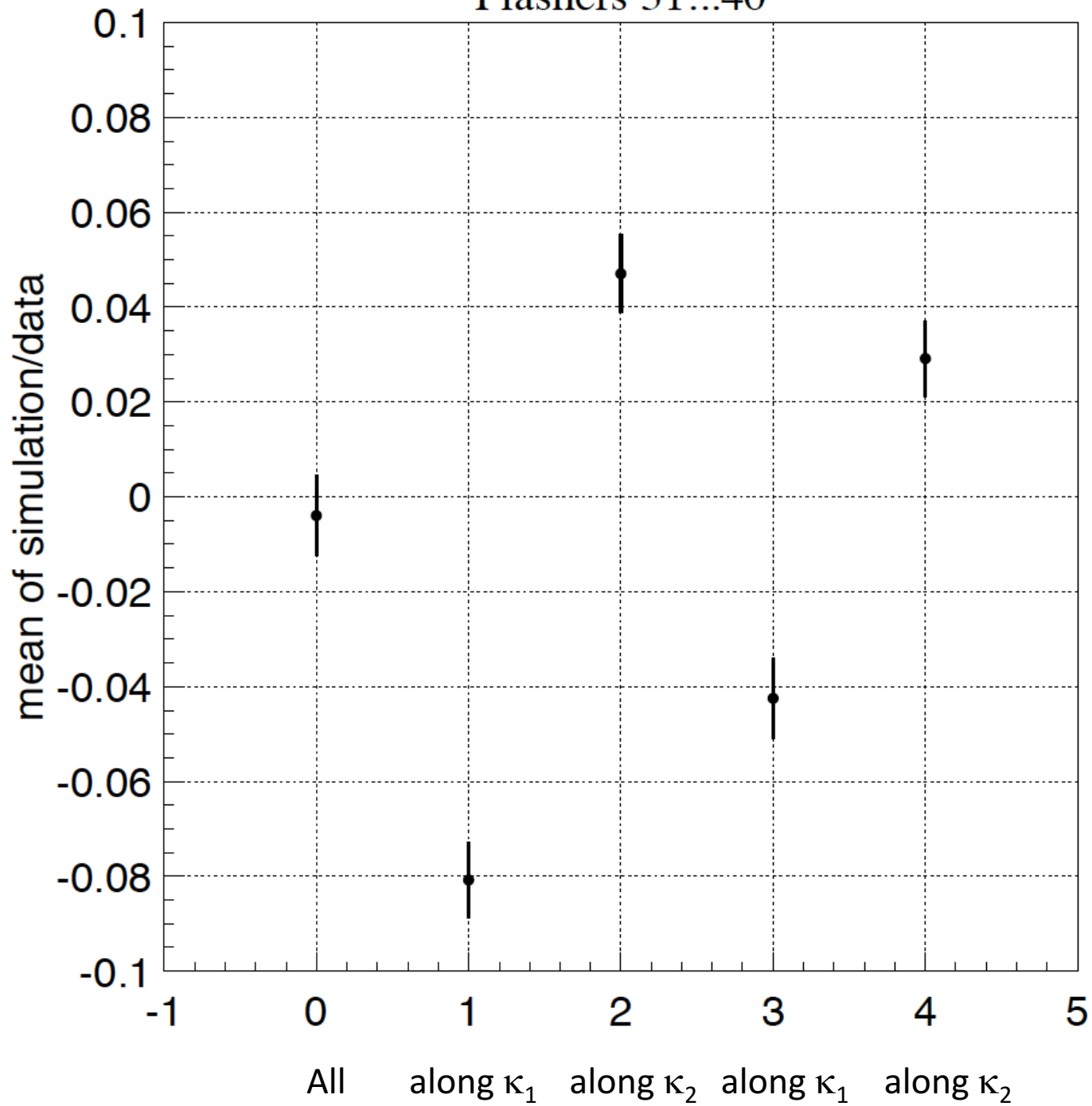
Flashers 11...20



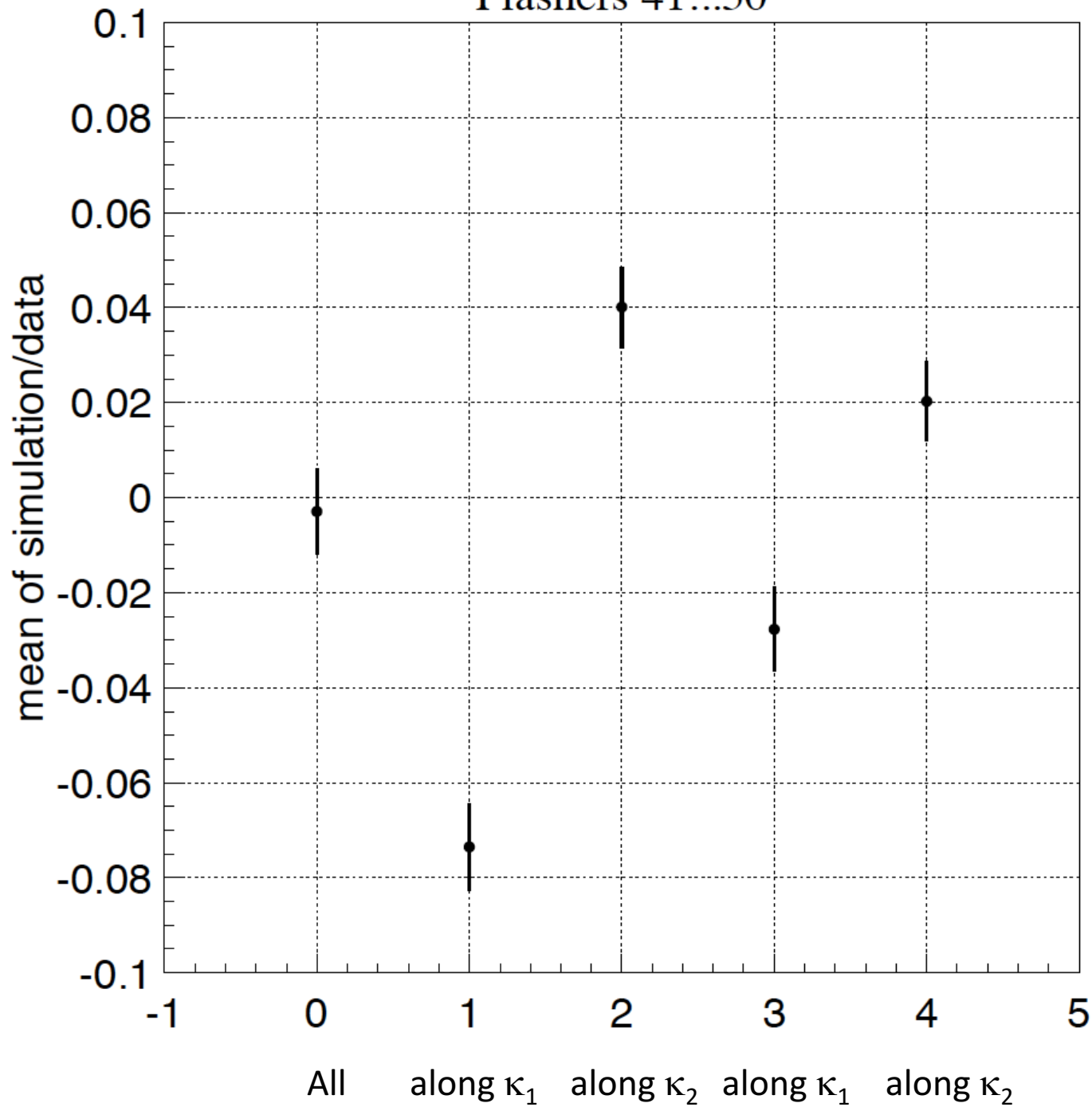
Flashers 21...30



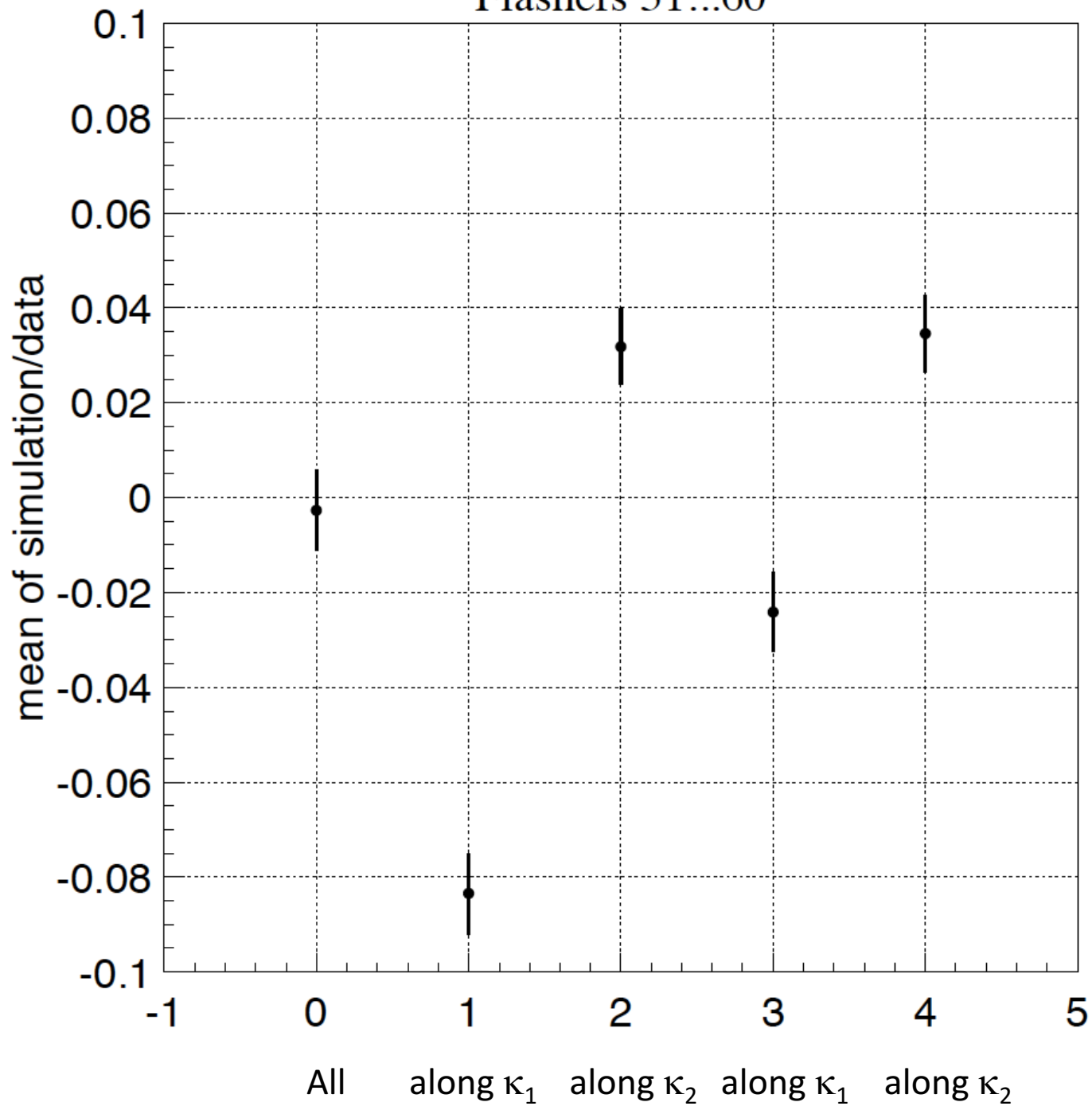
Flashers 31...40



Flashers 41...50



Flashers 51...60

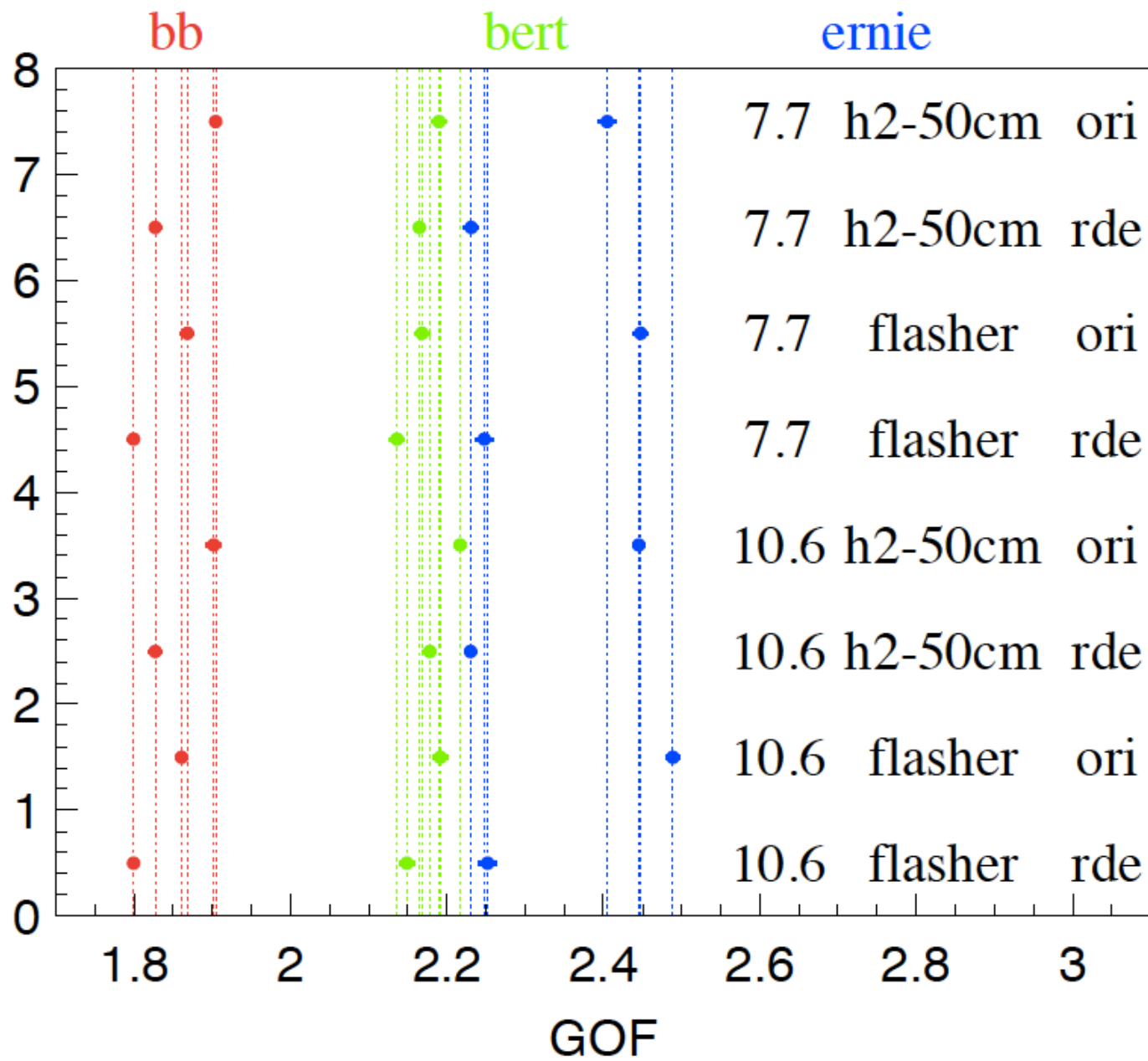


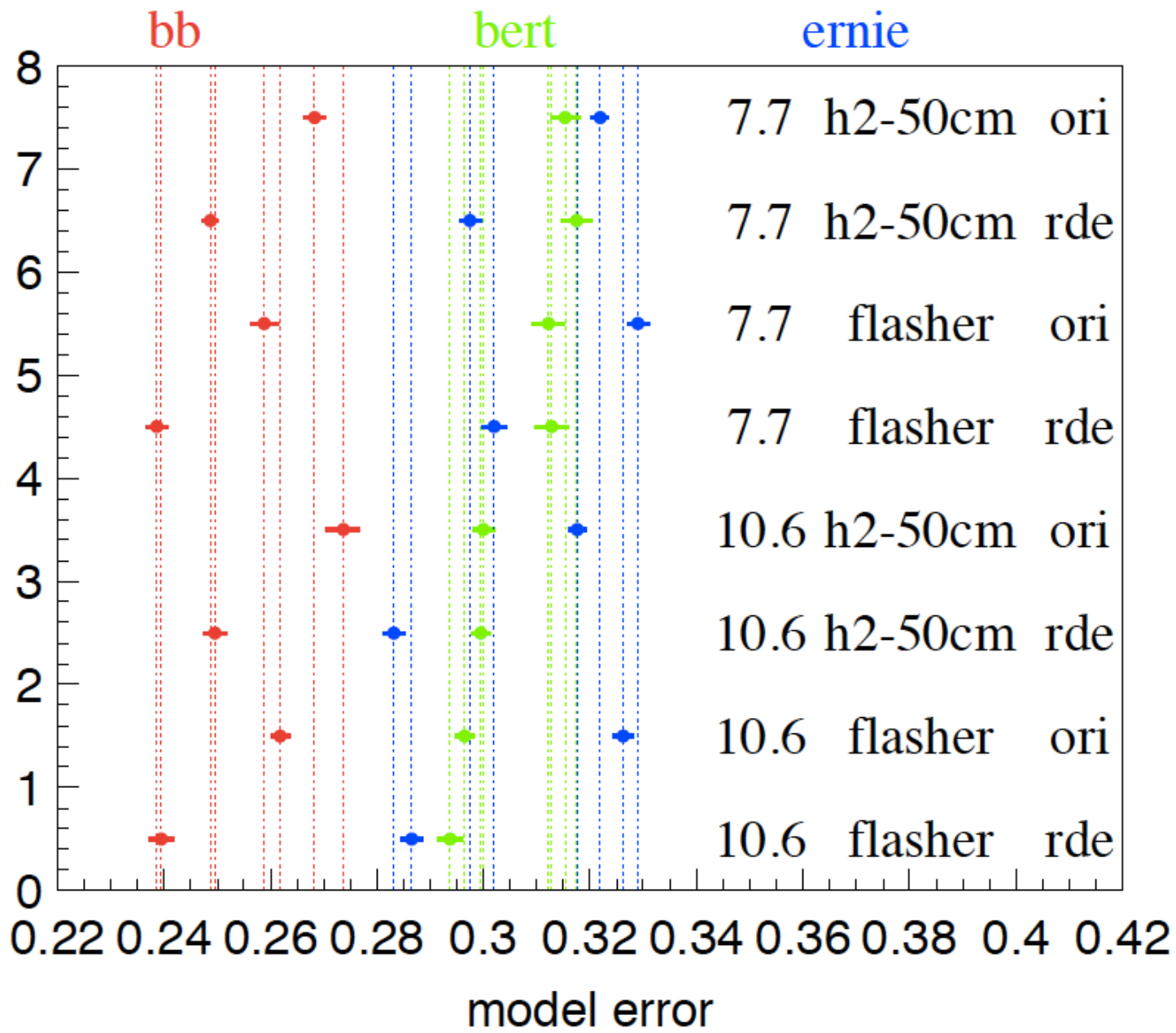
Ice variations with HESE cascades

Testing variations of the ice model (based on SPICE 3.2):

- anisotropy coefficient of 7.7% or 10.6%
- hole ice model: h2-50cm or flasher-derived (with $p=0.30$)
- DOM efficiency table: nominal (ori) or flasher-derived (rde)

25 events simulated in each case with 100x statistics ($s_{\text{rep}}=100$)





w/total WF charges above 5 p.e.

Flasher-derived DOM efficiencies improve both the GOF and model error for these cascade events in all cases!

10.6% anisotropy is slightly better than the 7.7% in the model error for Bert (for all combinations of the DOM efficiency and hole ice settings). It is also better for Ernie with the flasher-derived DOM efficiencies. Where 10.6% is worse the difference is smaller.